

CubeSat ADCS Simulation v3

1. Introduction
2. Code and Section Overview
3. Results
4. Next steps

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YUAA (Nov 2020)



Version history

- v1 in Fall 2020, based on propagator
- v2.2 with Sun, edited main file and Power function
- v3 checked Nov 2020, functions cleaned up

Objective

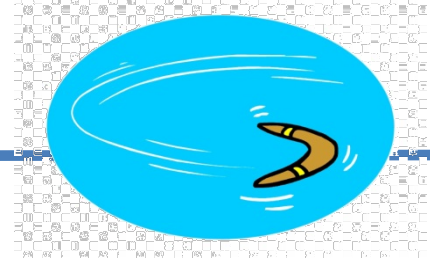
Simulate effectiveness of **various control strategies** to point CubeSat correctly

- *For a given initial spin, how fast can it stop?*
- **Tailored** to our orbit and mechanical design
- **Not flight software**

Currently assumes **perfect sensor, no lag** in controller execution
(can model B-dot implementation .etc if required)



Problem Description



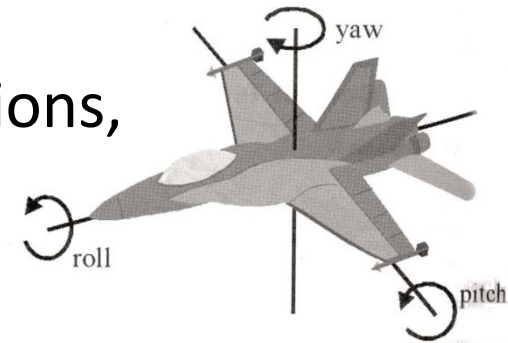
We need to consider both

- **Translational** motion along the orbit (ellipse)
- **Rotational** motion (tumbling in space)

While orbit is a “textbook” **2-body problem**, we need to determine how much torque to apply – given Earth’s varying **magnetic field vector** at different points



2-axis stabilization -> stop **roll** and **pitch** rotations, more efficient

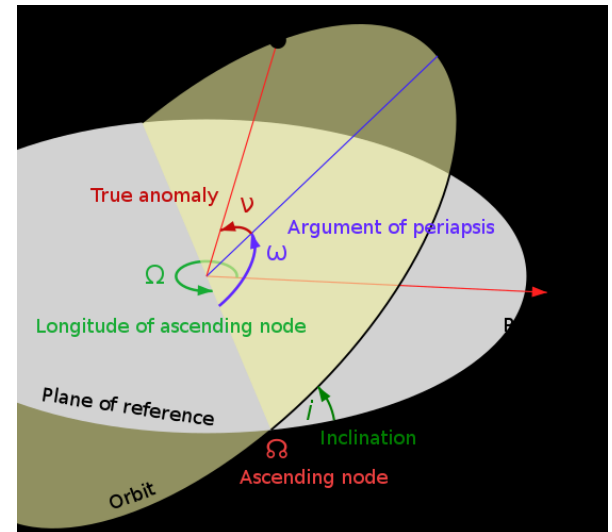


Outline of simulation

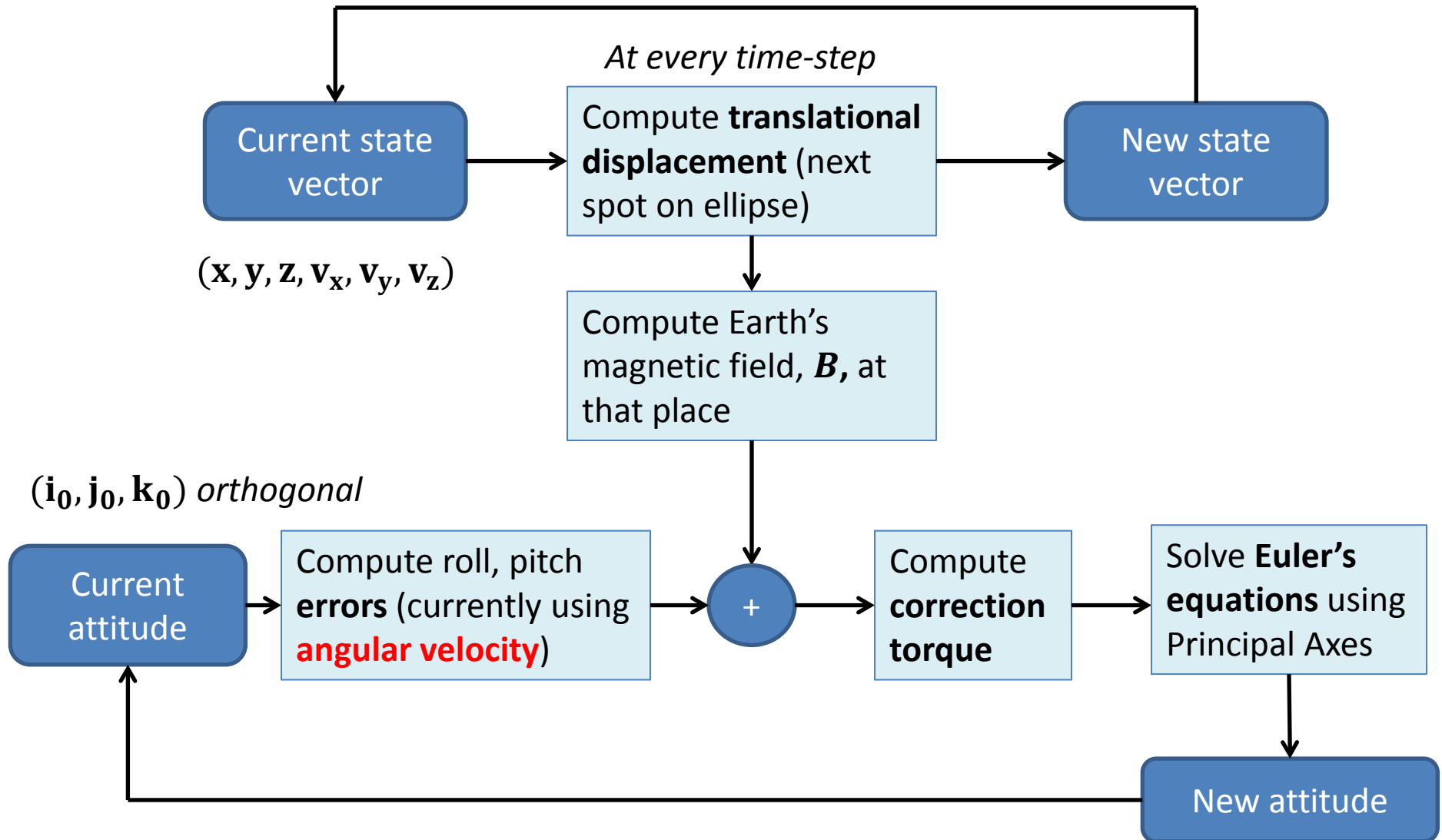
- Work out orbit position (propagation step) and velocity
- Find magnetic field at that point, inertial then transform into satellite frame (*a tricky matrix multiplication*)
- Determine what torque to apply (bang bang refers to On, Off only → no proportionality yet)
- Apply $\tau = I\alpha$ in **inertial frame** to find next step omegas (*don't want fictitious forces! More matrices*)
- Based on Sun's angle and satellite longitude angle, work out solar flux received for power.
- Back to propagation...

Propagation

- Orbit propagation is the basic step because we need to know where the satellite is and, in this model, what magnetic field is available for control
- Great to compare against STK here
- **Six** degrees of freedom so at least six inputs



Process Flowchart



Files: Attitude_Simulator_v1.1.py, Propagator.py, Controller.py, MagneticField.py, Solver.py

Key Equations

- 2-Body problem $\mathbf{F} = m\mathbf{a} \Rightarrow$

$$\frac{d^2 \mathbf{r}}{dt^2} = -\frac{GM}{r^3} \mathbf{r}$$
- Earth's magnetic field (tilted dipole model) from MIT notes
- Torque $\mathbf{T} = NIA \times \mathbf{B}$ with N turns of coils and current I
- In principal axes (frame of satellite), Euler's equations of motion

$$\begin{bmatrix} B_{north} \\ B_{east} \\ B_{down} \end{bmatrix} = \left(\frac{6378}{r_{km}} \right)^3 \begin{bmatrix} -C_\phi & S_\phi C_\lambda & S_\phi S_\lambda \\ 0 & S_\lambda & -C_\lambda \\ -2S_\phi & -2C_\phi C_\lambda & -2C_\phi S_\lambda \end{bmatrix} \begin{bmatrix} -29900 \\ -1900 \\ 5530 \end{bmatrix}$$

Where: C=cos, S=sin, ϕ =latitude, λ =longitude

Units: nTesla



$$M_x = I_{xx}\dot{\omega}_x - (I_{yy} - I_{zz})\omega_y\omega_z$$

$$M_y = I_{yy}\dot{\omega}_y - (I_{zz} - I_{xx})\omega_z\omega_x$$

$$M_z = I_{zz}\dot{\omega}_z - (I_{xx} - I_{yy})\omega_x\omega_y$$

Import functions

Approx. ISS orbit parameters; in **SI units**, refer to **Two Line Elements**

Tailor to our CubeSat inertia, coil geometry. **Time-step 1 s**

Initial **disturbance** (goal is to damp this)

x,y,z for roll, pitch, yaw

```

1  """
2  CubeSat Attitude Determination and Control System Simulation
3  Version 3: Bang Bang Control with Sun added (24 Nov 2020)
4
5  Orbit, omega, power outputs look reasonable
6  @author: Yu Jun
7  """
8  import math
9  import numpy as np
10 from numpy import linalg as LA
11 import csv
12 import Propagator
13 import MagneticField
14 import Controller
15 import Solver
16 import Power
17
18 '''=====Test conditions (change this part only)====='''
19
20 iniPos = np.array([-6719.400, 385.319, 2.669, -0.272368, -4.77507, 6.03443])
21 #ISS [x,y,z,vx,vy,vz] from STK, distances in km and velocity in km/s
22
23 dt = 1 # unified throughout
24 q = 10 # data record rate (every q frames)
25 Duration = 400*60 # seconds
26 i0 = np.array([1,0,0]) # initial attitude of spacecraft, in inertial coord
27 j0 = np.array([0,1,0]) # these '0' vectors must be orthogonal and unit mag.
28 k0 = np.cross(i0,j0)
29 omegaX = 0.2 # Starting test values in spacecraft frame. rad/s
30 omegaY = 0.1 # so correspond to roll/pitch/yaw
31 omegaZ = 0.3
32 '''=====

```

```

34
35     '''Satellite constants (input once the design is finalised)'''
36     Ix = 1    # moment of inertia along *principal* axes
37     Iy = 1
38     Iz = 1
39     turns = 10
40     area = 0.001
41     Kp = 0.02
42
43     '''=====Computation constants (don't need to change)====='''
44     totalSteps = int(Duration/dt)
45     Pos = iniPos          # initialize state vector
46     i = np.array([1,0,0]) # unit vectors in Earth non-rotating inertial frame
47     j = np.array([0,1,0])
48     k = np.array([0,0,1])
49     TestData = np.array([omegaX,omegaY,omegaZ,Kp]) # save initial test data
50     Jx = 0 # initialise current in x torque coil
51     Jy = 0
52     Jz = 0
53     M_old = [0.0, 0.0, 0.0] # initialize torque history
54     sunAngle = 0    # assume in ecliptic
55     History = []    # initialize records
56     i0data = []
57     Time = 0        # seconds
58
59     orbitDebug = [] # initialize empty list for testing
60

```

Data structures

Main code

```
61 '''*****=====Start Loop=====*****'''
62 print("Starting simulation... iniPos:", Pos)
63 for n in range(totalSteps):
64     """I. 2 Body Forward Propagation"""
65     newPos = Propagator.RK4(Pos[0],Pos[1],Pos[2],Pos[3],Pos[4],Pos[5],dt)
66     x = newPos[0]
67     y = newPos[1]
68     z = newPos[2]
69     vx = newPos[3]
70     vy = newPos[4]
71     vz = newPos[5]
72     Pos = newPos
73     #print(newPos)    # debug
74
75     r_vectorMag = (x**2 + y**2 + z**2)**0.5 # magnitude of radius vector, in km
76     lat = np.arcsin(z/r_vectorMag)         # latitude, radian
77     long = np.arctan2(y,x)                 # longitude; y=0 is Greenwich?
78
79     """II. Magnetic Field Calculation"""
80     BfieldGCI = MagneticField.TiltedDipoleXYZ(lat, long, r_vectorMag)
81     # in Earth non-rotating frame, nanoTesla, r_vectorMag in km
82
83     BfieldBFPA = MagneticField.GCItoBFPAtransform(i,j,k,i0,j0,k0,BfieldGCI[0],\
84                                                    BfieldGCI[1],BfieldGCI[2])
85
86     BfieldNED = MagneticField.TiltedDipoleNED(lat, long, r_vectorMag)
87     # transforms magnetic field to satellite principal axes frame
88     # sub 0's: actual spacecraft orientation, unit vectors
89
```

Compute translational displacement
in Geocentric Inertial Frame

Convert x,y,z to latitude/ longitude

```
90 """III. Magnetorquer Output"""
91 '''The B-field information is used to calculate torque by working out,
92 under nominal Bang Bang control. Then convert to BFPA frame. '''
93
94 NetTorque = Controller.nominalTorqueBFPA(omegaX,omegaY,omegaZ,BfieldBFPA)
95 #print(NetTorque)
96
97 """IV. Numerical Integration for omegas"""
98 nextOmega = Solver.EulerEqnSolver(omegaX,omegaY,omegaZ, NetTorque[0],
99                                   NetTorque[1], NetTorque[0],Ix,Iy,Iz,dt)
100
101 omegaX = nextOmega[0]
102 omegaY = nextOmega[1]
103 omegaZ = nextOmega[2]
```

Solve Euler's equations for
angular velocity

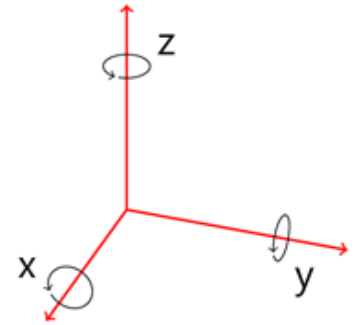
```

105     """V. Effecting omega""" # ensure unit vectors remain unit magnitude, orthogonal
106     # omegaX. rotate the principal axes accordingly. omegaX = Roll. x0 supposed to be forward facing
107     d0X = omegaX*dt # radian, small angle
108     i0new = i0
109     j0new = j0*math.cos(d0X) + k0*math.sin(d0X)
110     k0new = k0*math.cos(d0X) - j0*math.sin(d0X)
111
112     i0 = i0new
113     j0 = j0new
114     k0 = k0new # put in the new values
115
116     # omegaY = Pitch
117     d0Y = omegaY*dt # radian, small angle
118     j0new = j0
119     i0new = i0*math.cos(d0Y) - k0*math.sin(d0Y)
120     k0new = k0*math.cos(d0Y) + i0*math.sin(d0Y)
121
122     i0 = i0new
123     j0 = j0new
124     k0 = k0new # put in the new values
125
126     # omegaZ = Yaw
127     d0Z = omegaZ*dt # radian, small angle
128     k0new = k0
129     i0new = i0*math.cos(d0Z) + j0*math.sin(d0Z)
130     j0new = j0*math.cos(d0Z) - i0*math.sin(d0Z)
131
132     i0 = i0new
133     j0 = j0new
134     k0 = k0new # put in the new values
135

```

Main code

Updating the effect of omegas;
watch out for the **rotating** axes!



Save data at preset rate

Main code

```
137 """ VI. Power calculation with dark side"""
138 sunAngle = sunAngle + 2*np.pi/(24*60*60)*dt # sun moves
139 if sunAngle >= 2*np.pi:
140     sunAngle = sunAngle - 2*np.pi # keep to within 0, 2pi range
141 # because at ISS inclinations the sun's 23 deg tilt won't affect coverage
142 power = Power.flux(long, sunAngle, i0, j0, k0) # compute power
143
144 if (n//q)*q == n: # recording the data
145     Time = Time + q*dt
146     Data = [Time, Pos[0]/1000, Pos[1]/1000, Pos[2]/1000, Pos[3]/1000, Pos[4]/1000, Pos[5]/1000, BfieldNE
147             BfieldNED[1], BfieldNED[2], NetTorque[0], NetTorque[1], NetTorque[2], \
148             omegaX, omegaY, omegaZ, power]
149     History.append(Data) # B vectors experienced
150
151 orbitDebug.append([x,y,z,vx,vy,vz])
152
153 '''*****=====End loop=====*****'''
154 print("Simulation done. Timestep used: ", dt, "sec. Data recorded every", q, "frames.")
155 print("Total time:", dt*totalSteps, "sec")
156 print("[omegaX, omegaY, omegaZ, Kp]:", TestData)
157
158 # save data
159 with open('SimulationData.csv', 'w', newline='') as f:
160     writer = csv.writer(f)
161     writer.writerow(["Time (s)", "x (m)", "y (m)", "z (m)", "vx (m/s)", \
162                     "vy (m/s)", "vz (m/s)", "B_North (nT)", "B_East (nT)", \
163                     "B_Down (nT)", "RollTorque (Nm)", "PitchTorque (Nm)", \
164                     "YawTorque (Nm)", "OmegaX (rad/s)", \
165                     "OmegaY (rad/s)", "OmegaZ (rad/s)", "Power (Watt)"])
166     for row in History:
167         writer.writerow(row)
168 f.close()
```

Earth magnetic field

- Analytical tilted dipole model
 - Refer to notes from MIT 16.684 Space Systems Product Development
 - Need to convert to inertial frame

$$\begin{bmatrix} B_{north} \\ B_{east} \\ B_{down} \end{bmatrix} = \left(\frac{6378}{r_{km}} \right)^3 \begin{bmatrix} -C_\phi & S_\phi C_\lambda & S_\phi S_\lambda \\ 0 & S_\lambda & -C_\lambda \\ -2S_\phi & -2C_\phi C_\lambda & -2C_\phi S_\lambda \end{bmatrix} \begin{bmatrix} -29900 \\ -1900 \\ 5530 \end{bmatrix}$$

Where: C=cos , S=sin, ϕ =latitude, λ =longitude

Units: nTesla

 **flux**

Function: B field

```

2 """
3 Magnetic Field File (24 Nov 2020)
4 Contains two B field calculations: one in rotating Earth frame (North, East, Down)
5 and another in the x,y,z frame (with a vector transformation)
6
7 See MIT notes for tilted dipole mode (Slide 34):
8 https://ocw.mit.edu/courses/aeronautics-and-astronautics/16-851-satellite-engineering-fall-2003/Lectur
9 """
10 import numpy as np
11
12 def TiltedDipoleXYZ(lat,long, r_vectorMag):
13     """outputs magnetic field in ECI -> "x,y,z" frame"""
14     matrix = np.array([[ -np.cos(lat), np.sin(lat)*np.cos(long), np.sin(lat)*np.sin(long)],\
15                        [ 0,np.sin(long), -np.cos(long)],\
16                        [ -2*np.sin(lat), -2*np.cos(lat)*np.cos(long), -2*np.cos(lat)*np.sin(long)]] )
17     vector = np.array([ -29900, -1900, 5530])
18
19     Bfield = (6378/r_vectorMag)**3*matrix.dot(vector)
20     #In north, east and down currently. Use 6378 which is Earth radius in km.
21
22     Bfieldx = Bfield[0]*(-np.sin(lat)*np.cos(long)) + \
23               Bfield[1]*np.sin(long) + Bfield[2]*(-np.cos(lat)*np.cos(long))
24     Bfiel dy = Bfield[0]*(-np.sin(lat)*np.sin(long)) + Bfield[1]*np.cos(long) + \
25               Bfield[2]*(-np.cos(lat)*np.sin(long))
26     Bfieldz = Bfield[0]*np.cos(lat) - Bfield[2]*np.sin(lat)
27     # nanoTesla
28     return np.array([Bfieldx*10**(-9) , Bfiel dy*10**(-9) , Bfieldz*10**(-9) ])
29

```

Implement matrix

Function: B field

```
def GCIttoBFPatransform(i,j,k,i0,j0,k0,x,y,z):
    """Transforms vector x,y,z in coordinate frame with unit vectors i,j,k
    into vector x0,y0,z0 in coordinate frame with unit vectors i0,j0,k0.
    Use: transform B field from Geocentric Inertial Frame to spacecraft
    Body-Fixed Principal axes frame. See MIT Dynamics Lecture for math.
    Outputs: new vector x0,y0,z0."""

    x0 = np.dot(i0,i)*x + np.dot(i0,j)*y + np.dot(i0,k)*z
    y0 = np.dot(j0,i)*x + np.dot(j0,j)*y + np.dot(j0,k)*z
    z0 = np.dot(k0,i)*x + np.dot(k0,j)*y + np.dot(k0,k)*z

    return np.array([x0,y0,z0])

def TiltedDipoleNED(lat,long, r_vectorMag):
    """Debug function to compare against STK.
    Returns array of B field, in North, East, Down (NED) components
    in nanoTesla"""

    B_row1 = np.array([-np.cos(lat),np.sin(lat)*np.cos(long),
                       np.sin(lat)*np.sin(long)])
    B_row2 = np.array([0,np.sin(long),-np.cos(long)])
    B_row3 = np.array([-2*np.sin(lat),-2*np.cos(lat)*np.cos(long),
                       -2*np.cos(lat)*np.sin(long)])
    B_column = np.array([-29900,-1900,5530]) # from physics

    Mat1 = np.multiply(B_row1, B_column)
    B_north = Mat1[0] + Mat1[1] + Mat1[2] # matrix multiplication for B_north
    Mat2 = np.multiply(B_row2, B_column)
    B_east = Mat2[0] + Mat2[1] + Mat2[2] # B_east
    Mat3 = np.multiply(B_row3, B_column)
    B_down = Mat3[0] + Mat3[1] + Mat3[2] # B_down
    BfieldRot = np.array([B_north,B_east,B_down])*(6378/r_vectorMag)**3*10**(-9) # in nanoTesla

    return BfieldRot
```



About propagators...

- There are different levels of accuracy for this
 - 2 Body problem: textbook, six classical elements/ $[x,y,z,vx,vy,vz]$ state vector fully sufficient to describe orbit
 - J2, J4... : considers Earth “fatness” at the equatorial mass bulge. I am using J4 at the moment, relatively simple to implement under Runge Kutta 4 integration (no atmospheric drag)
 - SGP4 and above: much more advanced, considers other Earth mass distribution and other celestial bodies

In general for short durations (~days of orbit) no big deviation is expected; but for **long term** mission planning higher fidelity models are necessary.

Function: Propagator

```
24 def grad(p0,p1,p2,p3,p4,p5):      # RK4 gradient function
25     r = sqrt(p0**2 + p1**2 + p2**2) # Earth radius for J term calculations
26
27     Jx = 1 - J2*(3./2.)*(rE/r)**2*(5*p2**2/r**2-1) + \
28           J3*(5./2.)*(rE/r)**3*(3*p2/r-7*p2**3/r**3) - \
29           J4*(5./8.)*(rE/r)**4*(3-42*p2**2/r**2+63*p2**4/r**4) - \
30           J5*(3./8.)*(rE/r)**5*(35*p2/r-210*p2**3/r**3+231*p2**5/r**5) + \
31           J6*(1./16.)*(rE/r)**6*(35-945*p2**2/r**2+3465*p2**4/r**4-3003*p2**6/r**6)
32     Jz = 1 + J2*(3./2.)*(rE/r)**2*(3-5*p2**2/r**2) + \
33           J3*(3./2.)*(rE/r)**3*(10*p2/r-(35./3.)*p2**3/r**3-r/p2) - \
34           J4*(5./8.)*(rE/r)**4*(15-70*p2**2/r**2+63*p2**4/r**4) - \
35           J5*(1./8.)*(rE/r)**5*(315*p2/r-945*p2**3/r**3+693*p2**5/r**5-15*p2/r) + \
36           J6*(1./16.)*(rE/r)**6*(315-2205*p2**2/r**2+4851*p2**4/r**4-3003*p2**6/r**6)
37     thetaP = 0.00007292115
38     v = sqrt((p3+thetaP*p1)**2+(p4-thetaP*p0)**2+p5**2)
39     return [p3,p4,p5,-GM*(p0)/r**3*Jx,-GM*(p1)/r**3*Jx,-GM*(p2)/r**3*Jz]
40
41 def RK4(u0,u1,u2,u3,u4,u5,dt):    # standard RK4 implementation
42     k1 = grad(u0,u1,u2,u3,u4,u5)
43     k2 = grad(u0+k1[0]*dt/2, u1+k1[1]*dt/2, u2+k1[2]*dt/2, \
44             u3+k1[3]*dt/2, u4+k1[4]*dt/2, u5+k1[5]*dt/2)
45     k3 = grad(u0+k2[0]*dt/2, u1+k2[1]*dt/2, u2+k2[2]*dt/2, \
46             u3+k2[3]*dt/2, u4+k2[4]*dt/2, u5+k2[5]*dt/2)
47     k4 = grad(u0+k3[0]*dt, u1+k3[1]*dt, u2+k3[2]*dt, \
48             u3+k3[3]*dt, u4+k3[4]*dt, u5+k3[5]*dt)
49     res = [u0 + dt/6*(k1[0]+2*k2[0]+2*k3[0]+k4[0]), \
50           u1 + dt/6*(k1[1]+2*k2[1]+2*k3[1]+k4[1]), \
51           u2 + dt/6*(k1[2]+2*k2[2]+2*k3[2]+k4[2]), \
52           u3 + dt/6*(k1[3]+2*k2[3]+2*k3[3]+k4[3]), \
53           u4 + dt/6*(k1[4]+2*k2[4]+2*k3[4]+k4[4]), \
54           u5 + dt/6*(k1[5]+2*k2[5]+2*k3[5]+k4[5])]
55     return res
56
```

Gradient function for
Runge Kutta method



Controller

- Linearized calculation used in python
- Literally $\omega_{n+1} = \omega_n + \alpha \cdot \Delta t$
- Where α is the acceleration from the magnetorquer (no other torques considered **yet***, and bang bang control for now)
- Gravity boom can also be modelled for α

*eventually atmospheric drag, solar pressure will conspire to deviate the spacecraft :0

Controller

```
1  """
2  Controller File (24 Nov 2020)
3  Calculates output torque based on input error
4  Proposed Two Axis (pitch and roll) stabilization code assuming circular orbit
5  """
6  import numpy as np
7  import math
8
9  def nominalTorqueBFPA(OmegaX, OmegaY, OmegaZ, BfieldBFPA):
10     """
11     Input: current errors and Bfield to find corrective
12     torques (all in BPFA). Omegas are floats, BfieldBFPA is array
13     Output: Torque array in BFPA, to use directly in Solver module.
14
15     0.25 Amp*m2 is the nominal working dipole strength. This method is
16     BANG-BANG CONTROL : magnetorquer switches on or off only, at 0.25 Ampm^2
17     dipole strength. This value from Nanoavionics spec sheet, projeted to
18     consume 140 mW nominally.
19     """
20
21     TorqueX = np.array([0,0,0]) # initialize local variables
22     TorqueY = np.array([0,0,0]) # default bang bang buffer zone
23     TorqueZ = np.array([0,0,0])
24     dipoleX = np.array([0,0,0]) # vectors
25     dipoleY = np.array([0,0,0])
26     dipoleZ = np.array([0,0,0])
27     omega = np.array([OmegaX, OmegaY, OmegaZ])
```

Controller

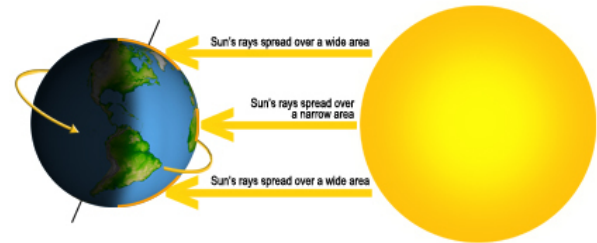
```
28
29     dipoleX = np.cross(np.array([1,0,0]),BfieldBFPA)
30     if dipoleX.dot(omega) < -10E-7:
31         TorqueX = 0.25*dipoleX          # run in +ive i0 direction
32     #     print("dipoleX dot:", dipoleX.dot(omega), "+i0")
33     elif dipoleX.dot(omega) > 10E-7:
34         TorqueX = -0.25*dipoleX         # run in -ve i0 direction
35     #     print("dipoleX dot:", dipoleX.dot(omega), "-i0")
36
37     dipoleY = np.cross(np.array([0,1,0]),BfieldBFPA)
38     if dipoleY.dot(omega) < -10E-7:
39         TorqueY = 0.25*dipoleY          # run in +ive i0 direction
40     #     print("dipoleY dot:", dipoleY.dot(omega), "+j0")
41     elif dipoleY.dot(omega) > 10E-7:
42         TorqueY = -0.25*dipoleY         # run in -ve i0 direction
43     #     print("dipoleY dot:", dipoleY.dot(omega), "-j0")
44
45     dipoleZ = np.cross(np.array([0,0,1]),BfieldBFPA)
46     if dipoleZ.dot(omega) < -10E-7:
47         TorqueZ = 0.25*dipoleZ          # run in +ive i0 direction
48     #     print("dipoleZ dot:", dipoleZ.dot(omega), "+k0")
49     elif dipoleZ.dot(omega) > 10E-7:
50         TorqueZ = -0.25*dipoleZ         # run in -ve i0 direction
51     #     print("dipoleZ dot:", dipoleZ.dot(omega), "-k0")
52
53     return TorqueX + TorqueY + TorqueZ  # do all at once
54
```

Solver

```
1  # coding: utf-8
2  """
3  v3 24 Nov 2020
4  Solver module
5  Goal: output accurate, fast omegax/y/z values after one timestep
6  input: this step's omega, dt
7  Design notes: At first we used single timestep forward march, now upgraded to
8  Runge Kutta 4 -> much faster when timestep is 1 sec instead of 1 ms before
9  IN SATELLITE BFPA
10 @author: user
11 """
12
13 Ixx = 1 # moment of inertia along *principal* axes
14 Iyy = 1
15 Izz = 1
16
17 def EulerEqnSolver(omegaX,omegaY,omegaZ, MomentX, MomentY, MomentZ, Ixx, Iyy, Izz, dt):
18     """Numerically solves Euler equations of motion
19     Inputs: current ang. velocity, torques, & moments of inertia in Principal Axes
20     Outputs: next timestep's angular velocities in Principal Axes frame
21     Direct forward march numerical integration
22     Linearize omega dot across one timestep; expect some error over time"""
23
24     omegaX0 = omegaX # old variable, nth step
25     omegaY0 = omegaY
26     omegaZ0 = omegaZ
27
28     omegaX = omegaX0 + (MomentX/Ixx + (Iyy-Izz)/Ixx*omegaY0*omegaZ0)*dt #(n+1)th step
29     omegaY = omegaY0 + (MomentY/Iyy + (Izz-Ixx)/Iyy*omegaZ0*omegaX0)*dt
30     omegaZ = omegaZ0 + (MomentZ/Izz + (Ixx-Iyy)/Izz*omegaX0*omegaY0)*dt
31     return [omegaX,omegaY,omegaZ]
```

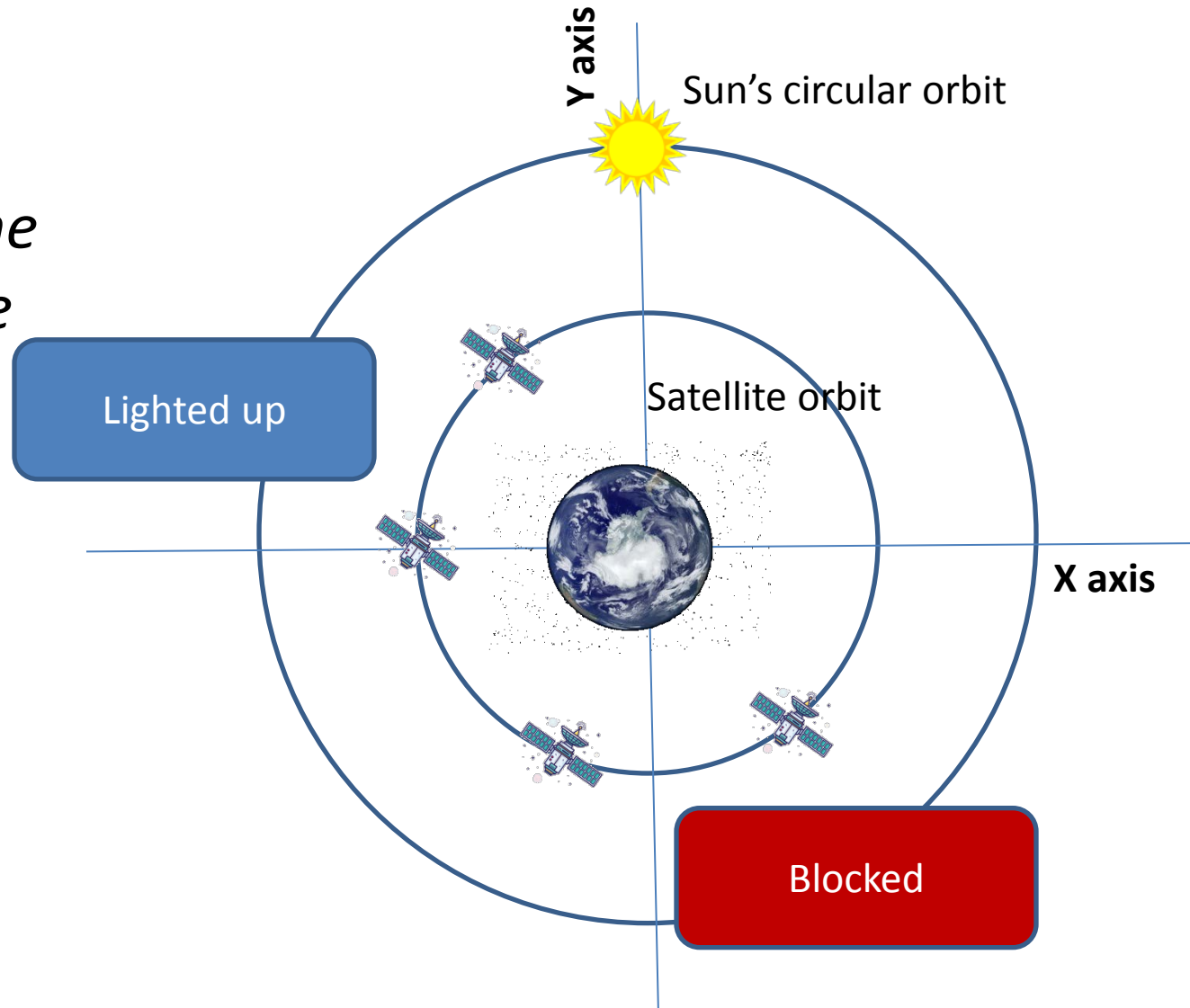
Solar power methodology

- Added a rotating sun
- Sunlight incident if and only if $|\theta_{Sun} - \theta_{Sat}| \leq \frac{\pi}{2}$
- Furthermore, $\theta_{Sun}, \theta_{Sat} \in [0, 2\pi]$ and considering the Equator/Ecliptic, $\theta_{Sat} = longitude = \tan^{-1} \left(\frac{y}{x} \right)$
- Can safely ignore 23 degrees tilt because our inclination (@ ~ISS) is less than $90 - 23 = 67$ at which this would be significant



Solar power illustration

We are looking down from the “North Pole”; the circles are in the plane of the equator



Power

```
7
8 def flux(long, sunAngle, i0, j0, k0):
9     """Calculates solar power received by satellite
10     Input: satellite longitude, sun angle and current satellite orientations in i0, j0, k0 vectors
11     Output: power (Watt)
12     k0 points upwards from top of satellite (solar panel exists there)"""
13
14     efficiency = 0.307 * 0.88 * (1 - (75-28) * 0.0022) #efficiency of solar panel
15     Area2U = 0.01076664 # Area of one 2U panel in m^2
16     # using arrays for inertial i and k unit vectors
17
18     phi = 1373*(math.cos(23)*np.array([1,0,0]) - math.sin(23)*np.array([0,0,1]))
19     # solar flux vector
20
21     powerTop = phi.dot(k0)*Area2U/2
22     if powerTop < 0: # top is sunlit
23         powerTop = efficiency*abs(powerTop)
24     else:
25         powerTop = 0
26
27     powerSidei0 = efficiency*abs(phi.dot(i0)*Area2U) # don't double count
28     powerSidej0 = efficiency*abs(phi.dot(j0)*Area2U)*0.75 #reduced to 1U on a side
29
30     powerAll = powerTop + powerSidei0 + powerSidej0 # before considering dark side
31
32     if abs(sunAngle - long) < np.pi/2:
33         powerAll = powerAll # lighted up
34     else:
35         powerAll = 0 # in shadow of Earth
36
37     return powerAll
```

Thoughts on Function Files

- Propagator.py: solves 2-body problem using Runge-Kutta 4 numerical method
 - We *could* consider orbit perturbations like Earth J2-6 harmonics, but I think **unnecessary** at this stage (errors from elsewhere + **different time-scale** of orbit vs tumbling)
- Controller.py: uses linear gain factor (P part of PID; this coefficient must be tested!)
- MagneticField.py: implements analytical tilted dipole model (see MIT AeroAstro notes)
 - Probably good enough for now; check SI **units**!
- Solver.py: uses **forward time march**
 - Suggest we start here to improve accuracy



Notes and some issues

- ~~1) Orbit duration is a little shorter than STK~~
- 2) The other set of values Jess and Michael plotted in STK on 10/11/20 surprisingly broke the same propagator. I have no idea why ☹️ so it's back to the ISS orbit here
- 3) Solar power considers Earth blockage (yay!)
- 4) Magnetic field looks sensible?
- 5) Bang Bang control is very crude and might give instabilities, though it works as a first approximation (omegas decrease as desired)

Simulation parameters

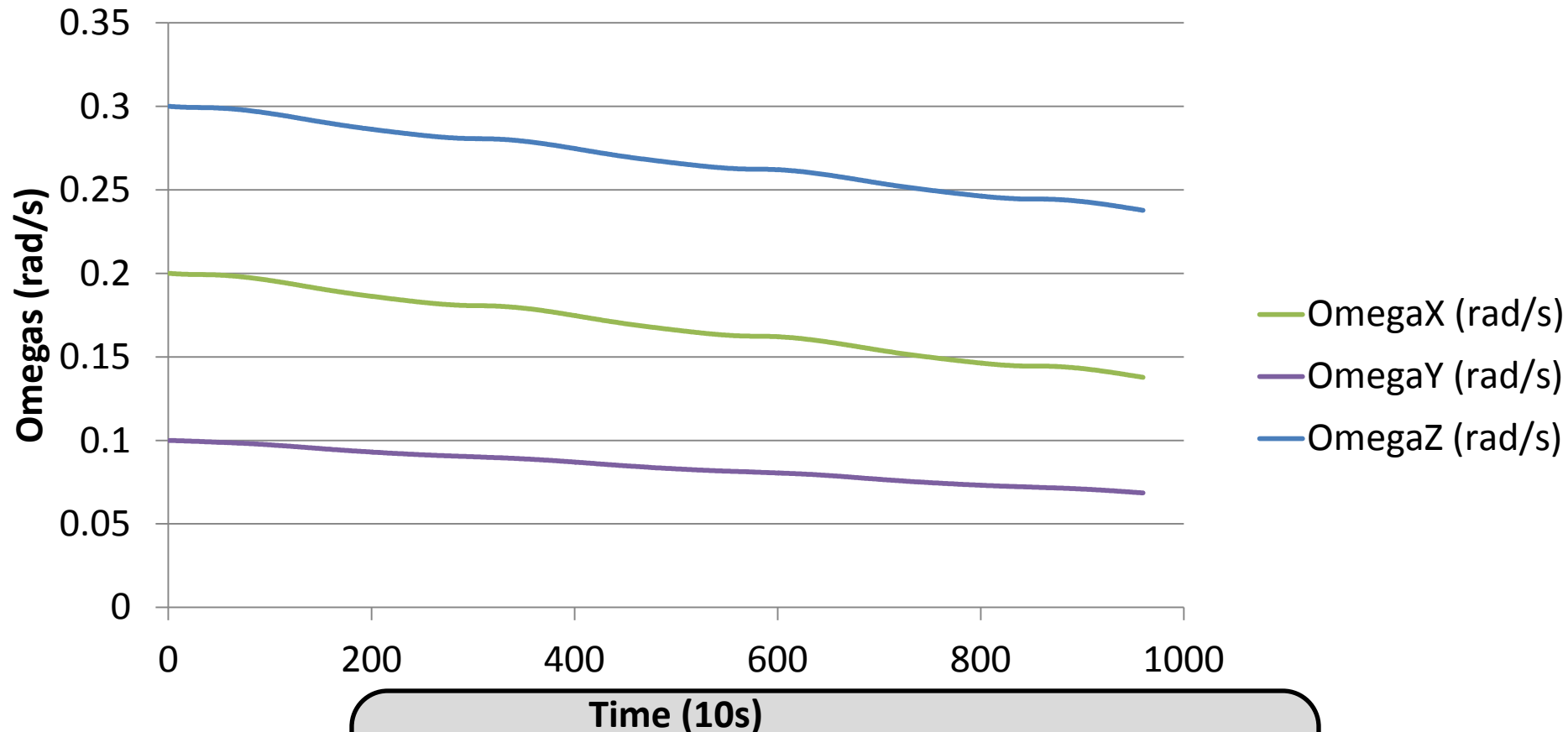
- Started off with (ISS?) orbit: (-5801232.89, 3520987.41, 6.751 [km], -2136.353, -3519.9, 6120.767 [km/s]) state vector (x,y,z,vx,vy,vz)
- Roll, Pitch, Yaw at 0.2, 0.1, 0.3 rad/s at first (ω_X , ω_Y , ω_Z)
- 1 second time step, data saved every 10 second (so as to see power fluctuation in rotating satellite)
- Bang Bang maximum-effort control

From STK

Start:	9 Feb 2020 12:00:00.000 LCLG	
Stop:	10 Feb 2020 12:00:00.000 LCLG	
Step Size:	60 sec	
Orbit Epoch:	9 Feb 2020 12:00:00.000 LCLG	
Coord Epoch:	9 Feb 2020 12:00:00.000 LCLG	
Coord Type:	Cartesian	
Coord System:	TrueOfDate	
Prop Specific:	Special Options...	
X:	2036.79 km	
Y:	3764.48 km	
Z:	5279.5 km	
X Velocity:	-6.17708 km/sec	
Y Velocity:	4.45762 km/sec	
Z Velocity:	-0.79389 km/sec	

Cancel Apply Help

Omega under control

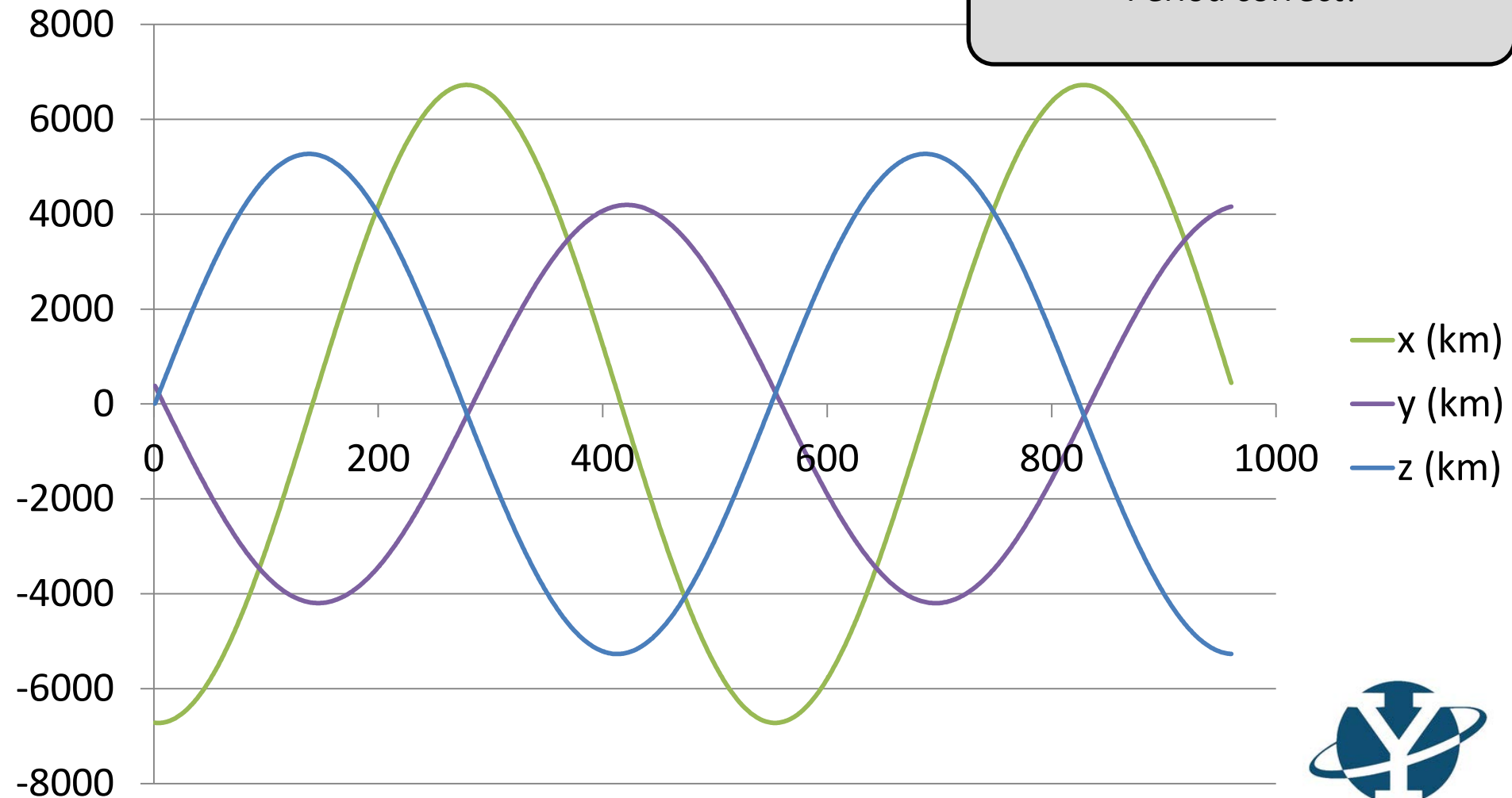


Exponential decay from **proportional** control $\frac{d\omega}{dt} = k\omega$
Greater K_p -> faster decay -> but power okay?

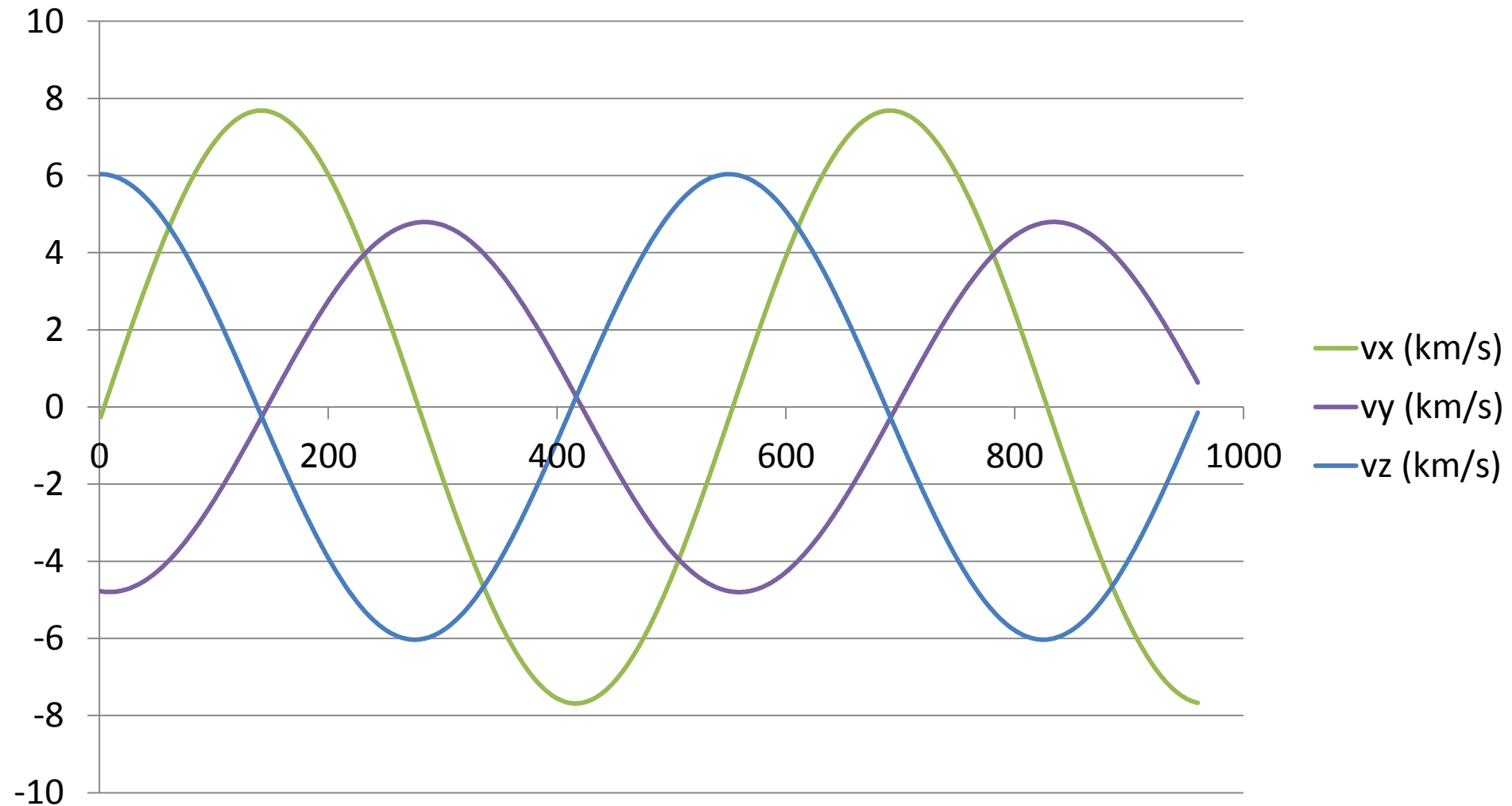


Orbital positions

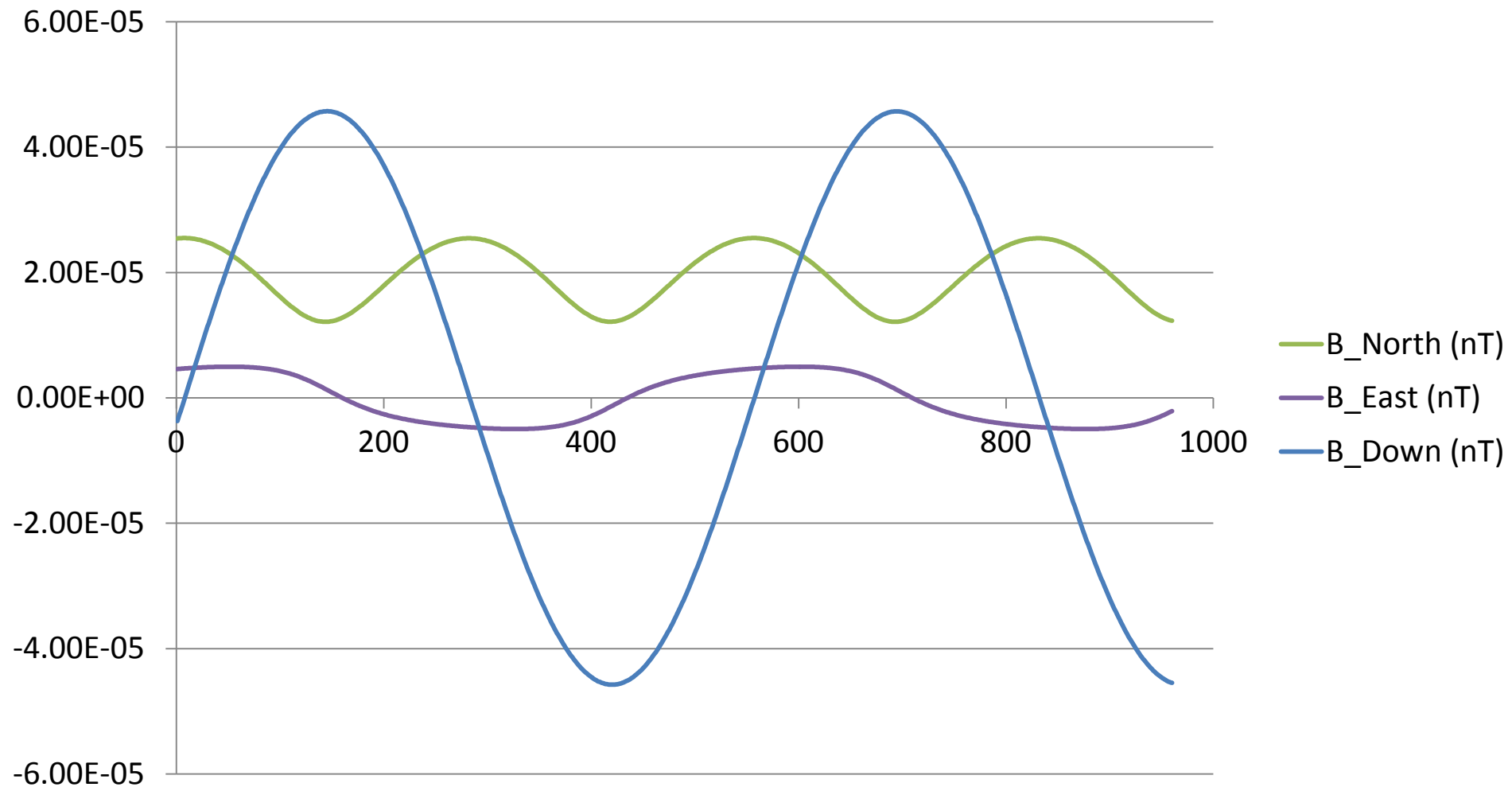
Period correct!



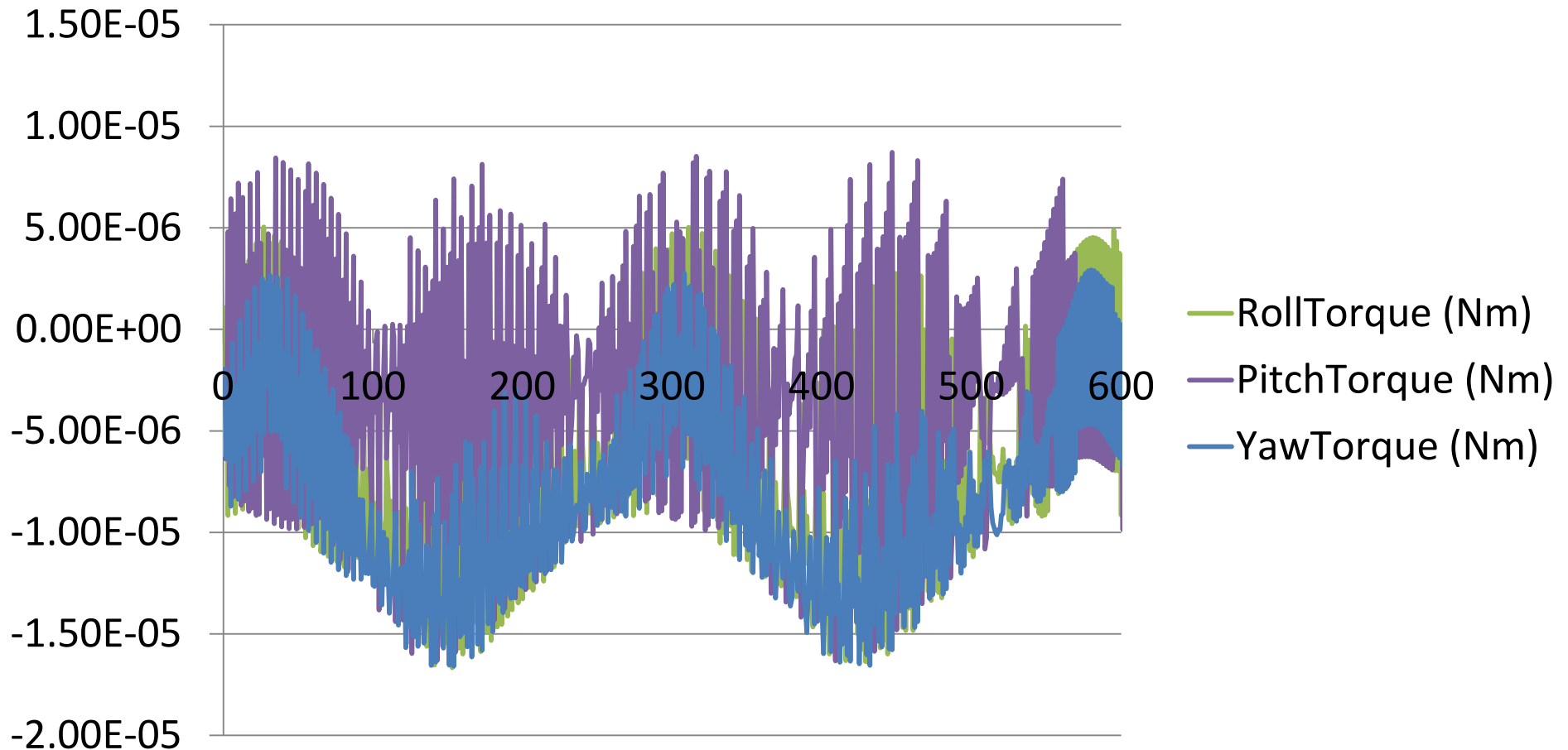
Orbital Velocities



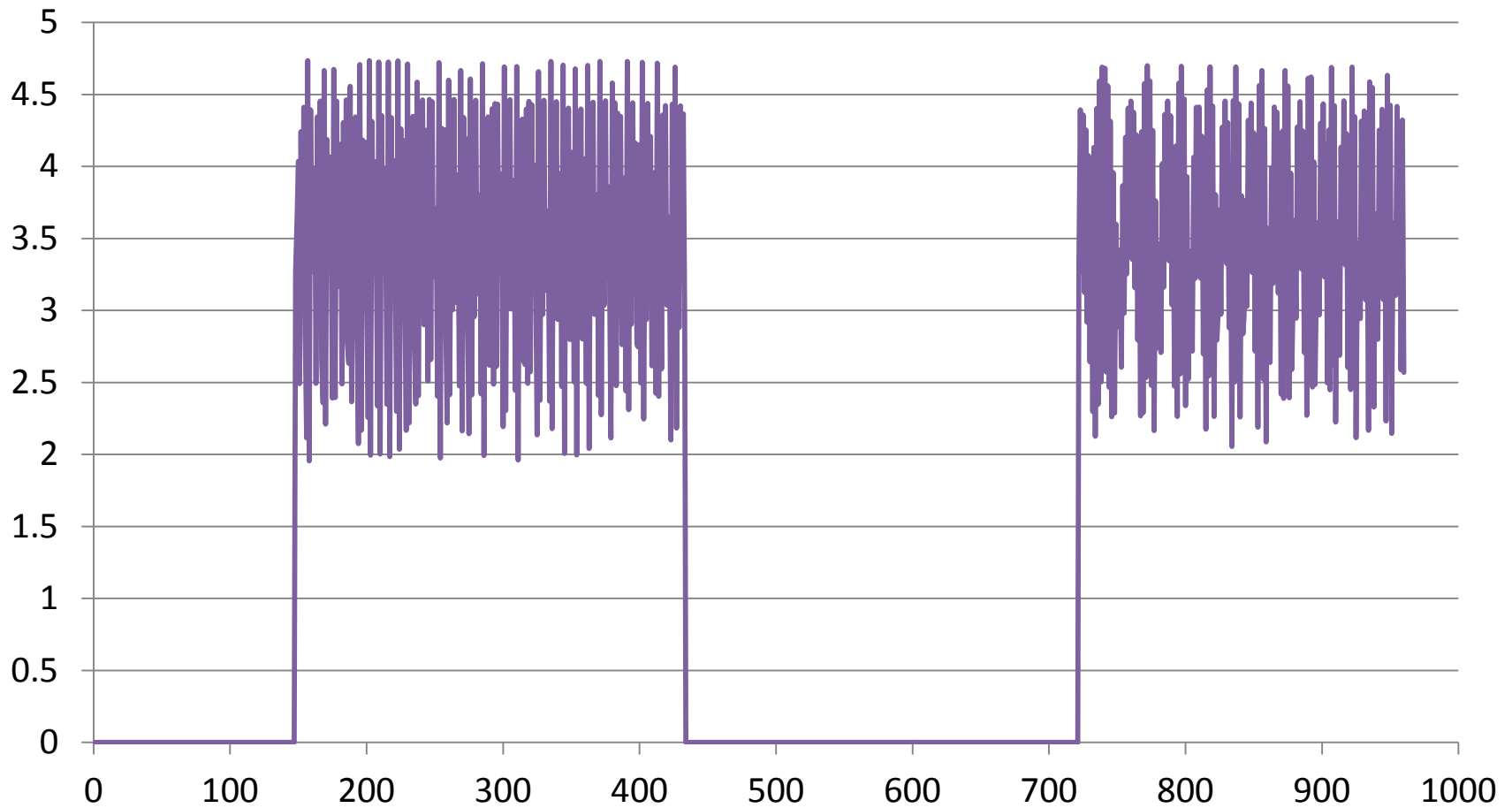
Earth magnetic field



Torques (Nm)



Power (Watt)



Suggested next steps

1. Check math and accuracy of results
2. Substitute our CubeSat values
3. Nicer way to get ω from Euler's equations?
4. If $\hat{\omega}$ are OK, determine solar flux available
5. Compare active control vs passive (gravity boom) control time needed



References

- MIT AeroAstro notes
 - <https://ocw.mit.edu/courses/aeronautics-and-astronautics/16-851-satellite-engineering-fall-2003/lecture-notes/> (esp Lecture 9 on ADCS)
- Numerical solutions
 - https://en.wikipedia.org/wiki/Runge%E2%80%93Kutta_methods
 - [https://en.wikipedia.org/wiki/Euler%27s_equations_\(rigid_body_dynamics\)](https://en.wikipedia.org/wiki/Euler%27s_equations_(rigid_body_dynamics))



From Lecture 3, we have that the transformation of a vector from a coordinate system to a coordinate system x'_1, x'_2, x'_3 is given by

$$\begin{pmatrix} H'_1 \\ H'_2 \\ H'_3 \end{pmatrix} = \begin{pmatrix} \mathbf{i}'_1 \cdot \mathbf{i}_1 & \mathbf{i}'_1 \cdot \mathbf{i}_2 & \mathbf{i}'_1 \cdot \mathbf{i}_3 \\ \mathbf{i}'_2 \cdot \mathbf{i}_1 & \mathbf{i}'_2 \cdot \mathbf{i}_2 & \mathbf{i}'_2 \cdot \mathbf{i}_3 \\ \mathbf{i}'_3 \cdot \mathbf{i}_1 & \mathbf{i}'_3 \cdot \mathbf{i}_2 & \mathbf{i}'_3 \cdot \mathbf{i}_3 \end{pmatrix} \begin{pmatrix} H_1 \\ H_2 \\ H_3 \end{pmatrix} = [T] \begin{pmatrix} H_1 \\ H_2 \\ H_3 \end{pmatrix} .$$

where we have introduced the symbol $[T]$ for the transformation matrix.