Expt. No	Page No
	Experiment No> 1
* Objective	
To determine the refrac	tive index of a material of the prism gth of light ($\lambda = 5893A$).
* Apparotus - [INDOSAW SH	(024 or skooj
* formula used	
Refractive index of pris	sm material is given as
lu = sin (A+Sm)/2
	A/2
when A %c the	anale of hulism
	of minimum deviation.
Imreduction	
An optical instrument w	sed to measure the wavelengths of lights
	sources is called spectoometer.
Spectrometer depends or	a phism on a groting, to separate t
	can be done with a spectrometer.
1 11 0	
The instrument consist	rs of a collimator, a prism table an
a telescope.	
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Date

> Paism Angle Angle of Incident Ray Paism Base in solut minimpal making to styles son in A , session Sm is the made of maximum definition. to before here are residence to have manuscraft to the rewhere y telled it is the reput with an and the terms you or police of the standard of the standard of the standard of or the resemble service and problems in which are come right the will make the works and a property of the state of

	Date
Expt. No.	Page No. 2
* Theomy	
Refraction through	a prism
The angle A betwoen the control of t	monochromatic light passes through a prism, it is ce as it enters and again as it leaves the prismeen the two surfaces where refraction takes and Angel. The intersection of the two refractions are Refractions of the two refractions.
Base is that side	ne Refracting Edge of the Prism. The Prism of the beam which is opposite to prism
N, and N2 are no between the incide of deviation. For a with the variation depends	owmals to the prism faces. The angle 8 ent and emergent way is defined as the angle of deviation varies in angle of incidence. The particular angle ands on the prism angle, the index of prism at that wavelength and the angle of
to the angle of it	minimum, if the angle of emergence is equal naidence. This condition is found using a sen the puism spectrameter is set at minimum given wavelength, we have
Ш	$= \frac{\sin(A+8m)/2}{\sin A/2}$
where, A is the	pe prism angle 8m is the angle of deviation for a particular wavelength.
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* OBSERVATIONS

- Table for angle of prism'A! [L.c - 1/180]

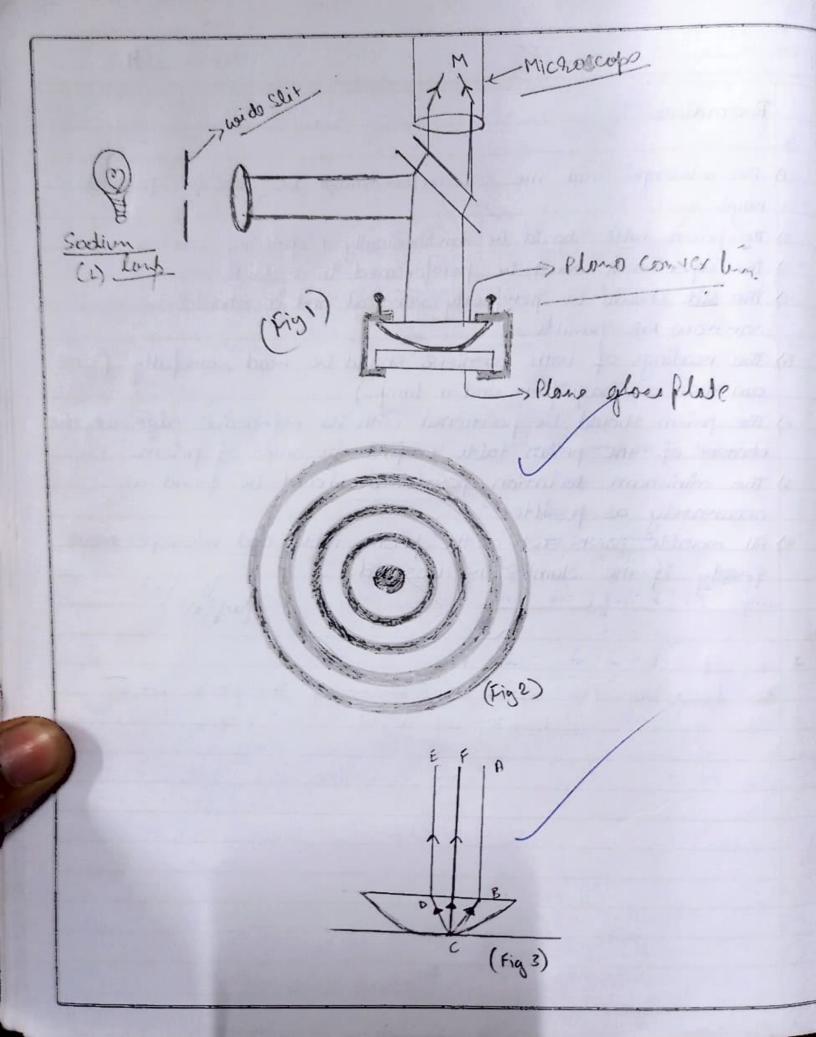
s. No. Vernice		1			Difference Mean of and=2A Value 2A	A(')
	r grade	· (.)	12 (°)			
		143.055	201.666	121.528	121.430	60.715
	FoceI		203.116	121.556	121.556	60.718

- Table for angle of minimum deviation Sm

cho	Vernier	Telescop Rea	ding	Difference	Meanvalue
3.1.0	TO THE	Minderiotion (a)	Direct image	a'-b'=om	o) Sn.
1.	V, V2	128.338	79.00	49.38	49.00
2.	V ₁ V ₂	30.333	79.00	48.667	((.50
	V ₂	•			

	Date
Expt. No	Page No
* Calculations	
Formula :-	
Sin (A + Sm)	
2	
$\mu \Rightarrow$	
Sin A	
2	
9 Sin (49 + 60, 746)	
2	
W 7	
Sin (60.746)	
2	
M => 1.617	
* Result:	
The sell-active index o	the brism for
the given wowelength of light i	c (1 (17)
The state of the s	- (100(f) o
	/
	050700355
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Date
Expt. No Page No
Precautions:
1) The telescope and the collimator must be set for parallel mays.
2) The prism table should be mechanically & optically leveled.
3) The experiment must be performed in a dark moon.
4) The slit should be perfectly vertical and it should be made narrow as possible.
5) The readings of both verniers should be read carefully. (one can use a magnifier and a lamp.)
6) The phism should be positioned with its refractive edge at the
center of the prism table to find the angle of prism.
The minimum deviation positioned should be found as
accurately as possible.
8) All movable parts such as the prism table and telescope move
freely if the clamp are loosened.
Juga 1/2/2
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Date	
Expt. No Page No.	
F+ periment No>2	
* AIM?-	
To determine the wavelength of Sodium light by newton	's rings
* Apparatus Required :-	
INDOSAW SKOOP.	
* Theory:- light form a monochromatic source (sodium lamp) is all	
a money nearly powallel beam.	ice it int
A film is said to be thin when its thickness about the	e order
Rings are fringes of equal thickness. They are absenved wis neflected from a plano - convex lens of a long focal lend	ngth placed
in contact with a plane glass plate. Then other film is foul the plate and the near. The thickness of air film varies	ed between
point of contact to some value t when the system is Il	betorion
with monochromatic light. Consecutive Bright and dark frin	igs are
The wavelength of monochromatic length can be determine	ned as
= Dm+2n - Dm2	
4xnxw	
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* Obcavation Table -

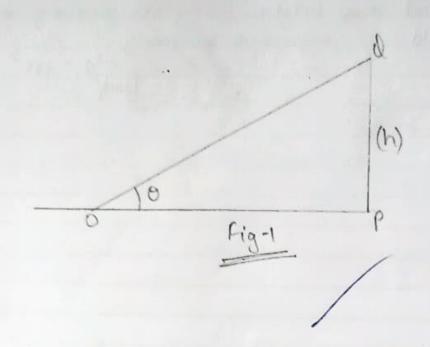
oWS	No.	(a)	Aigus (6)	Diamellon (a-b)	1 Cm	Dn m2 Dn2	hear
1.	1S 12	080.1	0.683	0.397	0.157	0.031	
3.	6	1.030	0.723	0.307	0.094	0.35	0,0325
.2	3	1.964		0.167	0.027	0.017	

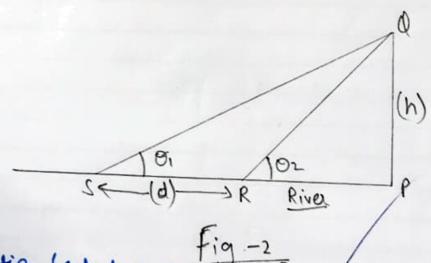
Calculations:
$$\frac{1}{6}$$
 - $\lambda = \frac{1}{6} \frac{1}{MP} - \frac{1}{5} \frac{1}{100}$
 $= \frac{0.325}{4 \times 3 \times 100}$
 $= 2.708 \times 10^{-5} \text{ cm}$

° \ \ = 2708 A°

	Date
Expt. No	Page No
of mth wing. Ris the warius of Curvature of	
* Procedure: - After the experimental aurangement, the glass on angle of 45° to the horizonate the glass plans from the source vertically downwards and for convex lens.	ate reflects lights
Newton's rings are seen using a long focus microscope is the film. The microscope is the first ring on the left side of the center main scale and Vortice scale of the microsco	s made tangently to whe he weadings of the
Then it is seen through the right side from observation are taken. Similarly readings from are taken from both left and wight side. The diameter of rings is found out by subtilled and right side. The square of diameter to make found out. Wavelength for and then mean wavelength is found out.	racting readings on the
* Result 2- wavelength of the given light source is * Are courtions 2-	
The lens used should be of large radius Tracker's Signa	of wavelength. The

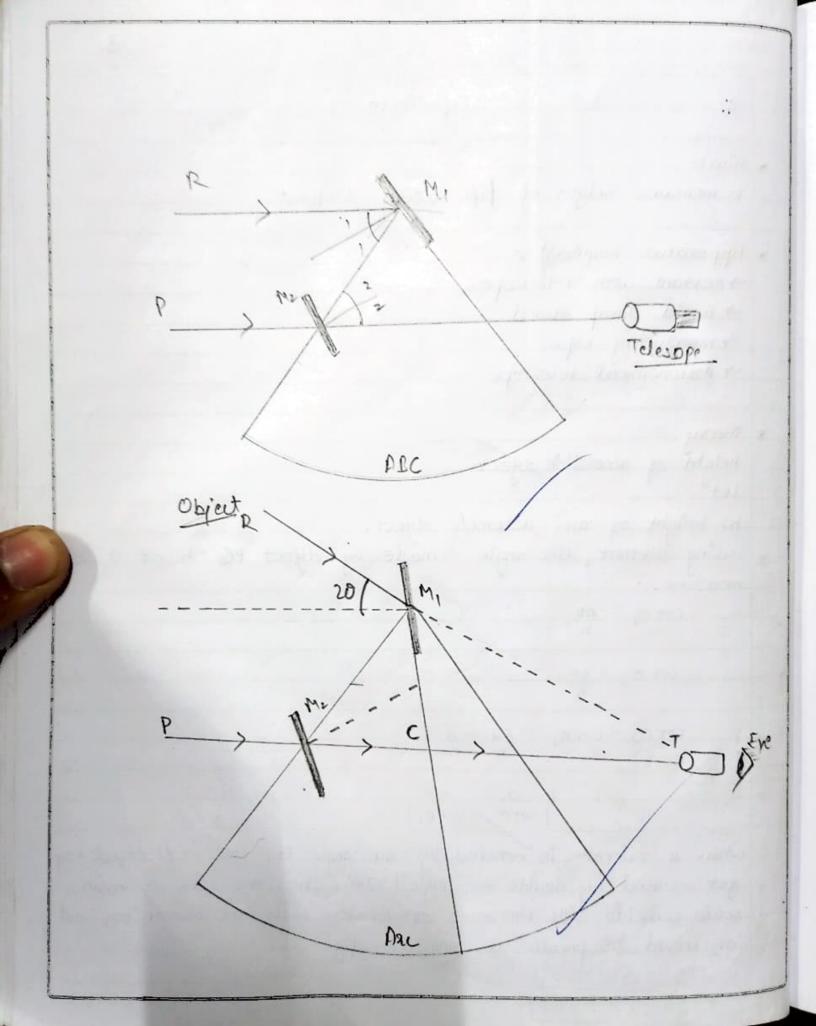
	Date	
	Page No	. 7
used should be	an extended one	wire
be measured o	Curately.	
	used should be on a bright use we measured a	



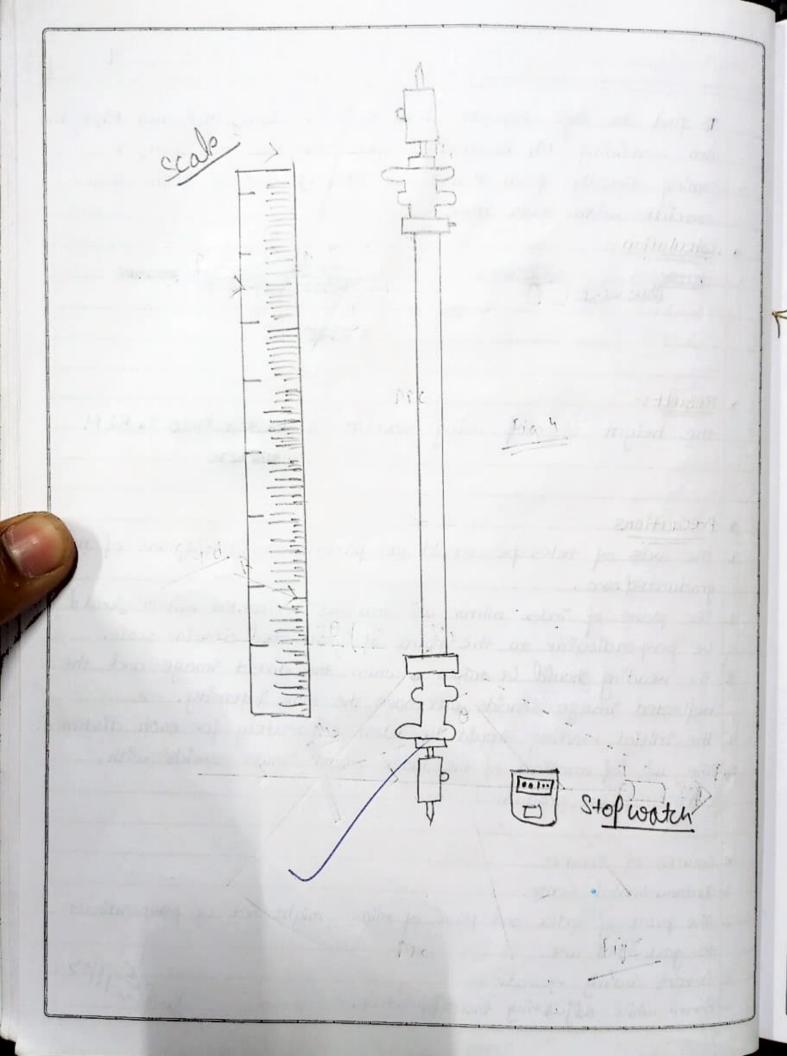


* Observation (calculation) A, = 91 8/60 + 1/300

	Date
Expt. No	Page No 8
Experiment -	3
* Aims-	
To measure hieight of Pipe using se	extant.
* Apparatus Required :-	
-> sextant with tellescope	
- Regid clamp stand	
- measuring take	
-> Astronomical telescope	
* Theory	
* Theory height of accessible object	
Let Let	
his height of any accemble object.	
his height of any accemble object. using sextant, the angle made by measure.	object PQ = h at 0 is
Cot O; = RP.	
Cot O2 = SP	
$\frac{\cot \sigma_2 - \cot \sigma_1 = sR = d}{h}$	
h=d co+oz co+o,	
when a mirrow is notated by an a	ngle 0, the neflectced vay
get wooted by double the angle (=20)	The state of the s
scale coincide with the zero of vernis	
me should be parallel as shown in f?	
Too	chor's Signature



	Date	
Expt. No.	Page No.	9
aim containing M1 is rotated such coming directly from P and ray BM concide with each other. * Calculation:	that the ray Pm2 T	
* Result: the height of roof using sextant * Precautions	is - 1.5.	2 M
1. The axis of telescope should be po- graduated are. 2. The plane of index mirror as well		•
s The reading should be taken a when reflected image concide and have the	graduated circular scale	e.
5. The axis of rotation of the index mentre of the graduated arc.	sapawately for each a	distance h
Sources of Error: To strumental Error. The plane of index and plane of minor to graduated anc. Tryons during recording	No.	
4. Error while adjusting the approxue	eacher's Signature	



	Date
Expt. No.	Page No
	Experiment - 4
* Arms- Determine the vale	u of 'g'.
* Apparatus =	THE RESERVE THE PARTY OF THE PA
Compound Pendulum,	Knipe edge, Meter Scale, Stopwate
	May 1
* Theory :-	
	es from the centre of gravity of the
pendulums from Ki and K2 co	ovvesponding to the time periods Ti or
T ₂ . Then	
$T_1 = 2\pi \sqrt{\frac{L_1^2 + K^2}{L_1 g}}$	(1)
-19	
2/2	
$T_2 = 2 \pi \sqrt{\frac{12 + K^2}{L_2 q}}$	(2)
	that obtained for a simple pendul
g T = 2K J g g. In fact a s	simple pendulum of length $L = K^2 + L$
and the court time b	eriod T. Such a pendulum is called
as equivalent simple pendulum	
where K = madius of gyration	
Subtract 1 from 2	
$\int \frac{1}{1} \int \frac{1}{1} + $	17
2 L1+L2 L1-L2	
$q = 8\pi^2$	(4)
$T_1^2 + T_2^2 - T_1^2 - T_2^2$	
L1+L2 L1-L2	Teacher's Signature

SANO	Knill gall	T. J.	7 = 5	T2 fc	T2 = 41
0	86.4	\$5.35	1.845	55.53	1.851
0	82.4	55.11	1.830	55.86	1.862
3	80.4	54.4	1.813	54.52	1.817

Date	
Expt. NoPage No	11
where L, and L2 can be made to have large difference suitably adjusting the masses M and m.	Lug
with a little efforts T_1 and T_2 can be made near by so the term $\left(\begin{array}{c} T_1^2 - T_2^2 \\ L_1 - L_2 \end{array}\right)$ can be neglected.	equal.
so formula which can used for calculating acc. due to	gravity
$\frac{8 \times^{2}}{(T_{1}^{2} + T_{2}^{2})}$ $g = L_{1} + L_{2}$	5)
2. Amplitude of escillations should be small. 3. The time period about K, and K2 should be nearly equ	حم امر
4. (L1+L2) must be measured accurately. 5. Avoid wind, for etc at the place of experiment.	
Result ?- The value of acceleration due to gravity "g" i	s 9.71 m
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Teacher's Signature	

Page No	12
ъ	
lases light so	ource using
n Neon bases	
	0
enter point on M and enter	n beam 5
of the Beam oce between ofference than few paths to	spilitten the two
	is calculated the shift of the shift of Mand enter point on at the Beam of the

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                       a anche e a mana, a anana - a ananti, a caras -
                         मामान माववात , विश्ववाद्य का विवादाय का विवादाय , विवादाय का अनिवादाय का अनुसार्थ म
                                                               A HEALE I A BARANT I A BARANTI I BARANZA I D. COODE I D. COODE
                                                                                                                                                                                                                                                                                                                                             = 674 (approx
```

function's highestern

* Observation : -

Initial (L1) am	(No) on friggs	final Reading	Distand L2-6, Cm	1 = 2d/2
0.45722 0.45840 0.45911 0.4500 0.46129 0.46190 0.46250	20 20 20 20 20 20 20 20 20 20 20 20 20 2	0.45940 0.45911 0.46000 0.46129 0.46129 0.46250 0.46309	0.00018 0.00099 0.000129 0.00061	0.0000.00 0.0000.00 0.0000.00 0.0000.00 0.0000.00

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* Result 4The wavelength of monochwamatic light source is 674 nm.

* Percentage Error "-

Standard value = 632 nm

(alculated = 674 nm

4 errar = 632 - 674 x 100 = 6.6%

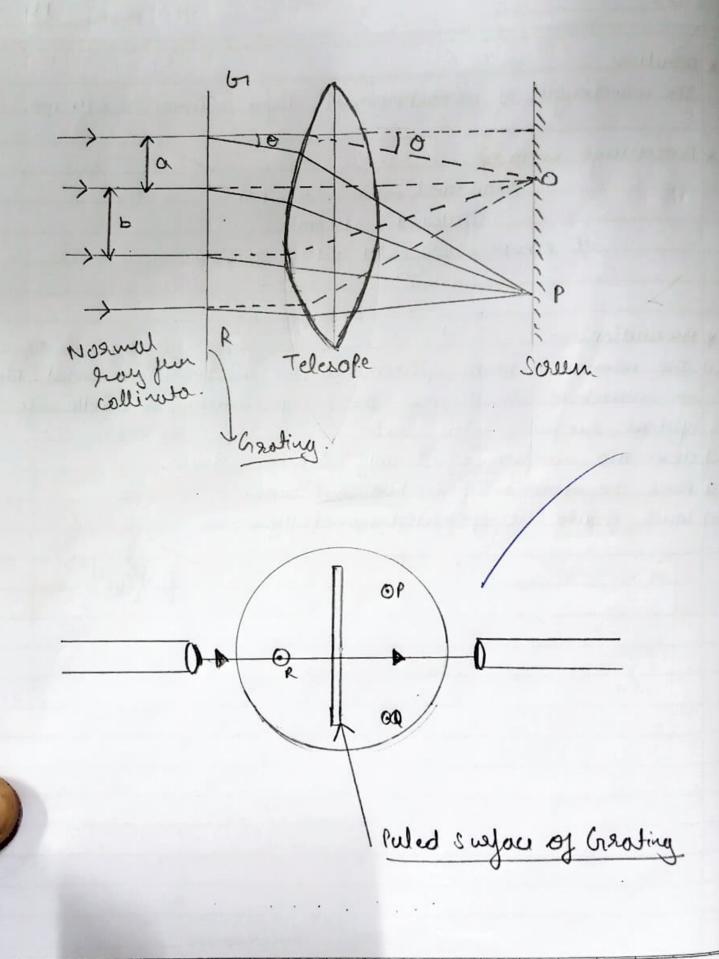
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* Pre cautions :-

- (1) The mirror & beam splitter surface are precision coated. Dirt or scratcher will distract the fringe. Pattern, so handle all optical surface with care.
- (2) clean the surface occasionally with lens tissue.
- (3) Place the whole setup in dust tree zone.
- (4) laws should be openated carefully.

Durge 611/13

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		Date	
Expt. No.		Page No	12
	Expeniment	No. 46	
* Aim :-			
To find the wo	welength of differen	nt colours of light	using
H	9.		
* Material Require	-s b	101	
Spectrometer.	grating element		
	1-1 1		
* Formula Used:		1000	
$(a+b)$ $\sin \theta = n$	\	III II relation	
where,	- 100 mm		
150	oo element The	each grating in grati	ng
	15000 lines pe		3
λ is	wavelength of light		
h is	onder of spectrum		
	angle of deviation.		
* Theory 3-			
it in gigune o re	presents the central	maxima due to u	n- 21441
mays. As the	width of the slit in	s comparable with +	he
to deposits of	ight, the diffwacted	beam produced by	1 arat
is focused by	lens in its focal	plane, to produce s	secondo
	of nth principle ma		
la direction	+b) sin 0 = n)	is given by	
when,	, , , , , , , , , , , , , , , , , , ,		
	= ±1,±2,±3,		
		acher's Signature	