

Year 1 – Relativity

Lecture 10

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Overview of lectures

- Lecture 1: Introduction, concepts and classical results
- Lecture 2: The postulates of Relativity
- Lecture 3: Length contraction and simultaneity
- Lecture 4: The Lorentz transformations
- Lecture 5: Space-time diagrams and world lines
- Lecture 6: Four-vectors and causality
- Lecture 7: Energy and momentum
- Lecture 8: Rest mass energy and particle decays
- Lecture 9: Particle reactions
- **Lecture 10: The relativistic Doppler effect**

Previously on Relativity

- Clarified invariant, conserved and constant
 - Some quantities are more than one of these
 - They tend to be the most useful
- Introduced the centre-of-mass frame
 - Often the easiest to do calculations in
- Looked at particle reactions
 - Energy and momentum conservation still holds
 - The mass sum can change from initial to final state
 - The centre-of-mass energy is critical in determining if a reaction can occur

Previously on Relativity

- Saw time slows down for moving objects
 - Time dilation: $T' = \gamma T$
- Looked at the energy and momentum four-vector for photons
 - Satisfies $E = |\underline{p}|c$
- Quantum mechanics gives
 - Planck-Einstein: $E = hf$
 - de Broglie: $\underline{p} = h/\lambda$

What we will do today

- Look at the Doppler shift in relativity
 - Two physical effects
 - Time dilation: only in Relativity
 - Wavelength squeeze and stretch: same as for the classical Doppler effect
- Reproduce the Doppler equations using photon four-momentum
 - Considering it as a Lorentz transform from the frame where the light source is stationary
- Combines both the space-time and the energy-momentum parts of Relativity

Classical Doppler shift

- Waves get “squeezed” as a sound source comes towards you and “stretched” as it moves away

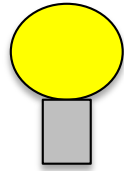
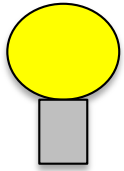


Lower
frequency



Higher
frequency

Time dilation effect



Speed u

Light pulse

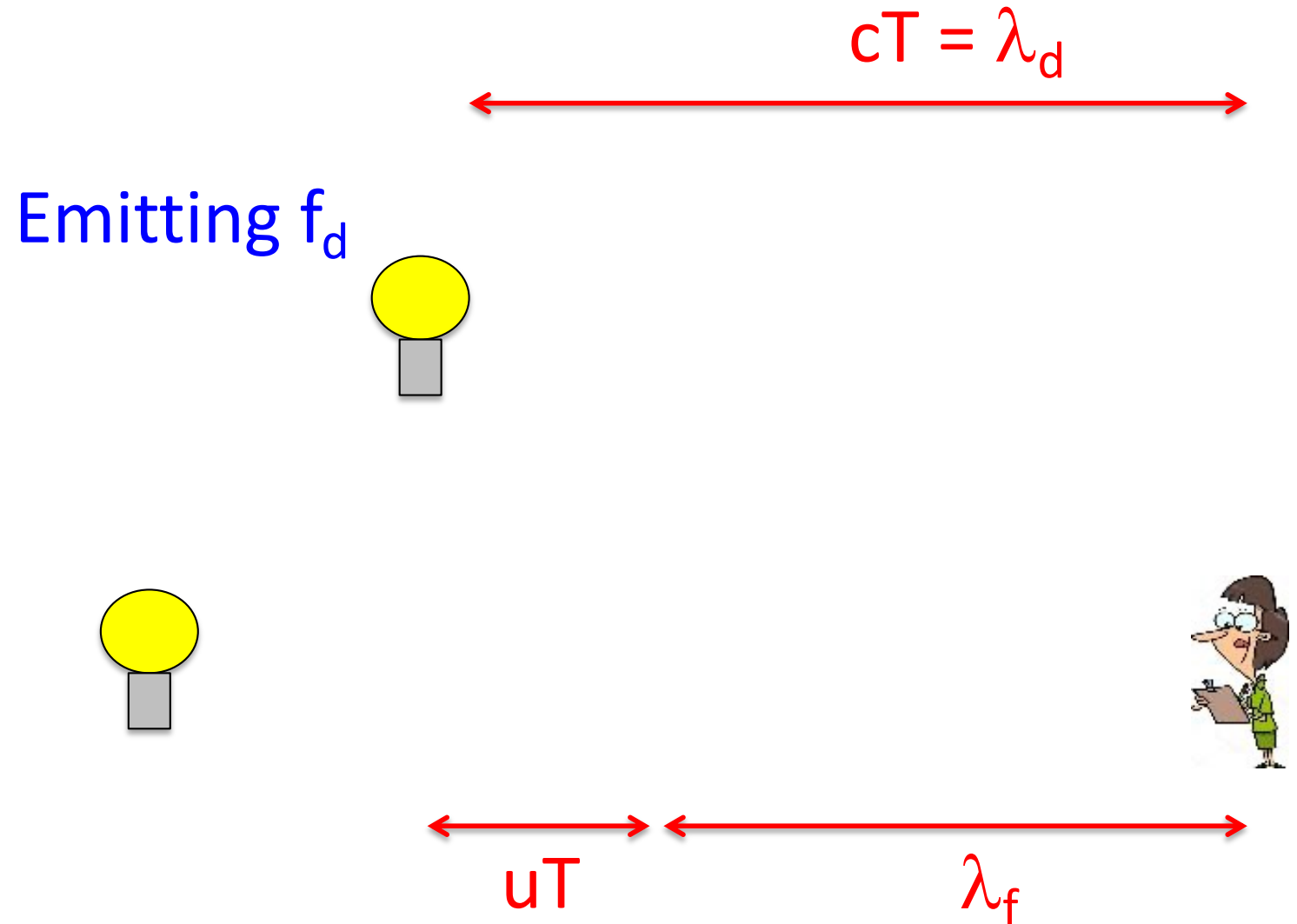


Frequency f



Frequency
 $f_d = f/\gamma_u$

Wavelength squeeze



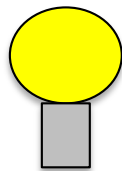
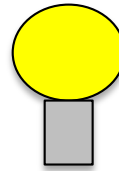
Derivation of resulting frequency

Wavelength stretch

$$cT = \lambda_d$$



Emitting f_d



Derivation of resulting frequency

Menti question

- Go to www.menti.com
- Two effects: time dilation always **decreases** f but forward wavelength squeeze always **increases** f
- Question: **Is total $f_f > f$ (blue-shifted)**
or $f_f < f$ (red-shifted)
 - A. Always blue-shifted
 - B. Always red-shifted
 - C. Depends on the value of β_u

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Some useful expressions

Expansion of the Universe

- Universe is expanding – galaxies further from us moving away faster than those closer to us



Expansion of the Universe

- Universe is expanding – galaxies further from us moving away faster than those closer to us
 - » Speeds the galaxies are moving away usually measured using the red shift of specific lines in spectra from atoms
- Partly related to the Doppler shift we have discussed above but also a GR effect of the actual space between us and the other galaxies being stretched by the expansion itself

Doppler shift from four-momentum

- A tank travelling at speed v that fires a shell with speed u will classically give a shell with speed $u+v$
- Shell “thrown forwards” by tanks motion
- If instead tank were firing photons, we could not increase their speed – always travel with c – but can still give extra momentum and hence energy

Lorentz transformation of photon

Rest frame of source



Moving frame of source = rest frame of observer

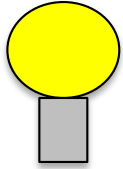


Requires Lorentz transformation in $-x$ direction

Lorentz transformation of photon

Time dilation effect again

Moving frame of source

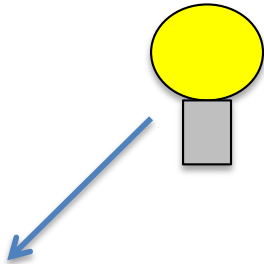


Photon
($E', \underline{p}'c$)



Rest frame of source

Photon
($E, \underline{p}c$)



Time dilation effect again

Change in angle - an Example

- A beam of light is emitted at an angle θ to the x-axis in a reference frame, S
- Show that in a reference frame S' moving with velocity v in the positive x-direction, the beam of light still has speed c
- Show that it is emitted at an angle θ' to the x' axis, where, $\cos \theta' = (\cos \theta - \beta) / (1 - \beta \cos \theta)$

What we did today

- Looked at the Doppler shift in relativity
 - Saw there were two physical effects
 - Time dilation: only in Relativity
 - Wavelength squeeze and stretch: same as for the classical Doppler effect
- Reproduced the Doppler equations using photon four-momentum
 - Is easier to do but does not make the separate physical effects as clear

OMG – we
survived the
whole lecture...

