Solving the Black Scholes PDE with Holger's reproducing kernel. In this example, MQ kernel is used. The two-dimensional PDE is approximated with collecting of order-one-functions.

For d dimensions, $\vec{x} \in [0 \ 1]^d$.

Define the desired evaluation points $X_E = \{\vec{x_1}, \vec{x_2}, \vec{x_3}, ..., \vec{x_{N_e}}\}$ and the set of anchored sample points $X_N = \{(x_1, a_y), (x_2, a_y), ..., (x_n, a_y), (a_x, a_y), (a_x, y_1), ..., (a_x, y_n)\}$ with $0 < x_i, y_i, < 1$. PDE, equation 1.

$$\frac{\partial u}{\partial t} + \frac{1}{2} \sum_{i,j}^{d} \mathbf{C}_{i,j} x_i x_j \frac{\partial^2 u}{\partial x_1 \partial x_2} + \sum_{i}^{d} r x_i \frac{\partial u}{\partial x_i} - r u = 0 \qquad \vec{x} \in [0 \, 1]^d, t \in [0, T]$$

$$u(\vec{x}, t) = \phi(\vec{x}, t) = \max(\frac{1}{d} \sum_{i}^{d} x_i - K e^{T - t}, 0) \qquad ||\vec{x}|| > 1, t \in [0, T]$$

$$u(\vec{x}, t) = 0 \qquad ||\vec{x}|| = 0, t \in [0, T]$$

$$u(\vec{x}, t) = \phi(\vec{x}, T) \qquad t = T$$

$$(1)$$

We have different sets of values.

- Evaluation points $X_{eval} = \{(x_i, y_i)\}_{i=1}^{N_e}$
- Anchored center points $X = \{(x_i, y_i)\}_{i=1}^{N}$
- Inner points $X_{inner} \subseteq X, \{(x_i, y_i)\}_i^{N_i nner}$
- Far and close points, where boundary condition apply $X_{far}, X_{close} \subseteq X, \{(x_i, y_i)\}_i^{N_{far}}, \{(x_i, y_i)\}_i^{N_{close}}$

We get that $N = N_{inner} + N_{far} + N_{close}$

Reproducing kernel

The reproducing kernel is for the 2D case defines as follows.

$$R(\vec{x}^1, \vec{x}^2) = 1 + k(x_1^1, x_1^2) + k(x_2^1, x_2^2)$$

$$k(x_1, x_2) = \sqrt{1 + \varepsilon^2 ||x_1 - x_2||_2^2}$$
(2)

From equation 2 we can define the following matrix.

$$\mathbf{A_0} = \left[R(\vec{x}^j, \vec{x}^k) \right]_{\vec{x}: \vec{x}_k \in X} \tag{3}$$

Followingly, the derived reproducing kernels with associated matrix may be built.

$$\partial_{x_i} R(\vec{x}^j, \vec{x}^k) = \frac{\varepsilon^2 (x_i^j - x_i^k)}{\sqrt{1 + \varepsilon^2 ||x_i^j - x_i^k||^2}} \Rightarrow \mathbf{A_1}, \mathbf{A_2}$$

$$\partial_{x_i x_i}^2 R(\vec{x}^j, \vec{x}^k) = \frac{\varepsilon^2}{(1 + \varepsilon^2 ||x_i^j - x_i^k||^2)^{3/2}} \Rightarrow \mathbf{A_{11}}, \mathbf{A_{22}}$$

$$\partial_{x_i x_j}^2 R(\vec{x}^j, \vec{x}^k) = 0 \Rightarrow \mathbf{A_{12}}$$
(4)

The projected approximation of a function $f(\vec{x})$ at a arbitrary set of evaluation points (may be the same as center points).

$$\underbrace{\hat{f}(X_{eval})}_{R^{Ne\times 1}} = \underbrace{\left[R(\vec{x}_i, \vec{x}_j)\right]_{\vec{x}_i \in X_{eval}, \vec{x}_j \in X}}_{R^{Ne\times N}} \cdot \underbrace{\mathbf{A_0}^{-1}}_{R^{N\times N}} \underbrace{f(X)}_{R^{N\times 1}} \tag{5}$$

From equation 5, the way to approximate the 2D Black Scholes operator becomes.

$$\mathcal{O}_{BS} = (rX_{1}^{1}\mathbf{B}_{1} + rX_{2}^{2}\mathbf{B}_{2} + \frac{1}{2}\sigma_{1}^{2}(X_{1}^{1})^{2}\mathbf{B}_{11} + \frac{1}{2}\sigma_{2}^{2}(X_{2}^{2})^{2}\mathbf{B}_{22} + \rho\sigma_{1}\sigma_{2}X_{1}^{1}X_{2}^{2}\mathbf{B}_{12} - r\mathbf{B}_{0})\mathbf{A}_{0}^{-1} \begin{cases} \mathbf{B}_{0} = [\mathbf{A}_{0}]_{i \in N_{i}, j \in N} \\ \mathbf{B}_{1} = [\mathbf{A}_{1}]_{i \in N_{i}, j \in N} \end{cases}$$
(6)

$$\begin{cases}
\mathbf{B_0} = [\mathbf{A_0}]_{i \in N_i, j \in N} \\
\mathbf{B_1} = [\mathbf{A_1}]_{i \in N_i, j \in N} \\
\mathbf{B_2} = [\mathbf{A_2}]_{i \in N_i, j \in N} \\
\mathbf{B_{11}} = [\mathbf{A_{11}}]_{i \in N_i, j \in N} \\
\mathbf{B_{22}} = [\mathbf{A_{22}}]_{i \in N_i, j \in N} \\
\mathbf{B_{12}} = [\mathbf{A_{12}}]_{i \in N_i, j \in N}
\end{cases}$$
(7)

Coordinate Transformation

We want to rotate the coordinate system and solve the problem for v_1, v_2 instead of x_1, x_2 . The transformation is linear and equation 9 & 10 is only applicable in the linear case.

$$\vec{v} = \begin{bmatrix} \frac{1}{2} & \frac{1}{2} \\ \frac{1}{2} & \frac{-1}{2} \end{bmatrix} \vec{s} + \begin{bmatrix} 0 \\ \frac{1}{2} \end{bmatrix}$$

$$\tag{8}$$

$$\frac{\partial^2 u}{\partial s_i \partial s_j} = \sum_{k,l=1}^d \left(\frac{\partial^2 u}{\partial v_k \partial v_l} \frac{\partial v_l}{\partial s_j} \frac{\partial v_k}{\partial s_i} \right) \tag{9}$$

$$\frac{\partial u}{\partial s_i} = \sum_{k}^{d} \left(\frac{\partial u}{\partial v_k} \frac{\partial v_k}{\partial s_i} \right) \tag{10}$$

With this, we can rewrite equation 1 in this new system.

$$\mathcal{O}_{BS}^{\mathcal{T}} = \left(r(V_{1}^{1} + V_{2}^{2} - \frac{1}{2}) \mathbf{B}_{1} + r(V_{1}^{1} - V_{2}^{2} + \frac{1}{2}) \mathbf{B}_{2} \right)
+ \frac{1}{2} \sigma_{1}^{2} (V_{1}^{1} + V_{2}^{2} - \frac{1}{2})^{2} \mathbf{B}_{11} + \frac{1}{2} \sigma_{2}^{2} (V_{1}^{1} - V_{2}^{2} + \frac{1}{2})^{2} \mathbf{B}_{22}
+ \rho \sigma_{1} \sigma_{2} (V_{1}^{1} + V_{2}^{2} - \frac{1}{2}) (V_{1}^{1} - V_{2}^{2} + \frac{1}{2}) \mathbf{B}_{12}
- r \mathbf{B}_{0} \right) \mathbf{A}_{0}^{-1}
\begin{cases}
X_{1}^{1} = (V_{1}^{1} + V_{2}^{2} - \frac{1}{2}) \\
X_{2}^{2} = (V_{1}^{1} - V_{2}^{2} + \frac{1}{2}) \\
\mathbf{B}_{0} = [\mathbf{A}_{0}]_{i \in N_{i}, j \in N} \\
\mathbf{B}_{1} = [\frac{1}{2} (\mathbf{A}_{1} + \mathbf{A}_{2})]_{i \in N_{i}, j \in N} \\
\mathbf{B}_{1} = [\frac{1}{2} (\mathbf{A}_{1} - \mathbf{A}_{2})]_{i \in N_{i}, j \in N} \\
\mathbf{B}_{11} = [\frac{1}{4} (\mathbf{A}_{11} + 2\mathbf{A}_{12} + \mathbf{A}_{22})]_{i \in N_{i}, j \in N} \\
\mathbf{B}_{12} = [\frac{1}{4} (\mathbf{A}_{11} - 2\mathbf{A}_{12} + \mathbf{A}_{22})]_{i \in N_{i}, j \in N} \\
\mathbf{B}_{12} = [\frac{1}{2} (\mathbf{A}_{11} - \mathbf{A}_{22})]_{i \in N_{i}, j \in N}
\end{cases}$$
(12)

Time Solver (BDF2)

From earlier equations, we can define the system that should be solved at each timestep. Note that the set each point belongs to (Inner, Far) is still defined in terms of x_1, x_2 . Reversing time $\tau = T - t$ gives:

$$U = \begin{bmatrix} u(\vec{x_i}, t) \\ \vdots \end{bmatrix}_{\vec{x_i} \in X}$$

$$\frac{\partial U}{\partial \tau} = \mathcal{O}_{BS}^{\mathcal{T}} U \qquad x_i \in X_{in}$$

$$U = \phi(\vec{x_i}, \tau) \qquad x_i \in X_{far}$$

$$U = 0 \qquad x_i \in X_{close}$$

$$(13)$$

Discretize the time $\vec{\tau} = \{\tau_m\}_{m=1}^M$ with time step $k_m = \tau_{m+1} - \tau_m$. The PDE is then expanded as:

$$U^{m+2} - \frac{4}{3}U^{m+1} + \frac{1}{3}U^m = \Delta \tau_m \mathcal{O}_{BS}^{\mathcal{T}} U^{m+2} \qquad x_i \in X_{in}$$

$$\Leftrightarrow (I - \Delta \tau_m \mathcal{O}_{BS}^{\mathcal{T}}) U^{m+2} = \frac{4}{3}U^{m+1} - \frac{1}{3}U^m \qquad x_i \in X_{in}$$
(14)

$$\Leftrightarrow (I - \Delta \tau_m \mathcal{O}_{BS}^{\mathcal{T}}) U^{m+2} = \frac{4}{3} U^{m+1} - \frac{1}{3} U^m \qquad x_i \in X_{in}$$
 (15)

$$\Leftrightarrow U^{m+2} = (I - \Delta \tau_m \mathcal{O}_{BS}^{\mathcal{T}})^{-1} (\frac{4}{3} U^{m+1} - \frac{1}{3} U^m) \qquad x_i \in X_{in}$$
 (16)

(17)

To eliminate the boundary conditions, the matrices are expanded to size $N \times N$. This new matrix, C, has identical rows to $(I - \Delta \tau_m \mathcal{O}_{BS}^{\mathcal{T}})$ for the indices corresponding with the interior points while being the identity matrix for the boundary rows. The boundary condition is then enforced by applying the respective condition to the boundary indices of the right-hand side. $U_{RHS} = (\frac{4}{3}U^{m+1} - \frac{1}{3}U^m)$. This is done before solving the linear system. The final system is then given by.

$$U^{m+2} = C^{-1}U_{RHS} \vec{x_i} \in X$$

$$\begin{cases} U_{RHS} = \frac{4}{3}U^{m+1} - \frac{1}{3}U^m, & \vec{x_i} \in X_{in} \\ U_{RHS} = 0, & \vec{x_i} \in X_{close} \\ U_{RHS} = \phi(\vec{x_i}, \tau_{m+1}), & \vec{x_i} \in X_{far} \end{cases}$$
(18)

Algorithm

Algorithm 1 Pseudo code for solving BS

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1: Scale X_{eval}, K to and define problem on [0, 1]^d
 2: Generate center points in transformed coordinates
 3: Obtain center points in standard system
 4: Find points corresponding to close and far boundary.
 5: Define A matrices, equation (3) - (4).
 6: With Matrices A, build the rotated local matrices B. Equation 12.
 7: Build the transform Black-Scholes operator \mathcal{O}_{BS}. Equation (11).
8: Extended \mathcal{O}_{BS} to form matrix C, eliminating BC. U_{m+1} = U_m = u_0
9: Apply initial condition
10: for \tau_m; m++ do
        U_{temp} = C^{-1} U_{RHS} \label{eq:Utemp}
11:
        U_{RHS} = \frac{4}{3} * U_{m+1} - \frac{1}{3} * U_m
Apply boundary conditions:
12:
13:
              U_{RHS}(\vec{x} \in X_{far}, \tau_{m+1}) = \phi(\vec{x}, \tau_m)
14:
              U_{RHS}(\vec{x} \in X_{close}, \tau_{m+1}) = 0
15:
        Move solution along:
16:
              U_m = U_{m+1}
17:
18:
              U_{m+1} = U_{m+2}
19:
              U_{m+2} = U_{temp}
20: end for
21: Build evaluation matrix and evaluate solution in X_{eval}. Equation (5)
22: Rescale problem
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