

# Weekly Report

LIU Honghao

Jan 27 2021

Information about input, output and questions.

## Input

### .ms folder

The input data is a .ms folder including information about a saved in tables.

## Visibility

In the code, visibility is a class with objects like uvw, vis, frequency, time, flag, weight, channel\_bandwidth, antenna\_1, antenna\_2 and image\_weight. See Table 1.

Object	Type	Shape	Description
uvw	numpy array	(n,3)	wavelength information in three different axes
vis	complex array	(n,1)	value of visibility
frequency	numpy array	(n,1)	?
time	numpy array	(n,1)	time of observation related to earth rotation
flag	numpy array	(n,1)	?
weight	numpy array	(n,1)	weight of each visibility?
image_weight	numpy array	(n,1)	?
channel_bandwidth	numpy array	(n,)	maximum frequency?
antenna_1	numpy array	(n,1)	index of first antenna
antenna_2	numpy array	(n,1)	index of second antenna
integration time	numpy array	(n,)	exposure time <sup>1</sup>

Table 1: Objects

channel\_bandwidth - number of times of detection in integration time?

## Output

### Image

The output are images saved in numpy or astropy.wcs format (frequency, polarisation, latitude, longitude) in angular coordinate.

---

<sup>1</sup>An observation is done by "integrating" over some integration time. i.e. average the power over time  $\tau$ .

## Questions

1. Why do different frequencies or channels have different images?
2. What is the content in the .ms file?
3. Questions related to Visibility class.