Computer-predicted ionization energy of carbon in agreement with experiment to within 1 cm^{-1}

Nike Dattani[*]

Oxford University, Hertford College, OX1 3BW, Oxford, UK, and
Kyoto University, Fukui Institute for Fundamental Chemistry, 606-8103, Kyoto, Japan,

Giovanni LiManni[†]

Department of Electronic Structure Theory, Max Planck Institute for Solid State Systems, Suttgart, Germany.

David Feller[‡]

Department of Chemistry, Washington State University, Pullman, Washington 99164-4630, USA,

Jacek Koput[§]

Department of Chemistry, Adam Mickiewicz University, 60–780 Poznan, Poland.
(Dated: 7th June 2017)

We show that with non-explicitly correlated basis sets, we can predict the first ionization energy in carbon to within 1 cm^{-1} of the experimental value. This is an improvement of more than a factor of 5 over the preceding best prediction in [Phys. Rev. A **81**, 022503] and comparable to the accuracies achieved using explicitly correlated wavefunctions for the smaller atoms up to boron.

In the last seven years, ionization energies (IEs) have been calculated with unprecedented precision for the Li atom [1], Be atom [2] and B atom [3]. Tight variational bounds for non-relativistic ground state energies assuming a clamped, point-sized nucleus have reached 49 digits in units of Hartree for He, 19 digits for Li, 12 digits for Be, and 11 digits for B (see Table I). Calculated IEs have been made in agreement with experiment to within 10^{-3} cm^{-1} for ^{7}Li, 10^{-1} cm^{-1} for ^{9}Be and 1 cm^{-1} for ^{11}B (see Table II).

For the C atom, before this present study, no high-precision calculation had been reported to predict an IE to within 1 cm^{-1} agreement with experiment. Table I shows that the method used for the calculation has a crucial impact on the precision obtainable: the best variational bound on the non-relativistic, clamped, point-nucleus (NR,CPN), ground state energy of Li is believed to be 8 orders of magnitude more accurate than that of the Be atom, because analytic expressions for the multi-center integrals of Hylleraas wavefunctions are only known for up to 3e$^-$, and it would be very slow to calculate them numerically for 4e$^-$. For 4e$^-$ and greater, ECGs (**e**xplicitly **c**orrelated **G**aussian wavefunction**s**) are used. ECGs mimic the theoretical shape of the wavefunction less accurately (e.g. they are rounded rather than cusp-like in the limit of vanishing electron-electron distance, and in the infinite electron-electron distance limit they decay as $1/r^2$ rather than $1/r$), so more parameters are needed than when using Hylleraas wavefunctions. However, the integrals of ECGs are known analytically, regardless of the number of electrons, which is why it has

been possible to do very high-precision calculations on Be [2] and B [3].

Table I demonstrates though, that with twice as many variationally optimizable parameters, the authors obtained convergence in 1 fewer digit for the B atom than for the Be atom. This suggests that it would be very difficult to use the ECG approach for atoms and molecules coming from most of the periodic table. The best known variational bound for the NR,CPN ground state energy for C was calculated in 2015 using fixed-node diffusion Monte Carlo (FN-DMC) with the nodes of the electronic wavefunction fixed at the locations of a CISD/cc-pV5Z wavefunction, and the statistical uncertainty on the Monte Carlo was ± 20 μE_{Hartree} [4]. However, the IE for C predicted by FN-DMC was in discrepancy with experiment by more than 40 cm^{-1}. In this paper, the approach we use to calculate the NR,CPN energy of the ground state of C is FCIQMC (full configuration interaction quantum Monte Carlo) in an aug-cc-pCV7Z basis set. Table I shows that our NR,CPN energy is at least 210 μE_h higher than the variational upper bound obtained from FN-DMC; but since in our approach, imperfections in the description of the wavefunction for the neutral atom are almost the same as in the cation, the errors are almost completely eliminated when taking the energy difference, therefore bringing us to agreement with experiment that is comparable to what has been seen with the explicitly correlated approach for atoms as big as (but not exceeding) boron.

After adding relativistic and quantum electrodynamics (QED) corrections, and corrections to the clamped nucleus approximation, we obtained an IE for the ground sate of C which is within 1 cm^{-1} disagreement with the best known experimental estimate. While this is not as impressive as the method of variationally optimizing parameters in an explicitly correlated wavefunciton ansatz has proven to be for Li and Be, the disagreement has

---

[*] dattani.nike@hertford.ox.ac.uk
[†] g.limanni@fkf.mpg.de
[‡] dfeller@owt.com
[§] koput@amu.edu.pl

the same order of magnitude as the latter approach for B (see Table II). We finally note that the approach used in this paper, of calculating FCIQMC on a basis set of non-explicitly correlated orbitals has successfully treated systems with far more electrons (transition metal atoms [5], diatomics [6], larger molecules such as butadiene [7], and even solid state systems [8]), so it is conceivable that the approach used in this paper may in the near future be able to determine (with fair accuracy) the IEs which at present remain experimentally elusive or poorly known. These include arsenic (whose experimental IE has an uncertainty of $\pm 2\,\mathrm{cm}^{-1}$), Pm, Pa, Fm, Md, No, Sg, Bh and Hs (whose IEs are only known based on extrapolations of other experimental data and have uncertainties between $\pm 140\,\mathrm{cm}^{-1}$ and $\pm 4000\,\mathrm{cm}^{-1}$), Rf and Db (whose IEs are only known from theoretical calculations), and Mt, Ds, Rg, Cn, Nh, Fl, Mc, Lv, Ts, and Og (for which no IE is given in NIST's most recent databases).

I. METHODOLOGY

Our calculations in this paper can be divided into five stages: (A) We develop larger core-valence (CV) basis sets than previously available for carbon, (B) we calculate the 1- and 2-electron integrals in these basis sets, (C) we attempt to solve the NR,CPN Schroedinger equation at the FCI level in finite basis sets, (D) we extrapolate the finite basis set results to estimate the energies at the complete basis set (CBS) limtis, and (E) we estimate the size of corrections due to special relativity, QED, and due to the atom having an unclamed nucleus.

A. Optimization of 'tight function' exponents for the aug-cc-pCV7Z and aug-cc-pCV8Z basis sets.

The largest basis sets known for C prior to this work were the (aug-cc-pVXZ, X =7,8,9) sets used by Feller in 2016 [16]. These basis sets did not contain 'tight' exponent functions for capturing the effects of the correlation between the core ($1s^2$, $2s^2$) electrons and the valence electrons ($2p^2$). The largest known basis set for carbon prior to this work including the CV (core-valence) correction was the aug-cc-pCV6Z [17] set. In this work we start by optimizing the 'tight' exponents for the CV correction to Feller's 2016 aug-cc-pV7Z and aug-cc-pV8Z basis sets, yielding the first aug-cc-pCV7Z basis set for carbon, and the first aug-cc-pCV8Z basis set known.

The final aug-cc-pCVXZ basis sets have X new tight functions of s-type, $X-1$ of p-type, $X-2$ of d-type, and so forth, up to the final i-type function for $X = 7$ and the final k-type function for $X = 8$. The ith exponent corresponding to a function of type L is named $\gamma_{X,L,i}$, and is assumed to follow an "even-tempered" model: $\gamma_{X,L,i} = \alpha_{X,L,i}\beta_{X,L,i}^{i-1}$.

In the non-linear optimization procedure to obtain $\alpha_{7,L,i}$ and $\beta_{7,L,i}$, the starting values were chosen to be the $\alpha_{6,L,i}$ and $\beta_{6,L,i}$ values that were already optimized

in [17]. These were then treated as free parameters to minimize the difference between the frozen core and all-electron CISD energies of the carbon atom with all other exponent functions fixed. The `MOLPRO` program [18] was used to calculate the CISD energies, and the `L-BFGS-B` program of [19] was used to optimize the free parameters. The s-type functions were added first, then once they were optimized they were held fixed while the p-type functions were added and optimized. Then both the s- and p-type functions were held fixed while the d-type functions were added, and so on up to the single i-type function. The procedure for $X = 8$ was the same, except the procedure continued to k-type functions, and the starting values came from the newly optimized $X = 7$ case rather than the $X = 6$ case from [17].

The tight exponents optimized in this work for aug-cc-pCV6Z and aug-cc-pCV7Z are presented in the Supplemental Material.

B. Calculation of 1- and 2-electron integrals including k- and l- functions

The calculation of the 1- and 2-electron integrals for (aug)-cc-p(C)VXZ basis sets with $X \geq 7$ is not possible with most quantum chemistry packages, since very few software packages support k- and l- functions, but for first row elements, k-functions appear in $X = 7$ basis sets and l-functions appear when $X = 8$. To calculate these integrals, we have used a locally modified version of Molcas 8.0 [] which supports larger basis sets. The MO-integrals were evaluated in the basis of the optimized CASSCF(6,5) and CASSCF(5,5) molecular orbitals (MOs), with the five active orbitals being the $1s$, $2s$, $2p_x$, $2p_y$ and $2p_z$ of the C atom. This active space is the minimal active space including all electrons, that is able to provide balanced orbitals for the three degenerate states of the ^3P state of the C atom

C. Calculation of NR-CPN energies in finite basis sets without truncating the possible excitation levels (with FCI and FCIQMC)

Deterministic FCI calculations were possible for the 5e$^-$ system C$^+$ in aug-cc-pCV6Z with about $130\,\mathrm{GB}$ of RAM to store the Davidson vectors, but were not possible for the 6e$^-$ neutral C atom in this basis set (which would require almost $9\,\mathrm{TB}$ of RAM), nor for 5e$^-$ in the aug-cc-pCV7Z basis set (which would require almost $55\,\mathrm{TB}$ of RAM).

We therefore use FCIQMC for all other NR-CPN calculations. The wavefunction of the relevant system is expanded as a sum of all possible Slater determinants given the basis set (full configuration interaction), and the coefficients are determined by the number of walkers standing on each determinant after a Monte Carlo sampling using the NR-CPN Hamiltonian. The method was

Table I. Tight variational upper bounds for atomic energies. For H, the listed energy is exact, since the Hartree energy unit is defined as twice this energy, and can be related to fundamental SI units through an analytic solution to the relavent Schroedinger equation. VO stands for variational optimization (parameters in a wavefunction ansatz are optimized in attempt to obtain the lowest possible energy). Hylleraas-Log indicates the use of Hylleraas functions supplemented with auxiliary log functions, and ECG(N) indicates the use of **e**xplicitly **c**orrelated **G**aussians with N variationally optimizable parameters. FN-DMC stands for fixed node diffusion Monte Carlo.

	Total non-relativistic clamped point-nucleus (NR,CPN) energy [Hartree]	Method/Ansatz type	Reference	
H 1	-0.5	Analytic	1926 Schroedinger	
He 2	-2.903 724 377 034 119 598 311 159 245 194 404 446 696 925 309 838	VO/Hylleraas-Log	2006 Schwartz	[9]
Li 3	-7.478 060 323 910 134 843	VO/Hylleraas	2011 Wang	[10]
Be 4	-14.667 356 494 9	VO/ECG(4096)	2013 Puchalski	[2]
B 5	-24.653 867 537	VO/ECG(8192)	2015 Puchalski	[3]
C 6	-37.844 48(2)	FN-DMC	2015 Yang	[4]
C 6	-37.844 251 5(08)	FCIQMC/aCV7Z	Present work	-
C 6	−37.843 333	VO/ECG(1000)	2013 Bubin	[11]

Table II. The most precisely calculated electronic excitation energies for the first 6 atoms, compared to the best known experimental measurements to date. The last column indicates that if aiming for the best precision, an experimental measurement is still the best way to obtain the energy for most atoms, but for Be, the energy has been obtained more precisely *in silico* than in any experiment to date. The value for carbon of 90 821.01 cm^{-1} was calculated in the present work.

	Transition	Experiment [cm^{-1}]		Theory [cm^{-1}]		Calc - Obs [cm^{-1}]	$\left|\frac{\text{Calc - Obs}}{\text{Uncertainty in obs}}\right|$	More precise
H								
He								
^{7}Li	$\text{Li}^+\left(1^1S\right) \leftarrow \text{Li}\left(2^2S\right)$	43 487.159 40(18)	[12]	43 487.159 0(8)	[13]	-0.000 4	57.1	Experiment
^{9}Be	$\text{Be}^+\left(2^2S\right) \leftarrow \text{Be}\left(2^1S\right)$	75 192.64(6)	[14]	75 192.699(7)	[2]	0.059	0.98	Theory
^{11}B	$\text{B}^+\left(2^1S\right) \leftarrow \text{B}\left(2^2P\right)$	66 928.036(22)	[15]	66 927.91(21)	[3]	-0.126	5.7	Experiment
C	$\text{C}^+\left(2^2P\right) \leftarrow \text{C}\left(2^3P\right)$	90 820.45(10)		90 821.01		0.31	3.1	Experiment

introduced in [20], and we use the initiator method first described in [21], and the semi-stochastic method as described in [22]. The calculations are performed using the developer version of the software `NECI`.

Within a given Hamiltonian (in this case the NR-CPN Hamiltonian) and basis set, there are three sources of error in the FCIQMC energy calculations:

1. Trial wavefunction error (ΔE_{trial}), which approaches zero in the limit where the number of determinants used in the trial wavefunction approaches the number of determinants in the FCIQMC wavefunction;

2. Initiator error ($\Delta E_{\text{initiator}}$), which approaches zero in the limit where the number of walkers N_{walkers} gets sufficiently large; and

3. Stochastic error (ΔE_{stoch}), which for a given number of walkers is estimated as the square root of the unbiased variance among different estimates E_i of the energy from their mean $\bar{E}$ after different

numbers N of Monte Carlo macro-iterations (determined using the Flyvbjerg-Petersen blocking analysis [23]) after the walkers have reached equilibrium:

$$\Delta E_{\text{stochastic}} \approx \sqrt{\frac{\sum_{i=1}^N \left(E_i - \bar{E}\right)^2}{N-1}} = \mathcal{O}\left(1/\sqrt{N}\right).$$

Our goal was to obtain all energies to a precision of $\pm\epsilon$ where $\epsilon \leq 1 \mu E_{\text{Hartree}} \approx 0.2$ cm^{-1} (within the aug-cc-pCV7Z basis set used). To ensure that $\Delta E_{\text{initiator}}$ can be neglected, we used a sufficiently large value of N_{walkers} for every energy calculation, so that the energy difference between using N_{walkers} and $\frac{1}{2}N_{\text{walkers}}$ was smaller than $1 \mu E_{\text{Hartree}}$. Likewise, to ensure that ΔE_{trial} can be neglected, we used a sufficiently large number of determinants in the trial wavefunction for every energy calculation, such that ΔE_{trial} would be also be smaller than $1 \mu E_{\text{Hartree}}$. We then ran every calculation for enough macro-iterations N such that ΔE_{stoch} was smaller than ΔE_{trial} and $\Delta E_{\text{initiator}}$.

D. Extrapolations to the complete basis set limit

We use two different families of formulas to extrapolate the correlation energies (FCIQMC energies with the ROHF energies subtracted out) from E_{X-1} and E_X to E_{CBS}:

$$E_{\mathrm{CBS}} = E_X - \frac{A}{X^n} \ , \qquad (1)$$

$$E_{\mathrm{CBS}} = E_X - \frac{A}{(X + 1/2)^n} \ . \qquad (2)$$

If we set $n = 3$ in Eq.(1), we recover the formula originally proposed in [24]. If we set $n = 4$ in Eq. (2), we recover the formula originally proposed in [25]. If we have values for E_X at two different X values, we can eliminate A in both cases, so Eq.(1) leads to Eq.(3) and Eq.(2) leads to Eq.(4):

$$E_{\mathrm{CBS}} = \frac{X^n E_X - (X-1)^n E_{X-1}}{X^n - (X-1)^n} \ , \qquad (3)$$

$$E_{\mathrm{CBS}} = E_X + \frac{(2X-1)^n (E_X - E_{X-1})}{(1+2X)^n - (2X-1)^n} \ . \qquad (4)$$

The values of E_{CBS} obtained from using $n = 3$ in Eq.(3) and $n = 4$ in Eq.4, for $X = 8$ were added to the ROHF energies for $X = 8$ and are presented in Table IV. As explained on page 5 of [16], using $n = 3$ in Eq.(1) tends to over-estimate the extrapolation (as we also observe in this present work, see Table IV). The value of $n = 3.5$ was therefore used in [16], and we have added a row in Table IV corresponding to this value. The final NR-CPN energy was taken as the mean of the three values obtained from extrapolating the correlation energy and adding it to the ROHF energy with $X = 8$, and the uncertainty in this mean was estimated as the unbiased variance of the three values.

E. Estimation of relativistic, QED, and finite nuclear mass corrections corrections

Scalar relativistic corrections were calculated by comparing the energies using the spin-free version of the 1e$^-$ X2C (exact 2-component) Hamiltonian, to the energies of the NR-CPN Hamiltonian. The spin-free 1e$^-$ version of the X2C Hamiltonian is equivalent to infinite-order DKH. The integrals of our X2C Hamiltonian with ROHF orbitals were done in the CFOUR program, and were calculated at various levels of coupled cluster theory with the MRCC program [26].

Further scalar relativistic effects were included by adding the Breit and QED corrections (including the vacuum polarization and the self-energy terms that together comprise the Lamb shift) from the state-averaged Dirac-Fock calculations done in [27]. The overall contribution to the IE for C was -0.48 cm^{-1}.

Table III. Basis set and correlation convergence of the X2C and DBOC corrections to the C(^3P) →C$^+$ (^2P) ionization energy in cm^{-1}. The difference between CCSDT and FCI for the smallest two basis sets is denoted by $\Delta_{\mathrm{correlaton}}$ and the values at aCV3Z are added to the CCSDT energies for all larger basis sets (resulting in FCI estimates presented in *italic* font). The numbers in parentheses for the CBS (complete basis set) estimates denote our assigned uncertainties in the last digits presented, which we believe to be very conservative. aCVXZ is short for aug-cc-pCVXZ-uncontracted.

	aCV2Z	aCV3Z	aCV4Z	aCV5Z	aCV6Z	CBS
			X2C			
CCSDT	-31.055	-30.144	-30.037	-30.004	-30.001	-29.999(003)
FCI	-31.074	-30.168	*-30.060*	*-30.028*	*-30.025*	*-30.023(050)*
$\Delta_{\mathrm{correlation}}$	-0.019	-0.024	*-0.024*	*-0.024*	*-0.024*	*-0.024(050)*
			DBOC			
CCSDT	-0.5	-0.5	-0.5	-0.5	-0.5	-0.5
FCI	-0.5	-0.5	*-0.5*	*-0.5*	*-0.5*	*-0.5*
$\Delta_{\mathrm{correlation}}$	0.0	0.0	*0.0*	*0.0*	*0.0*	*0.0*

Diagonal Born-Oppenheimer breakdown corrections (DBOC) to the clamped nucleus approximation were calculated using CFOUR, with MRCC used for the coupled-cluster part. Our value of -0.50 cm^{-1} is more than double the value of -0.08 cm^{-1} estimated in [27], due to including higher levels of correlation.

The basis set and correlation convergence of th X2C and DBOC corrections is shown in Table III

II. CONCLUSION

Table IV summarizes the various contributions to our value of the IE, and compares our final value to experiment and to three recent theoretical estimates. Our value is 0.01 cm^{-1} larger than the best experimental value, but the spread of ionization energies we get using the three different extrapolation formulas is almost two orders of magnitude larger than this, so calculations with an aug-cc-pCV9Z basis set might be necessary to confirm this seemingly marvelous agreement.

The best theoretical estimate of the IE before this work was in [27], which we believe to be mainly due to the approximations inherent to the F12 approach used for their NR-CPN energy, and the perturbative nature of their scalar relativistic corrections (the mass velocity and Darwin terms, rather than the X2C calculation used in the present work) and due to calculating the DBOC correction at CCSD level (as opposed to CCSDTQ).

Table IV. Summary

N_{walkers}	N_{trial}	Uncertainty	$E = \dfrac{\langle \psi_{N_{\text{trial}}}|H_{\text{NR−CPN}}|\psi_{\text{FCIQMC}}\rangle}{\langle \psi_{N_{\text{trial}}}|\psi_{\text{FCIQMC}}\rangle}$		$E_{\text{excitation}}$
			2^3P [E_{hartree}]	2^1D [E_{hartree}]	$2^3P \to 2^1D$ [cm^{-1}]
			aug-cc-pCV6Z		
64×10^6	1		-37.843 839 6(18)	-37.	
	1000		-37.843 840 5(05)		
		ΔE_{trial}	$< 0.000\ 001\ 1$		
128×10^6	1		-37.843 839 2(99)		
	1000		-37.843 840 5(05)		
		ΔE_{trial}	$< 0.000\ 001\ 3$		
		$\Delta E_{\text{initiator}}$	$< 0.000\ 000\ 4$		
			aug-cc-pCV7Z		
64×10^6	1		-37.844 251 5(15)	-37.430 345 0(03)	90 841.955
	1000		-37.844 251 5(08)	-37.430 345 0(01)	90 841.977
		ΔE_{trial}	$< 0.000\ 000\ 1$	$< 0.000\ 000\ 1$	< 0.02
128×10^6	1		-37.844 251 5(0?)	-37.430 345 0(02)	90 841.977
	1000		-37.844 251 5(05)	-37.430 345 1(01)	90841.955
		ΔE_{trial}	$< 0.000\ 000\ 1$	$< 0.000\ 000\ 1$	< 0.02
		$\Delta E_{\text{initiator}}$	$< 0.000\ 000\ 1$	$< 0.000\ 000\ 1$	< 0.02
			CBS Extrapolation		
$1/X^3$			-37.844 697 1	-37.430 741 9	90 852.665
$1/X^{3.5}$			-37.844 645 6	-37.430 696 0	90 851.416
$1/\left(X + \frac{1}{2}\right)^4$			-37.844 625 1	-37.430 677 8	90 850.925
		$\Delta_{\text{basis set}}$	$\approx 0.000\ 037\ 0$	$\approx 0.000\ 033\ 0$	0.90
Non-relativistic, clamped nucleus (FCIQMC)					90 851.69(90)
X2C Relativistic correction					-30.02(05)
Breit + QED (from)					-0.48
DBOC unclamped nucleus correction					-0.50(01)
Theory	(Present Work)	FCIQMC + X2C + Breit + QED + DBOC			90 820.46(95)
Experiment	(1990)				90 820.45(10)
Theory	(2010)	CCSD-F12 + HLC + MV + Darwin + Breit + QED + DBOC			90 826.09
Theory	(2017)	R/UCCSD(T) + HLC + DKH + Breit + QED + DBOC			90 827.50
Theory	(2015)	FN-DMC			90 774.

[1] M. Puchalski, D. Kędziera, and K. Pachucki, Phys. Rev. A **82**, 062509 (2010).

[2] M. Puchalski, J. Komasa, and K. Pachucki, Physical Review A **87**, 030502 (2013).

[3] M. Puchalski, J. Komasa, and K. Pachucki, Physical Review A **92**, 062501 (2015).

[4] Y. Yang, I. Kylänpää, N. M. Tubman, J. T. Krogel, S. Hammes-Schiffer, and D. M. Ceperley, The Journal of Chemical Physics **143**, 124308 (2015).

[5] R. E. Thomas, G. H. Booth, and A. Alavi, Physical Review Letters **114**, 033001 (2015).

[6] D. Cleland, G. H. Booth, C. Overy, and A. Alavi, Journal of chemical theory and computation **8**, 4138 (2012).

[7] C. Daday, S. Smart, G. H. Booth, A. Alavi, and C. Filippi, Journal of Chemical Theory and Computation **8**, 4441 (2012).

[8] G. H. Booth, A. Grüneis, G. Kresse, and A. Alavi, Nature **493**, 365 (2012).

[9] C. Schwartz, ArXiv Mathematical Physics e-prints (2006), math-ph/0605018.

[10] L. M. Wang, Z.-C. Yan, H. X. Qiao, and G. W. F. Drake, Phys. Rev. A **83**, 034503 (2011).

[11] S. Bubin, M. Pavanello, W.-C. Tung, K. L. Sharkey, and L. Adamowicz, Chemical reviews **113**, 36 (2013).

[12] B. A. Bushaw, W. Nörtershäuser, G. W. F. Drake, and H.-J. Kluge, Phys. Rev. A **75**, 052503 (2007).

[13] M. Puchalski and K. Pachucki, Physical Review A **78**, 052511 (2008).

[14] R. Beigang, D. Schmidt, and P. J. West, Le Journal de Physique Colloques **44**, C7 (1983).

[15] A. E. Kramida and A. N. Ryabtsev, Physica Scripta **76**, 544 (2007).

[16] D. Feller, The Journal of chemical physics **144**, 014105 (2016).

[17] A. K. Wilson, T. van Mourik, and T. H. Dunning, Journal of Molecular Structure: THEOCHEM **388**, 339 (1996).

[18] H.-J. Werner, P. J. Knowles, G. Knizia, F. R. Manby, M. Schütz, P. Celani, W. Györffy, D. Kats, T. Korona, R. Lindh, A. Mitrushenkov, G. Rauhut, K. R. Shamasundar, T. B. Adler, R. D. Amos, A. Bernhardsson, A. Berning, D. L. Cooper, M. J. O. Deegan, A. J. Dobbyn, F. Eckert, E. Goll, C. Hampel, A. Hesselmann, G. Hetzer, T. Hrenar, G. Jansen, C. Köppl, Y. Liu, A. W. Lloyd, R. A. Mata, A. J. May, S. J. McNicholas, W. Meyer, M. E. Mura, A. Nicklass, D. P. O'Neill, P. Palmieri, D. Peng, K. Pflüger, R. Pitzer, M. Reiher, T. Shiozaki, H. Stoll, A. J. Stone, R. Tarroni, T. Thorsteinsson, and M. Wang, "Molpro, version 2015.1, a package of ab initio programs," (2015), see http://www.molpro.net.

[19] C. Zhu, R. H. Byrd, P. Lu, and J. Nocedal, ACM Trans. Math. Softw. **23**, 550 (1997).

[20] G. H. Booth, A. J. W. Thom, and A. Alavi, The Journal of chemical physics **131**, 054106 (2009).

[21] D. Cleland, G. H. Booth, and A. Alavi, The Journal of Chemical Physics **132**, 041103 (2010).

[22] N. S. Blunt, S. D. Smart, J. A. F. Kersten, J. S. Spencer, G. H. Booth, and A. Alavi, The Journal of Chemical Physics **142**, 184107 (2015).

[23] H. Flyvbjerg and H. G. Petersen, The Journal of Chemical Physics **91**, 461 (1989).

[24] W. Kutzelnigg and J. D. Morgan, The Journal of Chemical Physics **96**, 4484 (1992).

[25] J. M. Martin, Chemical Physics Letters **259**, 669 (1996).

[26] .

[27] W. Klopper, R. A. Bachorz, D. P. Tew, and C. Hättig, Phys. Rev. A **81**, 022503 (2010).