

Year 1 – Relativity

Lecture 8

Mitesh Patel

Overview of lectures

- Lecture 1: Introduction, concepts and classical results
- Lecture 2: The postulates of Relativity
- Lecture 3: Length contraction and simultaneity
- Lecture 4: The Lorentz transformations
- Lecture 5: Space-time diagrams and world lines
- Lecture 6: Four-vectors and causality
- Lecture 7: Energy and momentum
- **Lecture 8: Rest mass energy and particle decays**
- Lecture 9: Particle reactions
- Lecture 10: The relativistic Doppler effect

Previously on Relativity

- Met the energy and momentum four-vector
 - Combined as $(E, \underline{p}c)$
 - Changes between frames using Lorentz transformations (like all four-vectors)
- Properties of energy and momentum

- When at rest:

$$E = mc^2, \quad \underline{p} = \underline{0}$$

- When moving with velocity \underline{u}

$$E = \gamma_u mc^2, \quad \underline{p} = \gamma_u m \underline{u}$$

- The length-squared invariant is $(mc^2)^2 = E^2 - p^2 c^2$
- Photons are special case with $E = |\underline{p}|c$

What we will do today

- Look at what mass means
 - Not the same as in classical physics
- Discuss energy and momentum conservation
 - Both are conserved, as is the case classically
 - But mass conservation is very different from the classical situation
- Look at the kinematics of particle decays
 - Use energy and momentum conservation to find what happens

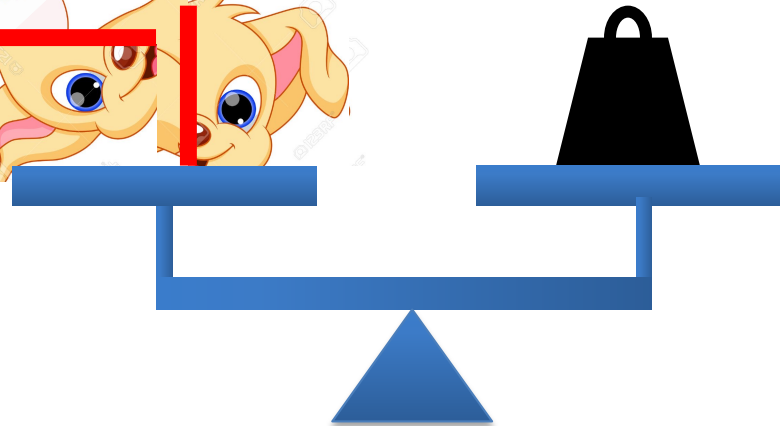
Scalars and vectors

- Before this course, you knew about scalars and vectors
 - These properties are defined by their rotations
 - These are often called “three-scalars” and “three-vectors” where “three” is shorthand for “property under rotation”
 - Often use the term “invariant” instead of “scalar”
- We now have four-scalars and four-vectors
 - Where “four” means “property under Lorentz transformations”
- Mapping one to the other is not obvious
 - Some three-scalars are also four-scalars; e.g. m
 - Some three-scalars are part of four-vectors; e.g. E , t
 - Some three-vectors are parts of four-vectors; e.g. \underline{r} , \underline{p}
 - Some three-vectors form other structures; e.g. \underline{v} , \underline{E} & \underline{M} fields

Units

- Almost all experimental tests of Special Relativity have involved particle physics
 - Mass and hence energy scales are very small
 - SI units are inconvenient; use eV energy units
 - An electron gains 1 eV of energy in moving through 1V, so $1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$
 - Also use keV (10^3 eV), MeV (10^6 eV), GeV (10^9 eV)...
- Dimensions: $[p] = [E/c]$ and $[m] = [E/c^2]$
 - Can also give p in units of eV/c, m in eV/c²
 - E.g. electron rest mass energy $mc^2 = 8.2 \times 10^{-14} \text{ J} = 0.511 \text{ MeV}$
 - Therefore electron mass $m = 0.511 \text{ MeV}/c^2$

Mass in classical physics



Sum of masses of all objects
is conserved classically

Mass in relativity



allen-electronic



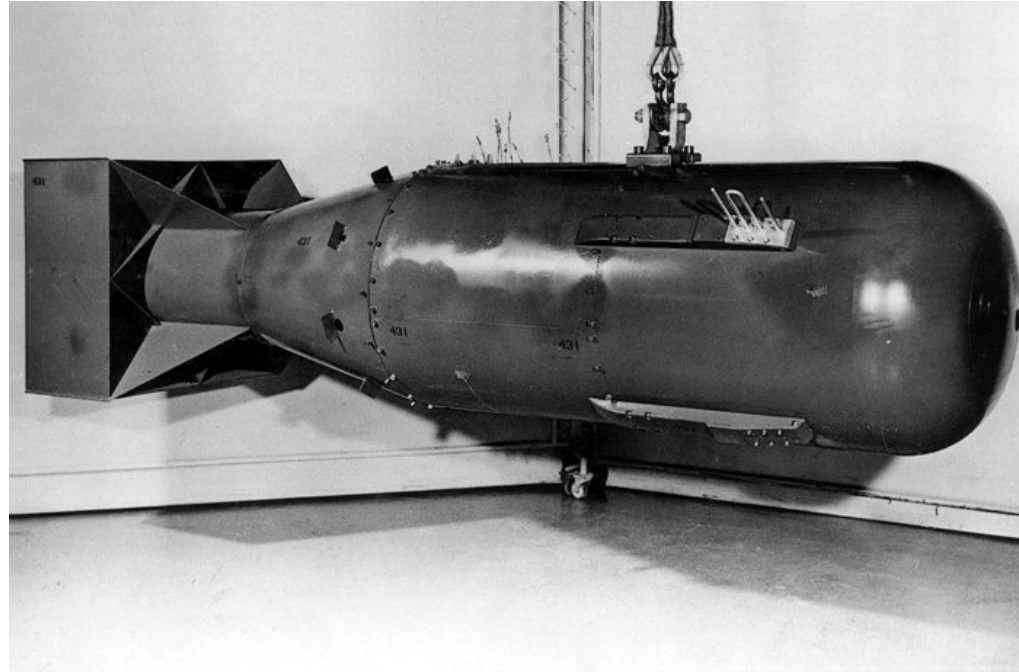
Classical vs Relativistic

- Bring e and p together to make a hydrogen atom
 - A 13 eV photon is radiated off a when e-p bind together
- Classical view
 - Mass of H atom = $m_e + m_p$
 - H atom has binding energy of 13 eV
- Relativistic view
 - Initial energy = $(m_e + m_p)c^2$
 - Energy lost through photon = 13 eV
 - Mass of hydrogen atom = $m_e + m_p - 13 \text{ eV}/c^2$
 - Atom is **lighter** than $m_e + m_p$
 - Reduction factor: $13 \text{ eV}/(m_e + m_p)c^2 \sim 10 \text{ eV}/10^9 \text{ eV} \sim 10^{-8}$
 - Very hard to detect as “chemical” energies $\ll m_p c^2$

Example – uranium

- Uranium-235 undergoes induced decay ('fission') when it is hit by a slow neutron
 - It usually splits into two other nuclei and several neutrons
 - The average mass change is $\sim 200 \text{ MeV}/c^2$
 - Uranium metal contains $\sim 5 \times 10^{28}$ nuclei per m^3
- Estimate the energy (in J) released if all the nuclei in a $10 \times 10 \times 10 \text{ cm}^3$ block of uranium fissioned

Little Boy: first atomic bomb



- Contained a cylinder of uranium
 - 16cm diameter x 18 cm long
 - Total energy release = 63 TJ = 6.3×10^{13} J
 - Inefficient as not all U-235, not all fissioned, etc.

Mass in relativity

- Both classically and relativistically, energy and momentum are conserved
- BUT – the mass of an object can change, so the sum of the masses will not in general be conserved in reactions
 - initial $\sum_i m_i \neq \text{final } \sum_i m_i$
 - Classically, mass appears to be conserved because we only rearrange electrons and nuclei – changes in mass then v small

- However, we know that

$$E_T^2 = p_T^2 c^2 + m_T^2 c^4 \quad \rightarrow \quad m_T = \sqrt{(E_T^2 - p_T^2 c^2)}/c^2$$

since E and p are conserved, m_T is conserved

Higgs boson decay at rest

Before decay

Higgs m_H



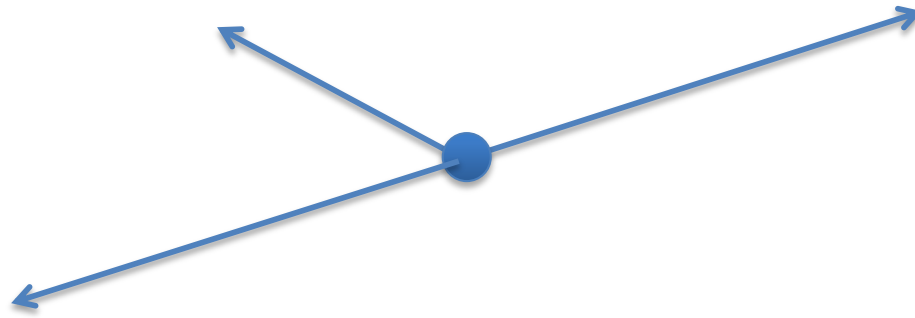
$$E_T = E_H = m_H c^2, \quad \underline{p}_T = \underline{p}_H = \underline{0}, \quad m_T = m_H$$

After decay

Photon 2

Photon 1

Photon 2



$$E_T = E_{\gamma 1} + E_{\gamma 2}, \quad \underline{p}_T = \underline{p}_{\gamma 1} + \underline{p}_{\gamma 2} = \underline{0} \quad \text{Momentum conservation}$$

Moving Higgs boson decay

Before decay

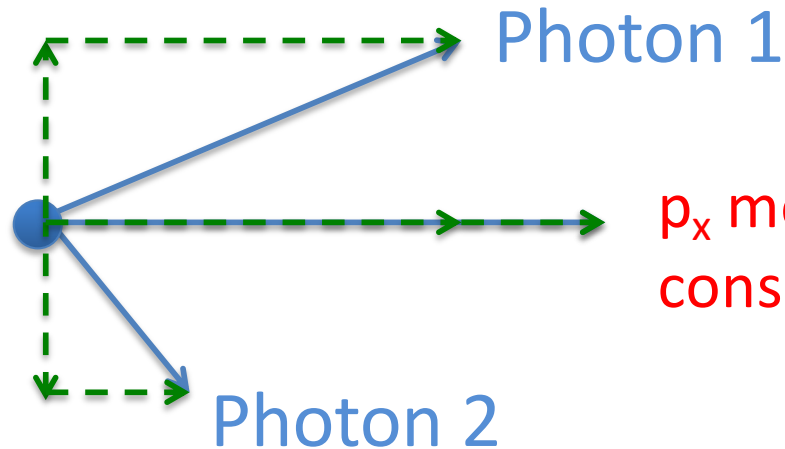
Higgs m_H



$$E_T = E_H, \quad \underline{p}_T = \underline{p}_H, \quad m_T = m_H$$

After decay

p_y momentum
conservation

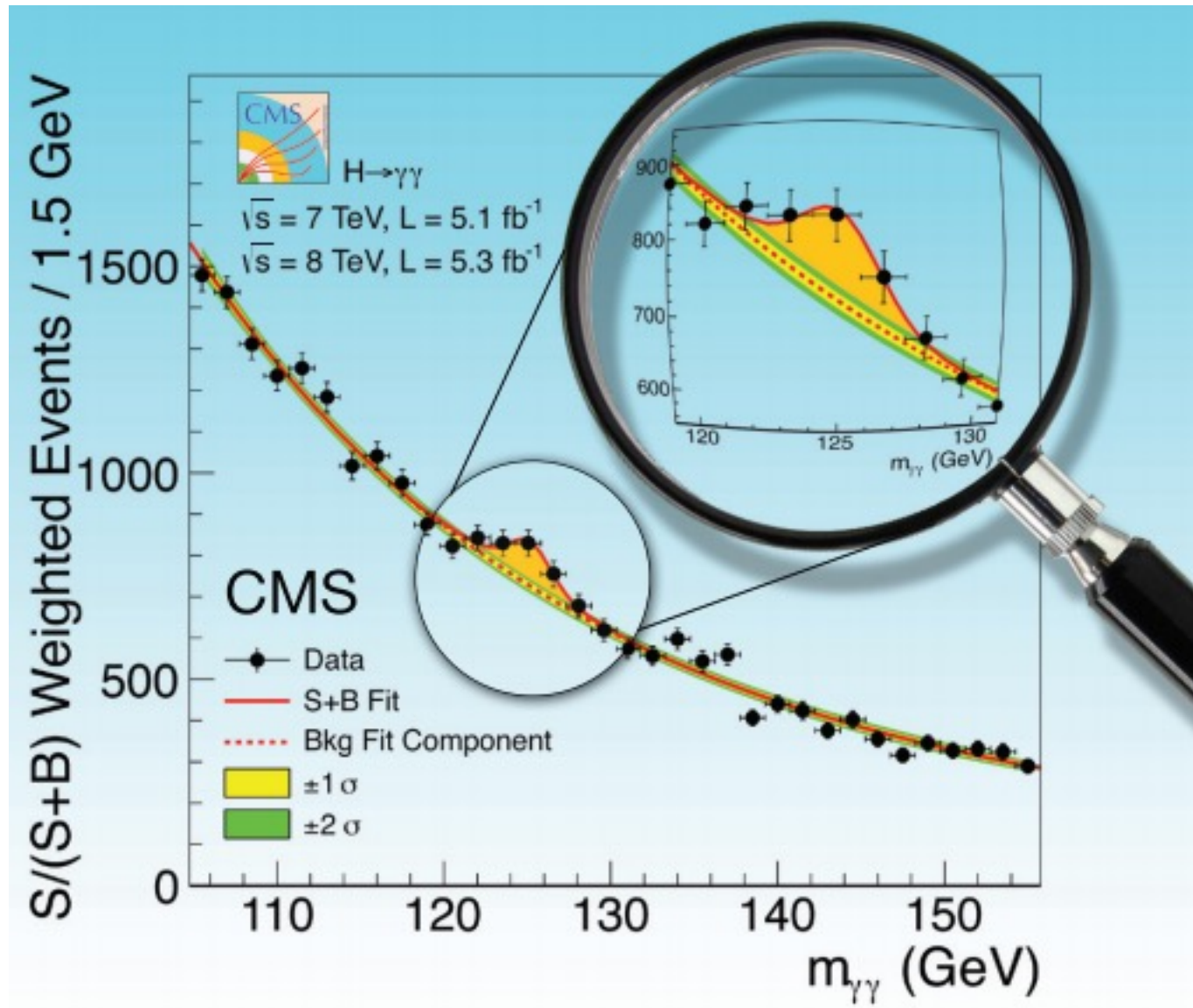


p_x momentum
conservation

$$E_T = E_{\gamma 1} + E_{\gamma 2}, \quad \underline{p}_T = \underline{p}_{\gamma 1} + \underline{p}_{\gamma 2}$$

Solving for the energies

The Higgs discovery



Example – pion decay

- A charged pion decays to a muon and a neutrino
 - Decay formula: $\pi^+ \rightarrow \mu^+ + \nu$
- The neutrino mass is negligibly small and the other masses are
 - $m_\pi = 139.6 \text{ MeV}/c^2$, $m_\mu = 105.7 \text{ MeV}/c^2$
- Find the change in the sum of the masses
- Find the total and kinetic energies of the muon and the neutrino in the pion rest frame

Solution for example – pion decay

- The masses are
 - $m_\pi = 139.6 \text{ MeV}/c^2$, $m_\mu = 105.7 \text{ MeV}/c^2$, $m_\nu = 0 \text{ MeV}/c^2$
 - so the mass change is $\Delta m = 139.6 - 105.7 - 0 = 33.9 \text{ MeV}/c^2$
- Using the formula, the total energy of the muon is
 - $E_\mu = [(139.6)^2 + (105.7)^2 - (0)^2]/(2 \times 139.6) = 109.8 \text{ MeV}$
 - so $K_\mu = E_\mu - m_\mu c^2 = 109.8 - 105.7 = 4.1 \text{ MeV}$
- Similarly, the total energy of the neutrino is
 - $E_\nu = [(139.6)^2 + (0)^2 - (105.7)^2]/(2 \times 139.6) = 29.8 \text{ MeV}$
 - so $K_\nu = E_\nu - m_\nu c^2 = 29.8 - 0 = 29.8 \text{ MeV}$
- Cross check: the total kinetic energy is
 - $K_\mu + K_\nu = 4.1 + 29.8 = 33.9 \text{ MeV} = \Delta mc^2$ as expected

What we did today

- Discussed the meaning of mass
 - Arises from all energy in the system
 - Total of individual object masses is not conserved
 - But total mass of system is conserved
- Discussed energy and momentum conservation
 - E and \underline{p} conservation allows us to solve kinematic problems in particle decays
 - The sum of the particle masses is not always conserved, but the total invariant mass is
 - This allows us to identify particles by only measuring their decay products