

# Introduction to lithospheric geodynamic modelling

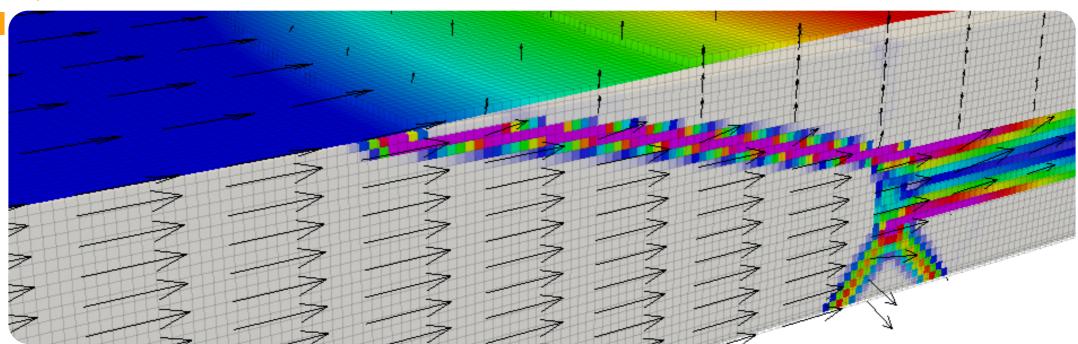
#### **Basic fluid mechanics**

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## Why fluid mechanics?



Velocities and strain rates in a lithospheric geodynamic model

- Most geodynamic models treat the Earth as a continuum such that there are no material gaps or voids at the macroscopic scale
  - Field variables such as pressure, velocity or stress are thus fully continuous
- In this context the Earth is a fluid with a very high viscosity (typically  $10^{18}$   $10^{23}$  Pa s)



## Fluids and the Earth

• Fluid: Any material that flows in response to an applied stress

Differences between solids and fluids

Solids	Fluids
Strain from being stressed	Continuous deformation under applied forces
Stresses related to strains	Stresses related to rates of strain
Strain result of displacement gradients	Strain result of velocity gradients

 Rheological (or constitutive) law: An equation relating stress to strain rates in a fluid



• Fluid mechanics is the science of fluid motion



• Fluid mechanics is the science of fluid motion

- Based on conservation of three basic physical property and their corresponding mathematical representations
  - Conservation of mass The continuity equation
  - Conservation of momentum The momentum equation
  - Conservation of energy The heat transfer equation



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 Conservation of mass, momentum and energy are combined with rheological laws to <u>describe fluid movement under an</u> <u>applied force</u>



## Roadmap

79.8 79 78 Tibe

Fundamental equations governing fluid flow

Calculation of fluid flow velocities/patterns for linear viscous materials



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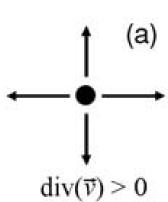
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Covered in lectures 2-3 → • Conservation of energy - The heat transfer equation



## Conservation of mass - Continuity equation



(b)

- Calculations in the continuum are performed by considering an infinitesimal volume of the material, the local volume
- The general form of conservation of mass for a local volume of a continuum in an Eulerian reference frame is

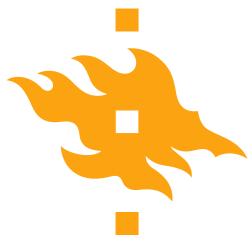
$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{V}) = 0$$

Change in local density Mass or volume flux (divergence of velocity)

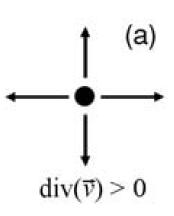
where  $\rho$  is the local density, t is time and V is the local velocity

$$\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \boldsymbol{V}) \quad \text{Alternative form}$$

Gerya, 2010



## Conservation of mass - Continuity equation

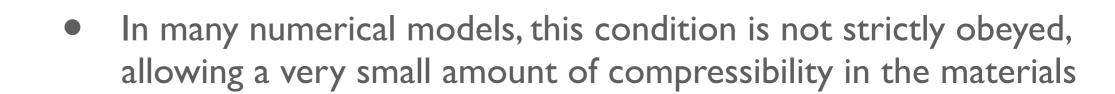


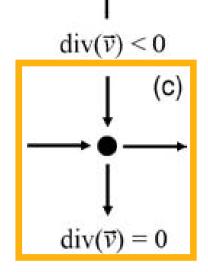
(b)

- It is common in geodynamic numerical models, particularly in the crust or lithosphere, to assume the <u>material is</u> <u>incompressible</u>
- In this case, the continuity equation simplifies to

$$\nabla \cdot \boldsymbol{V} = 0$$

stating simply that there is no divergence in the velocity field of the continuum





Gerya, 2010



Fluid mechanics is the science of fluid motion

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 Conservation of mass, momentum and energy are combined with rheological laws to <u>describe fluid movement under an</u> <u>applied force</u>



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## Conservation of momentum - Momentum eq.



Sir George Stokes

- The basic relationship that thus determines the dynamics of material in the continuum is conservation of momentum, the balance of internal and external forces acting on the material
- The conservation of momentum for a fluid subject to gravity is the Navier-Stokes equation

$$\nabla \cdot \eta (\nabla V + \nabla V^{\mathrm{T}}) - \nabla P - \rho g = \rho \dot{V}$$

Fluid velocity Fluid pressure

Body forces Acceleration

where  $\eta$  is the fluid shear viscosity, P is pressure, g is the acceleration due to gravity, and  $\dot{V}$  is the material time derivative of the fluid velocity (acceleration)



## Conservation of momentum - Momentum eq.



Sir George Stokes

 For highly viscous fluids with a very small Reynolds number the acceleration term of the Navier-Stokes equation can be ignored reducing to the equation of Stokes flow (and simplifying the solutions)

$$\nabla \cdot \eta (\nabla \boldsymbol{V} + \nabla \boldsymbol{V}^{\mathrm{T}}) - \nabla P = \rho g$$
 Fluid velocity Fluid pressure Body forces

 It is trivial to demonstrate that the Reynolds number of most geodynamic flows is extremely low (~10-20)

$$\mathbf{Re} = rac{
ho oldsymbol{V} L}{\eta}$$
 Inertial forces

Viscous forces

The Reynolds number



## Roadmap

Tibetan Plateau
India

Fundamental equations governing fluid flow

Calculation of fluid flow velocities/patterns for linear viscous materials

20 40 500



## Viscous flow - Newtonian (or linear) fluid

- A Newtonian fluid is a fluid in which there is a <u>linear</u> relationship between the rate of strain and the applied stress
  - What would this relationship look like as an equation?



## Viscous flow - Newtonian (or linear) fluid

Material	Approximate Viscosity [Pa s]
Air	I×I0 <sup>-5</sup>
Water	I×I0-3
Ice	I×10 <sup>16</sup>
Rock Salt	I×10 <sup>17</sup>
Granite	I×I0 <sup>20</sup>

- A Newtonian fluid is a fluid in which there is a linear relationship between the rate of strain and the applied stress
  - What would this relationship look like as an equation?

$$\sigma \propto \dot{\varepsilon}$$
 or  $\sigma = \eta \dot{\varepsilon}$ 

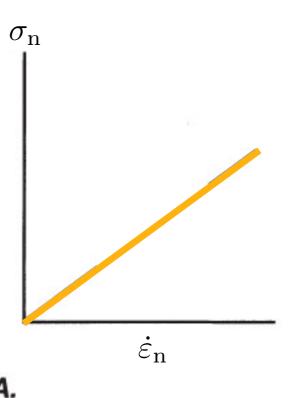
- The proportionality constant  $\eta$  is known as the dynamic (or shear) viscosity
  - Dynamic viscosity has units of [Pa s]



In simple shear,

$$au_s = \eta \dot{arepsilon}_s \qquad \eta$$
 Dynamic viscosity

Shear stress proportional to shear strain rate





In simple shear,

$$au_s = \eta \dot{arepsilon}_s$$

 $au_s = \eta \dot{arepsilon}_s \qquad \eta \; ext{Dynamic viscosity}$ 

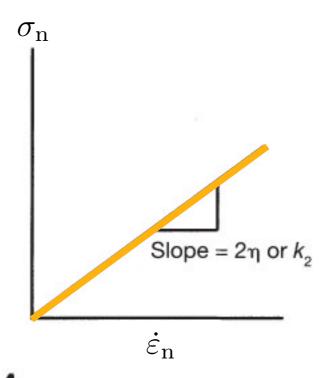
Shear stress proportional to shear strain rate

In general,

$$\sigma' = 2\eta \dot{\varepsilon}$$

deviatoric stress is proportional to strain rate

For linear viscous (Newtonian) materials,  $\eta$  is constant





In simple shear,

$$au_s = \eta \dot{arepsilon}_s \qquad \eta$$

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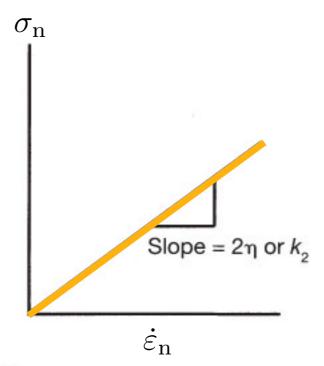
Shear stress proportional to shear strain rate

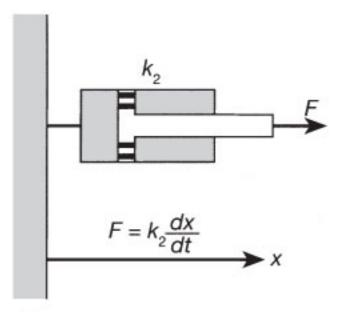
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In simple shear,

$$au_s = \eta \dot{arepsilon}_s \qquad \eta$$
 Dynamic viscosity

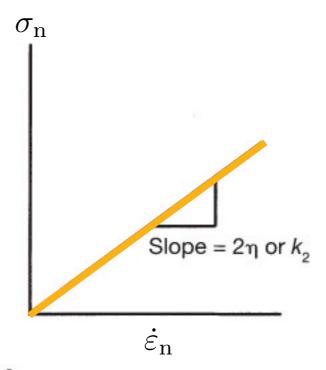
Shear stress proportional to shear strain rate

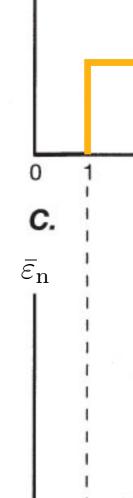
In general,

$$\sigma' = 2\eta \dot{\varepsilon}$$

deviatoric stress is proportional to strain rate

- For linear viscous (Newtonian) materials,  $\eta$  is constant
- Nonrecoverable



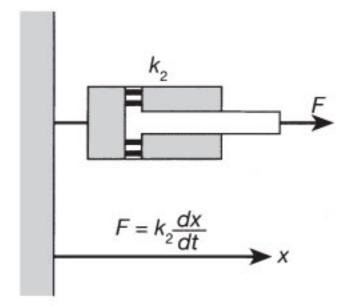


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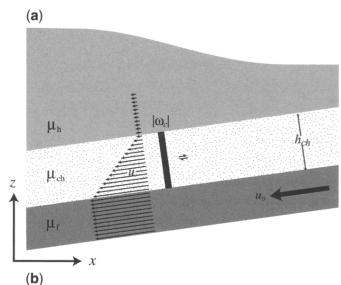
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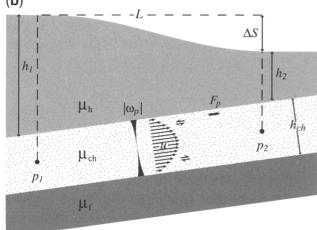
 $\sigma_{
m n}$ 

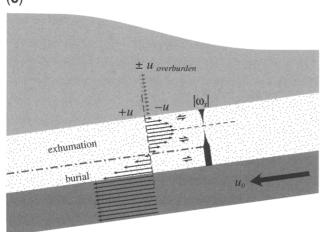




#### Channel flows in the Earth

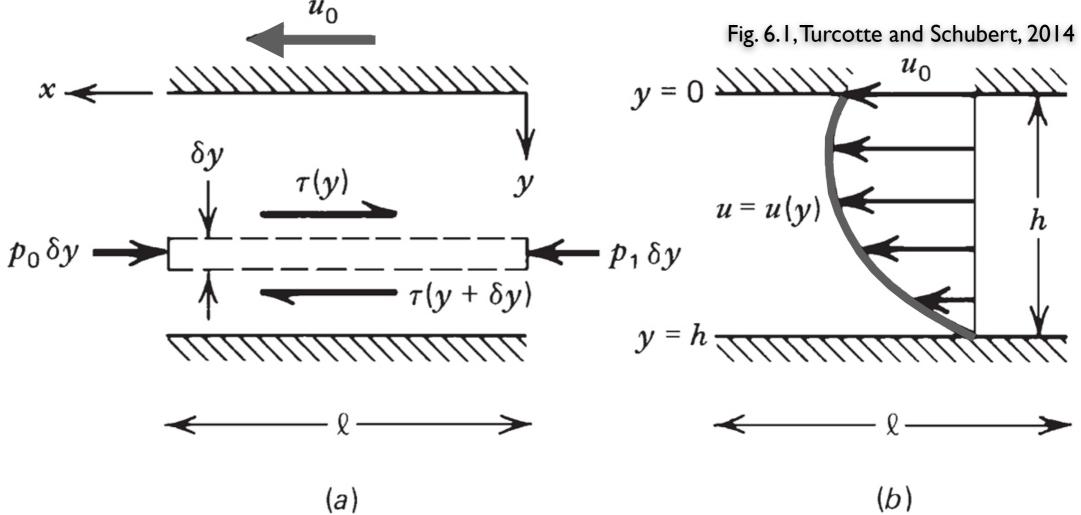






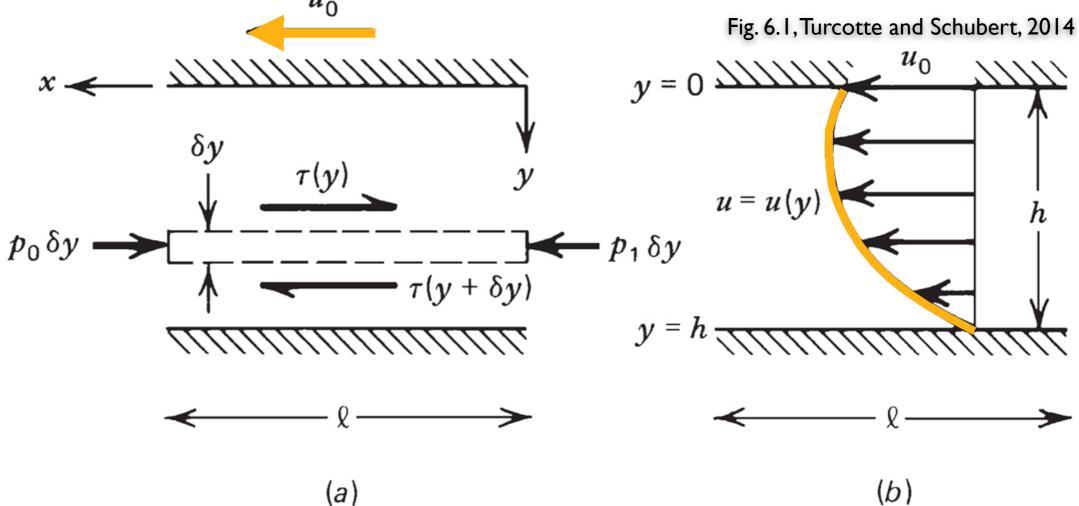
- Channel flows in the Earth occur when a fluid flows within a channel, between two solid "walls"
- Such channels can be found in a number of geological settings:
  - Counterflow in the asthenosphere
  - Lower crustal flow
  - Intra-crustal channels (figure on left)
  - Subduction channels
  - Salt tectonic channels





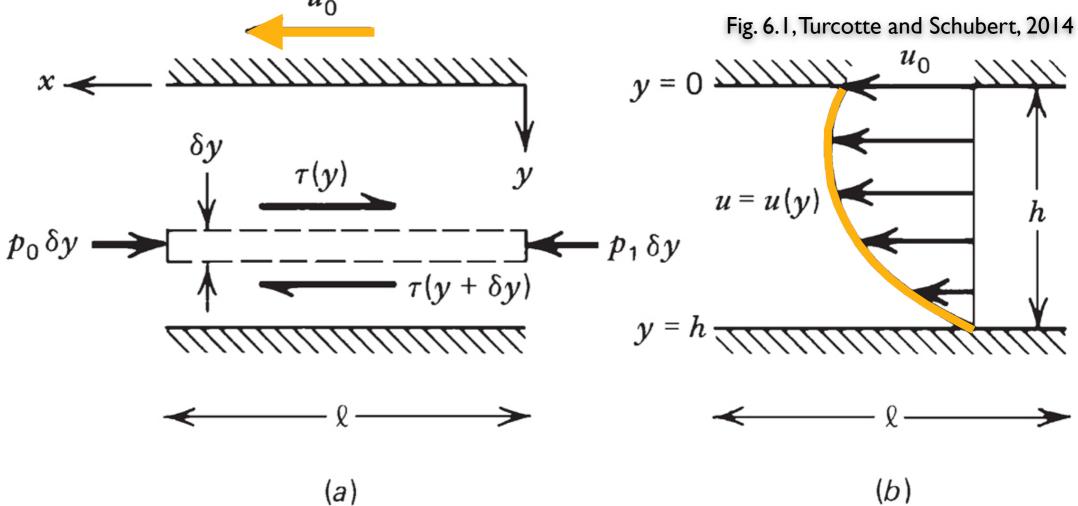
 The most simple fluid flow we can consider is <u>flow of a fluid in</u> one <u>direction within a channel of fixed width</u>





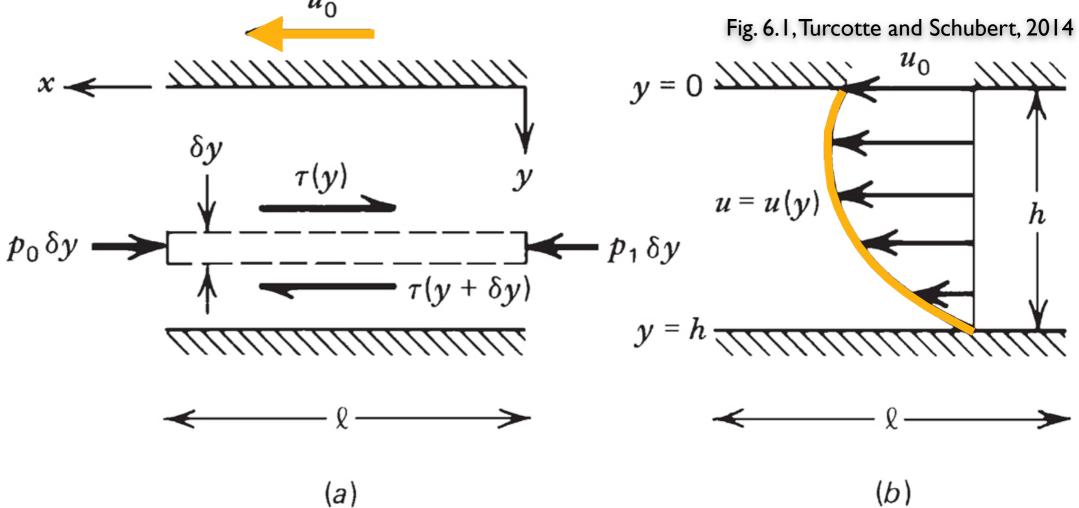
- Fluid is flowing with velocity u in the x direction, and the flow velocity u is a function of distance across the channel y
- Flow results from
  - a pressure gradient  $(p_0 p_1)/l$ , and/or
  - motion of the side wall of the channel  $u_0$





- Shear, or a gradient in the velocity, in the channel results in a shear stress  $\tau$  that is exerted on horizontal planes in the fluid
- For a Newtonian fluid with a constant dynamic viscosity  $\eta$  we can state  $\frac{du}{dt}$

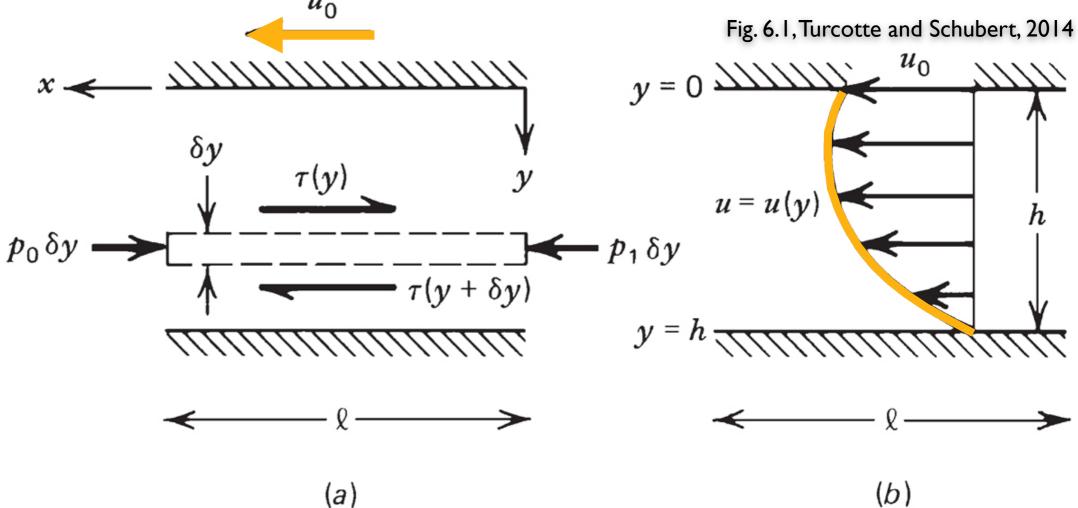




- Shear, or a gradient in the velocity, in the channel results in a shear stress  $\tau$  that is exerted on horizontal planes in the fluid
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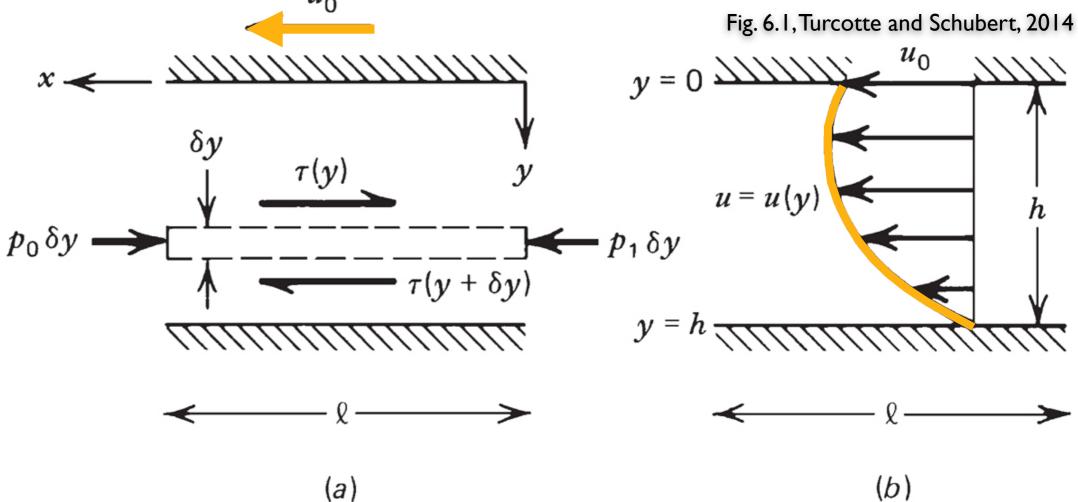




- We can now determine the flow in the channel using the equation of motion, based on the force balance on a layer of fluid of thickness  $\delta y$  and length l
- The net pressure force on the element in the x direction is

$$(p_1-p_0)\delta y$$

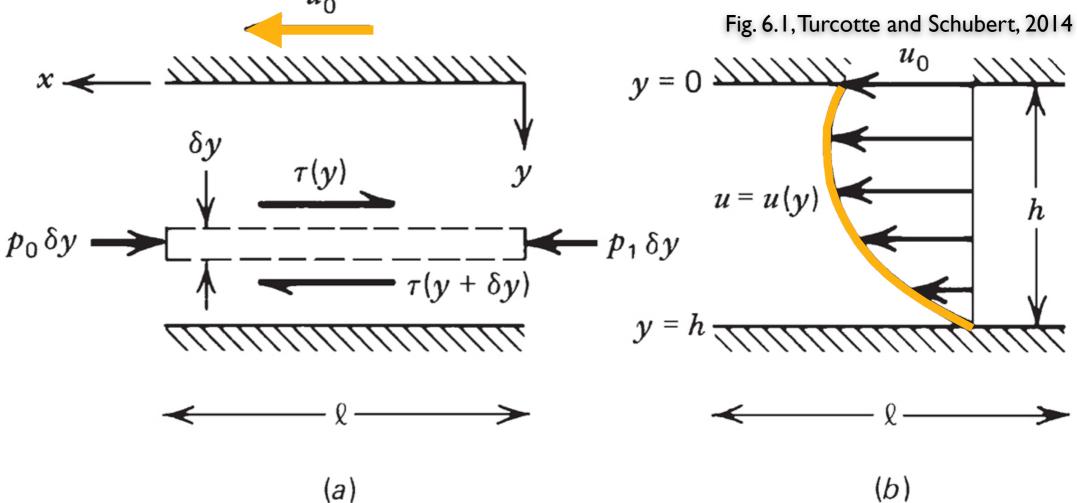




- Because the shear stress  $\tau$  and velocity u are both only a function of distance y, the shear force on the upper boundary of the element is  $-\tau(y)l$
- The equivalent shear force on the lower boundary is

$$\tau(y + \delta y)l = \left(\tau(y) + \frac{d\tau}{dy}\delta y\right)l$$





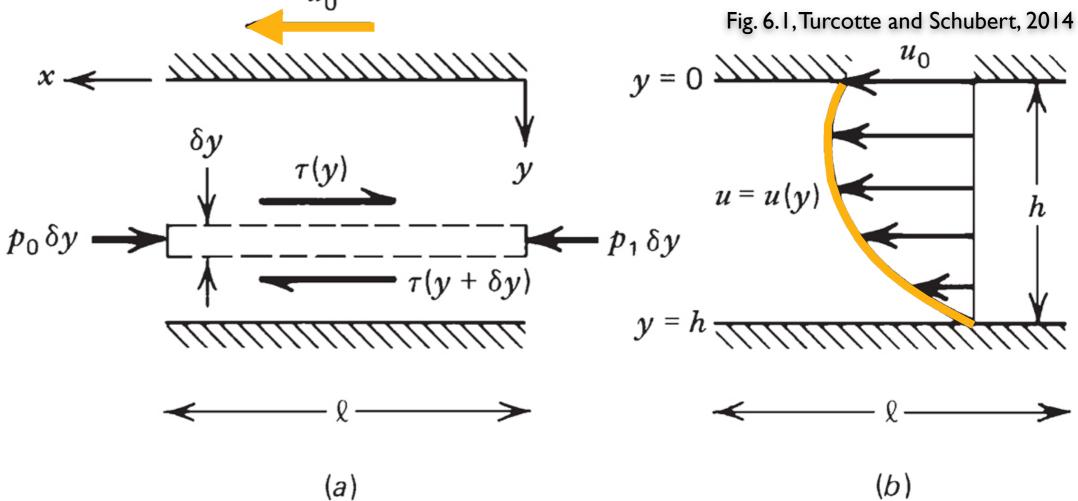
The net force (or sum of the forces) must be equal to zero, or

$$(p_1 - p_0)\delta y + \left[\tau(y) + \frac{d\tau}{dy}\delta y\right]l - \tau(y)l = 0$$

• As  $\delta y \to 0$ , the relationship above becomes

$$\frac{d\tau}{dy} = -\frac{(p_1 - p_0)}{l}$$





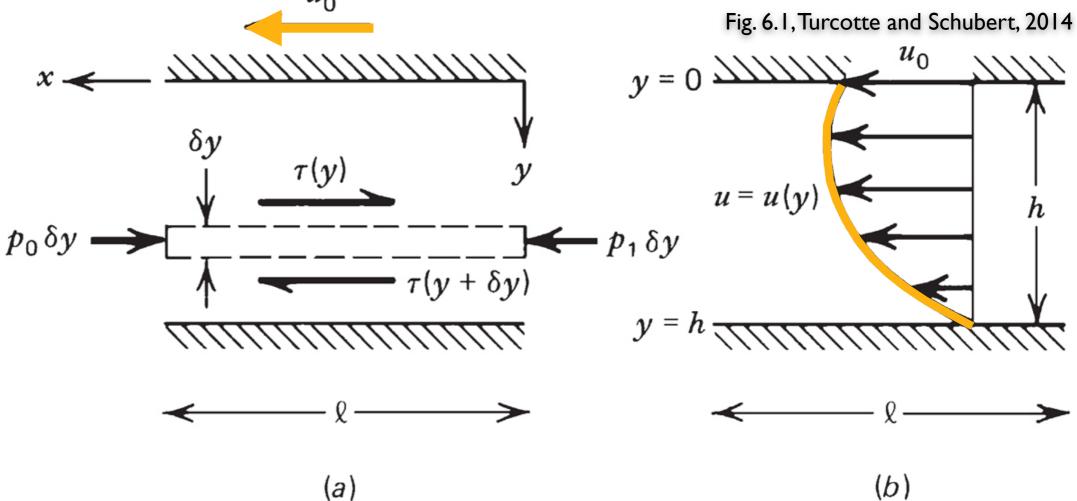
 The right side of the previous equation is the horizontal pressure gradient in the channel

$$\frac{dp}{dx} = -\frac{(p_1 - p_0)}{l}$$

• From which the equation of motion can be written

$$\frac{d\tau}{dy} = \frac{dp}{dx}$$





## Newtonian fluid

$$\tau = \eta \frac{du}{dy}$$

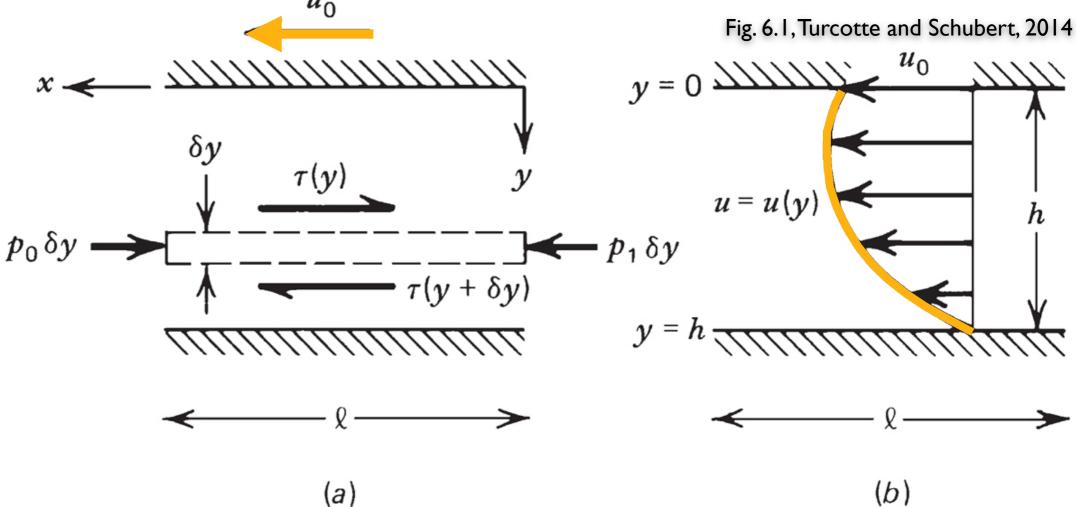
 Velocity in the channel is found by substituting the rheological law for a Newtonian fluid into the equation of motion

$$\frac{d\tau}{dy} = \frac{d}{dy}\eta \frac{du}{dy} = \eta \frac{d^2u}{dy^2} = \frac{d\eta}{dx}$$

Integrating the equation above twice yields

$$u = \frac{1}{2\eta} \frac{dp}{dx} y^2 + c_1 y + c_2$$

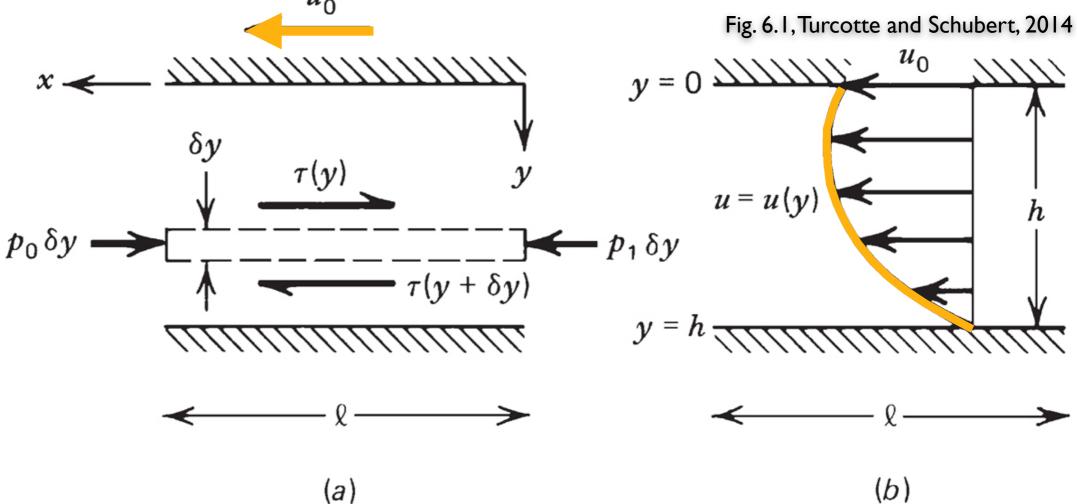




• The constants  $c_1$  and  $c_2$  can be found by applying the boundary conditions that u = 0 at y = h, and  $u = u_0$  at y = 0 (no slip)

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$





The constants  $c_1$  and  $c_2$  can be found by applying the boundary conditions that u = 0 at y = h, and  $u = u_0$  at y = 0 (no slip)

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$
ressure gradient
wall velocity



## Channel flow challenge #1

- Start by navigating to the directory
   NGWM2016-modelling-course/Lessons/04-Basic-fluid-mechanics/scripts
- Right-click on the Python script called 1D-channel-flow.py and choose "Edit with IDLE"
- The script cannot currently be run because it is missing the equation for the velocity in a ID channel

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$

your task is to input this equation into the script and then run the script and save plots for the following scenarios:

- No pressure gradient:  $u_0 = 1.0 \text{ mm/a}$ ; dp/dx = 0.0 Pa
- No wall velocity:  $u_0 = 0.0 \text{ mm/a}$ ; dp/dx = -2000.0 Pa
- Other cases: No pressure gradient/wall velocity, both, etc.



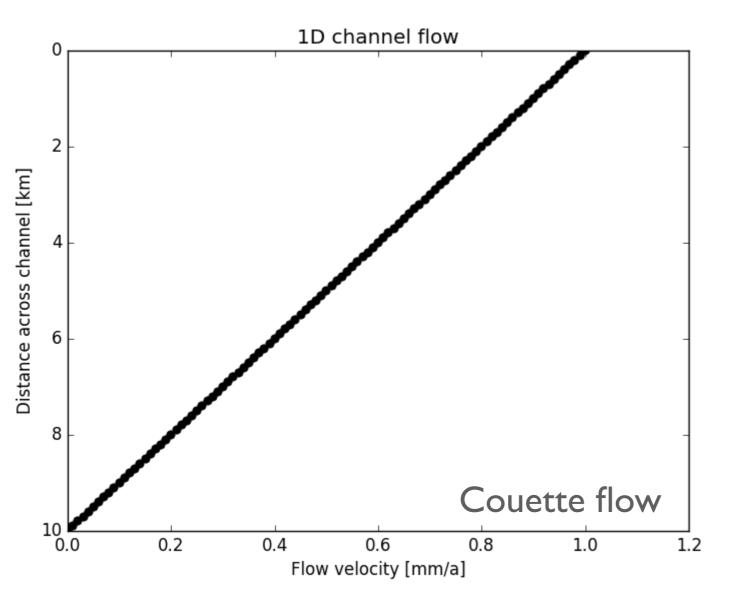
## What does this equation tell us?

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$

- We've now look at several simple fluid flow behaviors, including two important end members
  - Zero pressure gradient  $(\frac{dp}{dx} = 0)$
  - Zero boundary velocity ( $u_0 = 0$ )



#### Couette flow



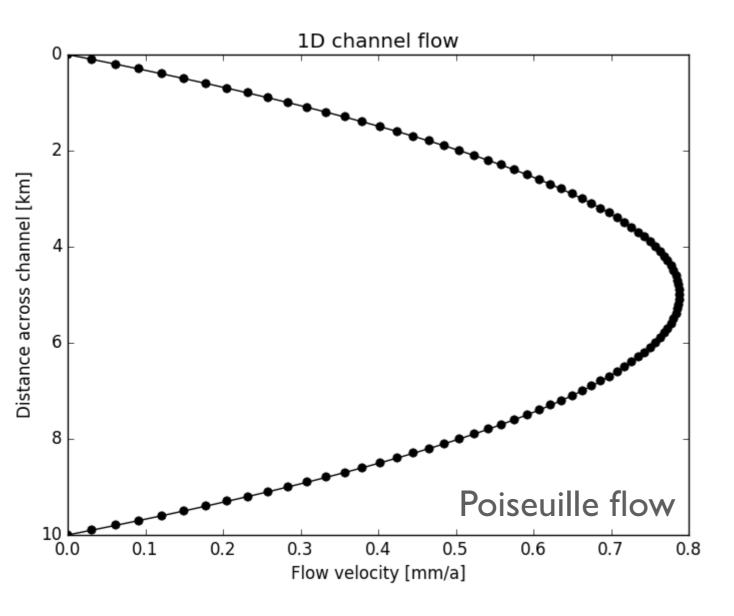
 A Couette flow has no pressure gradient, or dp/dx = 0, reducing the ID equation for velocity in the channel down to

$$u = u_0 \left( 1 - \frac{y}{h} \right)$$

Clearly, this predicts a linear increase in velocity from y = h to y = 0



#### Poiseuille flow

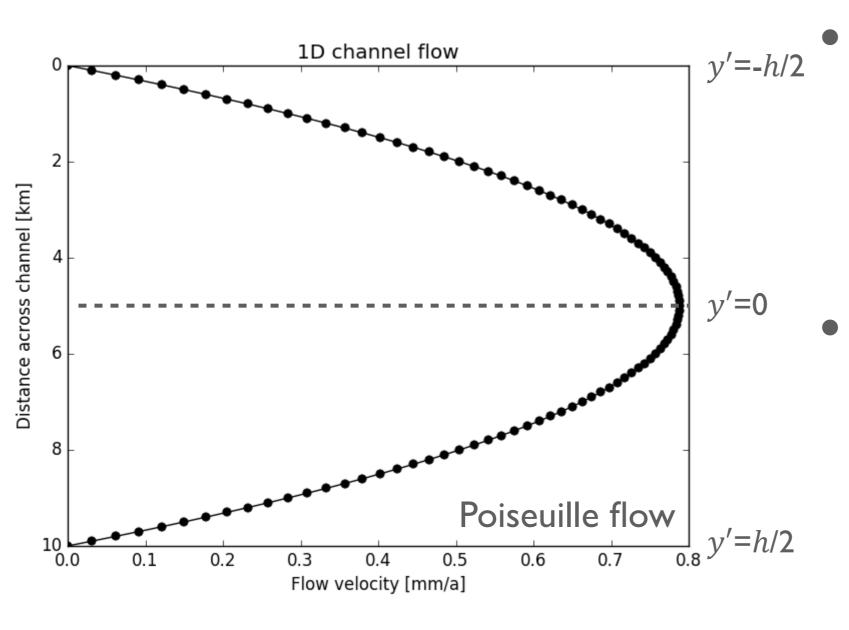


• Poiseuille flow is driven only by a pressure gradient in the channel with zero boundary velocities ( $u_0 = 0$ ), thus

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy)$$



#### Poiseuille flow



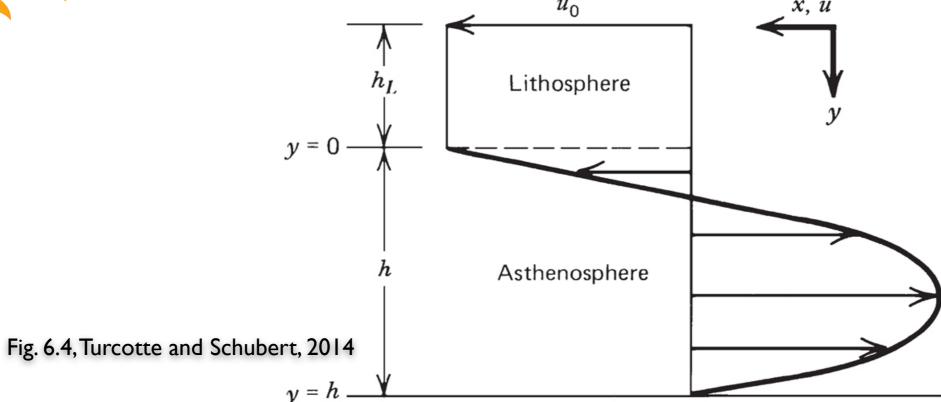
Poiseuille flow is driven only by a pressure gradient in the channel with zero boundary velocities ( $u_0 = 0$ ), thus

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy)$$

In a coordinate system with y' at the middle of the channel we can say y' = y - h/2, which results in the relationship

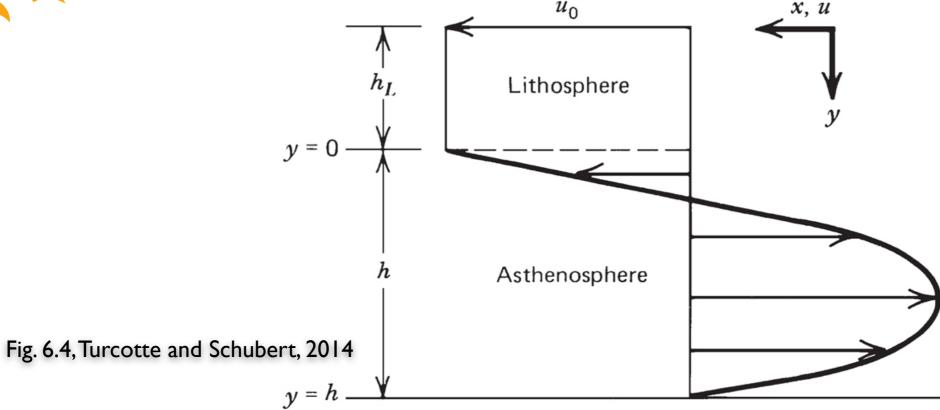
$$u = \frac{1}{2\eta} \frac{dp}{dx} \left( y'^2 - \frac{h^2}{4} \right)$$





- One model for mantle flow is that the motion of lithospheric plates on the Earth's surface produces a counterflow in the uppermost asthenosphere (upper ~100-200 km)
- If we assume the plate is rigid and moving at velocity  $u_0$ , and that the velocity at some depth y = h must be zero, it is clear that the counterflow is <u>opposite</u> in direction to the plate motion in order to <u>conserve mass</u>





Mathematically, we can state that as

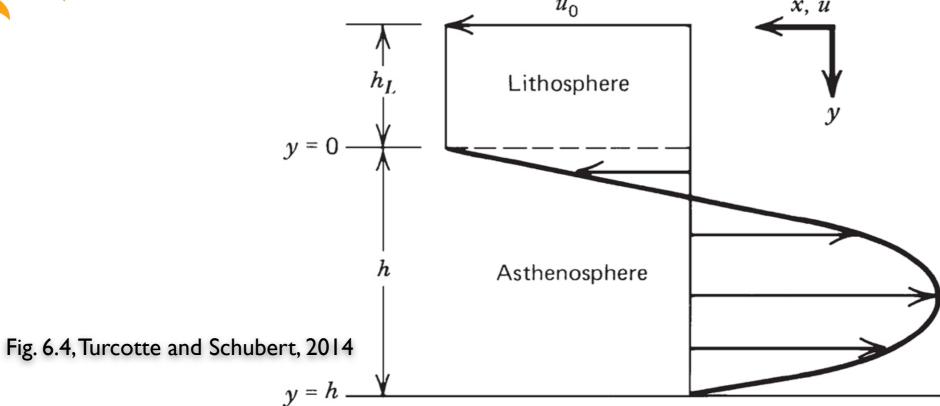
$$u_0 h_L + \int_0^h u \ dy = 0$$

where  $h_L$  is the thickness of the lithosphere and h is the thickness of the asthenosphere involved in counterflow

• If we insert our equation for ID channel flow in the second term, we get  $h^3 dn = u_0 h$ 

$$u_0 h_L + \frac{h^3}{12\eta} \frac{dp}{dx} + \frac{u_0 h}{2} = 0$$





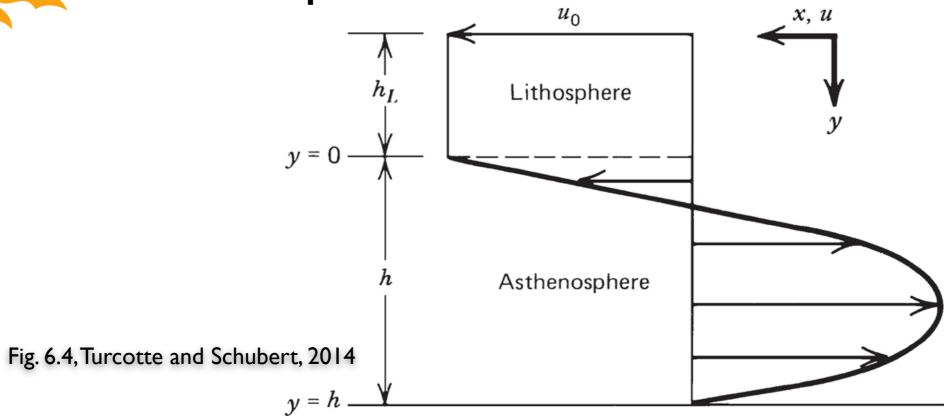
If we now solve for the pressure gradient, we find

$$\frac{dp}{dx} = \frac{12\eta u_0}{h^2} \left( \frac{h_L}{h} + \frac{1}{2} \right)$$

 And this can be inserted into the equation for ID channel flow to get the predicted velocity profile for a counterflow

$$u = u_0 \left[ 1 - \frac{y}{h} + 6 \left( \frac{h_L}{h} + \frac{1}{2} \right) \left( \frac{y^2}{h^2} - \frac{y}{h} \right) \right]$$





$$u = u_0 \left[ 1 - \frac{y}{h} + 6 \left( \frac{h_L}{h} + \frac{1}{2} \right) \left( \frac{y^2}{h^2} - \frac{y}{h} \right) \right]$$

 Looking at this equation for a moment, is there anything missing that you might expect to see?



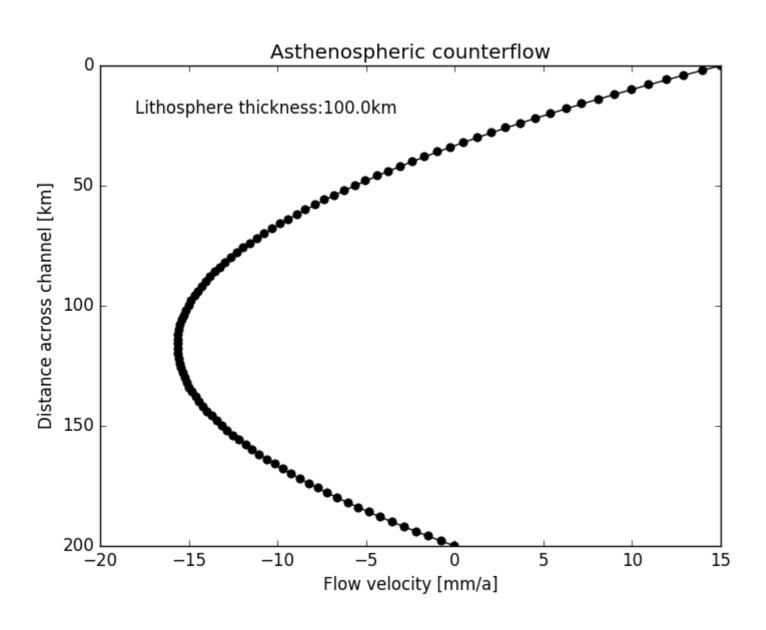
# Channel flow challenge #2

- In the same directory as before, right-click on the script 1D-asthenospheric-counterflow.py and choose "Edit with IDLE"
- Again, this script cannot be run because it is missing the equation for asthenospheric counterflow

$$u = u_0 \left[ 1 - \frac{y}{h} + 6 \left( \frac{h_L}{h} + \frac{1}{2} \right) \left( \frac{y^2}{h^2} - \frac{y}{h} \right) \right]$$

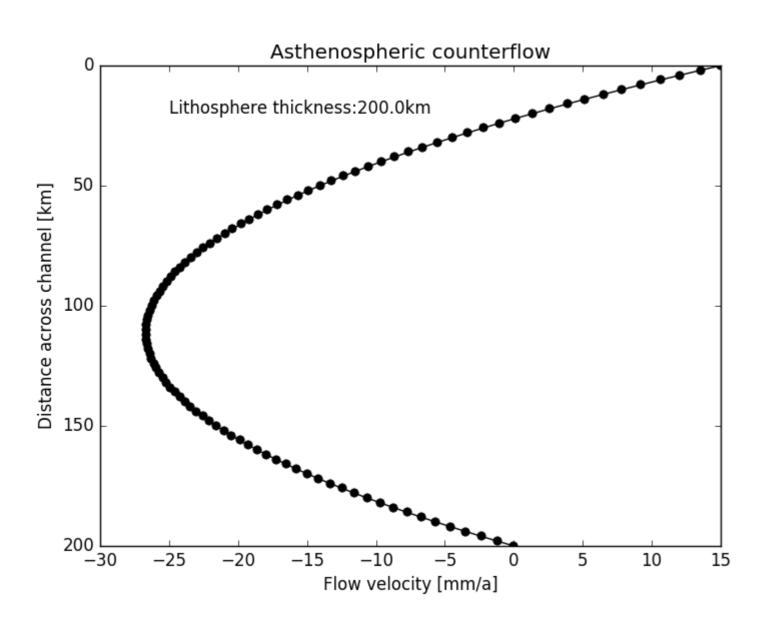
- In addition, you will need to add a text label to display the thickness of the lithospheric plate in [km] on the plot
  - You can find how to use the plt.text() function at http://matplotlib.org/api/pyplot\_api.html#matplotlib.pyplot.text
- After making the changes, run the script with the default values and save the plot
- How do the flow velocities change when you vary h and  $h_1$ ?





- As expected, movement of a 100-km-thick lithosphere at  $u_0 = 15$  cm/a results in counter flow in the 200-km-thick asthenosphere
  - The maximum
     counterflow velocity is
     roughly the same as the
     plate velocity in this case





- When the plate thickness is doubled, the counterflow velocity must also increase in order to conserve mass, as required by our counterflow equation
  - The maximum counterflow velocity here is about 1.8 times the plate velocity



### Limitations to our channel flow models

- Though we can predict the velocities in ID channels for a number of different scenarios, we're not able to handle a few things of geological importance
  - Nonlinear viscosity in the channel
    - As you'll see tomorrow morning, viscous flow in rocks is generally not linear
  - Spatial variations in the channel material
    - Perhaps we have a channel with two different fluids
  - Changes in the boundary conditions with time
    - What if the channel wall velocity varies with time?



## Summary

Fluid mechanics is the science of fluid motion

 Fluid motions are caused by internal and external forces, and modelled using simple formulations of the conservation of mass, momentum and energy

 For geological applications, we treat the Earth as a fluid with a high viscosity and model flow using the Stokes equation

 Analytical solutions for ID channel flows can provide insight into fluid flow in the Earth, but have <u>significant limitations</u>



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Turcotte, D. L., & Schubert, G. (2014). Geodynamics. Cambridge University Press.

Twiss, R. J., & Moores, E. M. (2007). Structural Geology, 2nd Edition. W.H. Freeman Co.