



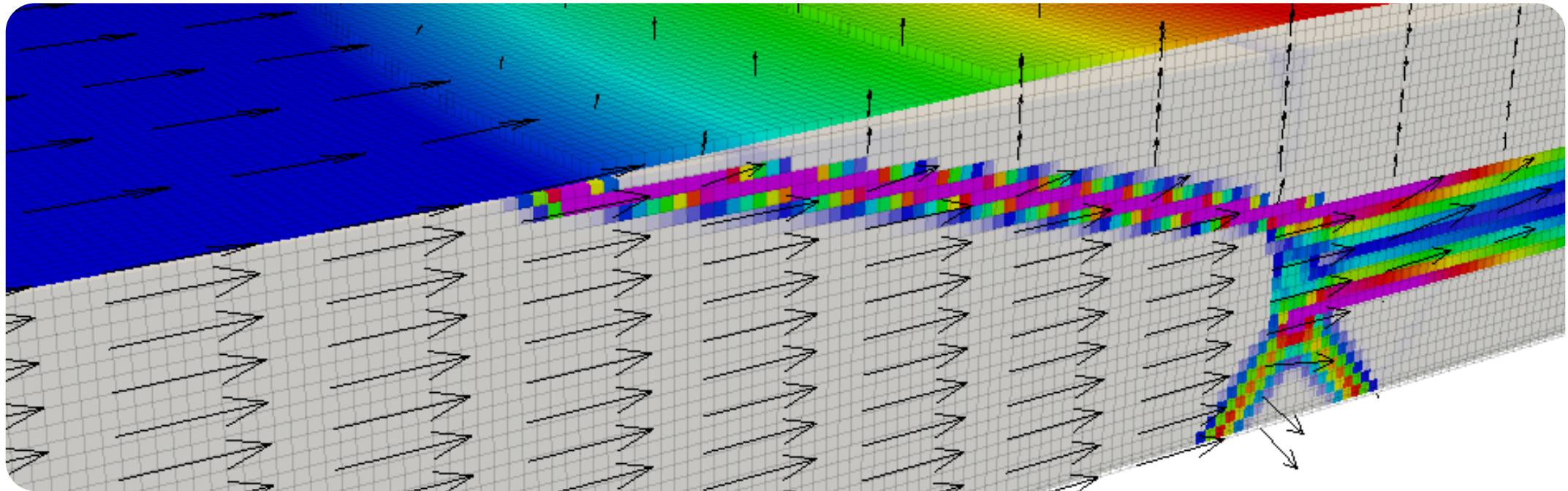
# Introduction to lithospheric geodynamic modelling

## Basic fluid mechanics

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Nordic Geological Winter Meeting 2016

# Why fluid mechanics?



*Velocities and strain rates in a lithospheric geodynamic model*

- Most geodynamic models treat the Earth as a **continuum** such that there are no material gaps or voids at the macroscopic scale
- Field variables such as pressure, velocity or stress are thus fully continuous
- In this context the Earth is a fluid with a very high viscosity (typically  $10^{18}$  -  $10^{23}$  Pa s)



# Fluids and the Earth

- **Fluid**: Any material that flows in response to an applied stress

- Differences between **solids** and **fluids**

Solids	Fluids
Strain from being stressed	<b>Continuous</b> deformation under applied forces
Stresses related to strains	Stresses related to <b>rates of strain</b>
Strain result of displacement gradients	Strain result of <b>velocity gradients</b>

- **Rheological** (or **constitutive**) **law**: An equation relating stress to strain rates in a fluid



# Fluid mechanics

- **Fluid mechanics** is the science of fluid motion



# Fluid mechanics

- **Fluid mechanics** is the science of fluid motion
- Based on conservation of three basic physical property and their corresponding mathematical representations
  - **Conservation of mass** - The continuity equation
  - **Conservation of momentum** - The momentum equation
  - **Conservation of energy** - The heat transfer equation



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- Conservation of mass, momentum and energy are combined with rheological laws to describe fluid movement under an applied force



# Roadmap

Model elevation (km)

79.8  
79  
78  
77  
76  
74.8

- **Fundamental equations** governing fluid flow
- Calculation of fluid flow velocities/patterns for **linear viscous materials**

Tibetan Plateau

India

0.0 20 40 50.0



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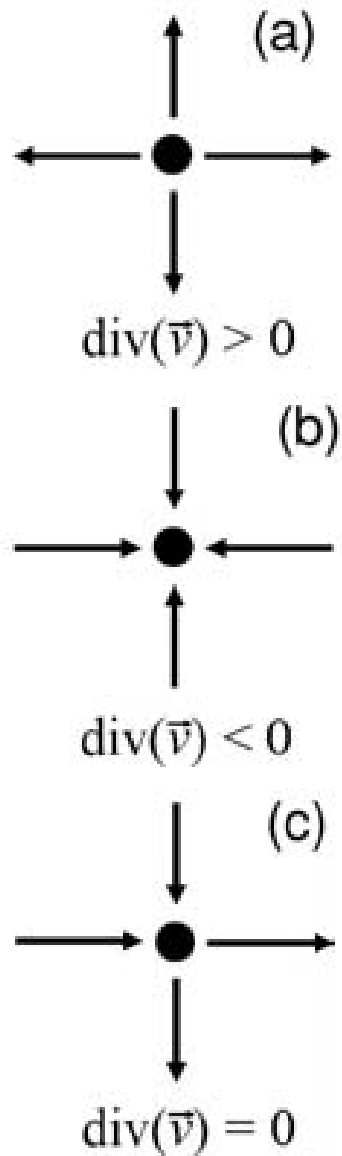




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    - **Conservation of momentum** - The momentum equation
- Covered in lectures 2-3 → ● ~~Conservation of energy - The heat transfer equation~~

# Conservation of mass - Continuity equation



Gerya, 2010

- Calculations in the continuum are performed by considering an infinitesimal volume of the material, the local volume
- The general form of conservation of mass for a local volume of a continuum in an Eulerian reference frame is

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{V}) = 0$$

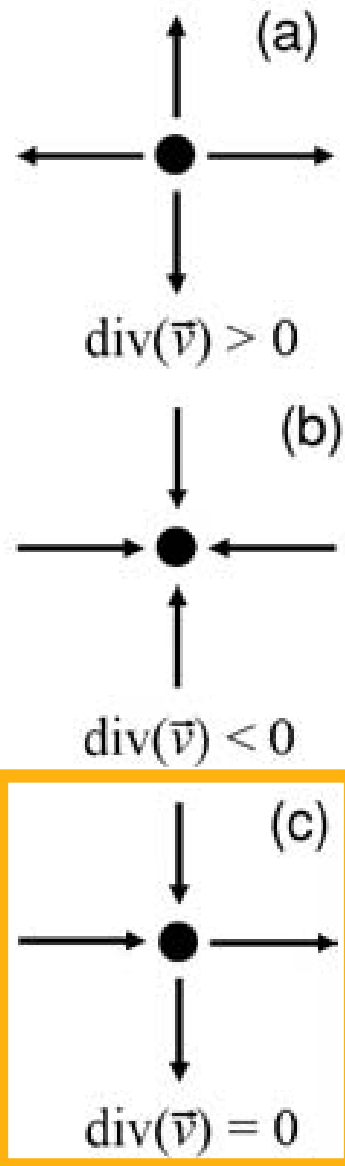
Change in local density      Mass or volume flux  
(divergence of velocity)

where  $\rho$  is the local density,  $t$  is time and  $\mathbf{V}$  is the local velocity

$$\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \mathbf{V}) \quad \text{Alternative form}$$



# Conservation of mass - Continuity equation



Gerya, 2010

- It is common in geodynamic numerical models, particularly in the crust or lithosphere, to assume the material is incompressible

- In this case, the **continuity equation** simplifies to

$$\nabla \cdot \mathbf{V} = 0$$

stating simply that there is no divergence in the velocity field of the continuum

- In many numerical models, this condition is not strictly obeyed, allowing a very small amount of compressibility in the materials



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**What kind of forces might we expect to have acting on a fluid?**



# Conservation of momentum - Momentum eq.



Sir George Stokes

- The basic relationship that thus determines the dynamics of material in the continuum is conservation of momentum, the balance of internal and external forces acting on the material

- The **conservation of momentum** for a fluid subject to gravity is the Navier-Stokes equation

$$\nabla \cdot \eta(\nabla \mathbf{V} + \nabla \mathbf{V}^T) - \nabla P - \rho \mathbf{g} = \rho \dot{\mathbf{V}}$$

Fluid velocity

Fluid pressure

Body forces

Acceleration

where  $\eta$  is the fluid shear viscosity,  $P$  is pressure,  $g$  is the acceleration due to gravity, and  $\dot{\mathbf{V}}$  is the material time derivative of the fluid velocity (acceleration)



# Conservation of momentum - Momentum eq.



Sir George Stokes

- For highly viscous fluids with a very small Reynolds number the acceleration term of the Navier-Stokes equation can be ignored reducing to the equation of **Stokes flow** (and simplifying the solutions)

$$\nabla \cdot \eta(\nabla \mathbf{V} + \nabla \mathbf{V}^T) - \nabla P = \rho \mathbf{g}$$

Fluid velocity

Fluid pressure

Body forces

- It is trivial to demonstrate that the Reynolds number of most geodynamic flows is extremely low ( $\sim 10^{-20}$ )

$$\text{Re} = \frac{\rho V L}{\eta} \quad \frac{\text{Inertial forces}}{\text{Viscous forces}}$$

The Reynolds number



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# Viscous flow - Newtonian (or linear) fluid

- A **Newtonian fluid** is a fluid in which there is a linear relationship between the rate of strain and the applied stress
- What would this relationship look like as an equation?



# Viscous flow - Newtonian (or linear) fluid

Material	Approximate Viscosity [Pa s]
Air	$1 \times 10^{-5}$
Water	$1 \times 10^{-3}$
Ice	$1 \times 10^{16}$
Rock Salt	$1 \times 10^{17}$
Granite	$1 \times 10^{20}$

- A **Newtonian fluid** is a fluid in which there is a linear relationship between the rate of strain and the applied stress

- What would this relationship look like as an equation?

$$\sigma \propto \dot{\epsilon} \quad \text{or} \quad \sigma = \eta \dot{\epsilon}$$

- The proportionality constant  $\eta$  is known as the **dynamic** (or **shear**) **viscosity**
- Dynamic viscosity has units of [Pa s]

# Viscous flow - (Linear) Viscous deformation

- In simple shear,

$$\tau_s = \eta \dot{\epsilon}_s \quad \eta \text{ Dynamic viscosity}$$

**Shear stress** proportional to **shear strain rate**



# Viscous flow - (Linear) Viscous deformation

- In simple shear,

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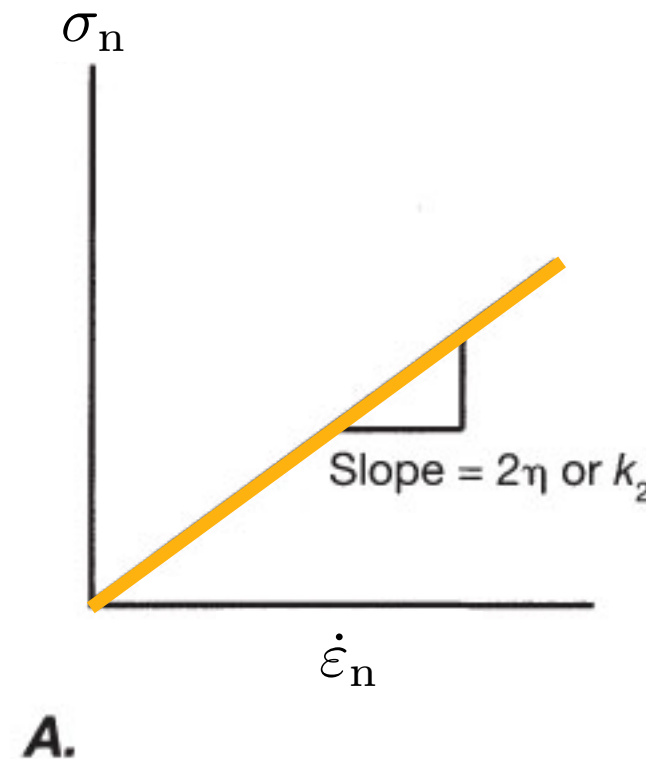
**Shear stress** proportional to **shear strain rate**

- In general,

$$\sigma' = 2\eta \dot{\epsilon}$$

**deviatoric stress** is proportional to **strain rate**

- For linear viscous (Newtonian) materials,  $\eta$  is constant



# Viscous flow - (Linear) Viscous deformation

- In simple shear,

$$\tau_s = \eta \dot{\epsilon}_s \quad \eta \text{ Dynamic viscosity}$$

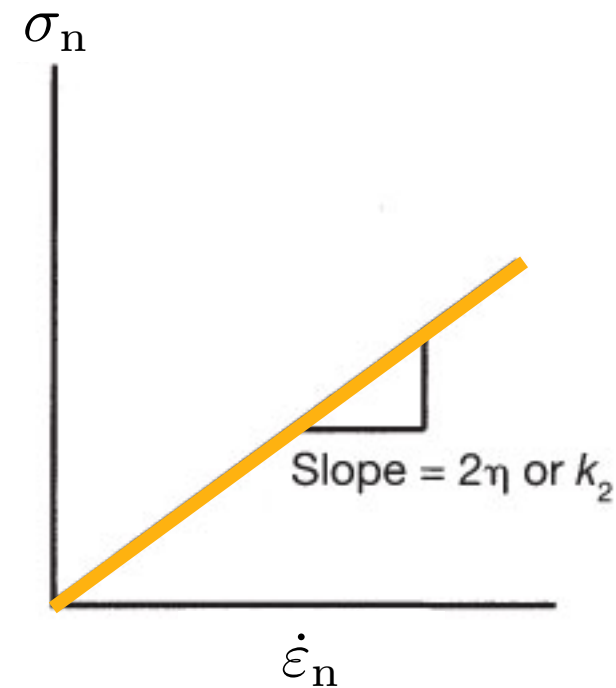
**Shear stress** proportional to **shear strain rate**

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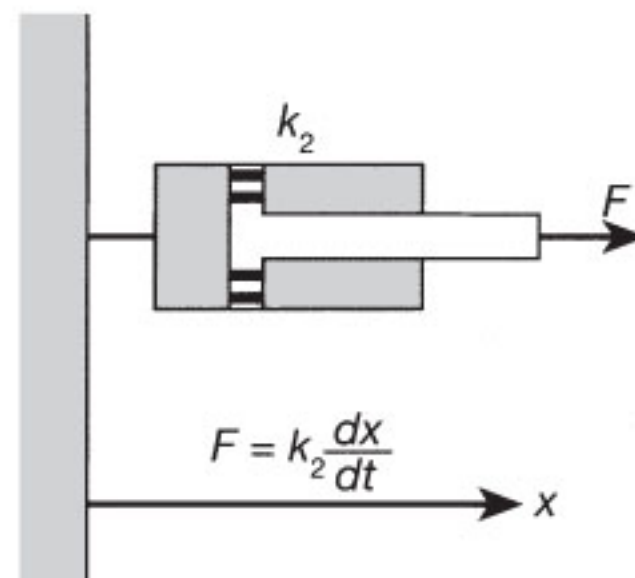
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**deviatoric stress** is proportional to **strain rate**

- For linear viscous (Newtonian) materials,  $\eta$  is constant



A.



B.

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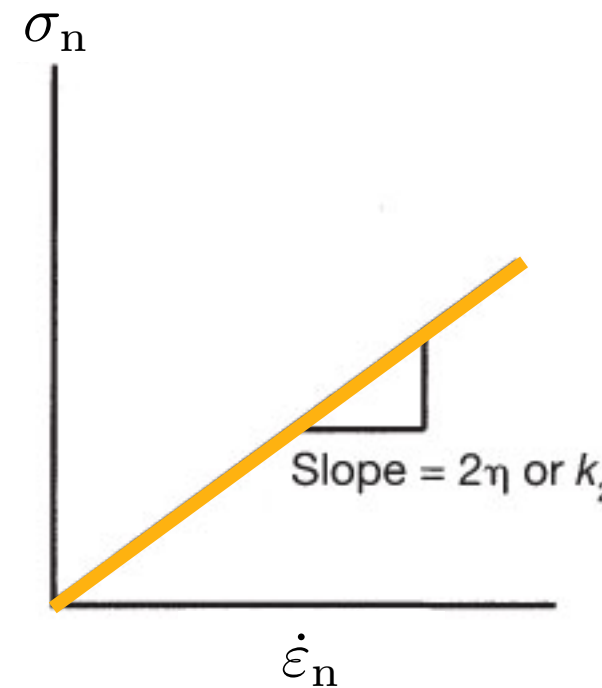
**Shear stress** proportional to **shear strain rate**

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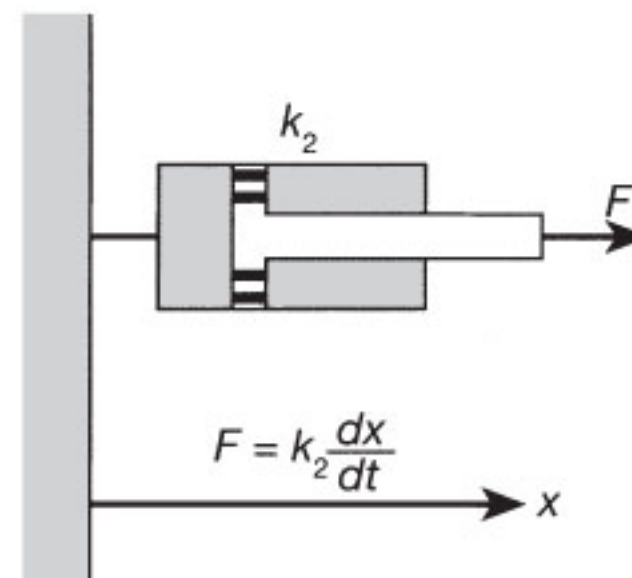
$$\sigma' = 2\eta \dot{\epsilon}$$

**deviatoric stress** is proportional to **strain rate**

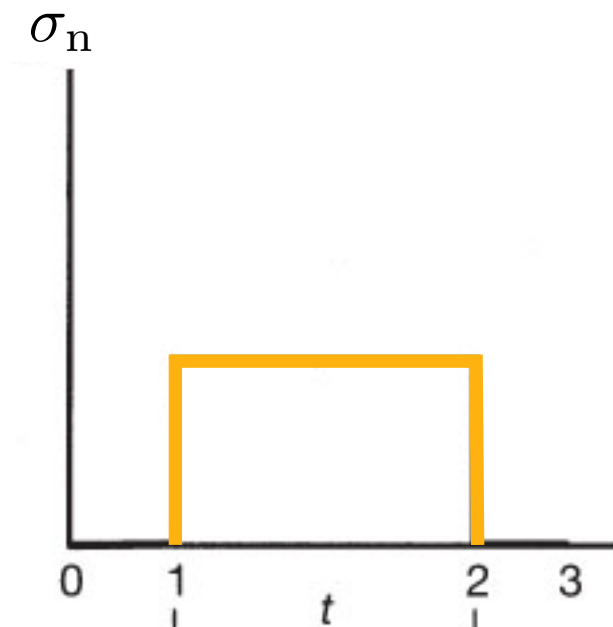
- For linear viscous (Newtonian) materials,  $\eta$  is constant
- Nonrecoverable



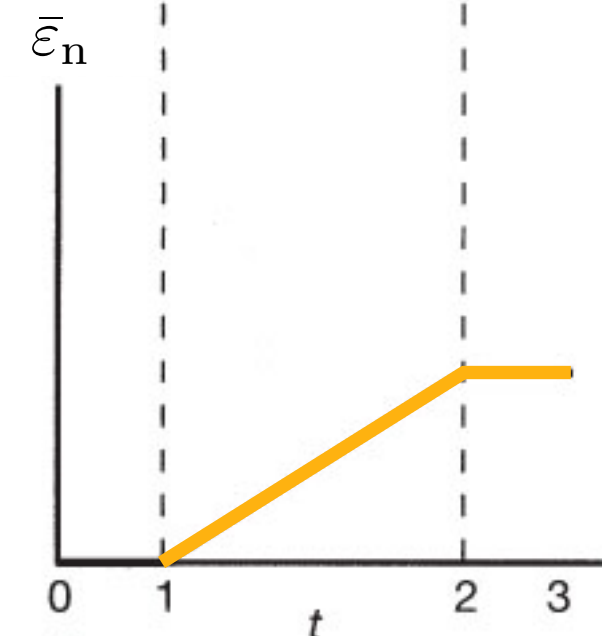
A.



B.

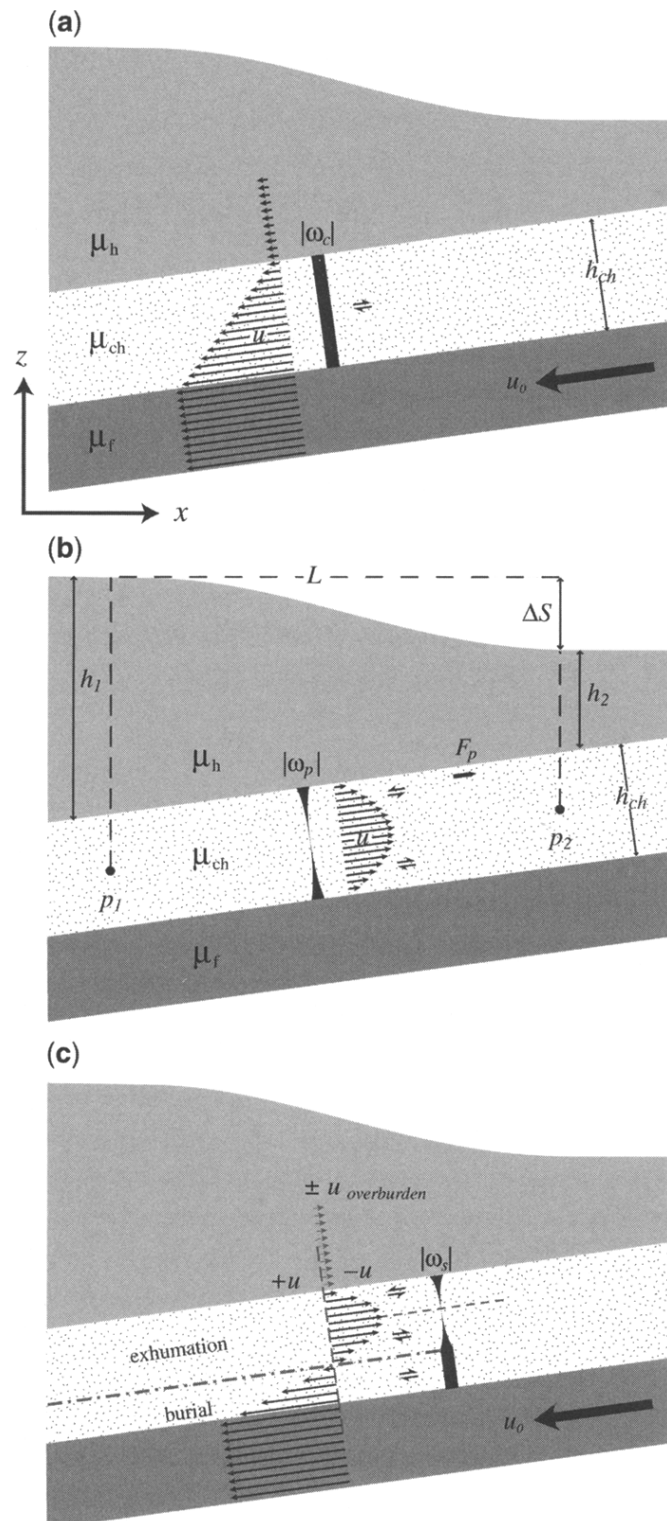


C.



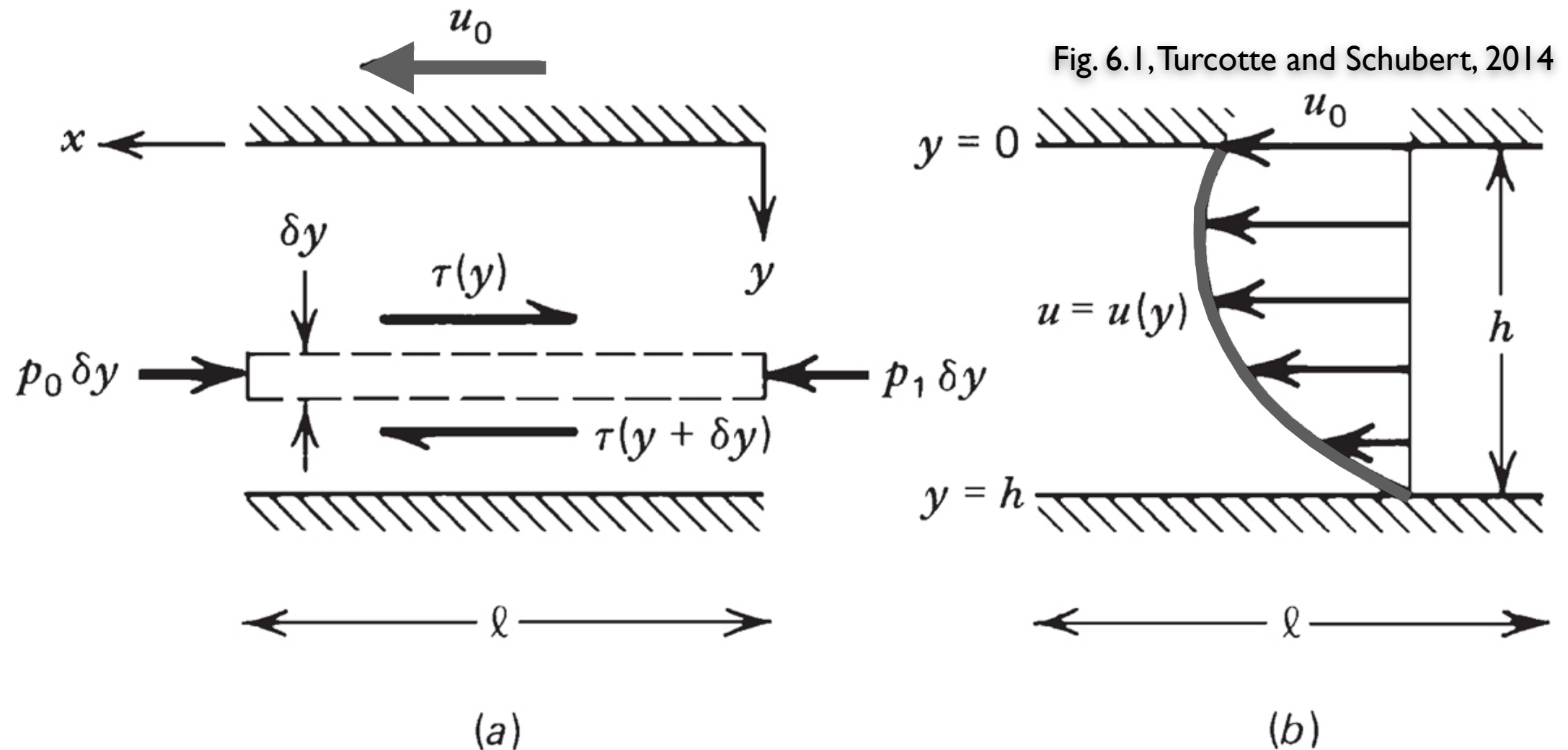
D.

# Channel flows in the Earth



- **Channel flows** in the Earth occur when a fluid flows within a channel, between two solid “walls”
- Such channels can be found in a number of geological settings:
  - Counterflow in the asthenosphere
  - Lower crustal flow
  - Intra-crustal channels (figure on left)
  - Subduction channels
  - Salt tectonic channels

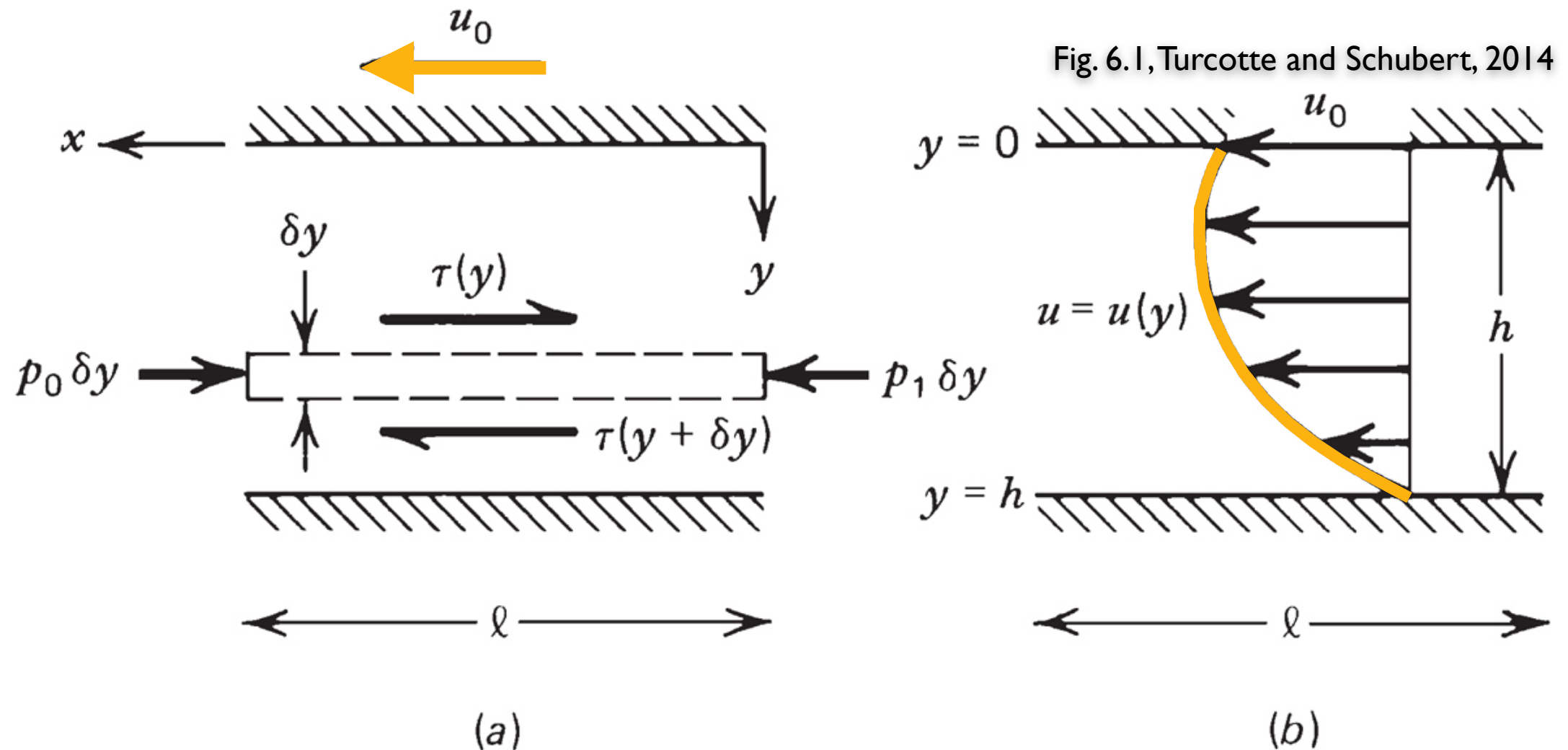
# 1D channel flows



- The most simple fluid flow we can consider is flow of a fluid in one direction within a channel of fixed width

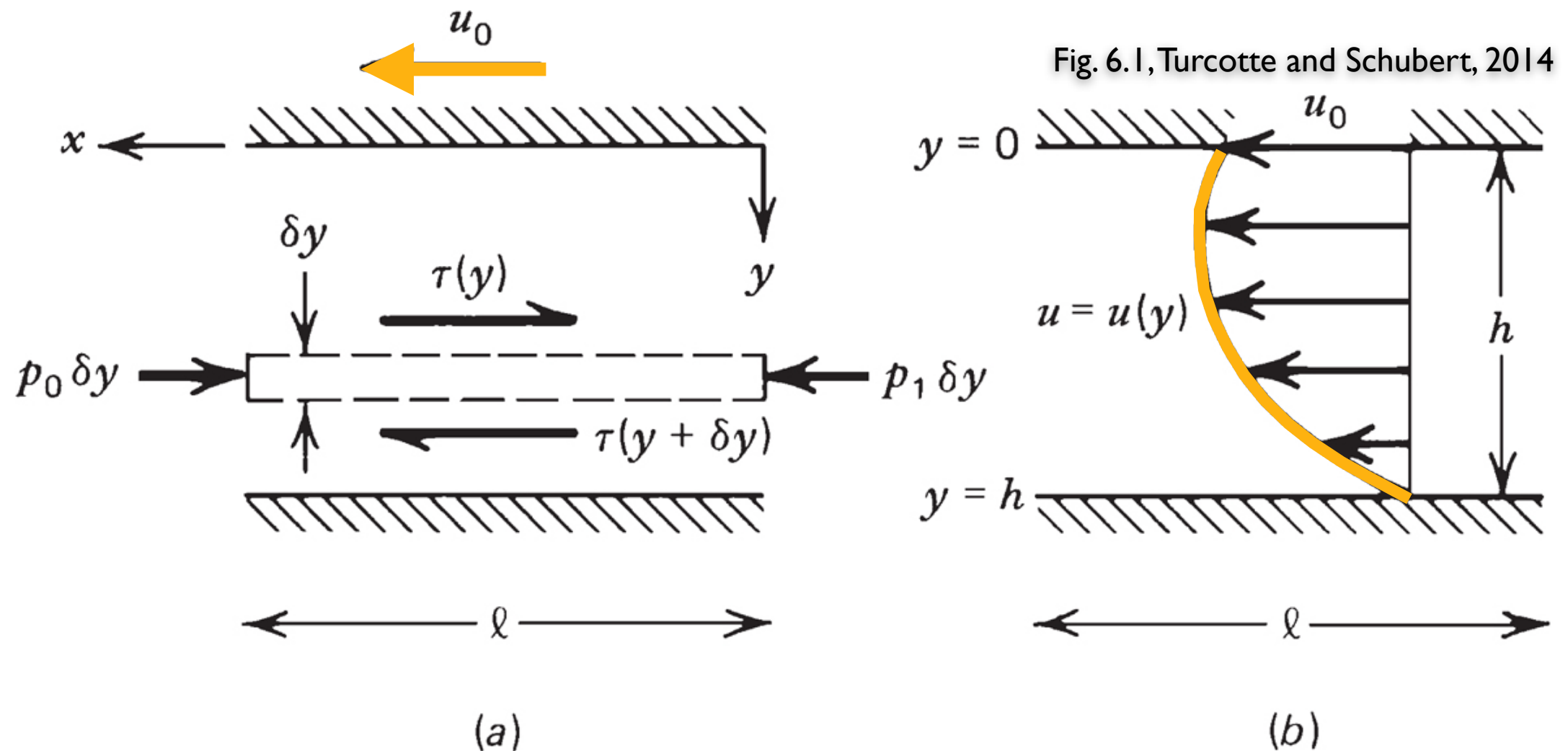


# 1D channel flows



- Fluid is flowing with velocity  $u$  in the  $x$  direction, and the flow velocity  $u$  is a function of distance across the channel  $y$
- Flow results from
  - a **pressure gradient**  $(p_0 - p_1)/l$ , and/or
  - **motion of the side wall** of the channel  $u_0$

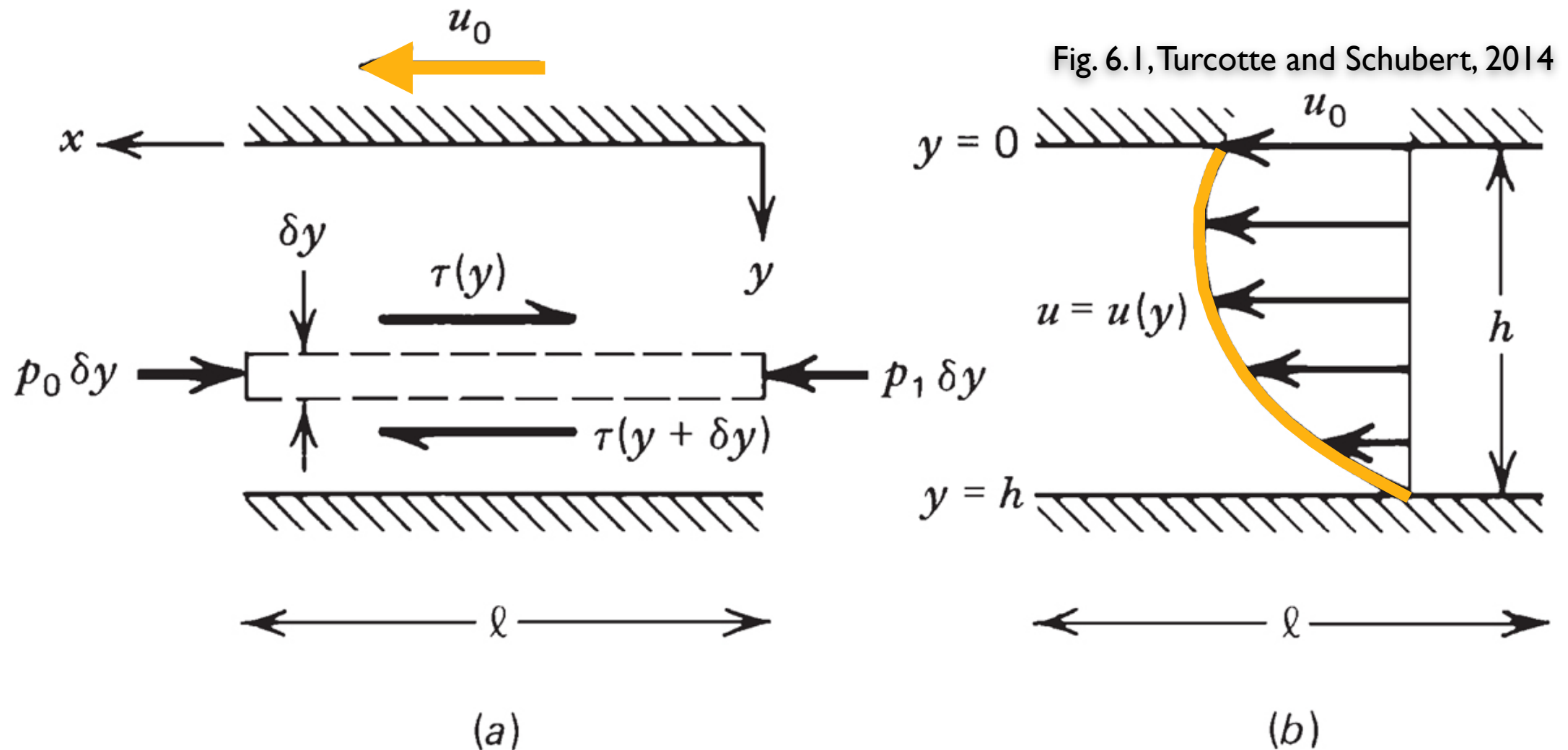
# 1D channel flows



- **Shear**, or a gradient in the velocity, in the channel results in a **shear stress  $\tau$**  that is exerted on horizontal planes in the fluid
- For a Newtonian fluid with a constant **dynamic viscosity  $\eta$**  we can state

$$\tau = \eta \frac{du}{dy}$$

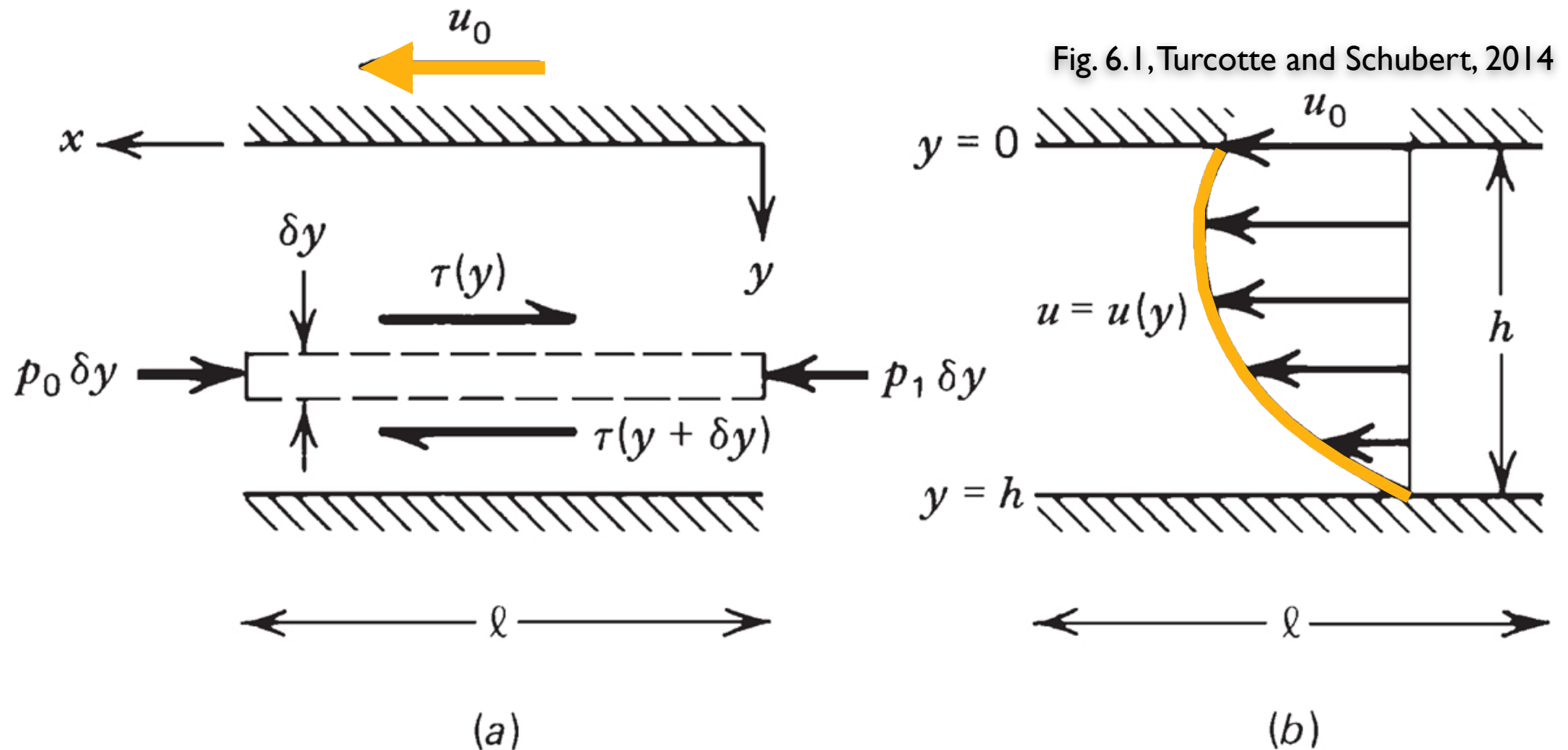
# 1D channel flows



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- For a Newtonian fluid with a constant **dynamic viscosity  $\eta$**  we can state

$$\text{stress} \rightarrow \tau = \eta \frac{du}{dy} \leftarrow \text{strain rate}$$

# 1D channel flows

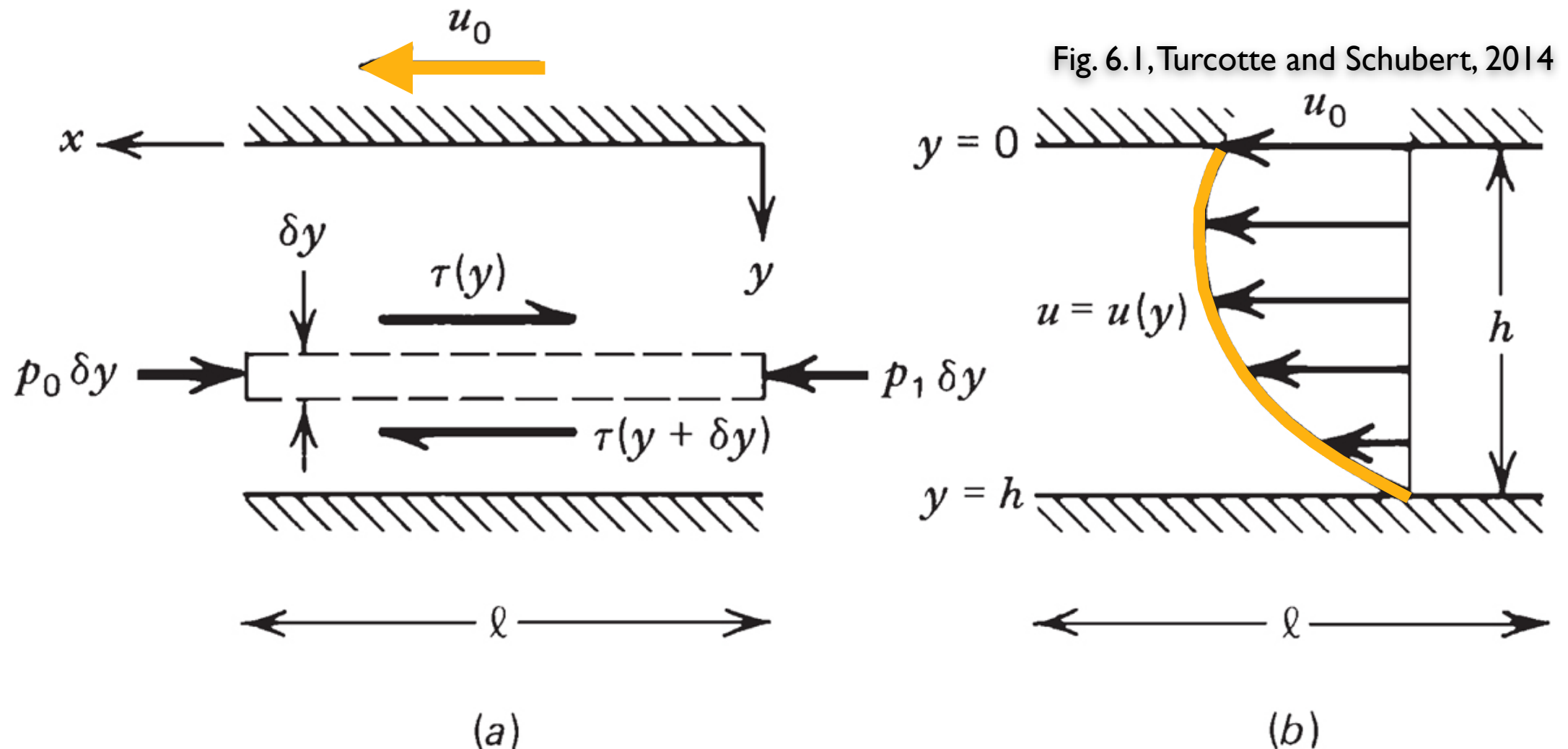


- We can now determine the flow in the channel using the **equation of motion**, based on the force balance on a layer of fluid of thickness  $\delta y$  and length  $l$

- The net **pressure force** on the element in the  $x$  direction is

$$(p_1 - p_0)\delta y$$

# 1D channel flows



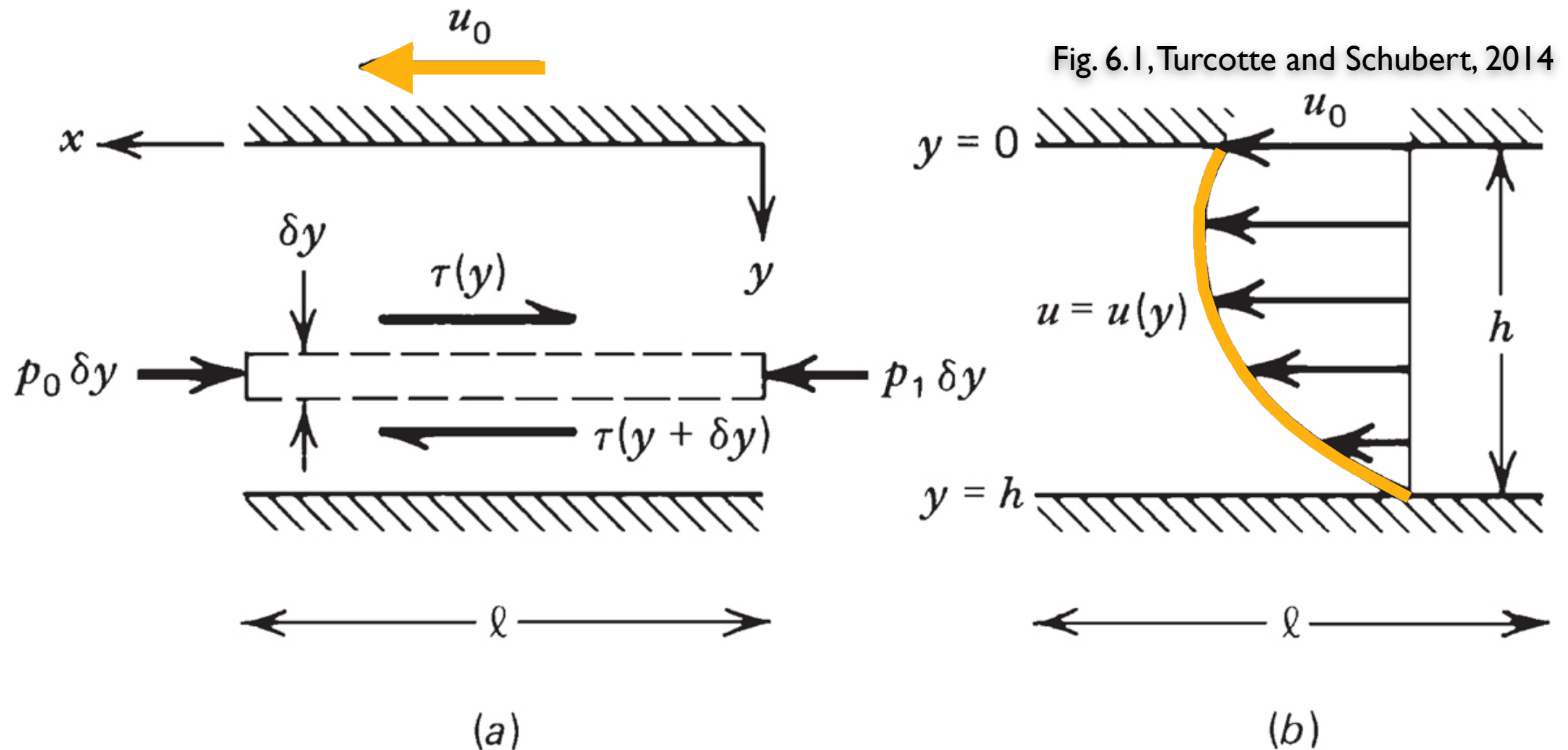
- Because the shear stress  $\tau$  and velocity  $u$  are both only a function of distance  $y$ , the **shear force** on the upper boundary of the element is

$$-\tau(y)l$$

- The equivalent **shear force** on the lower boundary is

$$\tau(y + \delta y)l = \left( \tau(y) + \frac{d\tau}{dy}\delta y \right) l$$

# 1D channel flows



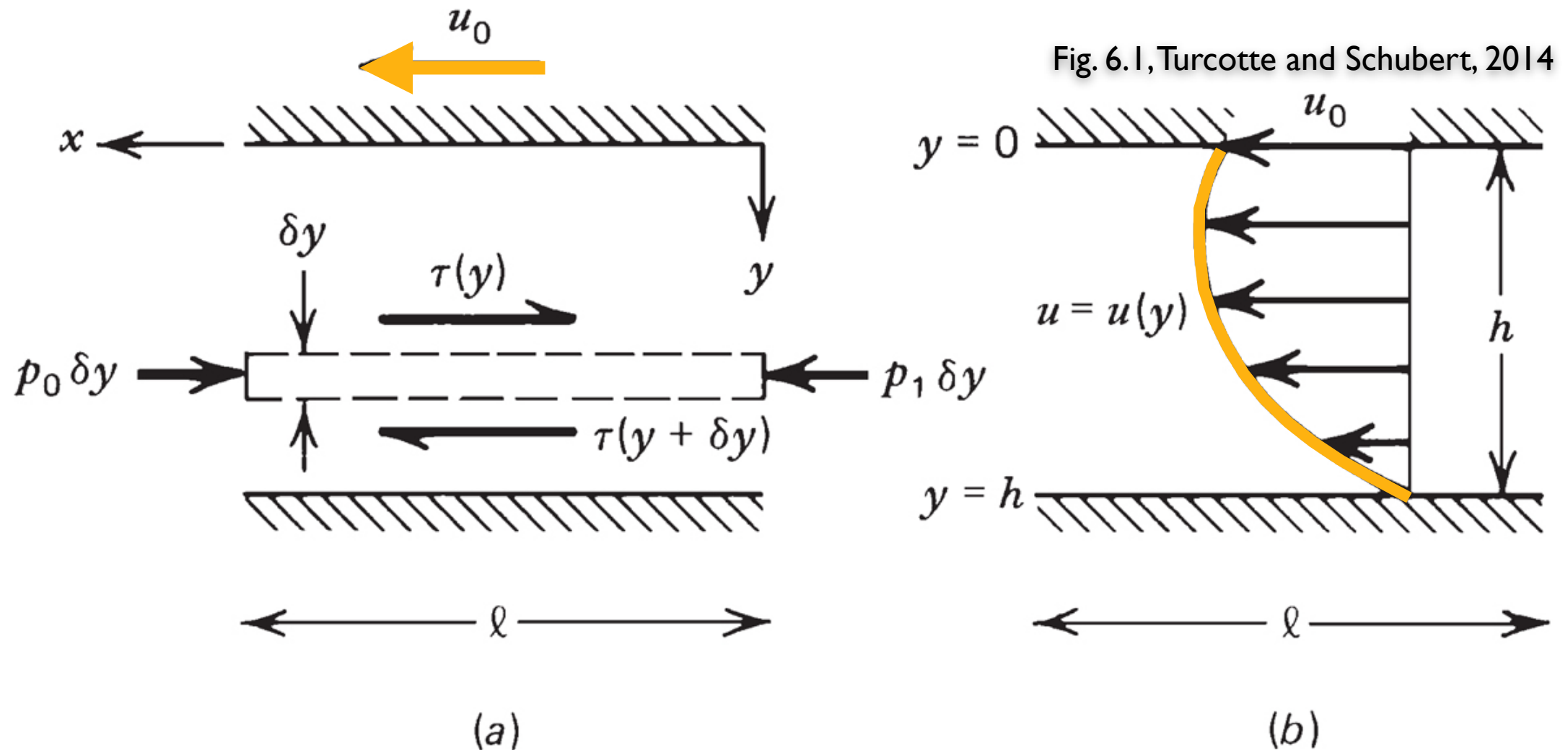
- The **net force** (or sum of the forces) must be equal to zero, or

$$(p_1 - p_0)\delta y + \left[ \tau(y) + \frac{d\tau}{dy}\delta y \right] l - \tau(y)l = 0$$

- As  $\delta y \rightarrow 0$ , the relationship above becomes

$$\frac{d\tau}{dy} = -\frac{(p_1 - p_0)}{l}$$

# 1D channel flows



- The right side of the previous equation is the **horizontal pressure gradient** in the channel

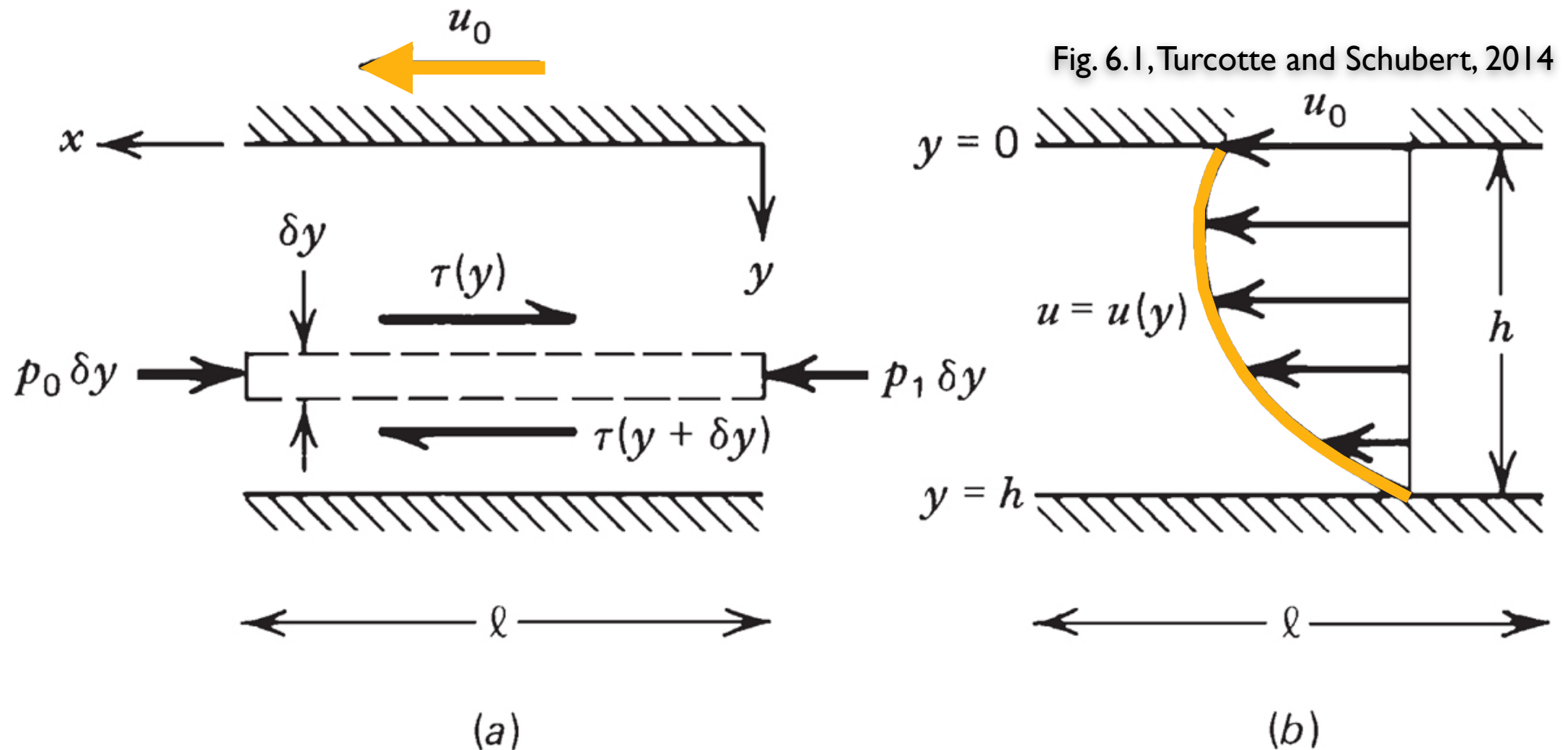
$$\frac{dp}{dx} = -\frac{(p_1 - p_0)}{l}$$

- From which the **equation of motion** can be written

$$\frac{d\tau}{dy} = \frac{dp}{dx}$$



# 1D channel flows



Newtonian  
fluid

$$\tau = \eta \frac{du}{dy}$$

- Velocity in the channel is found by substituting the rheological law for a Newtonian fluid into the equation of motion

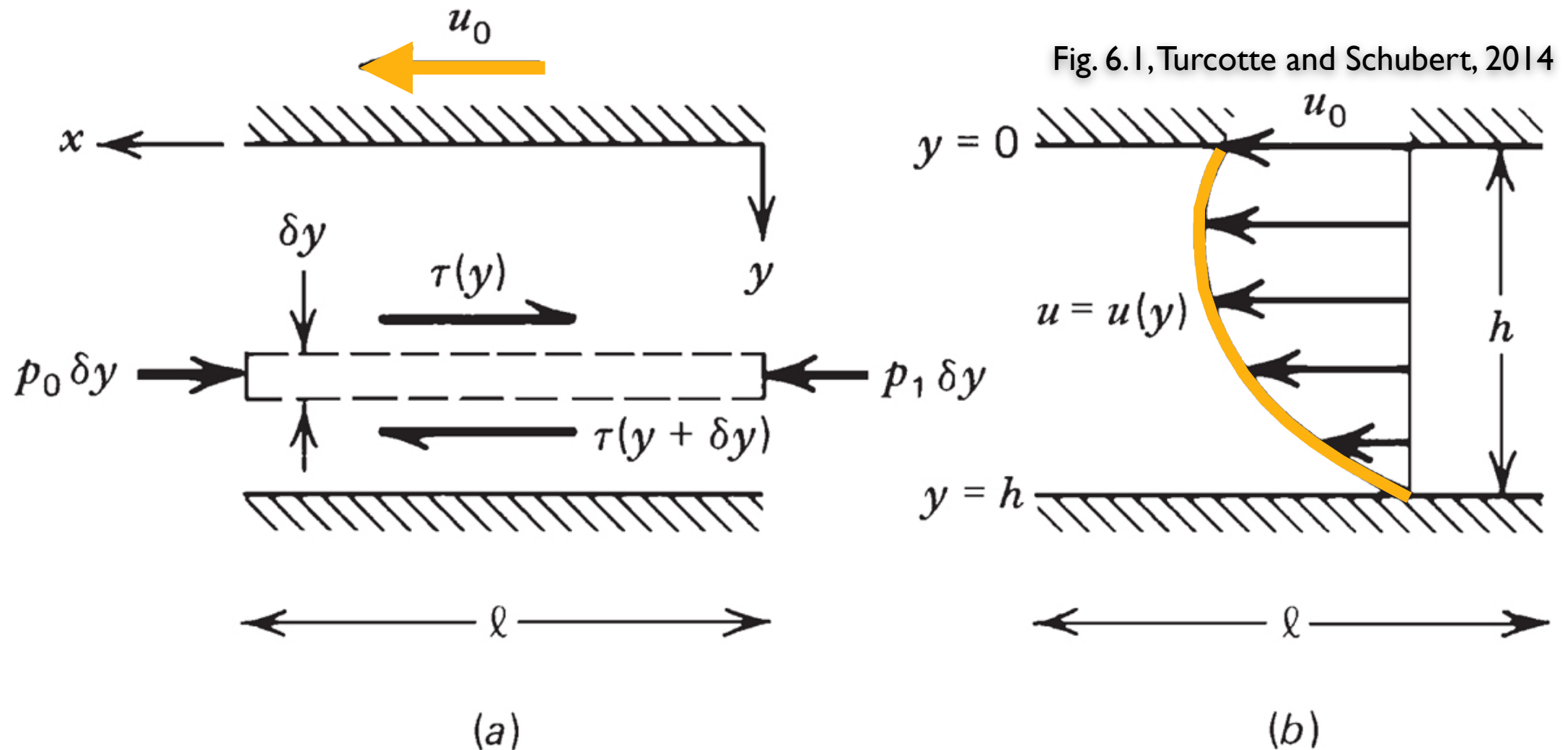
$$\frac{d\tau}{dy} = \frac{d}{dy} \eta \frac{du}{dy} = \eta \frac{d^2 u}{dy^2} = \frac{dp}{dx}$$

- Integrating the equation above twice yields

$$u = \frac{1}{2\eta} \frac{dp}{dx} y^2 + c_1 y + c_2$$



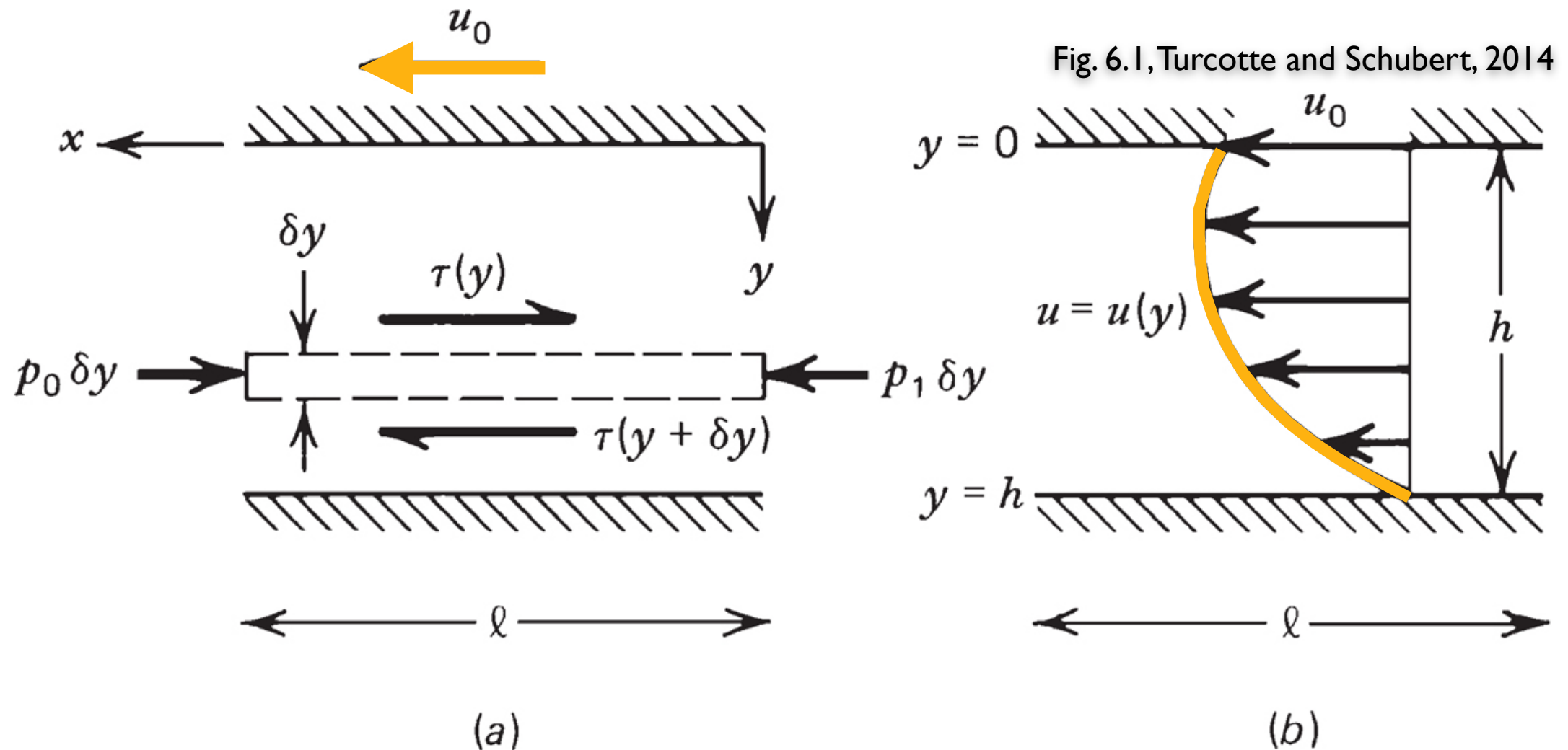
# ID channel flows



- The constants  $c_1$  and  $c_2$  can be found by applying the boundary conditions that  $u = 0$  at  $y = h$ , and  $u = u_0$  at  $y = 0$  (no slip)

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$

# 1D channel flows



- The constants  $c_1$  and  $c_2$  can be found by applying the boundary conditions that  $u = 0$  at  $y = h$ , and  $u = u_0$  at  $y = 0$  (no slip)

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$

pressure gradient

wall velocity



# Channel flow challenge #1

- Start by navigating to the directory [NGWM2016-modelling-course/Lessons/04-Basic-fluid-mechanics/scripts](https://github.com/NGWM2016-modelling-course/Lessons/04-Basic-fluid-mechanics/scripts)
- Right-click on the Python script called `1D-channel-flow.py` and choose “Edit with IDLE”
- The script cannot currently be run because it is missing the equation for the velocity in a 1D channel

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$

your task is to input this equation into the script and then run the script and save plots for the following scenarios:

- No pressure gradient:  $u_0 = 1.0 \text{ mm/a}$ ;  $dp/dx = 0.0 \text{ Pa}$
- No wall velocity:  $u_0 = 0.0 \text{ mm/a}$ ;  $dp/dx = -2000.0 \text{ Pa}$
- Other cases: No pressure gradient/wall velocity, both, etc.



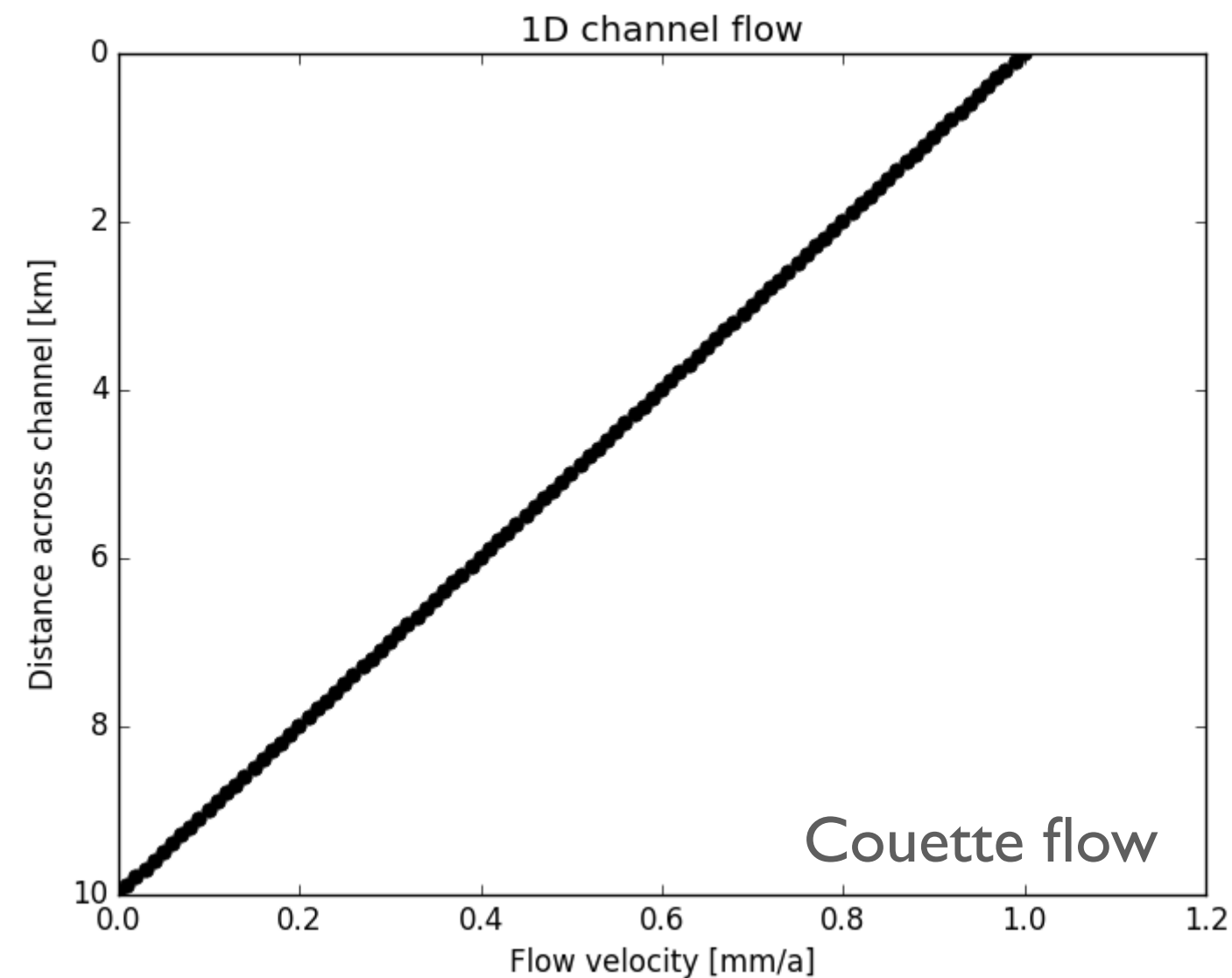
# What does this equation tell us?

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy) - \frac{u_0 y}{h} + u_0$$

- We've now look at several simple fluid flow behaviors, including two important end members
  - Zero pressure gradient ( $dp/dx = 0$ )
  - Zero boundary velocity ( $u_0 = 0$ )



# Couette flow



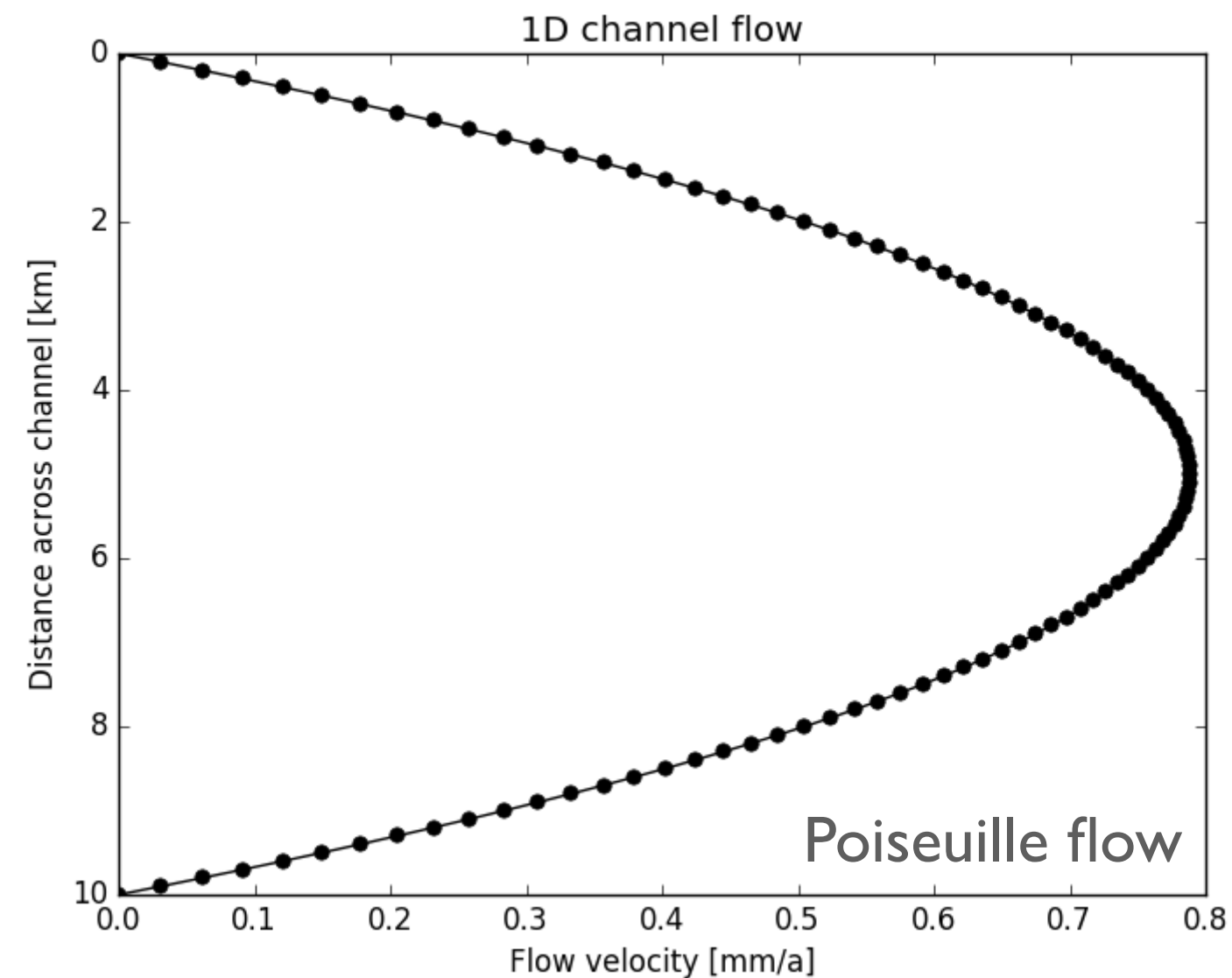
- A **Couette flow** has no pressure gradient, or  $dp/dx = 0$ , reducing the 1D equation for velocity in the channel down to

$$u = u_0 \left( 1 - \frac{y}{h} \right)$$

- Clearly, this predicts a linear increase in velocity from  $y = h$  to  $y = 0$



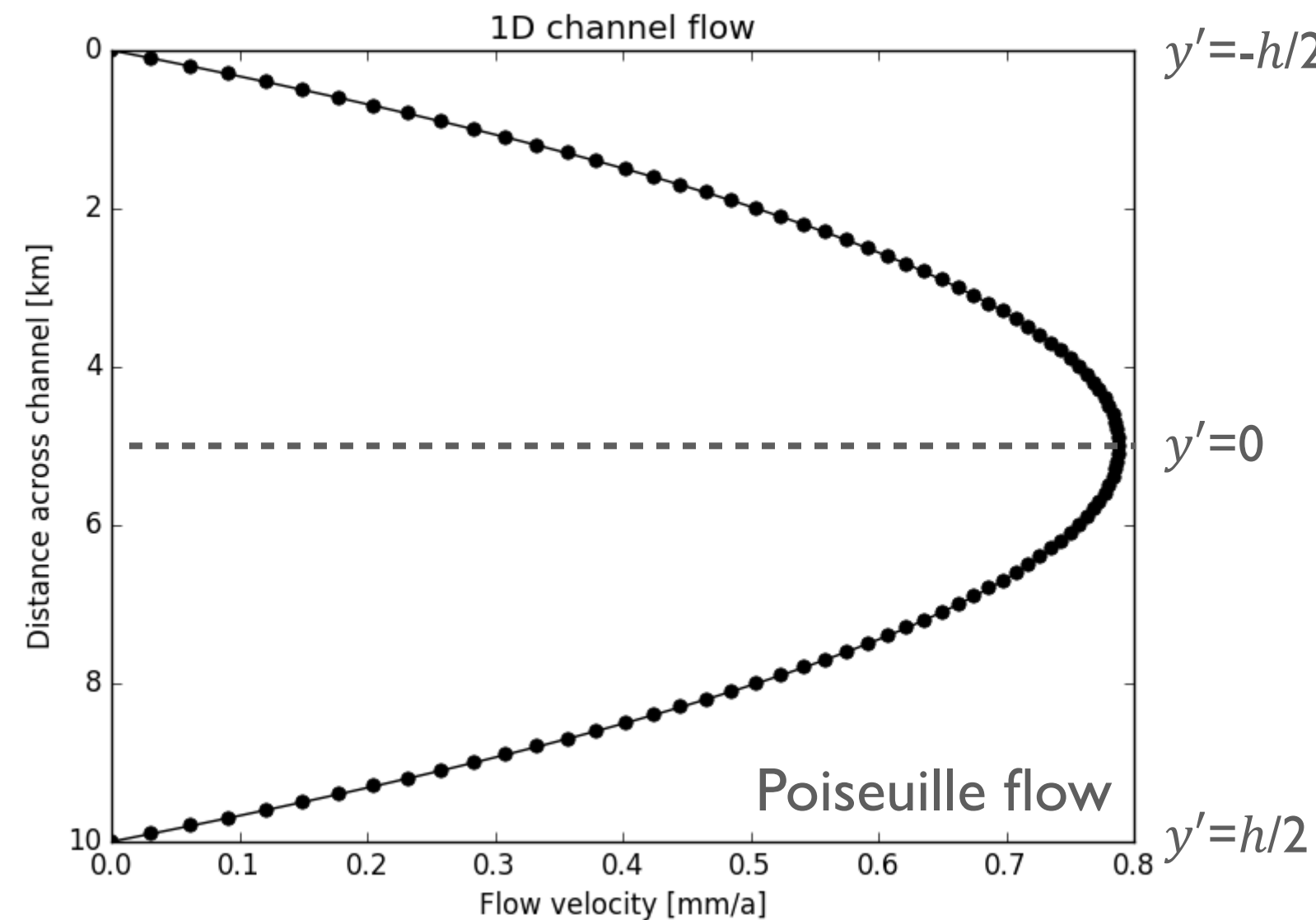
# Poiseuille flow



- **Poiseuille flow** is driven only by a pressure gradient in the channel with zero boundary velocities ( $u_0 = 0$ ), thus

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy)$$

# Poiseuille flow



- **Poiseuille flow** is driven only by a pressure gradient in the channel with zero boundary velocities ( $u_0 = 0$ ), thus

$$u = \frac{1}{2\eta} \frac{dp}{dx} (y^2 - hy)$$

- In a coordinate system with  $y'$  at the middle of the channel we can say  $y' = y - h/2$ , which results in the relationship

$$u = \frac{1}{2\eta} \frac{dp}{dx} \left( y'^2 - \frac{h^2}{4} \right)$$

# Asthenospheric counterflow

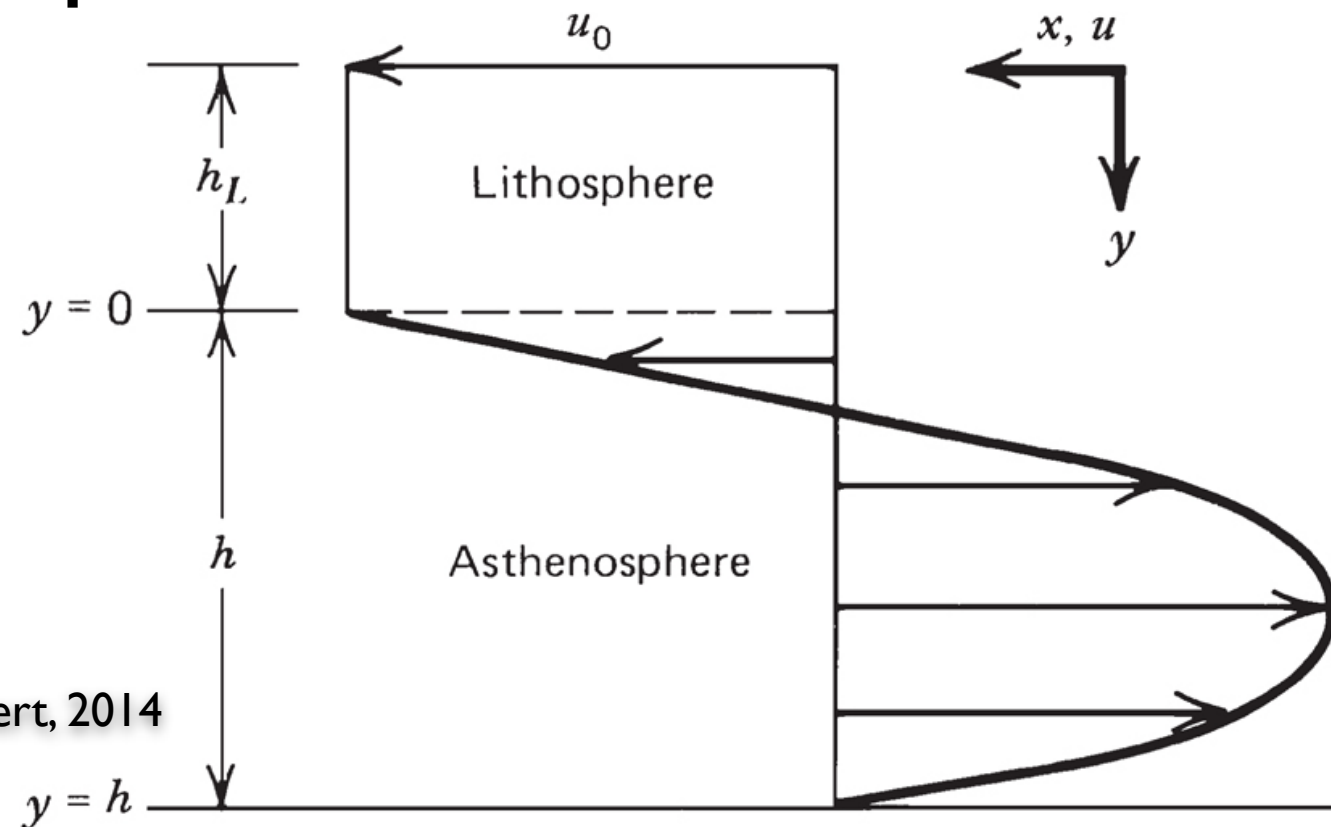


Fig. 6.4, Turcotte and Schubert, 2014

- One model for mantle flow is that the motion of lithospheric plates on the Earth's surface produces a **counterflow** in the uppermost asthenosphere (upper  $\sim 100$ - $200$  km)
- If we assume the plate is rigid and moving at velocity  $u_0$ , and that the velocity at some depth  $y = h$  must be zero, it is clear that the counterflow is opposite in direction to the plate motion in order to **conserve mass**



# Asthenospheric counterflow

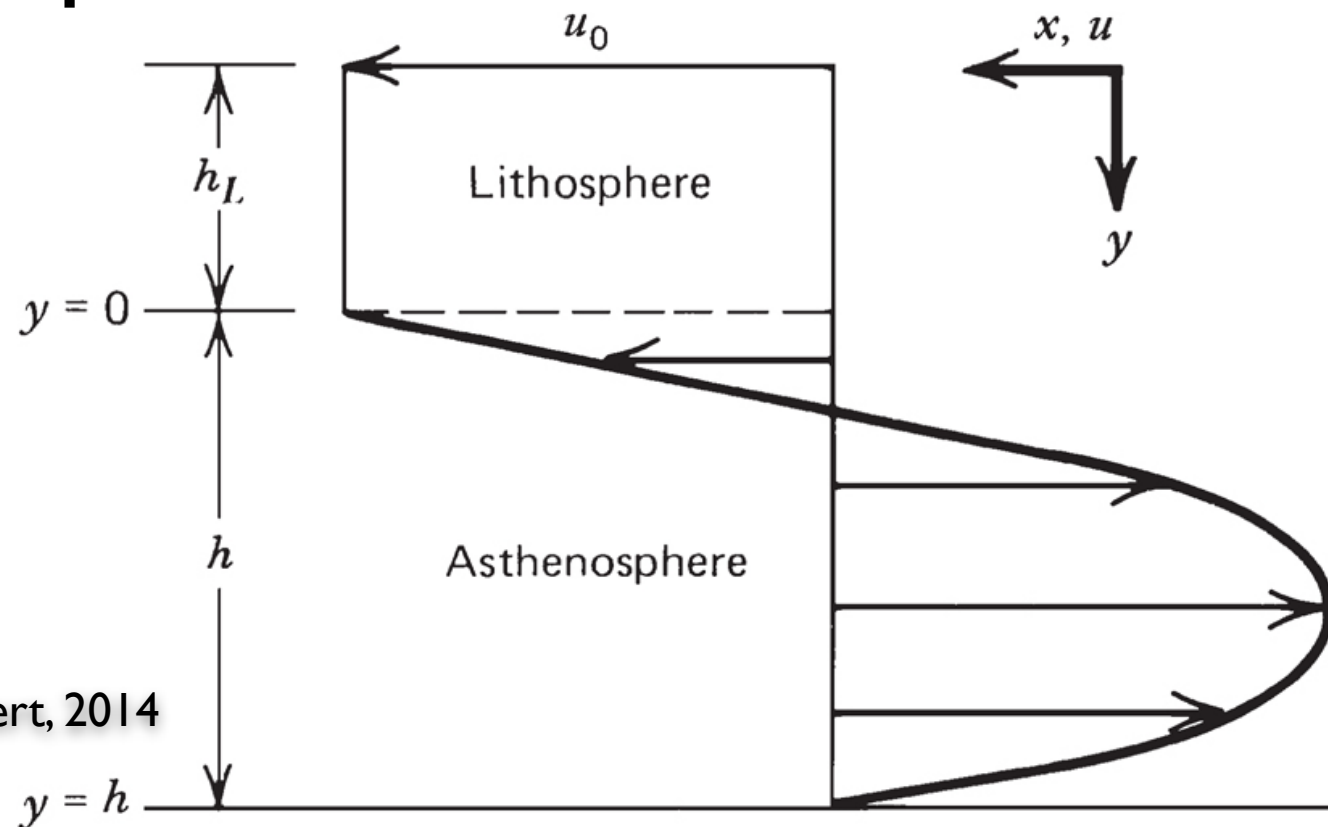


Fig. 6.4, Turcotte and Schubert, 2014

- Mathematically, we can state that as

$$u_0 h_L + \int_0^h u \, dy = 0$$

where  $h_L$  is the thickness of the lithosphere and  $h$  is the thickness of the asthenosphere involved in counterflow

- If we insert our equation for 1D channel flow in the second term, we get

$$u_0 h_L + \frac{h^3}{12\eta} \frac{dp}{dx} + \frac{u_0 h}{2} = 0$$

# Asthenospheric counterflow

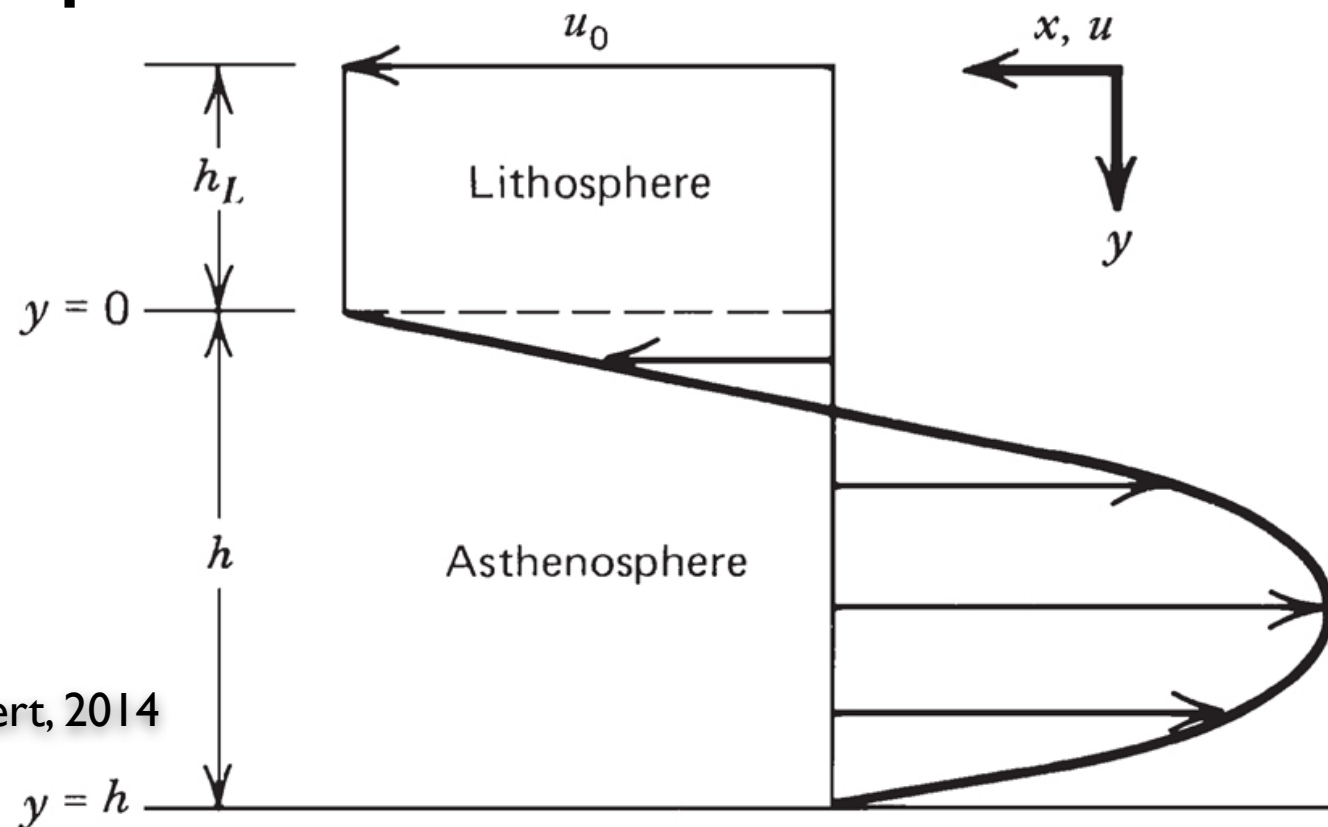


Fig. 6.4, Turcotte and Schubert, 2014

- If we now solve for the **pressure gradient**, we find

$$\frac{dp}{dx} = \frac{12\eta u_0}{h^2} \left( \frac{h_L}{h} + \frac{1}{2} \right)$$

- And this can be inserted into the equation for 1D channel flow to get the predicted velocity profile for a counterflow

$$u = u_0 \left[ 1 - \frac{y}{h} + 6 \left( \frac{h_L}{h} + \frac{1}{2} \right) \left( \frac{y^2}{h^2} - \frac{y}{h} \right) \right]$$

# Asthenospheric counterflow

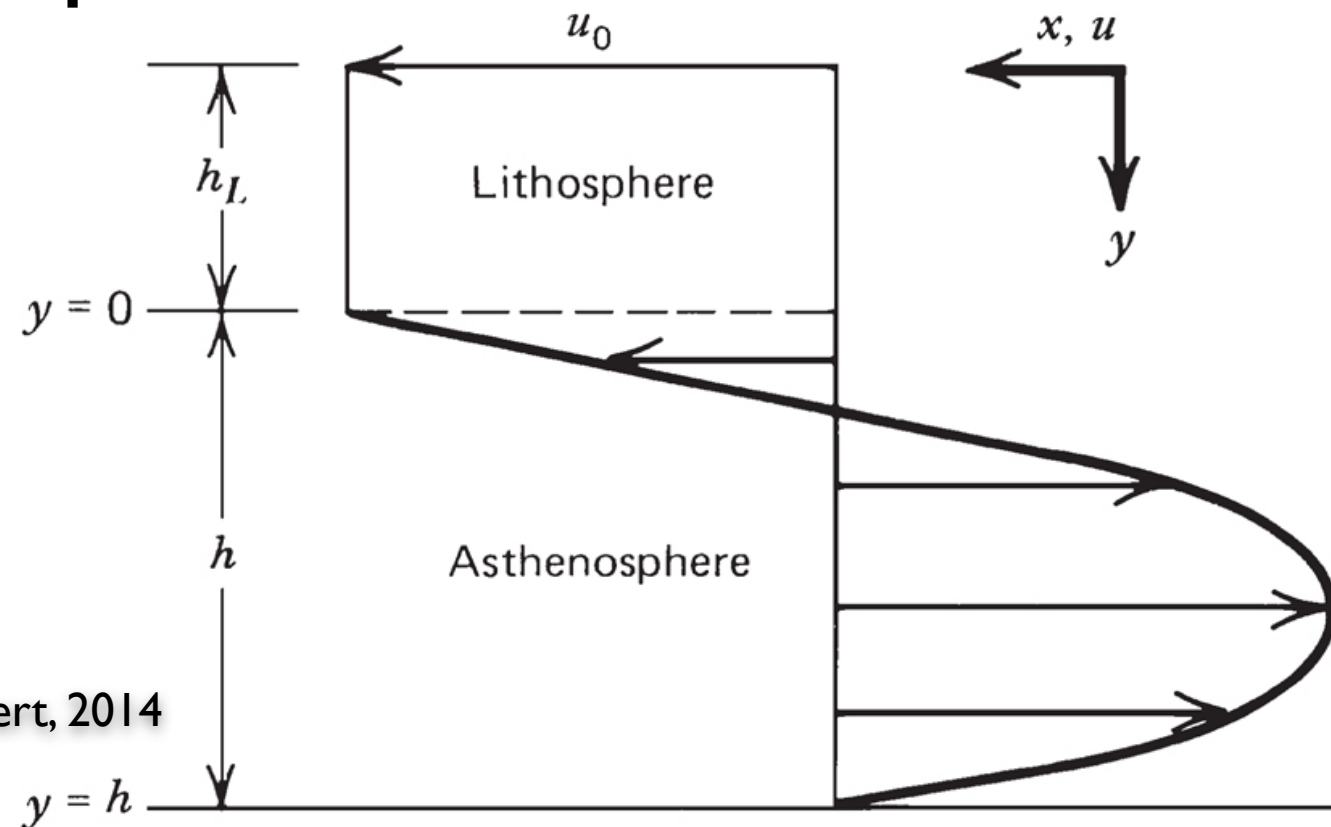


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$$u = u_0 \left[ 1 - \frac{y}{h} + 6 \left( \frac{h_L}{h} + \frac{1}{2} \right) \left( \frac{y^2}{h^2} - \frac{y}{h} \right) \right]$$

- Looking at this equation for a moment, is there anything missing that you might expect to see?



## Channel flow challenge #2

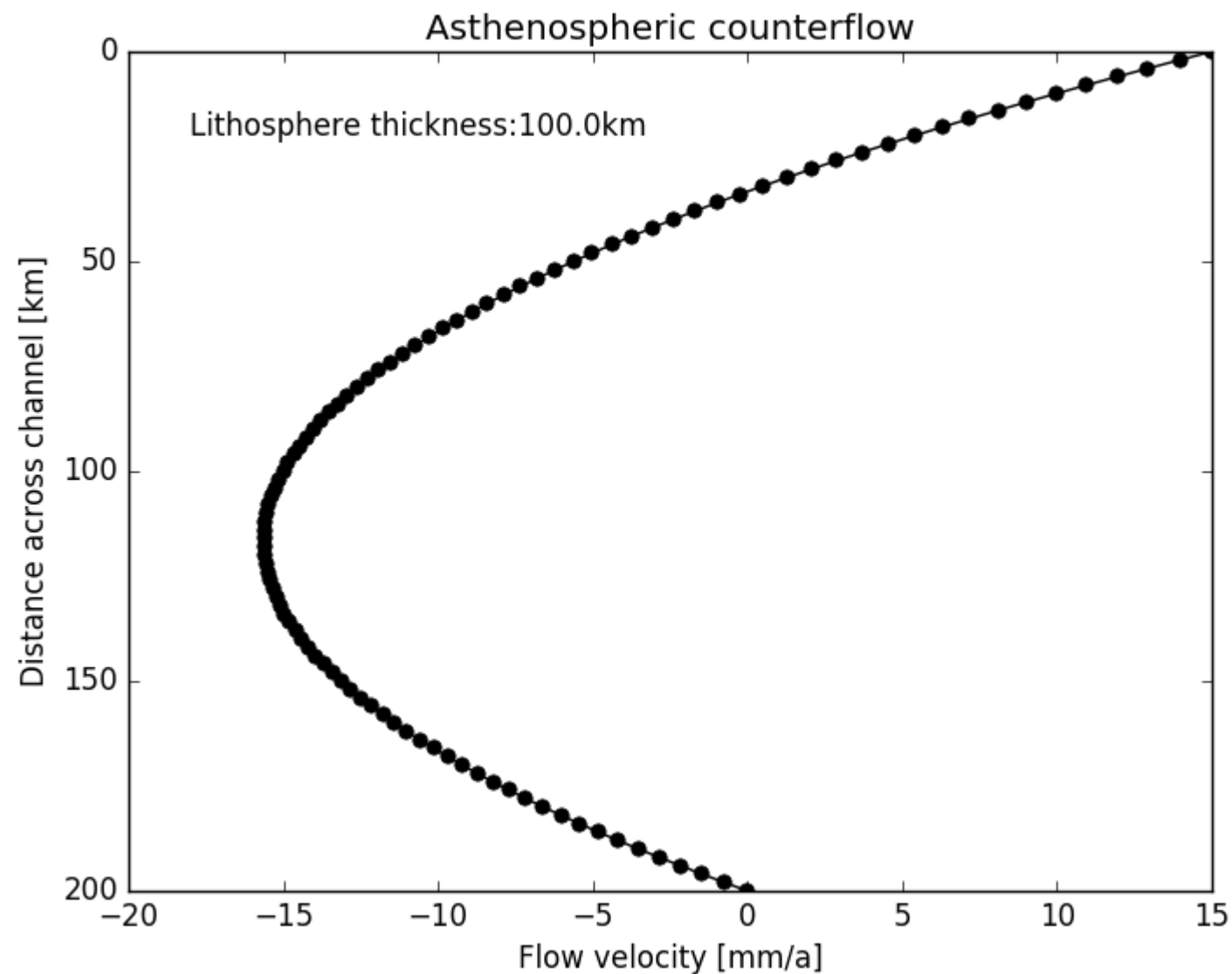
- In the same directory as before, right-click on the script `1D-asthenospheric-counterflow.py` and choose “Edit with IDLE”
- Again, this script cannot be run because it is missing the equation for asthenospheric counterflow

$$u = u_0 \left[ 1 - \frac{y}{h} + 6 \left( \frac{h_L}{h} + \frac{1}{2} \right) \left( \frac{y^2}{h^2} - \frac{y}{h} \right) \right]$$

- In addition, you will need to add a text label to **display the thickness of the lithospheric plate in [km]** on the plot
- You can find how to use the `plt.text()` function at [http://matplotlib.org/api/pyplot\\_api.html#matplotlib.pyplot.text](http://matplotlib.org/api/pyplot_api.html#matplotlib.pyplot.text)
- After making the changes, run the script with the default values and save the plot
- How do the flow velocities change when you vary  $h$  and  $h_l$ ?



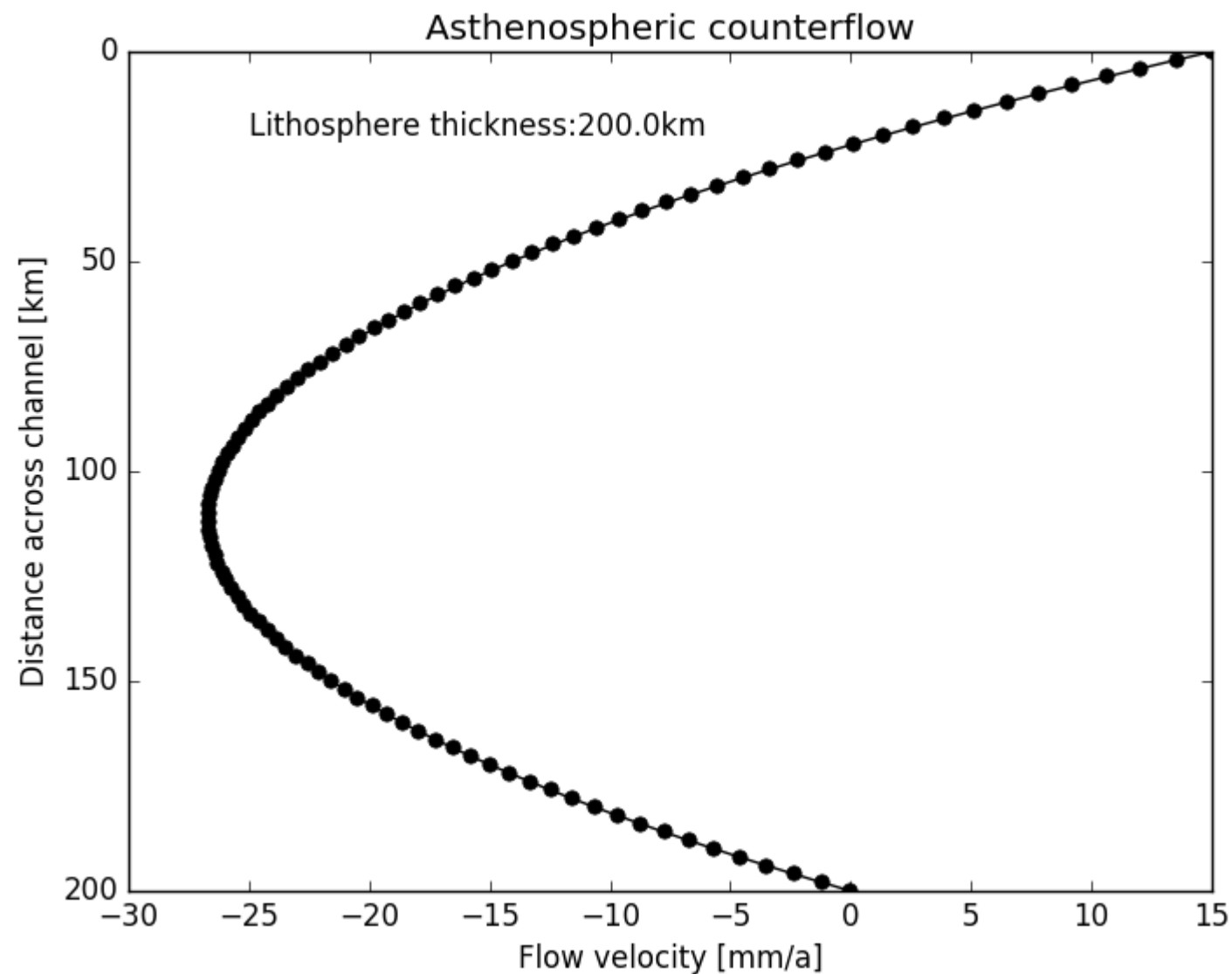
# Asthenospheric counterflow



- As expected, movement of a 100-km-thick lithosphere at  $u_0 = 15 \text{ cm/a}$  results in counter flow in the 200-km-thick asthenosphere
- The maximum counterflow velocity is roughly the same as the plate velocity in this case



# Asthenospheric counterflow



- When the plate thickness is doubled, the counterflow velocity must also increase in order to conserve mass, as required by our counterflow equation
- The maximum counterflow velocity here is about 1.8 times the plate velocity



# Limitations to our channel flow models

- Though we can predict the velocities in 1D channels for a number of different scenarios, we're not able to handle a few things of geological importance
  - Nonlinear viscosity in the channel
    - As you'll see tomorrow morning, viscous flow in rocks is generally not linear
  - Spatial variations in the channel material
    - Perhaps we have a channel with two different fluids
  - Changes in the boundary conditions with time
    - What if the channel wall velocity varies with time?



# Summary

- **Fluid mechanics** is the science of fluid motion
- Fluid motions are caused by **internal and external forces**, and modelled using simple formulations of the **conservation of mass, momentum and energy**
- For geological applications, we treat the **Earth as a fluid with a high viscosity** and model flow using the **Stokes equation**
- Analytical solutions for 1D channel flows can provide insight into fluid flow in the Earth, but have significant limitations





# References

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