

Nuclear excitation functions for medical isotope production:
Targeted radionuclide therapy via natIr(d,2n)^{193m}Pt

Hannah Lovise Okstad Ekeberg

June 2020

Contents

1 Theory	5
1.1 Treatment of cancer through radiative sources	5
1.1.1 Targeted radionuclide therapy	5
1.1.2 ^{193m}Pt as a potential therapeutic agent	10
1.2 Radioactive decay law and Gamma-ray spectroscopy	14
1.2.1 Radioactive decay law	14
1.2.2 High purity Germanium detector	15
1.2.3 Gamma-ray spectrum	16
1.3 General nuclear reaction theory	17
1.3.1 Nuclear reactions and reaction cross sections	18
1.3.2 Constraints in nuclear reactions	19
1.3.3 Nuclear cross sections	22
1.3.4 Expected products from deuterions on iridium, iron, copper and nc	23
1.4 From book	23
1.5 The decay of the compound nucleus	23
2 Experimental setup	25
2.1 Lawrence Berkeley National Laboratory's 88" Cyclotron	25
2.2 The experiment	25
2.2.1 Target design and foil characterization	27
2.2.2 Irradiation of target stack	27
2.2.3 Counting on high purity detectors	30
3 Analysis	33
3.1 Gamma-ray spectroscopy	33
3.1.1 Energy and peak-shape calibration	33
3.2 Efficiency calibration	35
3.3 End of beam activities	36
3.4 Energy and Beam current	37
3.4.1 Variance minimization	40
3.5 Cross section measurements	41
4 Results	45
5 Discussion	53
5.1 Beam current	53
5.2 Cross section products	53
5.2.1 Nickel products	53
5.2.2 Iridium Products	54
A Statistics	61

B Tables	63
B.1 Product nuclei, Q-values and gammarays	63
B.2 Production cross sections	70
B.2.1 $^{nat}Ir(d,x)$	70

In this thesis work, a production route of the medically potential therapeutic agent via $^{193}Ir(d,2n)^{193m}Pt$ has been investigated. In a stacked-target cross section experiment with a 33 MeV deuteron incident beam, cross sections of the products produced from iridium and the monitor foils nickel, iron and copper are reported.

Chapter 1

Theory

1.1 Treatment of cancer through radiative sources

Today, multiple options are available for treatment of cancerous tissue, where some of the most commonly used methods are chemotherapy, surgery, immunotherapy, external beam-therapy, brachytherapy and targeted radionuclide therapy citation. The latter three treatment types utilize ionizing particles to induce damage in the DNA of cancerous cells, so that they undergo apoptosis. In external beam therapy, photons, electrons or hadrons are accelerated and sent towards a tumor externally, and in brachytherapy a contained radioactive source (with for instance a β^- -emitter), is placed in proximity to tumor. Targeted radionuclide therapy is an emerging alternative to the traditional treatment options, and offers a patient-specific treatment dependent on choice of radiopharmaceutical which targets a type of tumor or cell correct? need citation. A radiopharmaceutical consists of a radionuclide and a cell-targeting molecule called a tracer. Targeted radionuclide therapy utilizes the biochemical pathways in the body, thus with an appropriate tracer, targeted tissue with a high uptake of the radiopharmaceutical will receive a high dose, and healthy tissue can be spared (Yeong2014).

1.1.1 Targeted radionuclide therapy

Write more detailed about what it is..

In treatment of diseases related to the use of radionuclides, there are multiple properties that play an important role, such as physical features like half-life, decay-mode, decay-products, radiation energy and linear energy transfer, and biochemical features like in tissue range, tissue targeting of tracer, uptake and retention in desired tissue, biological half-life and toxicity of radionuclide and decay-product (Yeong, therapeutic radionuclides in nuclear medicine). In addition, the availability, cost and quality is important, and that the production routes gives high isotopic purity of the product, along with high yield (Qaim, Nuclear Data for production of medical application of radionuclides: present status and future needs).

Here write about the various small things. Isotopic purity, half-life,

Particle interaction in matter

equation

$$Q_{\max} = \frac{4mME}{m + M}$$

can be from billiard physics, or inelastic collision physics. Also, bethe-block is proportional to

$$-\frac{dE}{dx} = 2\pi N_a r_e^2 m_e c^2 \rho \frac{Z}{A} \frac{z^2}{\beta^2} \left[\ln \left(\frac{2m_e \gamma^2 v^2 Q_{\max}}{I^2} \right) - 2\beta^2 - \delta - 2\frac{C}{Z} \right] \quad (1.1)$$

where r_e is electron radius, m_e electron mass, N_a avogadro, I mean excitation potential, Z & A atomic number and mass number of absorbing material, ρ density of absorbing material, z charge of incident

particle, $\beta = v/c$, $\gamma = 1/\sqrt{1-\beta^2} = 1/\sqrt{1-v^2/c^2}$, δ density correction, C shell correction, W_{\max} is max energy transfer per collision. In general,

$$\frac{dE}{dx} \propto \rho, \frac{z^2}{\beta^2}, \frac{Z}{A}$$

and logarithmic stuff which i dont know... so alpha particles are stopped quicker than protons if they have the same energy (Techniques for Nuclear and Particle physics experiments, p. 24). Particles interact differently in a given medium. Figure 1.1 shows how various particles interacts in a medium where the depth is on the x-axis and dose is on the y-axis. In a medium, there are positively charged nuclei and negatively charged electrons. We differ between neutral particles like neutrons and photons, and charged particles like electrons or protons, alpha-particles, and other heavy charged particles. Charged particles have a short range in comparison to neutral particles, as the Coulomb force forces the particle to interact continuously along the path, either either by scattering inelastically with the atomic electrons, or scattering elastically via the nuclei. Inelastic collisions is the dominant process, where the atomic electrons are either excited or ionized. If the now free electron had sufficient energy transferred from the incoming particle (called a δ -electron), this electron can further ionize along the particle track. Elastic scattering is the less dominant process, where the energy loss is small, as long as the nuclei in the medium are larger than the incoming particle (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 21). The energy loss of heavy charged particles are described by an equation called Bethe-Block, where the energy loss at high energies is small since very little energy is transferred per collision with atomic electrons (due to the large mass). However, when the particle reaches low velocity, the particle interacts at a much higher rate in a very specific peak called the Bragg peak. Can be seen on figure 1.1, for the proton. As we can see, the energy deposited behind the tumor is zero, along with the energy deposition in front of the tumor is low which is something which can utilized in radiation therapy. However, since the Bragg peak is so narrow, a broadening of the beam might be necessary, which is to send out a beam of multiple energies, so that it is spread out. The dose in front of the tissue gets larger, but the dose behind is still zero. The electrons/positrons interacts differently, since the mass is equal to the atomic electrons, the energy transfer can be up to half of its mass, and it scatters like crazy. Electrons can experience energy loss either from collisions like heavy charged particles, or via the electromagnetic radiation that arises when electrons are losing energy (bremsstrahlung). However, for energies up to a few MeV, the collision energy loss dominates (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 37). On figure 1.1, we see that the electrons and proton interacts differently, there is no clear peak, and the incoming electrons immediately interacts with the medium, lose energy fast and the secondary electrons interacts after and slowly decreasing. At 22 MeV, bremsstrahlung stopping power is also evident, so we see an exponential tail which?

Photons interacts differently than charged particles. As they are neutral, they do not interact via inelastic collisions, but interacts via three different processes called photoelectric effect, compton scattering and pair production. Thus the energy is not degraded, and the beam can only be attenuated as a function of thickness x (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 53)

$$I(x) = I_0 e^{-\mu x} \quad (1.2)$$

where μ is the attenuation coefficient. Photons can be gamma-rays or X-rays, where gamma-rays are produced in de-excitation of nuclei, and have ranges 0.1-10 MeV (Krane, p. 327), and X-rays can be either be produced in de-excitation of atomic electrons (where the energies are typically quite low), or be produced by de-accelerating fast electrons in an x-ray tube, for instance by hitting a material like tungsten, where the energy is high. As we can see on figure 1.1, the X-rays are produced in an x-ray tube (MV and kV is the voltage of the X-ray tube), and we can see that the 22 MV X-rays have a build-up effect, where ionized atomic electrons participates in increasing the dose, while the much lower 200 kV X-rays does only decrease. The gamma-rays from ^{60}Co behaves similarly to the 200 kV X-rays, but are slightly higher in energy deposition as the energy in general are slightly higher

than 200 keV. The tale is exponential as we can see in eq. 1.2, and is in fact a problem in radiation therapy as the dose to healthy tissue behind the tumor is quite high.

Photoelectric effect dominates at lower energies, and happens when a photon (energy) is completely absorbed an atomic electron. In gamma-ray spectroscopy, this process is desired as the gamma-rays from nuclei is specific, and we can identify the nuclei on a spectrum. Probability of photoelectric effect is

$$\tau \propto \frac{Z^n}{E_\gamma^{3.5}}, \quad n = 4 \text{ or } 5 \quad (1.3)$$

where Z is the atomic number of material ((Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 55).

In compton scattering, the photon has higher energy and can transfer parts of the energy to an atomic electron, be deflected and ionize further. The deflecting angle is dependent on energy, and the angle $\theta \in (0^\circ, 180^\circ)$ (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 55-56). Compton scattering thus gives photons in a spectrum of different energies. This process is not desired in gamma-ray spectroscopy, and contributes to noise in a Compton region. This process is however desired in external radiation therapy, as the process gives rise to high energetic secondary electrons, and photons in which ionize further. The probability of interaction is

$$\sigma \propto Z \quad (1.4)$$

Pair production demands a photon energy to be equal or higher than two electron rest masses, as the photon is converted to an electron/positron pair ($1.022 \text{ MeV} = 2m_0c^2$) (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 57). Probability of interaction is

$$\kappa \propto Z^2 \quad (1.5)$$

Finally neutrons interacts with the strong force with nuclei, and can be via absorption, or elastic or inelastic scattering. The interaction probability is dependent on energy, will not describe more. (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 64).

Whenever a charged particle like an electron or hadron enters a medium, the Coloumbforce will affect the charged particle and cause interactions with the atomic elecrtons. Neutral particles like photons and neutrons are not affected by the Coulomb force, thus interacts rarely. On figure ..., the 22 MV X-rays first has a build-up effect, where the initial photons induce ionizations and secondary electrons, which contributes to further ionization, and increases the dose. After, the dose decreases exponentially. The build-up effect for high-energetic photons which are only used in external beam therapy enables an energydependent maximum dose in tissue as a function of depth. However, the exponential tale gives dose to the healthy tissue behind. The gamma-rays in nuclei are not that energetic, and the lack of specificity does not make gamma-rays a suitable decay chain for therapeutic agents. Neutral particles interacts via absorption or

X-rays from an accelerator is produced by acceleration of electrons which hits a material like for instance wolfram or tungsten. The X-rays have energies ranging from zero to the max electron energy, and hence the X-ray energy is in MV instead of MeV. Why we are not using gamma-rays in treatment; range is too long, ionization are not frequent enough, and energy is too low?

The biological effect of ionizing radiation and linear energy transfer

in general in this paragraph refer to the radiation biology book.. Ionizing radiation are particles with sufficient energy to cause ionizations along the particle track, thus separating an atom and one or more electrons **citation, maybe the bio-book?**. The free electron(s) can ionize further, and the positive ion can cause undesired reactions **find a citation**. DNA is a large molecule with two strands bound in a double helix structure. Each strand is composed of sugar and phosphate groups, and nitrogenous bases which bind the two strands. These bases are called adenin & guanin and cytosine & thyamin (always bound pairwise), and are bound through weak hydrogen bonds which are exposed to strand breaks. The cell is equipped with an impressive repair mechanism, and unless both DNA strands are

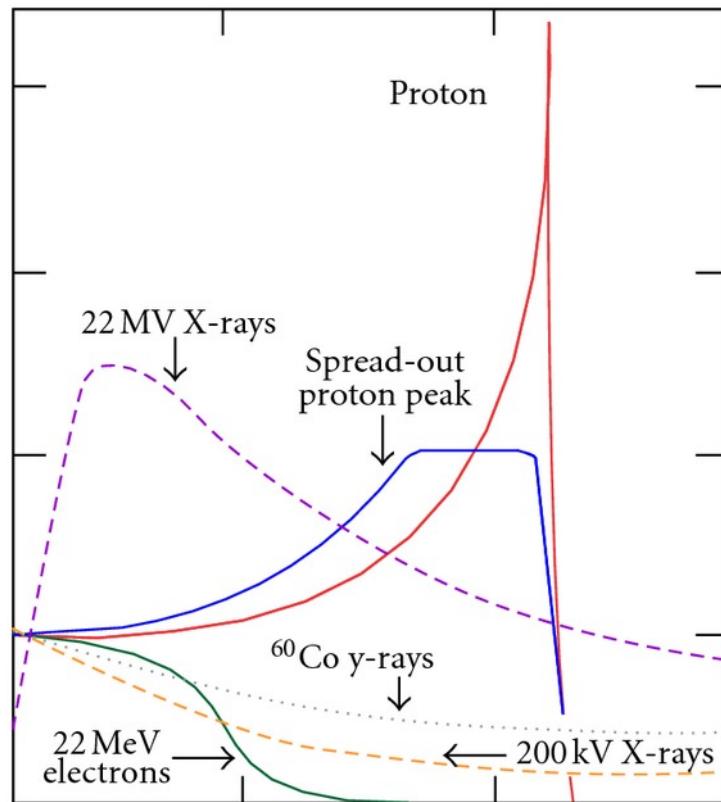


Figure 1.1: Medium depth along x-axis, energy deposition in tissue (or dose?) on y-axis. Find citation in special curriculum.

damaged, which is called a double stranded break (DSB), most damages are repaired [cite](#). Radiation damages in the DNA can be caused directly by the ionizing particle, or indirectly via free radicals which are subject to other ionizations. Since the body contains large amounts of water, ionization of water molecules giving for instance H^\bullet or OH^\bullet are important damaging factors. Damages induced in the DNA can be lethal to the cell and either cause apoptosis or mutation in which can be a cause of cancer. In therapy, the goal is to make the cancerous cells to undergo apoptosis, and choosing a particle which has a high probability of inducing damage will induce multiple double stranded breaks, if passing nearby.

Linear energy transfer (LET) describes the energy absorbed by the medium, and is defined as the average energy (typically measured in keV) deposited per unit length (typically measured in μm) of the density material (book, p. 101):

$$\text{LET} = \frac{dE}{dx} \quad (1.6)$$

LET thus describes the general frequency in which a particle will deposit energy along the particle track. A particle with a high LET-value will interact more frequently over a shorter distance than a low LET-particle. In targeted radionuclide therapy, it is important that the range of the particle is short-range so that the healthy tissue is spared, and that the dose to the cancerous tumor is large. For therapy, medical isotopes which emit radiation with high LET such as α , β and auger-electrons are desired since the ionization rate is high and the distance traversed is low (Yeong, therapeutic radionuclides in nuclear medicine). Gamma-radiation is not a desired process in therapy as photons does not degrade in energy, thus can have a long range unless attenuated. Figure ... shows how alpha, beta and auger electrons interacts on the scale of DNA.

Decay modes

Different decay modes have different ranges in the tissue. Dependent on the size of tumor, and critical tissue surrounding it, different emitters are used. However to reduce dose to healthy tissue, specific short range high LET particles are in general desired. Figure 1.2 shows the ranges of auger electrons, alpha-particles and β^- - particles on a cellular scale. The binding energy holds the nucleus together, and for unstable nuclei, the binding energy is so low that particles can be emitted to gain stability, lowering the binding energy. Beta and alpha decay changes the element, and the decay product is commonly in an excited state. Therefor gamma-decay commonly occurs after, and in gamma-ray spectroscopy, it is the gamma-rays of the de-excitation of the daughter which is detected. Knowledge about the energy and intensity of the radiation is very important. It is unrare that just one single process happens during a decay, and dependent on what the desired process is, there is a beta or alpha particle or gamma-rays, X-rays or auger electrons caused from electron capture or internal conversion. The ranges can be different.

Beta-decay happens whenever there is an overweight in number of protons or neutrons. For an overweight of neutrons, β^- -decay occurs, by transforming a neutron into a proton, a β^- -particle and an antineutrino



For an overweight in protons, β^+ decay occurs, by transforming a proton into a neutron, a β^+ -particle and a neutrino



β^+ -particles are typically used in diagnostic imaging in PET (positron emission tomography). Since neutron mass is higher than the proton mass (mass difference $1.022 \text{ MeV}/c^2$), there is an additional threshold energy needed to induce the reaction. If the energy is not high enough, electron capture dominates, which takes in an atomic electron (typically from K-shell), and with a proton in the nucleus, transforming to a neutron and a neutrino:



Since beta-rays are not monoenergetic, the energy is distributed between three particles and gives a spectrum of different possible energies. β^- -emitters used in therapy are stopped rather quickly in tissue due to negative charge and low mass (as mentioned in the above section, energy transfer can be up to half of its mass.),

Alpha-particles are ${}^4\text{He}$ -nuclei, consisting of two protons and two neutrons. It is emitted from heavy nuclei. Since the nuclear force is short range (1fm ?), and heavy nuclei tend to be in the limit between electromagnetic and nuclear force, an alpha particles (which carries large binding energy) lowers the binding energy sufficiently so that the nucleus is more stable. The binding energy of an alpha particle is very high (28.3 MeV) which is why an alpha-particle is emitted to lower binding energy.

Gamma-decay is emission of photons following de-excitation of the nucleus, as de-excitation from one state to another. It is specific for the nucleus, and is thus called the fingerprint of the nucleus, because we can observe it for instance in a gamma-ray spectrum. Gamma-decay commonly occurs after beta or alpha decay, or can happen if there is an isomer, then the decay mode is called isomer transition. However, gamma-rays are not well suited in therapy themselves (however for imaging in SPECT more relevant). But nuclei that emit electrons following gamma-decay are more relevant. Internal conversion is a process where the gamma-ray interacts electromagnetically with atomic electrons, and the electron is emitted instead of the photon. This electron is called an conversion electron and has energy in the order of the gamma ray (minus atomic binding energy). In order to fill the atomic shells, a cascade of X-rays and auger electrons are emitted (auger electrons have energies in order of X-rays).

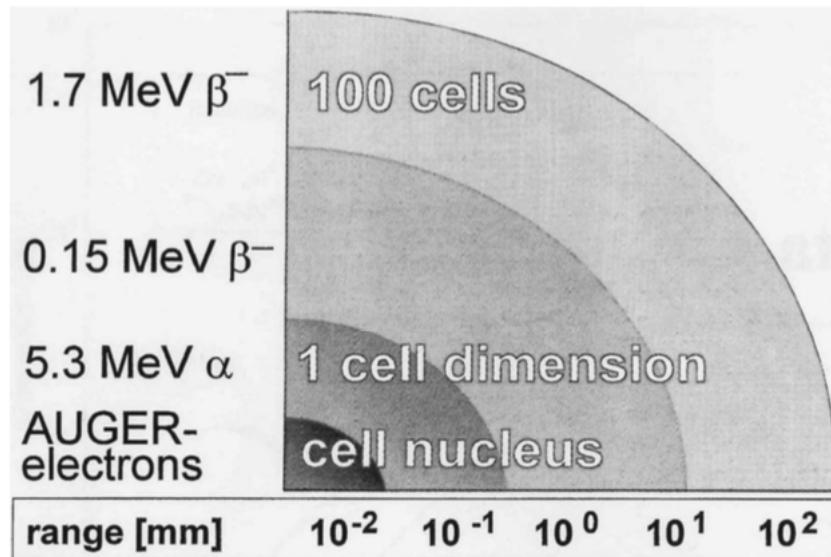


Figure 1.2: The figure illustrates the ranges of different particles on the scale of cells. Auger electrons does not exceed the cellular nucleus and alpha-particles does not exceed more than ca. 1 cell dimension which allows for very specific targeting of cells. Low energetic β^- emitters have ranges of ca. 1 mm. Picture is from (S M Qaim, R Capote, and F Tarkanyi. Nuclear Data for the Production of Therapeutic Ra-dionuclides. Trs 473, (473):395, 20), p. 2.

1.1.2 ^{193m}Pt as a potential therapeutic agent

^{193m}Pt ($t_{1/2}=4.33$ days) is an auger-emitting isomer which decays by isomeric transition (100%) to the long-lived ^{193g}Pt groundstate ($t_{1/2}=50$ years). Radionuclides produced from deuterons on natural iridium such as ^{191}Pt , ^{193m}Pt , ^{192}Ir and ^{194}Ir are believed to have potential in medicine, like chemotherapy, brachytherapy, radioimmunotherapy and imaging (Tarkanyi et.al 2006). Platinum radionuclides are of special interest, as platinum is the main element in chemotherapeutic agents such as cisplatin, which is a drug which is used clinically in treatment of testicular and ovarian cancer mainly, but also to treat esophagus, head and neck and bladder cancer¹. Cisplatin (cis-dichlorodiammine platinum(II)) is an inorganic molecule which contains one stable platinum atom surrounded by two chlorine atoms and two ammonia molecules (NH_3). The cisplatin-molecule enters the cell nucleus, and binds to the DNA, example-wise shown in figure 1.3, where the chlorine-atoms are de-attached and the platinum-atom binds through covalent bonds to the DNA base guanine (and in some cases adenine, *is that correct?*), and breaks the bonds between the DNA nitrogenous bases. Cisplatin thus targets the DNA. One of the major challenges with cisplatin is the chemical toxicity, but when auger-emitters such as ^{193m}Pt or ^{195m}Pt replace the stable platinum atom, the local auger-damage effect increases the chemical damage of cisplatin, suggesting that a smaller amount of the drug is required, and chemical toxicity can be avoided².

By replacing either of the stable nitrogen atoms with the PET-radionuclide ^{13}N ($t_{1/2}=9.965$ minutes), or by a radionuclide platinum, where ^{191}Pt ($t_{1/2}=2.83$ days, decay by electron capture (100%) to ^{191}Ir (stable)), ^{193m}Pt and ^{195m}Pt ($t_{1/2}=4.010$ days, decay by isomer transition (100%) to ^{195g}Pt (stable)) is of special interest, cisplatin can be used for imaging or therapy³, but therapy is most common.

As ^{191}Pt is electron-capture emitter, can be used in imaging, with for instance 129.4 keV (38.0%) or 172.19 keV (43.2%). Combining ^{191}Pt with a therapeutic agent might be possible for theranostic pair with either ^{193m}Pt or ^{195m}Pt ? Can be combined with therapy as it releases auger electrons?

¹https://www.sciencedirect.com/science/article/pii/S0969804399000822?casa_token=ZLJ8YPQzGZMAAAAAA:264QzKWpH8Kv6iHotiGMeoHTk8jKqmnoDgf709SrAD8BUWVwbRXriZbHgkYOtHg-2qyX3Hvt9E

²<http://citeseerx.ist.psu.edu/viewdoc/download?doi=10.1.1.987.2577&rep=rep1&type=pdf&page=506>, p. 493

³https://www.sciencedirect.com/science/article/pii/S0969804399000822?casa_token=ZLJ8YPQzGZMAAAAAA:264QzKWpH8Kv6iHotiGMeoHTk8jKqmnoDgf709SrAD8BUWVwbRXriZbHgkYOtHg-2qyX3Hvt9E

Decay mode: For ^{193}Pt , there are three states, the isomer state at 149.8 keV, with nuclear spin $13/2^+$ (4.33 d), a state at 14.3 keV with nuclear spin $5/2^-$ (2.52 ns), a state at 1.6 keV with nuclear spin $3/2^-$ (9.7 ns) and the ground state at 0.0 keV with nuclear spin $1/2^-$ (50 y)<https://www.nndc.bnl.gov/nudat2/getdecays>. V.S. Nuclear data sheets for A=193. Nucl. Data Sheets. 32, 593-679, 1981. **here write about gamma-decay and that the probability for M6 or whatever transition is improbable.** The populated isomer states decays from 149.8 keV to 14.3 keV releasing a 135.50 keV photon (0.1145475%), from 14.3 keV to 1.6 keV releasing a 12.634 keV photon (0.70%), and from 1.6 keV to the ground state releasing a 1.642 keV photon (0.0321). The photon abundance is thus low, and this isomer is not well suited for imaging. Due to the low intensity of the gamma-rays, it might be difficult to detect. There are X-rays too, but they overlap with other nuclei. Since the gamma-rays are weak, the IC probabilities are 99.89%, 99.33% and 99.99% for each state respectively, calculated by subtracting 100 - gamma-intensity ⁴. This also indicates that the phondon abundance is very low, as well high very high prob of low E electrons :D

In all decays, there are certain quantities in which needs to be conserved; angular momentum (ℓ) and parity (maybe ℓ should be written as L instead??). Krane says that a multipole of order ℓ transfers angular momentum $\ell\hbar$ per photon (Krane, p. 333). A nuclear state has a definite angular momentum (angular momentum and spin) and parity, and if a gamma-transition is to happen between two states, the photon must connect the two states by conserving angular momentum and parity. In order for the quantity ℓ to be conserved, the angular momentum can be integer values between

$$|I_i - I_f| \leq \ell \leq I_i + I_f \quad (1.10)$$

For the decay of ^{193m}Pt (E level=149.8 keV) to the excited state (E level=14.3 keV), the spin and parity change is from $13/2^+$ to $5/2^-$, so $\ell = 4, 5, 6, 7, 8, 9$. The parity decides the wether the radiation is electric multipole or magnetic multipole (equations from Krane p. 331),

$$\pi(ML) = (-1)^{\ell+1}, \quad \pi(EL) = (-1)^\ell \quad (1.11)$$

The electric decays have even parity when $\ell=\text{even}$, and magnetic has even when ℓ is odd. If parity is unchanged in the reaction ($\Delta\pi = \text{no}$), the electric multipoles are even and magnetic multipoles are odd. If the parity does change ($\Delta\pi = \text{yes}$), there would be odd electric and even magnetic multipoles. Hence the possible transition from $13/2^+$ to $5/2^-$ are whenever $\Delta\pi = \text{yes}$ and $\ell = 4, 5, 6, 7, 8, 9$, which gives possible M4, E5, M6, E7, M8 or E9.

In general, the lowest possible multipole dominates, and the emisssion of multipole of one order higher (L+1 than L), is reduced by a factor ca 10^{-5} (Krane p. 335, important!!). Thus, a multipole of order 4 or 5 has a low probability of occurring and thus the isomer has a long half life. In comparison to decay from isomer state, decay from $5/2^-$ to $3/2^-$ gives possible radiation, $\ell = 1, 2, 3, 4$, with no parity change, and $\Delta\pi = \text{no}$, gives possible M1, E2, M3, E4, and from $3/2^-$ to $1/2^-$ gives $\ell = 1, 2, 3, 4$, which also gives M1,E2, M3, E4.

Half life: the decay rate constant is the sum of the decay rates of all the populates states transitions, $\lambda = \lambda_{13/2^+} + \lambda_{5/2^-} + \dots$

Gamma-decay and isomeric transition

Gamma-decay is the lowering of the excitation energy by the release of a photon, with an energy ΔE equal to the energydifference in the two states. The typical half lives of gamma-emission are less than 10^{-9} seconds, however, longer lived states of a nucleus which is not the ground state is called an isomer, and the gamma-decay of an isomer state is called isomeric transition (Krane, p. 175). Whenever gamma-decay is possible, another process called internal conversion is competing. It is an electromagnetic process, where the nucleus interacts electromagnetically with the atomic electrons, and an electron is emitted instead of the photon (Krane, chapter 10, p. 341). The kinetic energy of the emitted electron is the transition energy minus the electron binding energy

⁴<http://citeseerx.ist.psu.edu/viewdoc/download?doi=10.1.1.987.2577rep=rep1type=pdfpage=506>, p. 496

$$T_e = \Delta E - B \quad (1.12)$$

where B is a positive number (even though bound states are negative??). The electron is called a conversion electron, and this electron is high in energy and matches the gamma-energy. The electron binding energy varies with the atomic orbital (Krane), and the electrons emitted following internal conversion are in a spectrum of different discrete energies. The transition energy must be higher than the electron binding energy, and as a consequence, the electron is labelled with the shell that it was emitted from. (remember, n=1=K, n=2=L, n=3=M, n=4=N)

In the case of the decay of ^{193m}Pt , internal conversion is highly favoured instead of gamma-decay (the intensity of the gammas are very weak). The total decay probability is the summed decay probability for gamma-decay and internal conversion

$$\lambda = \lambda_\gamma + \lambda_{IC} \quad (1.13)$$

and the internal conversion coefficient α can be defined as

$$\alpha = \frac{\lambda_{IC}}{\lambda_\gamma} \quad (1.14)$$

High values of α indicates high probability of internal conversion, relative to probability of gamma-emission, but the coefficient diverges towards infinity when λ_γ reaches towards zero, which for instance is when the gamma-transition is zero. In general, the coefficient increases as Z^3 , which will give a much greater coefficient for heavy nuclei than for lighter nuclei. In addition the coefficient decreases rapidly (ca. $E^{-2.5}$) with increasing transition energy. The multipole order also affects the coefficient, where a higher multipole order indicates a higher value. For higher atomic shells than the K shell ($n=1$), the coefficient decreases like n^{-3} (Krane, chapter 10, p. 346).

In therapy, the most important process is the process which occurs after the release of the conversion electron. There is a vacancy in the shell where the conversion electron was emitted, and an electron from a higher shell drops down to this energy level, with the release of an X-ray with an energy equal to the difference between the energy state of the two shells, ΔE . If the transition is an electron from an L shell drops to K shell, and an electron from the L shell is ejected, the process is called a KLL transition, and the energy of the auger electron is $E_{auger} = E_K - E_{LL}$ (Prasad A. Naik, in Encyclopedia of Spectroscopy and Spectrometry, 1999) <https://www.sciencedirect.com/topics/chemistry/auger-process>. If the vacancy is filled with an electron from the same shell (or subshell) but the ejected electron is from another shell, the electron is called a coster-kronig electron (like LLM, electron vacancy is moving from L to L and electron in M is emitted), and if the whole process occurs in the same shell, it is called a super coster-kronig process (MMM)

The energy of the X-rays are lower in energy than the gamma-rays, typically. If one of the X-ray photon interacts within the atomic electrons (via photoelectric effect), the electron (which is called an auger electron) will be emitted with the energy of the X-ray minus the atomic binding energy (Handbook of Nuclear chemistry, p. 390)

$$T_{a.e.} = \Delta E_{x-ray} - B \quad (1.15)$$

From the vacancy from the auger electron, a new electron can take this place and release another X-ray. The auger electron can cause further ionizations in the atom, either by interaction itself, or from X-rays following the de-excitation of another atomic electron by the vacancy. Thus it is possible to have a cascade of electrons and X-rays. The secondary electrons caused by the auger electron can lead to a cascade of new short-range electrons and X-rays, which are typically have ranges of nm in tissue (Handbook of nuclear chemistry p. 2203). Since the X-ray energy is in the low energy region, the auger electrons have low energies (from equation 1.15).

Since auger emitters are short range, they are very precise, and do only harm when bound to DNA or when incorporated into the cellular nucleus (handbook of nuclear chemistry, o. 2204), which means that no neighbouring cells will be affected.

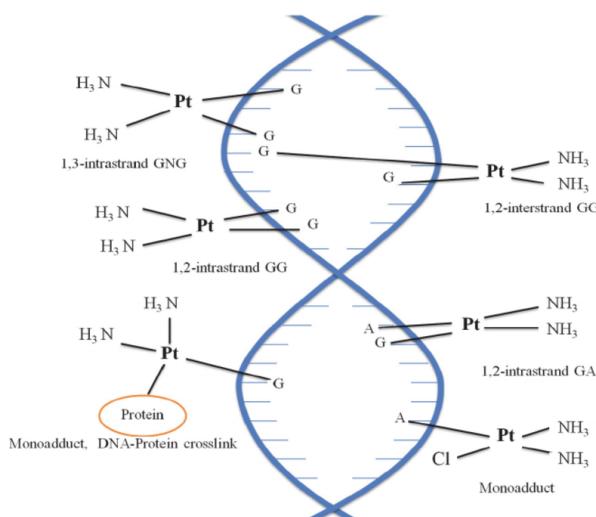


Figure 1.3: A DNA Repair Protein BRCA1 as a Potentially Molecular Target for the Anticancer Platinum Drug Cisplatin - Scientific Figure on ResearchGate. Available from: https://www.researchgate.net/figure/Common-cisplatin-DNA-adducts-and-functions-For-instance-the-platination-of-human-serum_fig2_21919257 [accessed 12 Apr, 2020].

After IC-process, vacancy is produced in an inner atomic shell (n) or subshell (like $l=spdf$). Vacancies in inner atomix orbitals are unstable, filled by electrons from higher energy levels. 4 processes, radiative X-ray transition, non-radiative transitions of auger, Coster-Kronig and super Coster-Kronig. move primary vacancies to higher shells or subshells. The non-radiative transitions involves multiplication of vacancies in the higher shells and subshells since two new vacancies are produced for each filled vacancy. Whenever energetically possible, super CS transitions dominate the other types. Thus the inner shell vacancies move upward to the valence and near valence shells of the atom, copious emission of electrons occur. Since the transition energies are very small for the higher shell transitions, the electrons ejected possess very small energies and is extremely short range (few nm) in biological matter, find a citation here, numb 8 in chapter.

Energy loss of low E auger electrons. In this energy region, is due to collision loss, not bremsstrahlung. Deflects frequently due to low mass, and the max energy loss is $T_e/2$ per collision.

General stuff ^{193m}Pt : Cellular nucleus is approximately $6\mu\text{m}$, while thickness of DNA is ca 2 nm (wikipedia). ⁵. Range of the electrons from the decay is between 3.29nm-231 μm , according to simulation done by Howell (1992), so well within cellular nucleus. In its decay, it emits 26.4 coster-konig and auger electrons (energy realeased per decay: 10.353 keV) and internal 3 conversion electons (energy released per decay: 126.738 keV). According to the simulation, an additative 12.345 keV is for X-ray energy deposition per decay.

Production: there are multiple ways that this isomer can be produced, either in a neutron field in a reactor, or in a charged particle accelerator like a cyclotron: $^{192}\text{Pt}(n,\gamma)$ or via $^{192}\text{Os}(\alpha,3n)$. One of the issues with production is that $193\text{m}\text{Pt}$ (and $195\text{m}\text{Pt}$) are difficult to produce with high specific activity (Qaim 2016), and are not well investigated. This study gives an examination of a new route. Many reasons, reactors are on their way out, and Osmium is a poisonous and difficult target to work with, so using iridium as target is easy, (expensive though?) and the production of radionuclides below iridium is evidently in this work and in papers tarkanyi et al (2006,2019) low.

By itself, not useful for imaging. 191Pt and $195\text{m}\text{Pt}$ can. Can replace stable N with 13N , but the half life is so short that the radionuclide can not image the distribution it self, so not as a theranostics pair?? or does cisplatin distribute so fast within the body?

Pt-poisoning

⁵<https://sci-hub.tw/https://doi.org/10.1118/1.596927>

1.2 Radioactive decay law and Gamma-ray spectroscopy

Gamma-ray spectroscopy is a method to identify and obtain information about radioactive nuclei present in a detector. As beta and alpha decay can result in an excited daughter product, the spectrum in fact shows the de-excitation of the daughter product. Since we know that these gamma-lines are transitions which happens right after a beta or alpha decay (or isomer transition), we identify the parent with gamma-ray spectroscopy. A detector has channels in which counts are registered. These channels are ... similar to the gamma-ray energy. Thus a spectrum has channels (which increases in energy) along the x-axis and counts along the y-axis. If a detector registers many counts, it means that the state is highly populated, and the intensity of the gamma is strong (Krane, p. 351).

1.2.1 Radioactive decay law

From here based on Krane chapter 6⁶

The activity of a nucleus is defined as the number of decayed nuclei per unit time of a radioactive product, which is equal to the radioactive decay rate

$$A = \frac{dN}{dt} = -\lambda N \quad (1.16)$$

where N is the number of nuclei, t is the time and λ is the decay constant. Solving equation 1.16 gives number of decayed products at time t

$$N(t) = N_0 e^{-\lambda t} \quad (1.17)$$

Since $N \propto A$, the relations $\frac{N_0}{A_0} = \frac{N(t)}{A(t)}$ are valid, and we can rewrite the equation 1.17 to

$$A(t) = A_0 e^{-\lambda t} \quad (1.18)$$

This accounts for single nucleus decaying into a daughter product, without anything first decaying into the parent nucleus. However it is common that a radioactive nucleus decays into another radioactive nucleus. Hence the daughter activity will increase due to feeding from the parent. For multiple decay, Bateman equation is used describing the activity in nucleus n of the decay chain ([Voyles2018, which article??](#))

$$A_n = \lambda_n \sum_{i=1}^n \left[\left(A_{i,0} \prod_{j=i}^{n-1} \lambda_j \right) \cdot \left(\sum_{j=i}^n \frac{e^{-\lambda_j t}}{\prod_{i \neq j} (\lambda_i - \lambda_j)} \right) \right] \quad (1.19)$$

where A_n is the activity of nuclei n in the decay chain, with the corresponding decay constant λ_n . The equation sums over all nuclei in the decay chain. $A_{i,0}$ is the initial activity of nucleus i , and j is the nucleus which is feeding into nucleus i .

If a target of stable nuclei is assumed, which is exposed to a particle beam which induces various nuclear reactions, the constant rate of production of a specific reaction is dependent on the number of target nuclei, the current of flux of the particle beam and the reaction cross section

$$R = N_T \Phi \sigma \quad (1.20)$$

where R is the production rate, N_T is the number of target nuclei, Φ is the beam current or flux and σ is the reaction cross section. In the assumption of the production rate being a constant value, the number of transformed target nuclei is small in comparison to the total number during the irradiation time. The number of produced nuclei from a specific reaction per unit time is thus the produced nuclei minus the decayed nuclei (activity)

$$dN = R dt - \lambda N dt \quad (1.21)$$

which has the solution

⁶<https://faculty.kfupm.edu.sa/phys/aanaqvi/Krane-Ch-6.pdf>

$$N(t) = \frac{R}{\lambda}(1 - e^{-\lambda t}) \quad (1.22)$$

From equation 1.16, the total activity produced during irradiation time t is thus

$$A(t) = R(1 - e^{-\lambda t}) = N_T \Phi \sigma (1 - e^{-\lambda t}) \quad (1.23)$$

At the end of beam, the activity is denoted as A_0 , and t is the irradiation time:

$$A_0 = N_T \Phi \sigma (1 - e^{-\lambda \Delta t_{\text{irr}}}) \quad (1.24)$$

When a target is irradiated, the activity of the product nucleus will increase until secular equilibrium is achieved, which is when the product rate and decay rate are constant. Hence it is not necessary to irradiate a target for more than 2-3 half lives.

If a spectrum is counted at a delay time Δt_d after end of beam with a counting time Δt_c the total number of decayed products are

$$N_D = \int_{\Delta t_d}^{\Delta t_d + \Delta t_c} A(t) dt \quad (1.25)$$

Using equation 1.18 for $A(t)$, the solution to the above equation is

$$N_D = \frac{A_0}{\lambda} e^{-\lambda \Delta t_d} (1 - e^{-\lambda \Delta t_c}) \quad (1.26)$$

which again is equal to

$$N_D = \frac{A(t)}{\lambda} (1 - e^{-\lambda \Delta t_c}) \quad (1.27)$$

We can only know the number of decayed products which are detected. This is dependent on the efficiency of the detector, the intensity of the gamma-rays and the true number of decayed products

$$N_C = N_D \epsilon I_\gamma \quad (1.28)$$

where N_C is the number of observed/collected gamma-rays, ϵ is the efficiency of the detector and I_γ is the gamma-ray intensity.

Thus, we can obtain an expression for $A(t)$ after a delay time:

$$A(t) = \frac{N_C \lambda}{\epsilon I_\gamma (1 - e^{-\lambda \Delta t_c})} \quad (1.29)$$

Again using 1.18 for $A(t)$, the above expression can be rewritten using A_0 and the delay time Δt_d

$$A_0 = \frac{N_C \lambda}{\epsilon I_\gamma (1 - e^{-\lambda \Delta t_c}) e^{-\lambda \Delta t_d}} \quad (1.30)$$

1.2.2 High purity Germanium detector

High purity Germanium detector is a type of semiconductor, which is a material where the energy required to remove an electron from the valence band (in the outer atomic shell) to the conduction band is small. The germanium atom has atomic number 32, and 4 valence electrons in the outer p4 shell (need citation?). The atoms in the detector are bound through covalent bonds in a crystal structure. The main mechanism of a semiconductor is creation of electron-hole pairs after energy deposition of an ionizing particle in the crystal. If an electron is excited to the conduction band, a hole is left. This hole can move as a neighboring electron fills this spot, and it can cause a chain reaction, and the hole will move in the crystal. Both the electron in the conduction band and the hole in the valence band contributes to an electric current. Under influence of an electric field, the electron-hole pairs will be collected and we can measure the incident as a count. The major

advantage with semiconductor detector is that the average energy to create an electron-hole pair is very low, which results in a superior energy resolution in comparison to other detectors like gas and scintillation detectors. High energy resolution advantageous in gamma-ray spectroscopy which makes it possible to separate gamma-ray peaks within less than a keV. At room temperature, thermal energy can excite the electron from the valence to the conduction band and cause noise in spectra. Therefore, Germanium detectors are operated at 0 Kelvin. Write about recombination and trapping, noise, np semiconductor junction, depletion depth?? (Techniques for Nuclear and Particle Physics Experiments, William R. Leo, p. 215-216).

Ideally, for all gamma-rays with the same energy, should be detected in the same channel giving a step function. However, realistically, the resolution of a detector is not that good, and instead of seeing a delta peak, the peak is typically gaussian shape with a finite width. The full width half maximum ΔE of the peak tells us how well the relative resolution at gamma-energy E ,

$$\text{resolution} = \frac{\Delta E}{E} \quad (1.31)$$

The energy resolution is important, as it tells us how well it can distinguish two close lying peaks from each other (Techniques of Nuclear and particle Physics.. , p. 117). The resolution of a germanium detector very good (0.1% for a 1 MeV gamma-ray) in comparison to for instance NaI detector (8-9% for a 1 MeV gamma-ray) (Techniques of Nuclear and particle Physics.. , p. 117). **explain why, prob in semiconductor chapter!**

The peak itself is not directly gaussian. Ionizing radiation statistics is based upon Poisson statistics, where the probability of observing N events is a discrete value

$$P(N) = \frac{\mu^N e^{-\mu}}{N!} \quad (1.32)$$

where μ is the mean value. This distribution counts when the probability is a small (eg decay prob?) value and that the total number of trials are large (number of decays) (Techniques of Nuclear and particle Physics.. , p. 85). For poisson distribution, the average is equal to the variance; $\sigma^2 = \mu$. From there, the standard deviation (σ) is thus equal to the square root of the average.

The distribution is not symmetric, but as μ increases in value, the peak approxes a gaussian shape. The total number of counts is the area of the peak. The total peak is a Gaussian assumption but with an exponential skew towards low E caused by incomplete charge collection, and a step function for taking compton background into account.

In calculation of the peak area, there are two uncertainties of relevance, the relative statistical uncertainty in the counting from the Poisson statistics,

$$\sigma N_i = \sqrt{N_i} \quad (1.33)$$

If number of counts $N_i = 10000$, the relative uncertainty ($\frac{\sigma N_i}{N_i} = \frac{1}{\sqrt{N_i}} = 1\%$). Therefore we say that a good number of counts is 10000 or more to reduce the statistical uncertainty. The other is systematic in the detector, and can for instance be due to a process called annealing, which is heat damage to the detector. Can fix by taking a blanket of resistor wrap crystal in, rise to high temp, let it sit and slowly deheat to room temp, traps will diffuse and detector is repaired (this is notes from Andrew).

Also write about deadtime!

1.2.3 Gamma-ray spectrum

Spectrum: consists of photopeaks, a compton continuum, compton edge, backscatter peak, single escape double escape. In cases where positrons exist, chances of having a broad fat 511 keV peak.

Germanium detectors, highest resolution for gamma-rays, from few keV to 10 MeV. The peak to Compton ratio is much greater due to the higher photoelectric cross section of Germanium . The largest challenges are with signal to noise ratio, it is important to shield very well to minimize background radiation (Techniques for Nuclear and particle..... William R. Leo, p. 241).

here from another citation: "Practical Gamma-ray Spectroscopy". Gordon R. Gilmore. Nuclear Training Services Ltd Warrington UK. (can be find under articles in masterthesis). This book can also be used in particle interaction in matter check!! In a detector, the particles interacts as the photons described in particle interaction, via photoelectric, compton scattering and pair production. Photoelectric absorption where the photon is completely absorbed by atomic electron is desired because all of the energy is deposited within the detector. For a compton scattering event, if the resulting photon's energy is also deposited in the detector (for a large detector), then the total energy would add up. Same for pair production. The photon must interact in the detector volume, and the resulting electron and positron energy is deposited in the detector volume. However when the positron slows down, it annihilates with one atomic electron, releasing two 511 keV photons. If both annihilation photons's energy is deposited in the detector volume this will also contribute to a full width peak. If one 511 photon escape and the other is deposited, there will be a peak at $E_{\gamma} - 511$ keV, and if both peaks escape, there will a double escape peak at $E_{\gamma} - 1022$ keV. The "degree of incomplete absorption" depends upon the size of the detector and the gamma-ray energy. As previously discussed photoelectric effect dominates at low energies, and the less compton scattering and of course pair production (for E gamma higher than the threshold.). The detector size also matters because the larger the more room for the photon to scatter in and lose energy before escaping. (p. 32)

The total spectrum can be seen on p. 33 in the book. Pile-up is done because of random summing, determined by the statistical probability of two gamma-rays being detected at the same time and therefore on the sample count rate.

Interaction with detector shielding: Photoelectric effect can be followed by emission of characteristic X-ray of the absorbing medium. X ray can escape the shielding and be detected by the detector. Compton scattering: most gamma rays are scattered through a large angle by the shielding, BACKSCATTERED. Whatever the initial energy was (if scattered by more than 120 degrees) are within 200-300 keV. Peak appears as broad. Pair production: annihilation peak (511 peak) caused by the escape of one of the 511 keV photons from the shielding following annihilation of the pair production positron. Analogous to the single and double escape mechanisms within the detector but only on 511 keV photons can ever be detected since they are emitted in the opposite direction. So in order to have a 511 peak, energy of gamma ray must be more than 1022 keV. (p. 34-35).

The 511 peak can also be expected when positron emitters are present since beta + particle interacts with electron.

Since Compton scattering can be in a spectrum of energies, the gives rise to a Compton continuum, before the gamma-ray escapes the detector.

The shape of the peak: The peak is a histogram that approximate a Gauss curve (p. 186). Peak searching (SAMPO) using first and second order derivatives to search for peaks (p.185) Due to incomplete charge collection (that electron or holes are not collected) no matter how caused moves counts from the centre of the Gaussian distribution to lower channels, creating a low energy tail to the peak (p.135).

Include a picture of peak shape and gamma-ray spectrum!! from the same book

1.3 General nuclear reaction theory

paragraph based on special curriculum Medical radionuclides can be produced directly using charge particle (cyclotron) or neutron beams (reactors), or indirectly using radionuclide generators or fission (reactor). Medical radionuclides are typically produced in reactors, cyclotrons or by a longer lived-parent decaying into a short-lived daughter in a radionuclide generator system. In general, the production should be cheap, available. Today many radionuclides are only produced in reactors, which is the main source of neutrons, and with reactors aging (Chai Hong Yeong, Mu hua Cheng, and Kwan Hoong Ng. Therapeutic radionuclides in nuclear medicine: Current and future prospects. Journal of Zhejiang University: Science B, 15(10):845–863, 2014.), we need alternative routes to produce critical

radionuclides. Cyclotrons have many benefits, like size so that it can be produced directly at the site of usage. One of the major disadvantages is that there is a need to enriched targets to get the desired reaction, and those can be very expensive. Along with high beam intensity the melting of the target can give challenges, so target cooling techniqueis need to be there.

In order to create isotopes, nuclear reactions need to occur. There are many different production routes available for a single radionuclide, which is dependent on multiple factors such as choice of target, incident particle-beam and beam energy. To each reaction route, there is an corresponding excitation function which tells us how probable the reaction channel is at various energies. The nuclear reaction data is very important for the optimization of the product, achieving minimal level of isotopic impurities and maximum yield (S M Qaim, R Capote, and F Tarkanyi. Nuclear Data for the Production of Therapeutic Radionuclides. Trs 473, (473):395, 2011., p. 3).

Isotopic purity is important as it is impossible to separate isotopes of the same element (Syed M. Qaim. Nuclear data for production and medical application of radionuclides: Present status and future needs. Nuclear Medicine and Biology, 44:31–49, jan 2017.). An undesired radionuclidel can lead to undesired dose to healthy tissue, and a non-radioactive nuclide may lead to poisoning (if large amounts injected), but it will not have any therapautic effect. This is especially important when working with poisoeneos elements such as platinum. The only option to minimize isotopic impurities is to choose an appropriate energy window.

Using charged particles instead of neutrons allows for measurement at multiple energies as the particle energy degrades in the foils. The neutron energy is not degraded in the same way, due to electric neutrality, thus can only give cross section at one single energy.

1.3.1 Nuclear reactions and reaction cross sections

A nuclear reaction occurs when a collision between two nuclei or a nucleus and a subatomic particle takes place. Collision between an accelerated subatomic particle or small nucleus and target nuclei is common in isotope production. A nuclear reaction is denoted as

$$X(a, b)Y \quad (1.34)$$

where X is the target, a is the incoming projectile, b is the outgoing decay channel and Y is the product of the nuclear reaction (Krane, chapter 11.1). There are multiple processes which can occur, radiative capture is the process where a particle is captured and a γ -ray is emitted in a (x, γ) process. If the incoming and outgoing particle is the same, it is a scattering process, where elastic scattering leaves the target nucleus in the energy same state, and inelastic if the target nucleus is in an excited state. In these type of experiments however, we are interested in emission of particles to create products in which we can measure the reaction cross section.

In a nuclear reaction, the total energy and linear momentum, proton and neutron number, angular momentum and parity are conserved quantities (assuming no meson formation) (Krane, p.380). In the low energy-region in which isotope production typically takes place (~~j80 MeV?~~), compound nucleus reactions take place, where an incoming particle and target nucleus merges by sharing the kinetic energy, and particle emission takes place to reduce the excess energy. The contrary are direct reactions, where an incoming particle interacts with the only the surface of the particle.

A nuclear reaction in the energy ranges of isotope production via the Compound nucleus conserves quantities such as mass-energy, linear momentum, the total number of nucleons, angular momentum and parity ~~cite, Krane?~~.

The cross section for a reaction can be divided into the cross section of the formation of the compound nucleus via interaction with the incoming projectile a, and the probability that the compound nucleus decay by decay channel b. The total reaction cross section is thus the sum of all the different reaction channels,

$$\sigma = \sum_b \sigma(a, b) \quad (1.35)$$

where b can be multiple particles. The general equation which is used to calculate cross sections in this experiment (solving equation 1.30) is the following equation

$$\sigma(E) = \frac{A_0 \cdot t_{\text{irr}}}{N_T \cdot \Phi(E)(1 - e^{-\lambda t_{\text{irr}}})} \quad (1.36)$$

where A_0 is the end of beam activity of the resulting product nucleus (Y), t_{irr} is the irradiation time, N_T is the number of target nuclei (X), $\Phi(E)$ is the projectile flux or current (a), and λ is the decay constant of the product nucleus.

The compound nucleus model (Bohr, 1936) is a model which describes the formation of a compound nucleus by absorption of an incoming projectile by a nucleus close enough to interact with the strong nuclear force, and the decay of the compound nucleus. The kinetic energy shared between the incoming projectile and the nucleon which was struck leads to multiple collisions with other nucleons and rapid exchange of energy. The energy is distributed throughout the nucleus, leaving the original nucleus in an highly excited state. The average energy per nucleon is not sufficient to overcome the binding energy of the nucleus, but due to the statistical distribution in energies there is a probability that one or more nucleons may get sufficient energy to escape the nuclear potential (Krane, chapter 11.10, p. 416). This is decay of the compound nucleus, and this will lower the excitation energy. We can include the formation of the compound nucleus in the nuclear reaction as



where C^* is the excited compound nucleus (Krane, chapter 11.10, p. 416)

For each possible decay channel of the compound nucleus, there is an associated probability or cross section, which is dependent on the energy of the incoming projectile. A function which evaluates the various cross sections at different energies is called an excitation function. In figure 1.4, the excitation function of the reactions channels for the platinum isotopes $^{188,189,191,193m}\text{Pt}$ resulting from deuterons on natural iridium is plotted (using TENDL nuclear reaction code [cite](#)). Natural iridium consists of two stable isotopes, ^{191}Ir (37.3% abundance) and ^{193}Ir (62.7% abundance). ^{193m}Pt can only be produced from ^{191}Ir , ejecting 2 neutrons in the process, which can be denoted as $^{193}\text{Ir}(d,2n)^{193m}\text{Pt}$ (^{193}Pt is the compound nucleus formation of deuteron on ^{191}Ir , which has a low production cross section). The other platinum isotopes can be produced as $^{191}\text{Ir}(d,2n)^{191}\text{Pt}$ or $^{193}\text{Ir}(d,4n)^{191}\text{Pt}$, $^{191}\text{Ir}(d,4n)^{189}\text{Pt}$ or $^{193}\text{Ir}(d,6n)^{189}\text{Pt}$ and $^{191}\text{Ir}(d,5n)^{188}\text{Pt}$ or $^{193}\text{Ir}(d,7n)^{188}\text{Pt}$. For each reaction route possible, there is a local maximum for the specific route, hence, ^{193m}Pt has only one maxima, and the other platinum isotopes has two. The desired particle emission is energy dependent, and the higher energy provided to the compound nucleus, the probability that more particles will be emitted is higher (Krane, chapter 11.10, p. 419). When a specific isotope is desired, the excitation function can tell us which energy window that maximizes the production and most importantly minimizes particularly other isotopes of the same element, due to the difficulty of separating same chemical elements.

1.3.2 Constraints in nuclear reactions

The nucleus is bound together by the strong nuclear force in the nuclear potential well, with a radius up to a few femtometer. The nuclear potential can be illustrated in figure 1.5. The Coulomb barrier is a repulsive barrier for charged particles. For a charged particle induced nuclear reaction, the energy should exceed to potential, or there will be an elastic scatter. However, there is a chance of tunneling, which drops with a factor $1/r$ where r is the distance from the center of the nucleus (Handbook of Nuclear Chemistry, chapter 3 - Nuclear Reactions, section, 3.2.3). The barrier also constraints the emission of particles for a decay channel of the compound nucleus, as the energy for an outgoing decay channel of positive particles must exceed the barrier. There is also a centrifugal barrier, which is dependent on the orbital angular momentum of the the nucleus.

The height of the Coulomb barrier is dependent on the radius and charge of the incoming or outgoing particle a and the target nucleus b.

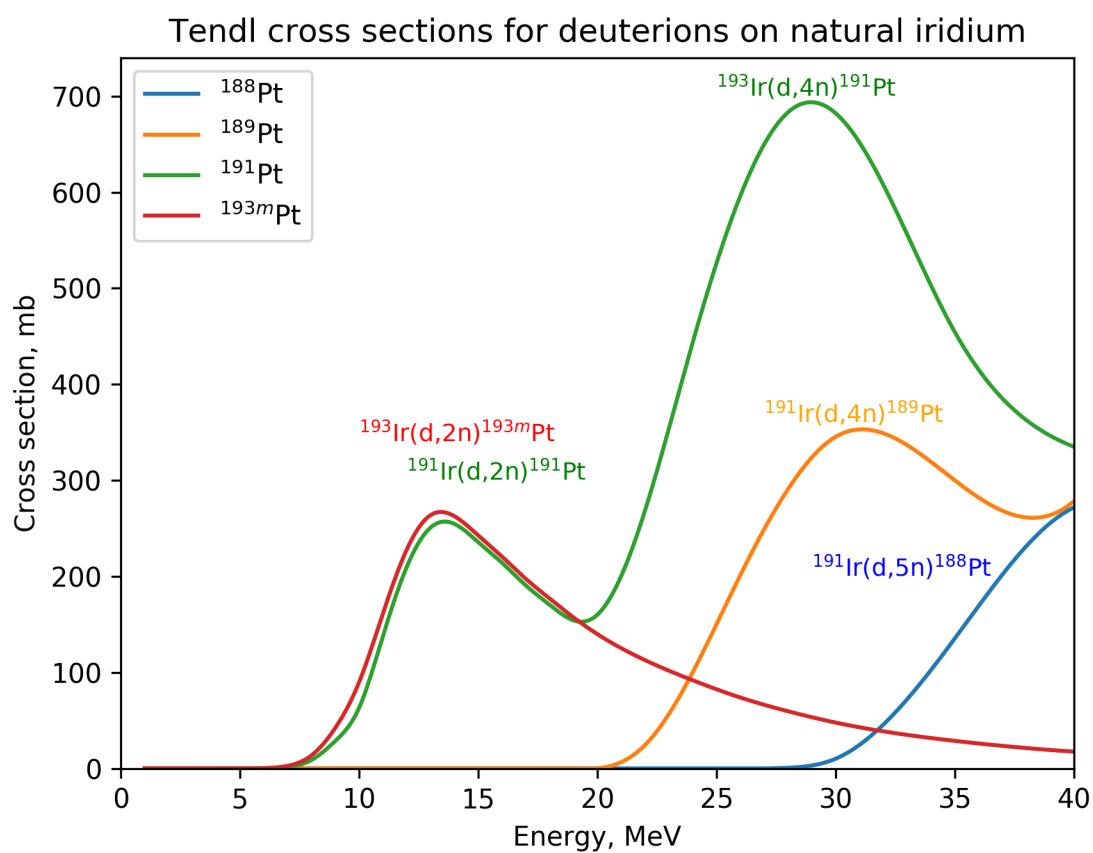


Figure 1.4: Reaction cross sections provided by Tendl for the reactions ${}^{\text{nat}}\text{Ir}(d,x)^{188,189,191,193m}\text{Pt}$

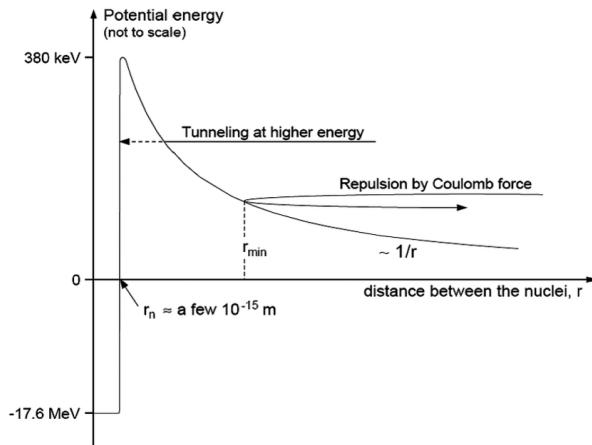


Figure 1.5

$$U_C = \frac{1}{4\pi\epsilon_0} e^2 \frac{Z_a Z_b}{r_a + r_b} \quad (1.38)$$

In addition

In a nuclear reaction, the mass-energy is conserved, which is denoted as the Q-value. The reaction Q-value is the difference in masses between before and after the nuclear reaction occurred (Krane, chapter 11.2). It is defined as

$$Q = (m_i - m_f)c^2 = (m_X + m_a - m_Y - m_b)c^2 \quad (1.39)$$

where m_i is the initial mass, m_f is the final mass and c is the speed of light. If $Q > 0$, then the reaction is exoergic, which means that energy is released in the reaction. There is no threshold energy of the projectile required for the reaction to occur, if only the projectile is present the reaction can occur. If $Q < 0$, then the reaction is endoergic, which means that the kinetic energy of the incoming projectile is converted into nuclear mass or binding energy. For endoergic reactions to occur, there is a minimum threshold energy of the projectile in order for the reaction to happen, which is defined as (Krane, 11.2, p. 382)

$$E_{\text{threshold}} = (-Q) \cdot \frac{m_Y + m_b}{m_Y + m_b - m_a} \quad (1.40)$$

The energy threshold thus depend on the Q-value, the Coulomb barrier for charged particles, and the centrifugal barrier if angular momentum $\ell \neq 0$. The parity though depend, even numbers of ℓ mix with even, and odd with odd (Handbook of Nuclear Chemistry, chapter 3Nuclear Reactions, section, 3.2.3). This gives an indication on when a reaction can energetically occur, but does not tell us how probable the reaction is. Equation 1.40 indicates that a higher mass for the particle in the outgoing channel m_b , will lower the energy threshold.

The nuclear binding energy is a number which tells us how tightly bound the nucleus is, and much energy is required to separate nucleons from the nucleus. The nuclear binding energy is related to the mass of the nucleus, through the famous equation

$$E = mc^2 \quad (1.41)$$

The binding energy is the mass-difference between the nucleus as a whole, and the number of protons and neutrons added

$$B = c^2(z \cdot m_p + n \cdot m_n - m_N) \quad (1.42)$$

where z is the number of protons, n is the number of neutrons, m_p is the proton mass, m_n is the neutron mass, M_N is the mass of the nuclide, which is the number of nucleons A minus the number of electrons, $M_P = m_A - z \cdot m_e$ (the electronic binding energy per electron is excluded). From Krane's

derivation of the nuclear binding energy (Krane, chapter 3.3, p. 65), the equal number of protons and neutrons make up $z \cdot {}^1 H$, thus we can rewrite equation 1.42 to

$$B = z \cdot m_1 H + n \cdot m_n - m_A \quad (1.43)$$

Separation energy of a particle is the energy required to be removed from the nucleus. This is defined as the difference in binding energy between the nucleus with the particle and the nucleus without the emitted particle. Since heavier particles such as tritium has a lower binding energy, more energetically easy to separate. n and p require more.

Write about the semi-empirical mass formula, and what terms that affects the decay of compound. Especially the parity term? Odd A or even A. If A odd, decay by proton, since then A is even. If A is even, then decay by alpha for A to remain even. But why is n,p favoured over for instance t??

1.3.3 Nuclear cross sections

The cross section for a reaction can be divided into the cross section of the formation of the compound nucleus via interaction with the incoming projectile a, and the probability that the compound nucleus decay by decay channel b. The total reaction cross section is thus the sum of all the different reaction channels,

$$\sigma = \sum_b \sigma(a, b) \quad (1.44)$$

where b can be multiple particles. The general equation which is used to calculate cross sections in this experiment is the following equation, which is solved from equation 1.24 (referenced to in the later parts of the paper)

$$\sigma(E) = \frac{A_0}{N_T \Phi(E) (1 - e^{-\lambda \Delta t_{\text{irr}}})} \quad (1.45)$$

not, need to check if this is the correct one..

$$\sigma(E) = \frac{A_0 \cdot t_{\text{irr}}}{N_T \cdot \Phi(E) (1 - e^{-\lambda t_{\text{irr}}})} \quad (1.46)$$

where A_0 is the end of beam activity of the resulting product nucleus (Y), t_{irr} is the irradiation time, N_T is the number of target nuclei (X), $\Phi(E)$ is the projectile flux (a), and λ is the decay constant of the product nucleus.

Why we are interested, why is there a need for nuclear production cross sections.

The Q-value for nuclear reactions

Compound nucleus, what happens. Q value, binding energy, particle emission, probabilities (why does n, p/alpha emission require higher energy, but is more favoured than T emission if both possible.)

1.3.4 Expected products from deuterions on iridium, iron, copper and nc

Include figure of different nuclear charts, and write about Q values, and reaction channels, which makes it more probable that we observe it (eg. when neutron, proton/alpha channels opens).

1.4 From book

In a nuclear reaction, the energy, linear momentum, the total number of protons and neutrons, angular momentum and conservation of parity is conserved.

The compound nucleus; the collisions are described by a model called Multistep Compound Model (Feshback, Kerman and Koonin, 1980), which is the model that reaction modeling code EMPIRE and TALYS rely on. **find out what ALICE does ???**

Products of a nuclear reaction can be anything than conserve the various variables. Mass energy conservation is the Q-value, which is the mass difference in the reactant nuclei (projectile and target in this case) and product nuclei (decay channel particles and product). Binding energy (nuclear mass energy) often denoted as $\Delta = (M_0 - Au)c^2$, where A is atomic mass number, u is atomic mass unit. Q-value can be rewritten:

$$Q = \Delta_{\text{projectile}} + \Delta_{\text{target}} - \sum \Delta_{\text{products}} \quad (1.47)$$

Charged particle reactions with heavy nuclei which is the case here, can have positive or negative Q-value. If the Q-value is negative, energy is required to induce the reaction. The reaction is negative if the sum of products is less than the sum of reactants. The excitation energy of the product nucleus is the extra energy which is deposited in the product nucleus. The decay of this is often times γ -rays (?). If the excitation energy is lower than the binding energy of the nucleus? Threshold is the minimum energy required for the incoming projectile to form a product nucleus, which is then in its ground state.

$$\text{mass-energy : } \Delta_t + \Delta_p + E_p \quad (1.48)$$

1.5 The decay of the compound nucleus

At the energy ranges in which this work is operating in, the nuclear reaction is caused by a fusion between the incoming projectile to create an highly excited compound nucleus.

The decay favoured decay channel is dependent on multiple properties such as the charge and angular momentum. The coulomb potential is defined as

$$U_c = K \frac{z_1 z_2 e^2}{A_1^{1/3} + A_2^{1/3}} \quad (1.49)$$

where z is the the number of charges, e is the electrical charge and $A^{1/3}$ is the nuclear radius which is equal to $r = r_0 A^{1/3}$. K is a constant value.

The centrifugal barrier (which is the amount of energy which is conserved in rotational energy) depends on the angular momentum

$$U_s = \frac{\hbar \ell (\ell + 1)}{2 \mu r^2} \quad (1.50)$$

where μ is the reduced mass $u \cdot \frac{A_1 A_2}{A_1 + A_2}$.

Table 1.1: The table shows the most common decay channels for the decay routes available in this energy region. Δ is the binding energy of the alpha particle, which can be calculated using equation 1.42, where the mass of the proton is $m_p = 938.28 \text{ MeV}/c^2$ and $m_n = 939.57 \text{ MeV}/c^2$. Δ is thus the difference in energy-mass between the particle composed of protons and or neutrons, and a equal number of protons and neutrons independent. All parities are positive due to angular momentum $\ell = 0$.

Particle	$\Delta(\text{MeV})$	J^π	Charge
p	-	$\frac{1}{2}^+$	+1
n	-	$\frac{1}{2}^+$	0
d	2.2	1^+	+1
t	8.5	$\frac{1}{2}^+$	+1
${}^3\text{He}$	7.7	$\frac{1}{2}^+$	+2
α	28.3	0^+	+2

Target	Proton	neutron	deuteron	triton	${}^3\text{H}$	α
Fe (Z=26)						
Ni (Z=28)						
Cu (Z=29)						
Ir (Z=30)						

Table 1.2: The Coulomb barrier in MeV for each of the targets. Can be used to explain why we see what we see.

Table 1.1 might be wrong in the sense J^π . Might just be the total spin which is of interest? But since $J = l + s$, the spin will give a higher intrinsic angular momentum. Since $|I_i - I_f| \leq \ell \neq I_i + I_f$, ℓ can be multiple values, and the parity is thus dependent on the value here. (Krane, chapter 8.5, p. 257). The centrifugal potential is thus

$$U_{0,s} = \frac{\ell(\ell+1)\hbar}{2mr^2} \quad (1.51)$$

and the Coulomb potential is

$$U_{0,c} = K \frac{z_1 z_2 e^2}{A_1^{1/3} + A_2^{1/3}} \quad (1.52)$$

The

The total angular momentum J is the sum of the orbital and intrinsic angular momentum (spin), thus

$$\mathbf{J} = \ell + \mathbf{s} \quad (1.53)$$

The spin is the sum of unpaired paired protons and neutrons. The total nuclear spin is thus $I = \sum_i J_i$. The value of the angular momentum is

$$|I_i - I_f| \leq \ell \geq |I_i + I_f| \quad (1.54)$$

Thus,

Chapter 2

Experimental setup

2.1 Lawrence Berkeley National Laboratory's 88" Cyclotron

Lawrence Berkeley National Laboratory (LBNL) is a national research laboratory on behalf of the U.S. Department of Energy through its Office of Science, and is operated by University of California, Berkeley. LBNL was founded by Ernest Orlando Lawrence, the inventor of the cyclotron ¹.

The 88" Cyclotron has many purposes, and can accelerate both light and heavy ions up to Uranium, with a cyclotron number $K=140$ ². The cyclotron number is the maximum kinetic energy which can be reached for protons (with no relativistic factors taken into account). The maximum kinetic energy a particle can gain is found from the cyclotron number:

$$\frac{E_k}{A} = K \left(\frac{Q}{A} \right)^2 \quad (2.1)$$

For deuterons with mass number $A=2$ and charge $Q=1$, the maximum kinetic energy is $E_k = 70$ MeV. There are multiple programs that takes place in the facility; chip testing and space effects testing, super heavy element searches, fundamental nuclear structure measurements, novel scintillation characterization, fission yield and neutron inelastic scattering measurements (GENESIS) (from Andrew). Medically motivated isotope production goes under fundamental nuclear structure experiments.

A cyclotron is a device that accelerates positively charged particles. It is operated by an alternating electric field, and a perpendicular magnetic field, which by the Lorentz Force forces the particle to accelerate in an outward spiral. The facility is figured in figure 2.1, which consists of a cyclotron vault, and experimental caves in which the beam can be bent to with bending magnets. Faraday cups (not in figure) can measure the beam current at different steps along the tube, which makes it possible to measure the transition efficiency of the beam. Faraday cups are dense metal block, usually 6-7 cm broad Copper and Tantilum. It works as a beam stopper, and can be lowered into the beam line to measure the current. It is electrically isolated, which makes it possible to measure the current, since we know the number of initial particles accelerated. Due to electrons close to surface might be scattered off, it can read off higher positive charge than what is correct. Therefor, a magnet surrounds the cup to bend the electrons back to the Faraday cup in what is called magnetic suppression. Cave 0 is used mainly for neutron beam, chemistry, and isotope production, and was used for irradiation of the target stack.

2.2 The experiment

The main motivation of this experiment was to measure production cross sections of the products produced after irradiation of a stack of thin iridium foils along with thin monitor foils Nickel, Copper and Iron foils, with a 33 MeV incident deuteron beam, as shown in figure 2.2.

¹<https://www.lbl.gov/about/>

²<http://cyclotron.lbl.gov/home>

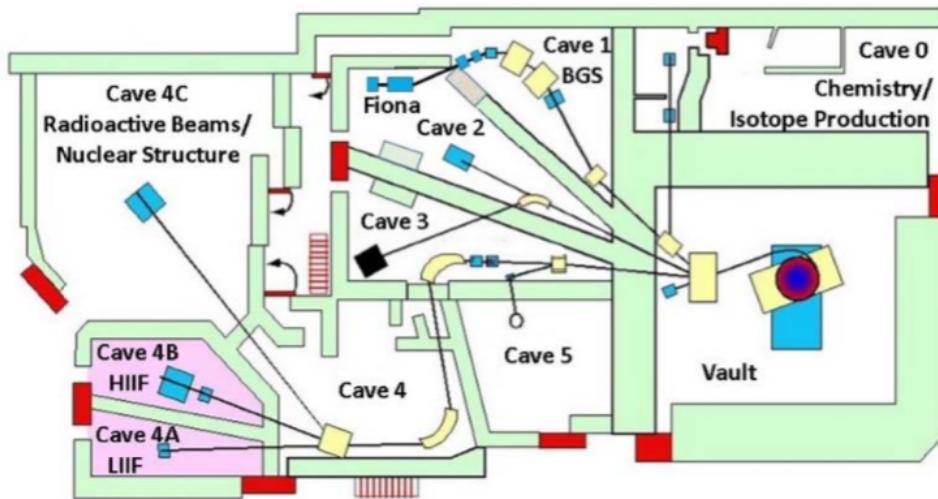


Figure 2.1: An overview of the 88" Cyclotron facility. <https://cpb-us-e1.wpmucdn.com/sites.usc.edu/dist/7/89/files/2018/04/133-18q03um.pdf>

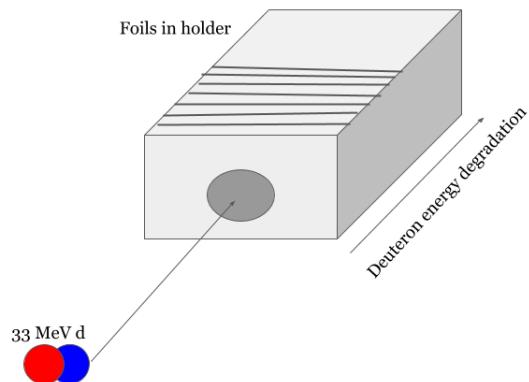


Figure 2.2: The fundamental idea of the experiment where a stack of targets are placed in a target holder, and irradiated with accelerated 33 MeV deuterons. As the energy degrades through the beam stack, it is possible to have multiple cross section measurements at different energies.

The beam was ca. 1 cm in diameter, and with each target foil being ca. 25 by 25 mm, the beam was underfilled. As the energy was degraded through the stack, multiple cross sections at different energies were possible to measure for the different induced reactions. For production cross section data experiments, thin targets (foils) are used, in the other of a few μm^3 ³ are used, since the induced activity is low, meaning that the deadtime of the detector and the dose to human will be low. Equation 1.45 is the equation which is used in the calculation of the cross sections. The end of beam activity goes into the equation, which is given in equation 1.29. The activity from spectra measured at different delay times after end of beam can be found, with respect to number of counts, efficiency of detector, intensity of gamma-rays, the decay constant of the nucleus λ and the counting time Δt_c . As we know that radioactive decay curves follows equation 1.19, and dependent on how many step the decaychain consists of, the end of beam activity can be estimated by extrapolating backwards in time with a curve fit.

In order to calculate the cross section of a product, end of beam activity, number of target nuclei and beam current must be found, where for the end of beam activity, the detector efficiency need to be estimated. The nickel, copper and iron deuteron-induced cross section are well-established, and can be used to determine the beam current throughout the stack.

2.2.1 Target design and foil characterization

In this experiment, the target stack was composed of ten natural iridium foils (99.9%) foils, along with ten natural nickel foils (..%), ten natural copper foils (..%) and three natural iron foils (..%) serving as monitor foils. Along with two stainless steel foils in the front and the back of the stack, a proton degrader (a 6061 aluminum alloy), and an extra nickel neutron monitor foil was used to obtain production cross sections at multiple energies, using one incident deuteron beam. The full order of the stack can be seen in table 2.1.

Each foil were cut into approximately 25 by 25 mm squares, and each foil was characterized using a caliper to measure the length across each side, a gauge caliper to measure the thickness and a weight to measure the mass of each foil which was prewashed with isopropanol. For each measurement, the unit was measured 4 times, and the values states in table 2.1 are averaged values. The length and mass were used to measure the mass density. The thickness was not used in the calculation of the mass density, but was a good indication that the foil thicknesses were more or less consistent. **For underfilled beams, the mass density of the foil is used to find the number of nuclei per cm^2 , by using the area of each foil.** The mass density was calculated using the mass of each foil divided by the area

$$\rho\Delta r = \frac{m}{A} \quad (2.2)$$

The uncertainty in each parameter was calculated using the standard deviation (equation A.2) of the four measurements per unit, and the total uncertainty was calculated using the approximation of uncorrelated variables used in equation A.13. The conversion from mg per cm^2 to nuclei per cm^2 was done numerically, by multiplying the mass density with Avogadro's number N_A and dividing by the mol-mass of the target atoms.

After the characterization, each foil was mounted on a plastic frame and attached with kapton tape along the edges (from previous experiments, kapton tape have shown to be a large **proton?** degrader, so it was important that the tape was not in the beamline **cite article? but which?**). The target frames can be seen in figure 2.3.

2.2.2 Irradiation of target stack

The irradiation included tuning of the beam and one hour of radiation over the target stack. Whenever the beam was turned on, the beam tube had to be pumped down to a vacuum, to not attenuate the

³(Syed M. Qaim. Nuclear data for production and medical application of radionuclides: Present status and future needs. Nuclear Medicine and Biology, 44:31–49, jan 2017.)

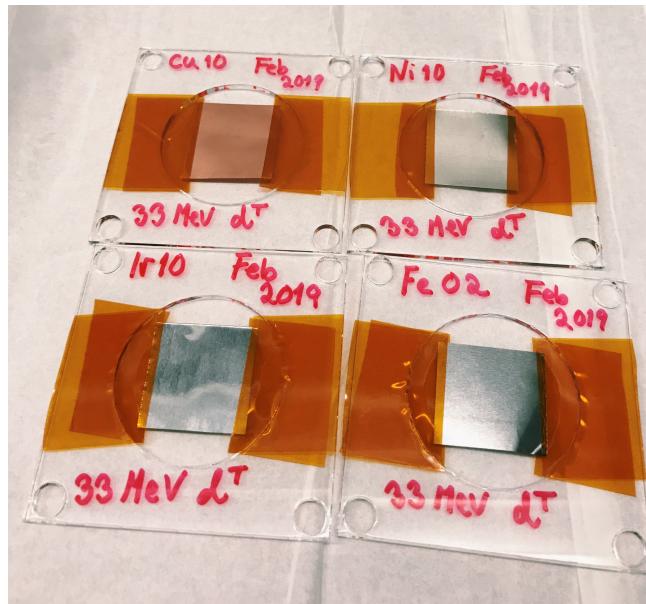


Figure 2.3: The figure shows the four different targets mounted on plastic frames with capton tape attached along the edges of the foils.

beam.

Tuning the beam

The cyclotron was tuned for a 33 MeV deuteron beam, and needed to have the correct beam spot. First, the beam spot was visualized using a ca. 2.5 cm thick borosilicate glass, painted with a mixture of phosphor powder and vacuum grease (so that the paint does not evaporate as the tube was pumped down to vacuum). When ionizing radiation strikes the phosphor, the phosphor is excited and emits light in the de-excitation, called phosphorescence. The glass is placed on the end of the beam tube. With a camera placed in cave 0, from the control room, the beam could be steered to be centered and ca. 1 cm in diameter. Secondly, visualization of the beam throughout the beam stack was important to see that the beam did not diverge converge/diverge or move in the wrong direction over the target stack. Gafchromic films which change color if struck by ionizing radiation was placed in the front and the back of the target holder. The films were exposed for a brief second, and the blue spot was evaluated. This was done until the beam gave good results both in the front and in the back of the stack.

The beam efficiency transmission was calculated by measuring the current at the Faraday cup right after the cyclotron vault (BS-02) and right before cave 0 (FC-01). BS-02 was measured to be 420 nA and FC-01 was measured to be 285 nA. This gave beam efficiency of transmission

$$\frac{FC - 01}{BS - 02} = 67\%$$

Irradiation of the target stack

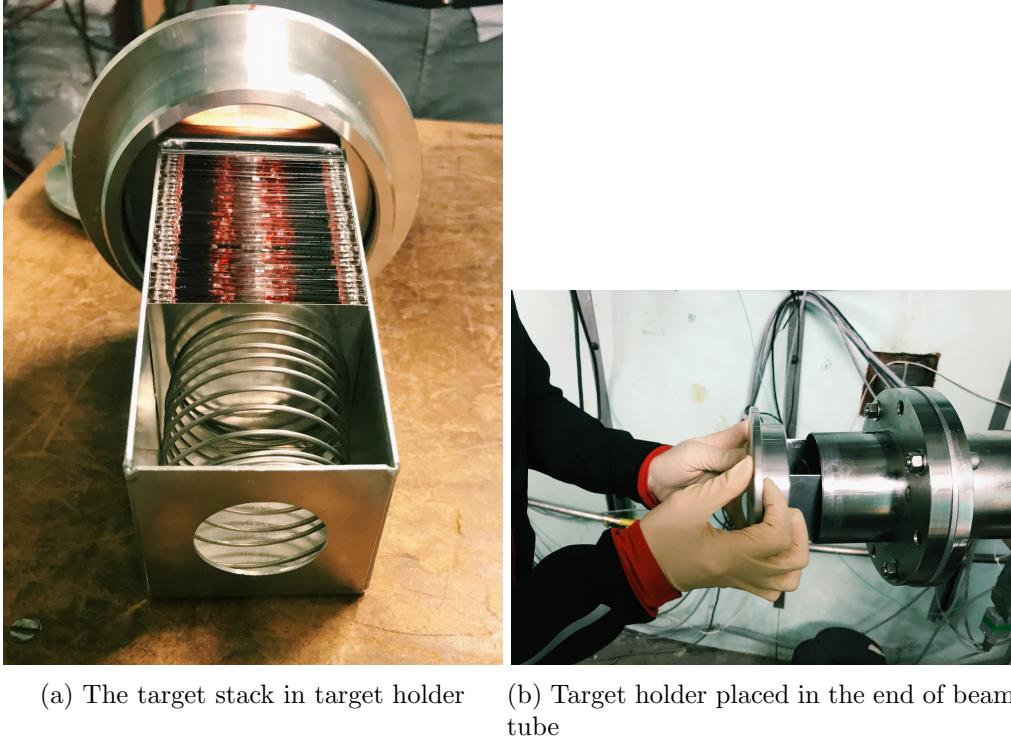
The targetstack was irradiated for exactly an hour, and the current was read of the beam integrator evenly, to assure that the it increased constantly. The beamcurrent from the beam integrator read 128.5 nA. Right after end of beam, the targets were sealed in plastic bags to avoid contamination. The foils were counted at the seven different detectors for the following 4 weeks after end of beam, first short counts to get as many observations as possible of the short-lived activities and longer and longer counts as the times since end of beam passes, so that the counting statistics for the longer lived activites are good.

Table 2.1: Characterization of each foil, along with calculated mass density. Each length is measured in mm, and mass in grams.

Foil	Length1 (mm)	Length2 (mm)	Thickness (mm)	Mass (g)	Mass density (mg/cm ²)
SS1	...				
Ni01	25.228	25.293	0.0285	0.1453	22.772 ± 0.138
Ir01	24.943	24.968	0.0295	0.3436	55.174 ± 0.053
Cu01	25.553	24.883	0.0341	0.1420	22.338 ± 0.048
Fe01	24.400	26.068	0.0278	0.1274	20.030 ± 0.110
Ni02	25.288	25.428	0.0295	0.1487	23.118 ± 0.096
Ir02	24.923	25.005	0.0278	0.3465	55.601 ± 0.238
Cu02	25.443	25.550	0.0348	0.1451	22.325 ± 0.028
Fe02	25.525	23.800	0.0274	0.1216	20.017 ± 0.034
Ni03	25.295	25.210	0.0270	0.1425	22.338 ± 0.066
Ir03	24.885	24.983	0.0243	0.3459	55.643 ± 0.121
Cu03	25.560	25.508	0.0343	0.1455	22.313 ± 0.043
Fe03	26.113	25.235	0.0310	0.1315	19.948 ± 0.114
Ni04	25.303	24.888	0.0273	0.1304	20.704 ± 0.068
Ir04	24.960	24.833	0.0261	0.3471	56.000 ± 0.109
Cu04	25.153	25.603	0.0333	0.1435	22.284 ± 0.027
Ni05	25.325	25.495	0.0263	0.1406	21.768 ± 0.045
Ir05	24.948	24.958	0.0256	0.3435	55.161 ± 0.081
Cu05	25.213	25.573	0.0334	0.1447	22.443 ± 0.028
Ni06	25.530	25.195	0.0285	0.1471	22.861 ± 0.123
Ir06	24.760	24.960	0.0240	0.3444	55.731 ± 0.088
Cu06	25.343	25.513	0.0340	0.1448	22.396 ± 0.012
Ni07	25.338	25.278	0.0268	0.1479	23.092 ± 0.078
Ir07	24.955	25.008	0.0278	0.3538	56.685 ± 0.085
Cu07	25.625	25.248	0.0326	0.1444	22.320 ± 0.014
Ni08	25.205	24.950	0.0256	0.1409	22.409 ± 0.124
Ir08	24.723	24.985	0.0281	0.3585	58.030 ± 0.130
Cu08	25.370	24.885	0.0333	0.1414	22.401 ± 0.033
Ni09	25.220	25.378	0.0257	0.1392	21.741 ± 0.073
Ir09	24.670	24.993	0.0273	0.3494	56.669 ± 0.043
Cu09	25.390	26.455	0.0331	0.1506	22.425 ± 0.041
Ni10	25.285	24.405	0.0271	0.1425	23.093 ± 0.024
Ir10	24.973	24.980	0.0270	0.3435	55.065 ± 0.055
Cu10	25.470	25.338	0.0355	0.1440	22.314 ± 0.047
SS2	...				
P-degrader	...				
Ni neutron monitor	...				

Intensity profile of the beam

After irradiation, gafchromic film were attached to the activated stainless steel in the front and the back of the stack, to obtain an intensity profile of the beam. The radius of the activity from stainless on gafchromic film is used in the imaging process program Image-J, which can be seen on figure 2.5. The gafchromic films were scanned, and the intensity data (**x and y arrays**) were obtained by inverting the scanned image, and drawing a line segment along the beam spot that automatically created an position dependent intensity array. The intensity profile can be fitted to a Gaussian, which is shown examplewise in figure 2.6, which is the horizontal beam profile in the front and the back of the stack. By measuring the full width half maximum (FWHM) of the beam profile, we were able to build confidence in the size of the beam spot, which was ca. 1 cm in diameter. In the assumption that the beam was underfilled, it was important to build confidence in that we are calculating correctly. The FWHM over SS1 was 1.2017 cm horizontally ($\sigma^2 = 0.2604 \text{ cm}^2$) and 1.1420 cm vertically ($\sigma^2 = 0.2352 \text{ cm}^2$). The FWHM over SS2 was 0.6706 cm horizontally ($\sigma^2 = 0.0811 \text{ cm}^2$) and 0.5783 cm vertically ($\sigma^2 = 0.0603 \text{ cm}^2$).



(a) The target stack in target holder (b) Target holder placed in the end of beam tube

Figure 2.4: Figure shows the target stack and how it was placed in the beam tube.

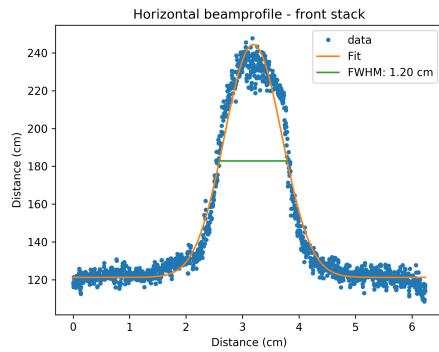


Figure 2.5: The gafchromic film on the activated SS2 foil.

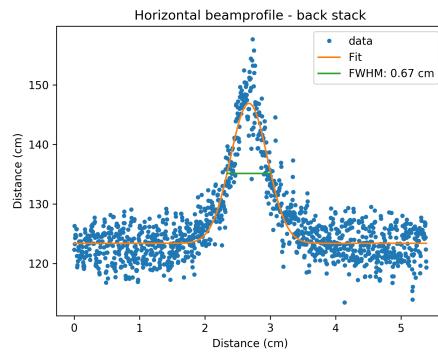
Normally the beam broadens throughout the stack due to scattering. As we can see, this is not the case, since the beam is stopped in the target stack, and therefore we do not know how much the beam truly scatters. This gives a higher uncertainty. The stainless steel (which consists of ..) has fast decay time. However since it emits beta-particles, the radius will slightly increase, and the true beam spot is slightly smaller. Thus the estimated FWHM values for SS1 seem to be within the criterion for underfilled targets.

2.2.3 Counting on high purity detectors

Seven different detectors were used, six IDM Ortec detectors (detectors 1-6) with detector diameter 85 mm, detector length 30 mm and hole depth 15 mm, and one Germanium detector (detector 7) with detector diameter 64.9 mm, detector length 57.8 mm and hole depth 48.6 mm [from detector diagrams](#). Besides, IDM detectors were located in cave 4c (see figure 2.1), which have previously been used as radiation chamber. Thus, background radiation was present. For detector 7, there is lead shielding around the detector. Spectra taken on the Germanium detector is preferred. In order to visualize the



(a) Horizontal intensity profile of SS1



(b) Horizontal intensity profile at SS2

Figure 2.6: Figure shows the intensity profile of the deuteron beam in the front and in the back of the stack horizontally.



Figure 2.7: The calibration point sources that were used in the efficiency calibration of the detector. (^{22}Na was excluded because it was difficult to work with.)

Table 2.2: The calibration point sources along with gamma lines used in the calibration of the detectors. * indicates that the value has been averaged over two peaks with similar energy, less than 1 keV. For the intensity its just added together.

^{137}Cs		^{133}Ba		^{152}Eu	
E_γ	I_γ	E_γ	I_γ	E_γ	I_γ
32.005*	5.63*	53.1622	2.14	121.7817	28.53
36.3405*	1.02*	80.9979	32.9	244.6979	7.55
661.657	85.10	160.6120	0.638	295.9387	0.440
		223.2368	0.453	344.2785	26.5
		276.3989	7.16	367.7891	0.859
		302.8508	18.34	411.1165	2.237
		356.0129	62.05	244.4853*	3.125*
		383.8485	8.94	503.467	0.1524
				586.2648	0.455
				678.623	0.473
				688.670	0.856
				719.353*	0.345*
				778.9045	12.93
				810.451	0.317
				867.380	4.23
				963.712*	14.65*
				1112.076	13.67
				1212.948	1.415
				1299.142	1.633
				1408.013	20.87

signal from the detector, Maestro (Multichannel Analyzer Emulation Software⁴) was used.

The detectors were calibrated using ^{137}Cs ($t_{1/2} = 30.08$ years[?]), ^{133}Ba ($t_{1/2} = 10.551$ years[?]) and ^{152}Eu ($t_{1/2} = 13.517$ years[?]) point sources, using the gammalines listed in table 2.2. The calibration was done at different distinct distances from the detector surface. The point sources can be seen on figure 2.7. The calibration was done in FitzPeakz which is described in section 3.1.1.

The foils were counted for ca. four weeks following end of beam, with short counts in the beginning to have good statistical data for the short-lived activities, and longer and longer counts as the shorter and medium-lived activities decayed out, to have good statistics (enough counts).

⁴<https://www.ortec-online.com/products/application-software/maestro-mca>

Chapter 3

Analysis

3.1 Gamma-ray spectroscopy

The spectra were analyzed in FitzPeakz¹. The mathematic algorithm in which FitzPeakz is based upon is SAMPO80². The peaks are assumed Gaussian with an exponential tail on both sides of the peak. The exponential tail and Gaussian function are joined so function and first derivative are continuous. The algorithm searches for peaks by using the smooth second difference (derivative?) Particularly good for detecting small peaks on a high or low background. The peak areas are calculated by fitting the precalibrated modified Gaussian to the data with a weighted least squares formula using a parabolic background. Fitting intervals are determined automatically by the program. Peaks separated by less than 4 times the average fwhm are fitted together.

For each spectra, a report file containing peak energy, centre channel, full width half maximum, significance, goodness of fit, peak area, uncertainty in peak area, gammas per second, uncertainty in gammas per second and a background estimation for each peak was provided. The most important parameters were the energy, the peak area N_C and uncertainty in peak area. Peak area was needed for the activity calculation in equation 1.30 which is an important parameter in the calculation of the cross section (equation ??), and in the calculation of the efficiency for the calibration sources (equation 3.1). Gammas per second (also called countrate) was used to get an indication if the rate of gammas, which were used as a critical tool to evaluate for instance background peaks, or if the activity was too high in comparison to other spectra measured at other detectors.

Figure 3.1 shows an example of a gamma ray spectrum for one of the iridium spectra (Ir05) approximately 35 hours after end of beam. Figure 3.2 shows the X-ray region and gamma region of ^{193m}Pt .

3.1.1 Energy and peak-shape calibration

Used calibration sources ^{137}Cs ($t_{1/2} = 30.08$ years[?]), ^{133}Ba ($t_{1/2} = 10.551$ years[?]) and ^{152}Eu ($t_{1/2} = 13.517$ years[?]), which can be seen on figure 2.7.

The calculated peak locations and areas are finally corrected with energy and efficiency calibration data to yield peak energies and intensities. For the energy calibration, linear interpolation on a linear scale and for the efficiency calibration linear interpolation on a log-log scale are used in this code. Calibration errors are added to the peak location and intensity errors to give the final result (p.94). The peak shape calibration uses 7 parameters; two background peaks, peak height and location, peak width, distance from peak centroid to the starting point of exponential on either side. The minimization of the least-squares expression to solve for the peak parameters is done by a subroutine package with an iterative gradient algorithm utilizing the variable metric method. Minimization is terminated when all components in the next step change by less than 10^{-8} , if four succeeding values of χ^2 are the same or if 100 iterations have been completed. The performed shape calibration can be checked with a few parameters, goodness of fit, χ^2 per degree of freedom, sigma and error correlation. Sigma

¹<https://www.jimfitz.co.uk/fitzpeak.htm>

²[https://sci-hub.tw/https://doi.org/10.1016/0029-554X\(81\)90209-3](https://sci-hub.tw/https://doi.org/10.1016/0029-554X(81)90209-3)

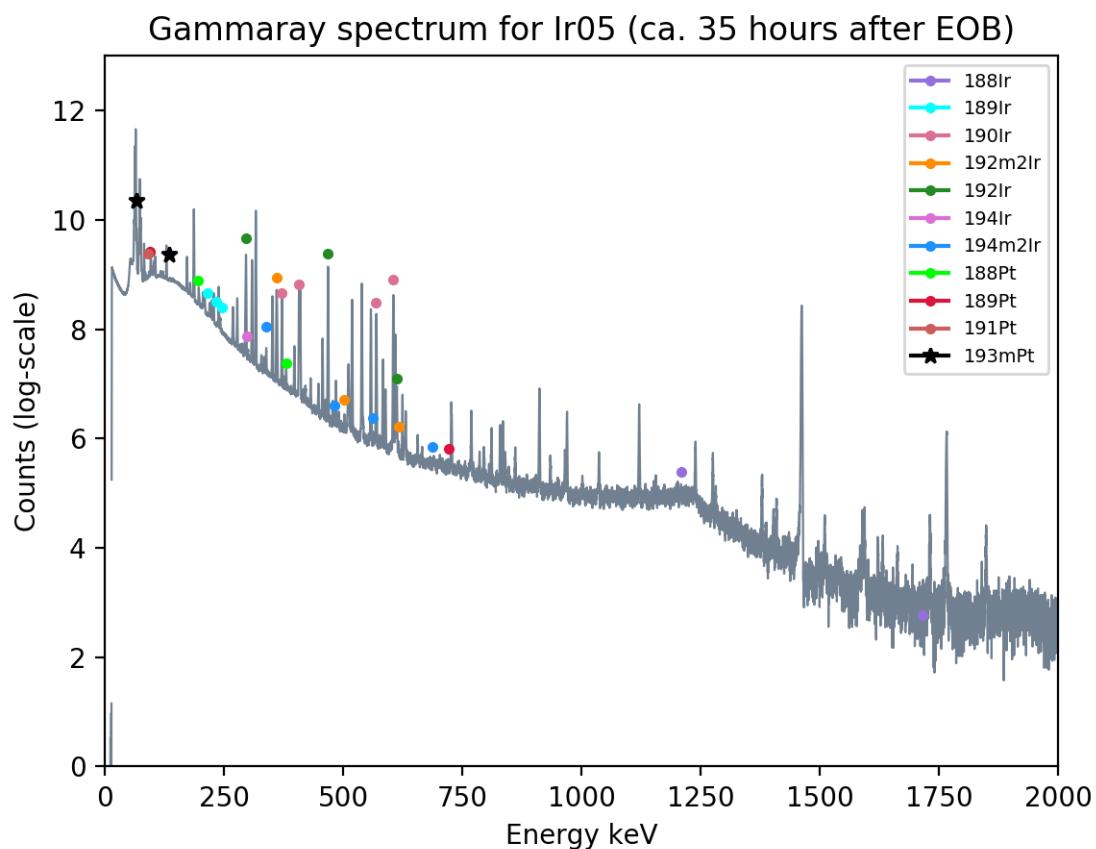


Figure 3.1: Gammaray spectrum for Ir05 taken approximately 35 hours after end of beam. Nuclei does not necessarily represent what is in the peak, but where the peak would have been. Hard to include all since there are different decay times.

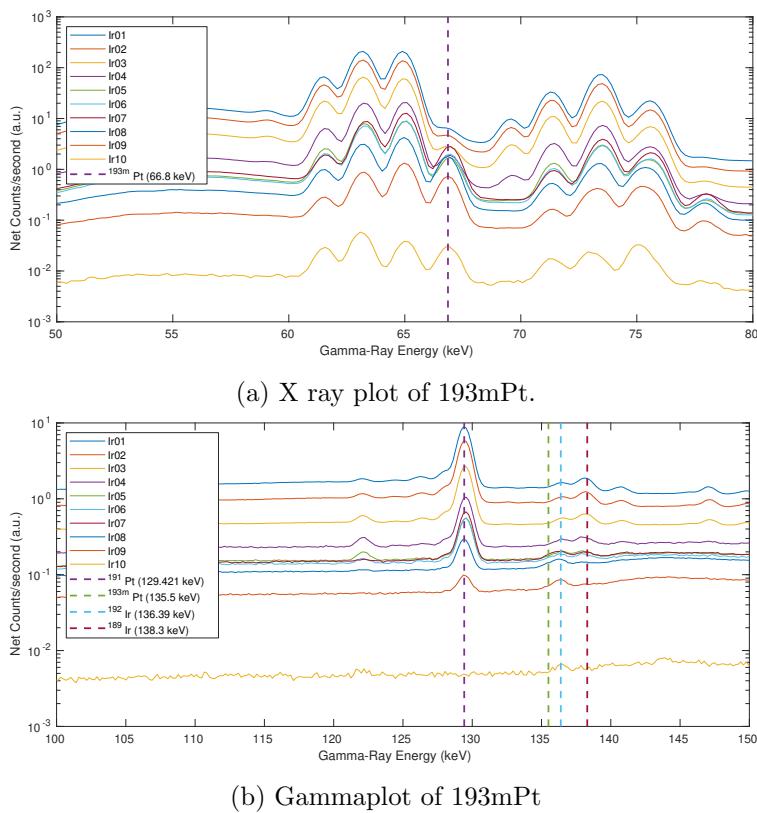


Figure 3.2: Which spectra are these??

below 5 and error correlation between -1 and 1 are acceptable values. (p.90)

From the webpage jim-fitzy.com/calib.html: Each detector was calibrated with peak shape and energy for the calibration sources. Fitzpeakz takes in energy (.enc) and peak shape (.shp) calibration source files, containing the energies listed in table2.2. For the peak shape, the program determines the parameters of width and the amount of low energy tailing. The energy calibration and peak shape calibration was estimated to a 1st order function.

3.2 Efficiency calibration

The efficiency calibration is an important factor in the calculation of the cross section in equation ???. The detector efficiency is the number of events registered divided by the events emitted by the source. The absolute efficiency can be divided into intrinsic and geometrical efficiency, where the intrinsic efficiency is the number of events registered divided by the number of events hitting the detector. The intrinsic efficiency thus depends on the interaction cross section between incident particle and detector material. For neutral particles, the size of the detector affects the intrinsic efficiency, the larger crystal the larger the probability of interaction is. The geometrical efficiency is the radiation emitted by the source which hits the detector. (Techniques for Nuclear and Particle Physics Experiments. William R. Leo. Second Revised Edition. Springer Verlag Berlin Heidelberg GmbH, New York (1994). p. 121-122)

The efficiency was measured using calibration point sources ^{137}Cs ($t_{1/2} = 30.08$ years[?]), ^{133}Ba ($t_{1/2} = 10.551$ years[?]) and ^{152}Eu ($t_{1/2} = 13.517$ years[?]). Figure 2.7 shows the calibration points sources (^{22}Na was excluded during the data-analysis since it only contains a single gamma-line and gave poorer results). On each calibration source, a reference date is given with an activity, which here is referenced to as A_0 of the calibration sources.

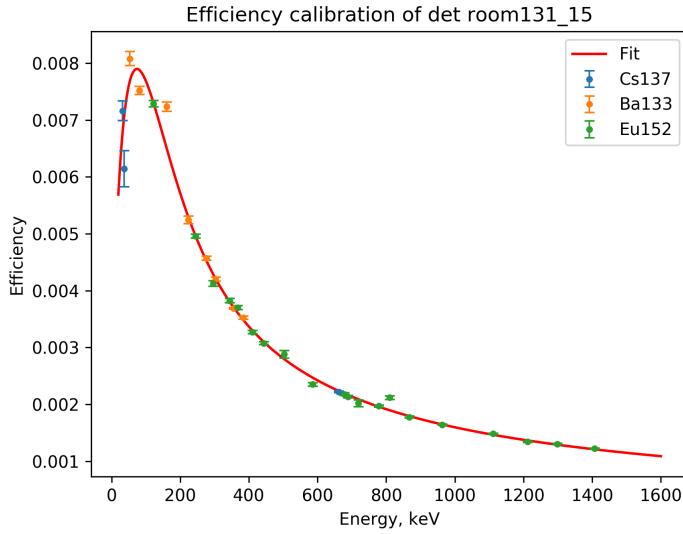


Figure 3.3: An example of an efficiency curve with exact points calculated from equation 3.1 and a curve fit from equation 3.2.

Solving Equation 1.30 for efficiency, ϵ , the analytical efficiency as a function of gamma-ray energy and intensity is

$$\epsilon(E_\gamma) = \frac{N_C \lambda}{A_0 I_\gamma (1 - e^{-\lambda \Delta t_c}) e^{-\lambda \Delta t_d}} \quad (3.1)$$

where λ is the decay constant and N_C is the number of counts in the measured spectra, and Δt_d is the delay time since the reference date. The analytical efficiency gives one single value for the efficiency at energy E_γ , but we want a continuous function which gives the efficiency at any gamma-energy. A model based upon Gallagher, W. J., Cipolla, S.J. (1974) was applied which takes the probability of penetration through the deadlayer of the detector and the probability of interaction in the detector volume into account

$$\epsilon(E_\gamma) = B_0 + \underbrace{(e^{-B_1 E_\gamma^{B_2}})}_{\text{dead layer}} \underbrace{(1 - e^{-B_3 E_\gamma^{B_4}})}_{\text{interacting with volume}} \quad (3.2)$$

where B_i is optimum parameters minimizing the χ^2 in the scipy optimizing curve fit function³). Figure 3.3 shows an example of an efficiency curve for a detector at a specific distance from the detector. The uncertainty of the efficiency was estimated using equation A.6 numerically. For each source, the gamma-lines with the intensities which were used to calculate the efficiency points for each source is listed in table 2.2.

3.3 End of beam activities

The end of beam activities were estimated by extrapolating backwards in time, with the measured activities at various timepoints after the end of beam. The activities as a function of time sine EOB was calculated using equation 1.29, along with a self-attenuation correction:

$$A(\Delta t_d) = \frac{N_C \lambda}{\epsilon_\gamma (1 - e^{-\lambda \Delta t_d}) e^{-\mu \rho \Delta r / 2}} \quad (3.3)$$

where μ is the attenuation coefficient⁴, and $\rho \Delta r$ is the mass density of the foil. The gammas which were used are listed in tables B.2, B.1, B.3 and B.4 for iron, nickel, copper and iridium respectively.

³https://docs.scipy.org/doc/scipy/reference/generated/scipy.optimize.curve_fit.html

⁴<https://www.nist.gov/pml/xcom-photon-cross-sections-database>

The self-attenuation correction is based on the assumption that all activity that is made is located midway in the foil thicknesses. In reality however, the activity profile will follow the same shape as the excitation function over the energy range that expands over the foil (since activity and cross section are proportional). We do not know the excitation function ahead of time, and the excitation function does not change much either, since the foil thicknesses are so thin. So instead, this simplification is done, assuming that the average attenuation is through half of the foil thickness.

The equation describing the shape of the decay curve is given in equation 1.18 for single decay or 1.19 for multiple decay. Decay chains of single and two-step decay ($n=1,2$) was sufficient in this analysis;

$$A = A_0 e^{-\lambda \Delta t_d}, \quad \text{single step decay} \quad (3.4)$$

and

$$A_2(t) = \lambda_n \left[A_{1,0} \lambda_1 \frac{(e^{-\lambda_1} + e^{-\lambda_2})}{\lambda_1 - \lambda_2} + A_{2,0} e^{-\lambda_2 t} \right], \quad \text{two step decay} \quad (3.5)$$

where subnumber 1 is the parent nucleus, and subnumber 2 is the daughter nucleus. Parent activity is calculated from single step decay. The uncertainty was treated as **covarianced variables?**

The way in which the extrapolation was done was the scipy optimize curve fit function, where the A_0 of the daughter was the optimizing parameter. Since there is only one optimized parameter, there was no covariance and the uncertainty was calculated using equation A.13. In the cases where neither parent or daughter activity were known, which were the case for the monitor reaction ^{58}Co with ^{58m}Co decaying into the ground state by internal conversion, both parent and daughter activity were optimizing parameters which are very correlated and thus the uncertainty in end of beam activity was calculated A.6. Figure 3.4 shows two examples of the two different activity curves; one step decay for ^{193m}Pt ($t_{1/2}=4.33$ days) and two step decay for the monitor product ^{58}Co ($t_{1/2}=70.86$ days) with feeding from the isomer ^{58m}Co ($t_{1/2}=9.10$ hours).

3.4 Energy and Beam current

For the equation for cross section (equation 1.45), the beam current $\Phi(E)$ must be known. The beam integrator measured 128.5 nA, which is the current entering the stack. However, due to large energy degradation in the energy stack, there will be a certain spread of the beam, following scattering. In addition, there have not been an experiment with deuterium on a target stack before, so we also needed to see how much deuteron break up affected the current throughout the stack. Monitor reactions are reactions with well-known cross sections⁵. The IAEA recommended cross sections for $^{nat}Fe(d,x)^{56}Co$, $^{nat}Ni(d,x)^{61}Cu$, $^{nat}Ni(d,x)^{56,58}Co$ and $^{nat}Cu(d,x)^{62,63,65}Zn$ (**write about Q value, half life**) were used to estimate a more sensitive deuteron beam current throughout the stack. By solving 1.45 for beam current, the beamcurrent throughout the stack can be estimated

$$\Phi(E) = \frac{A_0}{N_T(1 - e^{-\lambda \Delta t_{irr}})\sigma(E)_{mon}} \quad (3.6)$$

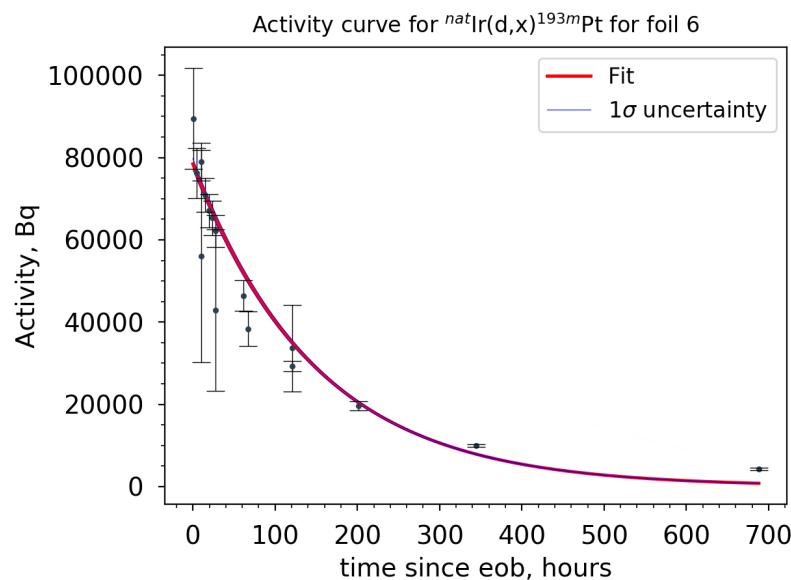
In cross section experiments using thin targets⁶, the suggested value is a flux average cross section, which implies that the cross section is dependent on the flux-weighted average beam energy. One single foil thus provides one cross section measurement, with the uncertainty in energy only being dependent on the energy distribution in each foil. For thin targets, a stopping power $dE/dx = 0$ is assumed which is a very good approximation for targets which are less than 50 mg/cm² **cite?**.

Normalized differential beam current (**need some help understanding this in detail**)

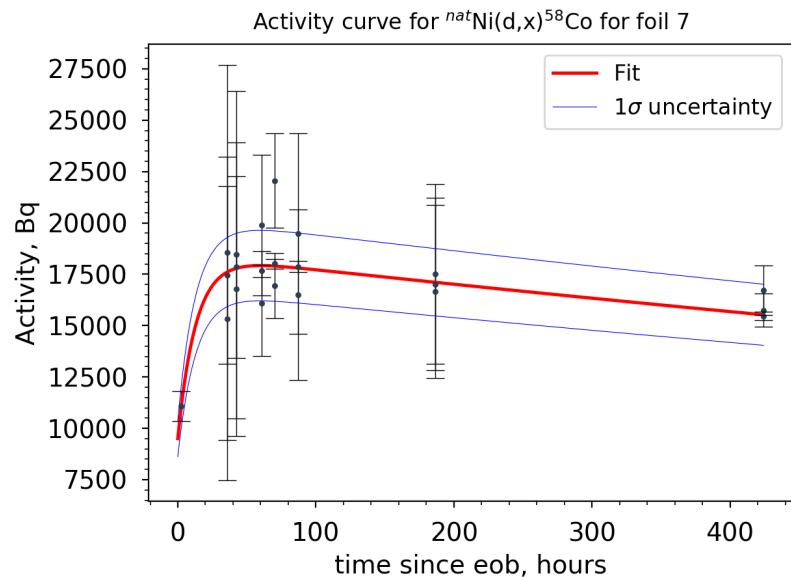
$$\sigma(E)_{mon} = \frac{\int \sigma(E) \frac{d\phi}{dE} dE}{\int \frac{d\phi}{dE} dE} \quad (3.7)$$

⁵https://www-nds.iaea.org/medical/monitor_reaction_article.pdf

⁶Special curriculum p. 14



(a) Activity of ^{193m}Pt ($t_{1/2}=4.33$ d) produced from iridium. The end of beam activity was estimated using a one step decay (equation 3.4)



(b) Activity of ^{58}Co ($t_{1/2}=70.86$ d) produced from nickel. The end of beam activity is estimated using a two step decay (equation 3.5. The feeding is from ^{58m}Co ($t_{1/2}=9.10$ h.)

Figure 3.4: Two examples of activity curves.

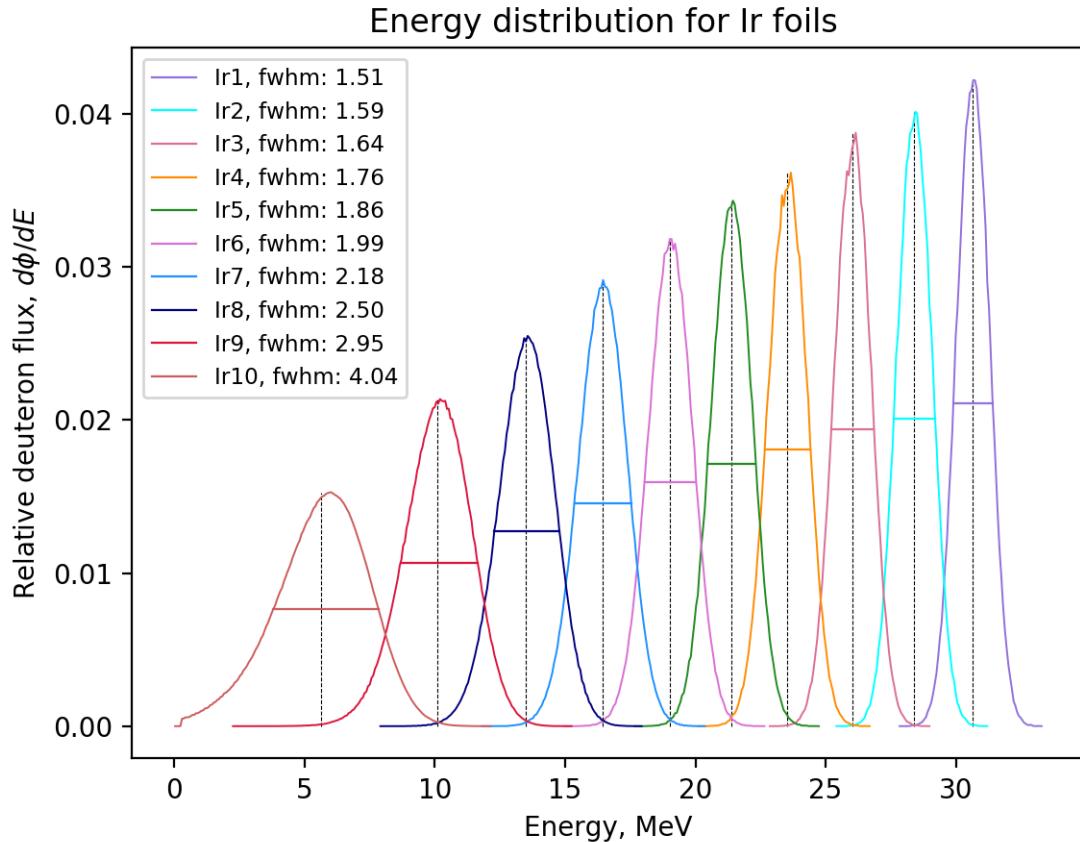


Figure 3.5: Iridium energy flux distribution over the 10 foils. As the energy degrades, skewed and larger full width half max. The vertical line in each peak is the mean value. This indicates that at lower energies, the right uncertainty is greater than the left uncertainty in the peak.

From niobium paper: Since stopping power is inversely proportional to beam energy, the low energy tail of the energy distribution is degraded more in each stack element than the high energy tail. This effect compounds towards the rear of the stack, creating a significantly broadened low energy tail, and a progressively larger net shift of the centroid to a lower energy. To account for this increasing energy uncertainty, a suitably representative energy must be established for each foil of the target stack. The flux weighted average deuteron energy in each foil $\langle E \rangle$ represents the energy centroid for deuterons in a target stack using the energy distributions from $d\phi/dE$ from the Ziegler stopping power deuterion transport

$$\langle E \rangle = \frac{\int E \frac{d\phi}{dE} dE}{\int \frac{d\phi}{dE} dE} \quad (3.8)$$

where $\sigma(E)$ is the IAEA recommended cross section, $\frac{d\phi}{dE}$ is the energy dependent deuteron flux through each foil. The deuteron flux (or energy degradation) was estimated using a code called NPAT's (nuclear physics analysis tool) Ziegler simulation⁷. NPAT uses the Anderson & Ziegler formalism for calculating charged-particle stopping powers in matter in a stack with targets. **write a few sentences about this in the theory!**. The code thus simulates the deuteron flux as a function of energy in the beam stack (assigned to a foil).

Figure 3.5 shows an overview of the flux-energy distribution in each foil for Iridium. The other monitor foils have the same functions. **Write about recommended data, the spline function etc etc..**

need some info from andrew here

⁷<https://pypi.org/project/npat/>

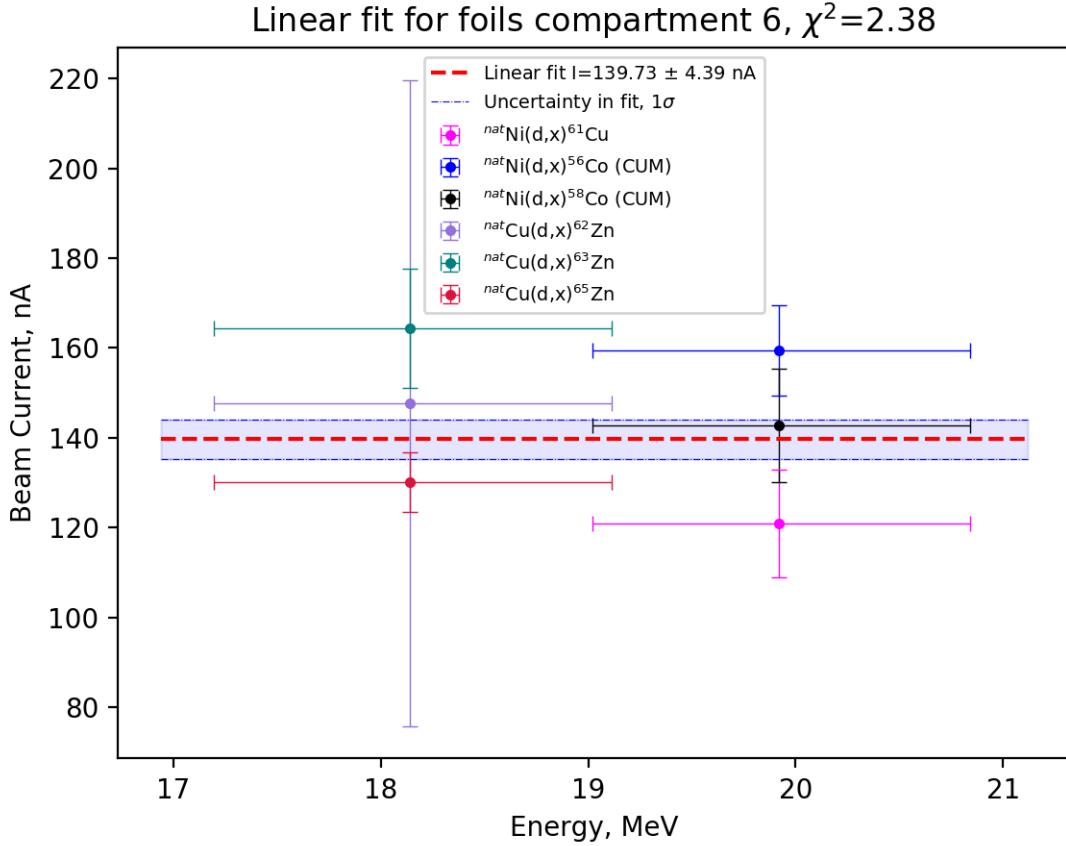


Figure 3.6: The estimated (uncertainty weighted) beamcurrent over compartment 6.

3.4.1 Variance minimization

From Niobium paper p. 57: In theory, the beam current should be more or less constant through the stack, even though the deuterons lose energy. Variance minimization is a technique to reduce the uncertainty in the deuteron beam energies. Non-consistent values for the beam current further back in the stack can be wrong energy bin assignments in the modeled energy distribution (ziegler), or a systematic error in the areal density, which gets progressively worse further back in the stack, "due to the compound effect of systematic uncertainties in stack areal densities". The areal density and the beam energy was varied with 20% increase and decrease, and the reduced χ^2 (equation A.4) was estimated over compartment 3, 6 and 9. For compartment 3 ($E_d \simeq 25$ MeV) all seven monitor reactions were above threshold, thus 6 degrees of freedom. However, early in the target stack, the scattering was low, and the χ^2 does not tell how well the energy bin assignment work further back in the stack. For compartment 6 ($E_d \simeq 18$ MeV), all the six possible monitor reactions (from nickel and copper) were above threshold, and it gave a good estimate of how the beam current was developing throughout the stack. In compartment 9 ($E_d \simeq 10$ MeV), five monitor reactions are above threshold (except for ^{62}Zn). At the very end it is possible to see the full effect of the scattering.

The beamcurrent loss is assumed zero in one compartment, so a linear beam current fit would have a slope equal to zero. The estimated beam current in each compartment was estimated using the scipy curve fit function, with a straight line as model. Figure 3.6 shows the uncertainty weighted linear fit over compartment 6.

Figure 3.7 shows the beam current before and after variance minimization, and the weighted average beam currents are listed in table 3.1 estimated before and after the variance minimization. After variance minimization, the beam current estimated in each compartment (stabled lines) were similar, and meanwhile the weighted χ^2 was about the same in compartment 6, it has improved in

Table 3.1: The weighted average beam current before and after variance minimization in each compartment. The beam current on the 88-Inch Cyclotron beam integrator was 128.5 nA.

Compartment	Before	After
01	131.56 \pm 3.64	134.08 \pm 3.70
02	132.23 \pm 3.74	136.42 \pm 3.83
03	133.81 \pm 3.64	138.02 \pm 3.75
04	134.89 \pm 4.21	138.88 \pm 4.31
05	136.85 \pm 4.21	139.67 \pm 4.29
06	137.40 \pm 4.53	138.85 \pm 4.58
07	139.55 \pm 4.37	139.77 \pm 4.37
08	133.60 \pm 4.27	134.96 \pm 4.32
09	133.16 \pm 5.04	143.59 \pm 5.67
10	108.49 \pm 5.80	121.75 \pm 6.65

compartment 3 and very visible in compartment 9. In general the points are also more aligned.

For cross section calculations, equation ?? is used, with the estimated weighted average beam current. Figure 3.8 shows the estimated cross sections for the monitor reactions, using the weighted average beam current over all monitor foils. The recommended monitor cross section data for the monitor reactions are also plotted, which was used in the cross section calculation.

3.5 Cross section measurements

With all variables for cross sections, cross sections can be calculated using equation ???. The value that is provided as cross section is a flux-averaged cross section, which means that the cross section is dependent on the flux-weighted average beam energy. An accurate measure of the cross section as a function of deuteron energy was possible, as the thin foils provides smaller average beam energy intervals, and it makes it possible to have more measurements if thick foils are replaced with several thinner (one single foil represents a single measurement). Thin foils also produce minimal amounts of radioactivity, thus the deadtime of the detector and the dose to humans is low. For charged particles, the stopping power is inversely proportional to their energy⁸, and therefor the energy degradation in thicker foils will be large. For thin targets, we can however assume that the stopping power $dE/dx \simeq 0$, and the cross section can be replaced with a differential (normalized) cross section

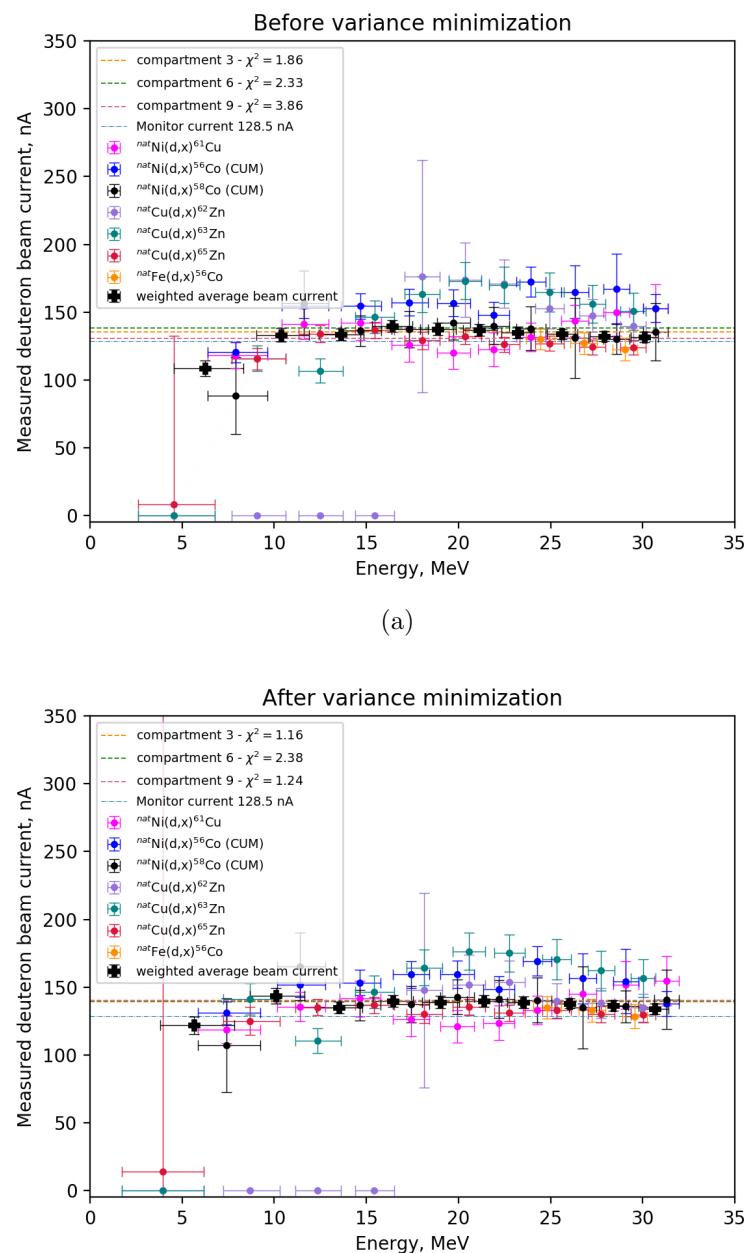
$$\sigma(E) = \frac{\int \sigma(E) \frac{d\phi}{dE} dE}{\int \frac{d\phi}{dE} dE} \quad (3.9)$$

The energy-limits in the integral can be minimized and the error in the energy for charged particles will be small.

Thin foils decreases the energy width, making a more precise measurement dependent on energy. However the reported cross sections are flux-averaged over the energy distribution subtended by each foil.

The cross sections are reported as independent if there is nothing decaying into it (beta feeding or from isomer transition), which means that the first observed element in a decay chain is reported as cumulative unless it is the first possible element (which are the nuclei with one more proton more than the target nuclei, which for this experiment is Platinum (from Ir), Zink (from Cu), Copper (from Ni) and Cobalt (from Fe)). If there is feeding, and the half life is much shorter or much longer than the specific nuclei, can choose appropriate timewindow and report as independent, when we know that the feeding has either died out or is very low!

⁸Andrew S. Voyles, Lee A. Bernstein, Eva R. Birnbaum, Jonathan W. Engle, Stephen A. Graves, Toshihiko Kawano, Amanda M. Lewis, and Francois M. Nortier. Excitation functions for (p,x) reactions of niobium in the energy range of $E_p = 40\text{--}90$ MeV. Nuclear Instruments and Methods in Physics Research Section B: Beam Interactions with Materials and Atoms, 429:53–74, aug 2018.



(b) A 2% increase in beam current and a 4.25% increase in areal density gave the overall most consistent beam current, with reasonable values for the weighted .

Figure 3.7: Beam current before and after variance minimization.

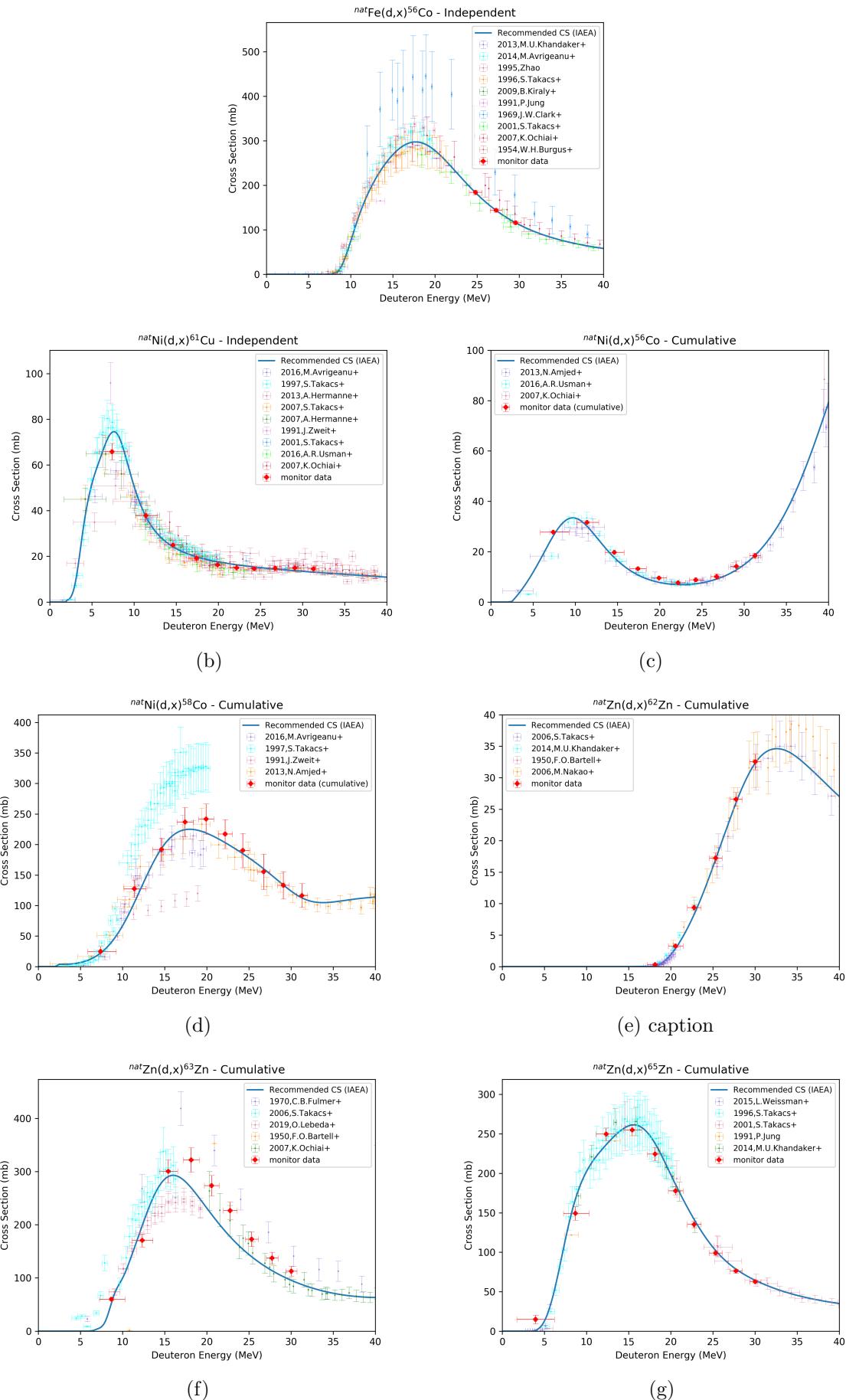


Figure 3.8: Figure shows the estimation of monitor cross section using the estimated weighted average beam current. It is compared along with the recommended (IAEA) monitor data, and other experimental data

The measured cross sections in this work was compared to previous experimental data, along with reaction modelling codes TALYSh^s<https://sci-hub.tw/https://doi.org/10.1016/j.nds.2012.11.002>, TENDL, ALICE20 and CoH.

Optical model calculations performed first,

Talys takes in projectile, target element (specific isotope or all stable target isotopes), energy array with desired spacing and upper limit energy.

For COH: To get both 191Ir and 193Ir to run, we had to adjust the parameter "tweakSD", which adjusts the effective single-particle state density for a particular particle emission channel. In the end, we ran with tweakSD=0.25 for both outgoing alphas and neutrons (protons were unaffected). In other words, we set the single-particle state density for outgoing alphas and neutrons [(d,xa) and (d,xn) reactions] to be 25% of what they normally are, which is a HUGE change.

from talys cite p. 2843-2844: TENDL is developed from talys (TALYS evaluated Nuclear data Library). This library consists of a complete set of nuclear reaction data for incident neutrons, photons, protons, deuterons, tritons, Helium-3 and alpha particles, from 10^5 eV up to 200 MeV, for all 2430 isotopes from ^6Li to ^{281}Ds that are either stable or have a half-life longer than 1 second. All data are completely and consistently evaluated using a software system consisting of the TALYS nuclear reaction code, and other software to handle resonance data, experimental data, data from existing evaluations, and to provide the final ENDF-6 formatting, including covariance information. The result is a nuclear data library with mutually consistent reaction information for all isotopes and a quality that increases with yearly updates. To produce this library, TALYS input parameters are adjusted for many nuclides so that calculated cross sections agree with experimental data, while for important nuclides experimental or evaluated data are directly included. Also feedback from integral measurements is processed into the data libraries. For nuclides for which (almost) no experimental data exists, default TALYS calculations based on global models and parameters are used.

Dont understand this part.....

Chapter 4

Results

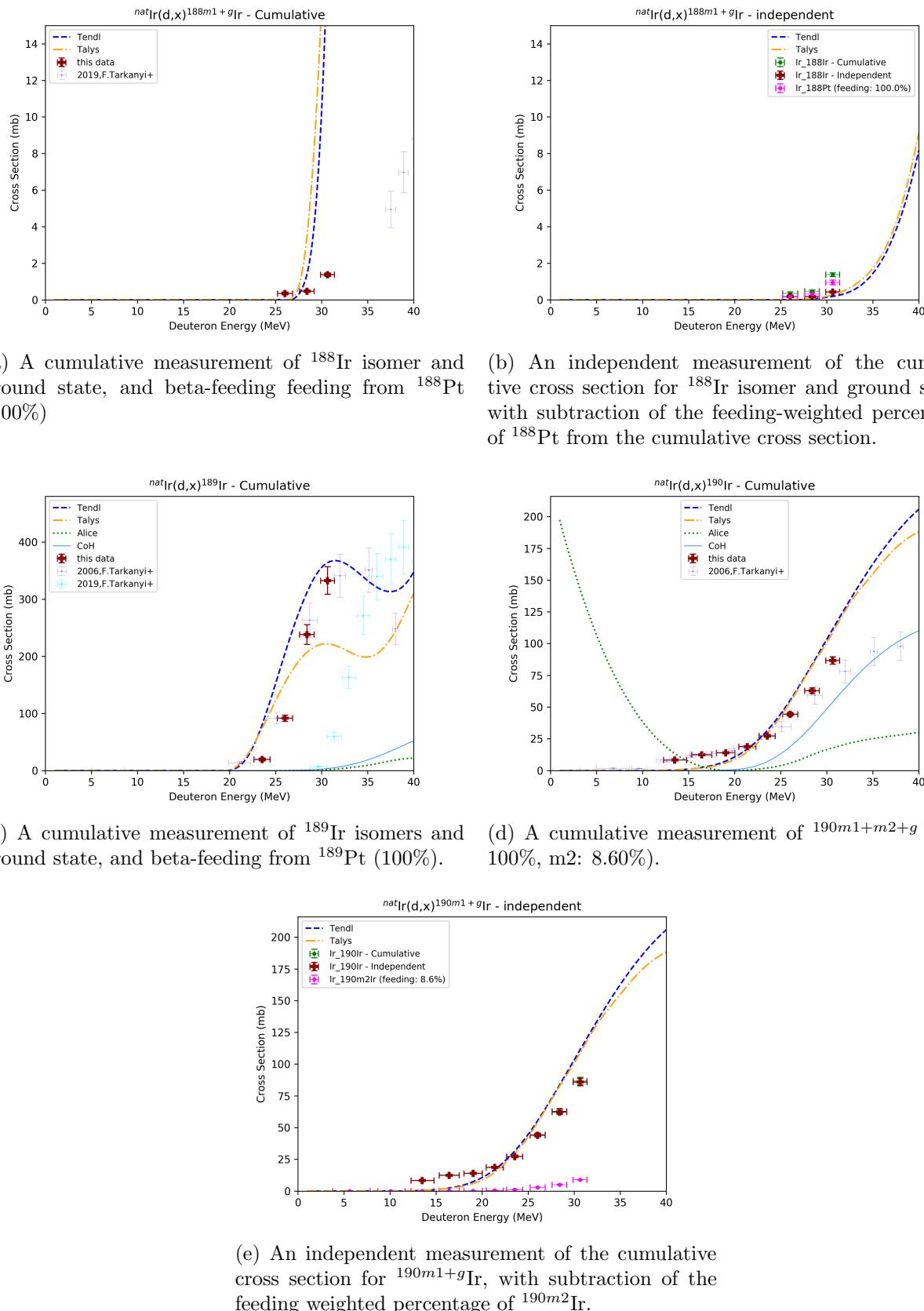


Figure 4.1: Excitation functions for $^{188m1+g}\text{Ir}$ (independent and cumulative) and ^{189}Ir (cumulative)

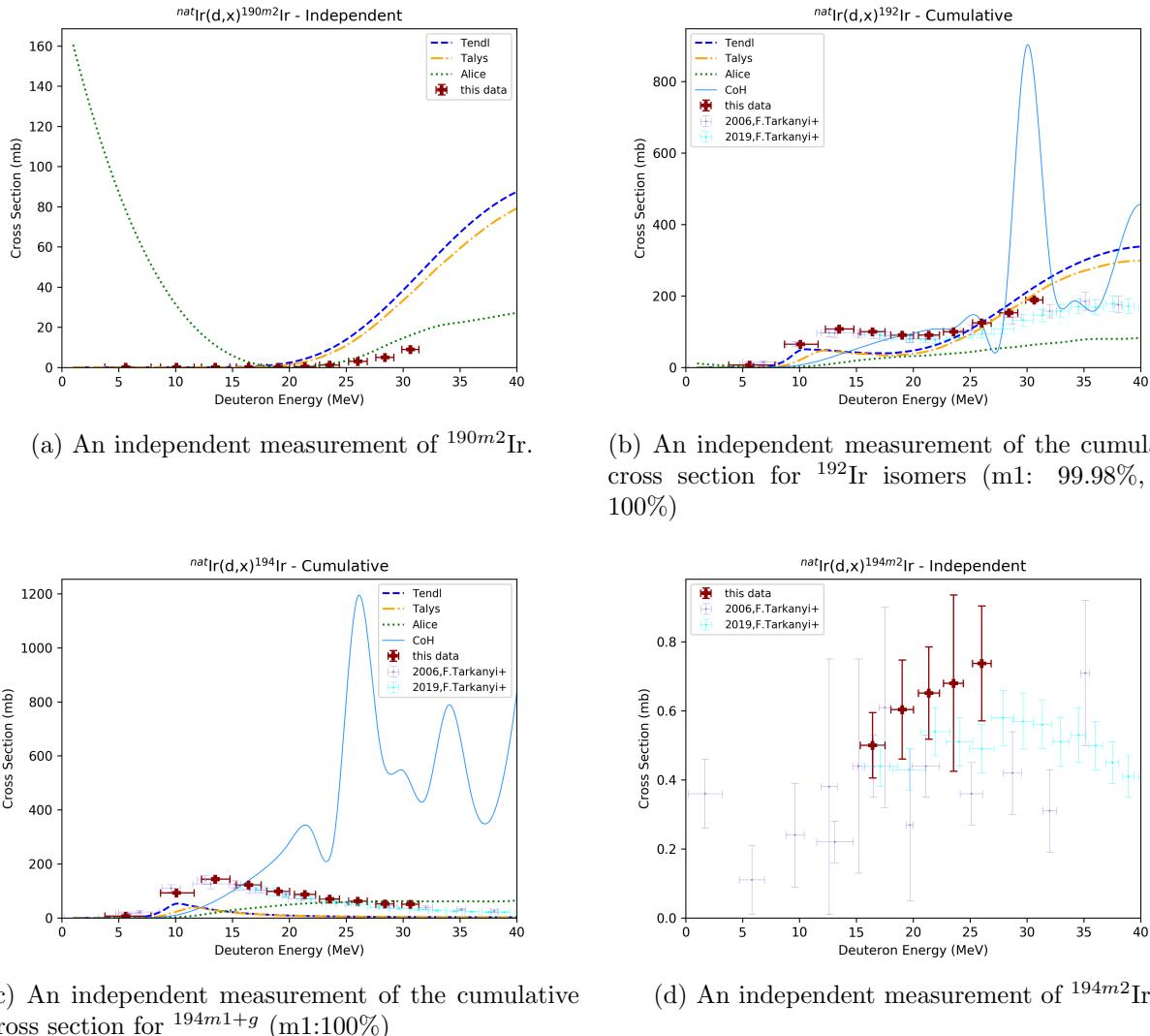
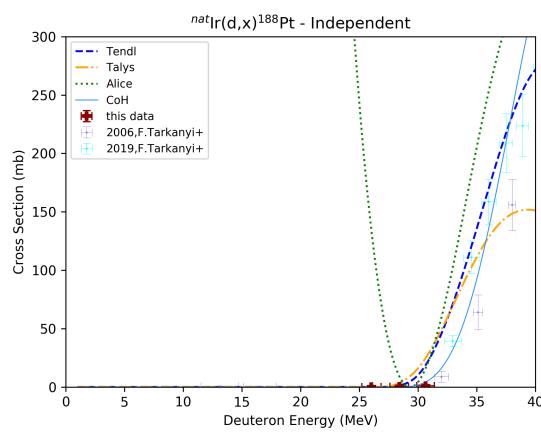
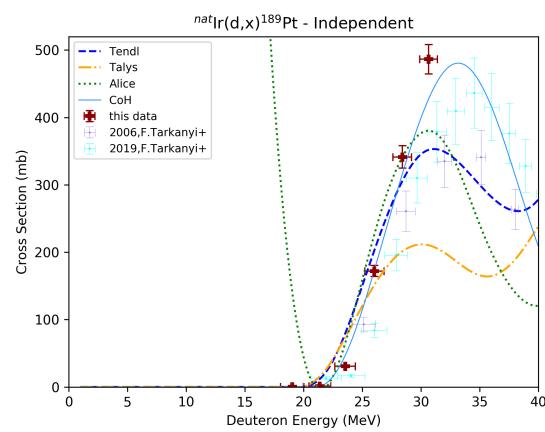


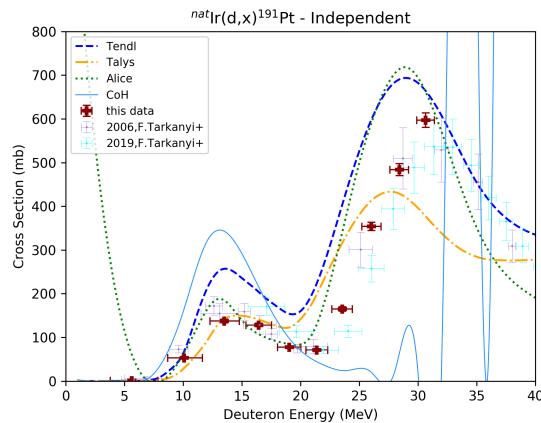
Figure 4.2: Excitation functions for $^{190m^1+m^2+g}\text{Ir}$ (cumulative), $^{190m^1+g}\text{Ir}$ (cumulative), $^{190m^2}\text{Ir}$ (independent), $^{194m^1+g}\text{Ir}$ (cumulative) and $^{194m^2}\text{Ir}$ (independent)



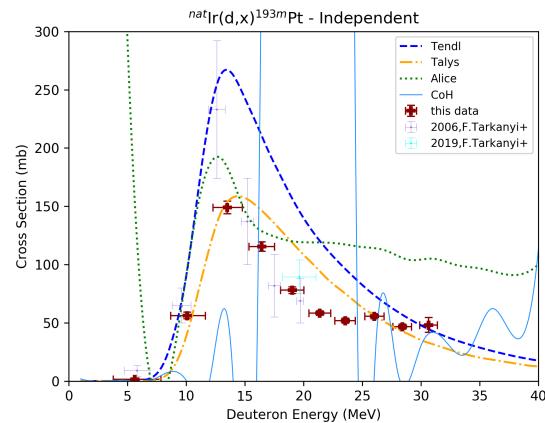
(a) An independent measurement of ^{188}Pt .



(b) An independent measurement of ^{189}Pt .



(c) An independent measurement of ^{191}Pt .



(d) An independent measurement of ^{193m}Pt .

Figure 4.3: Excitation functions for Platinum radionuclides

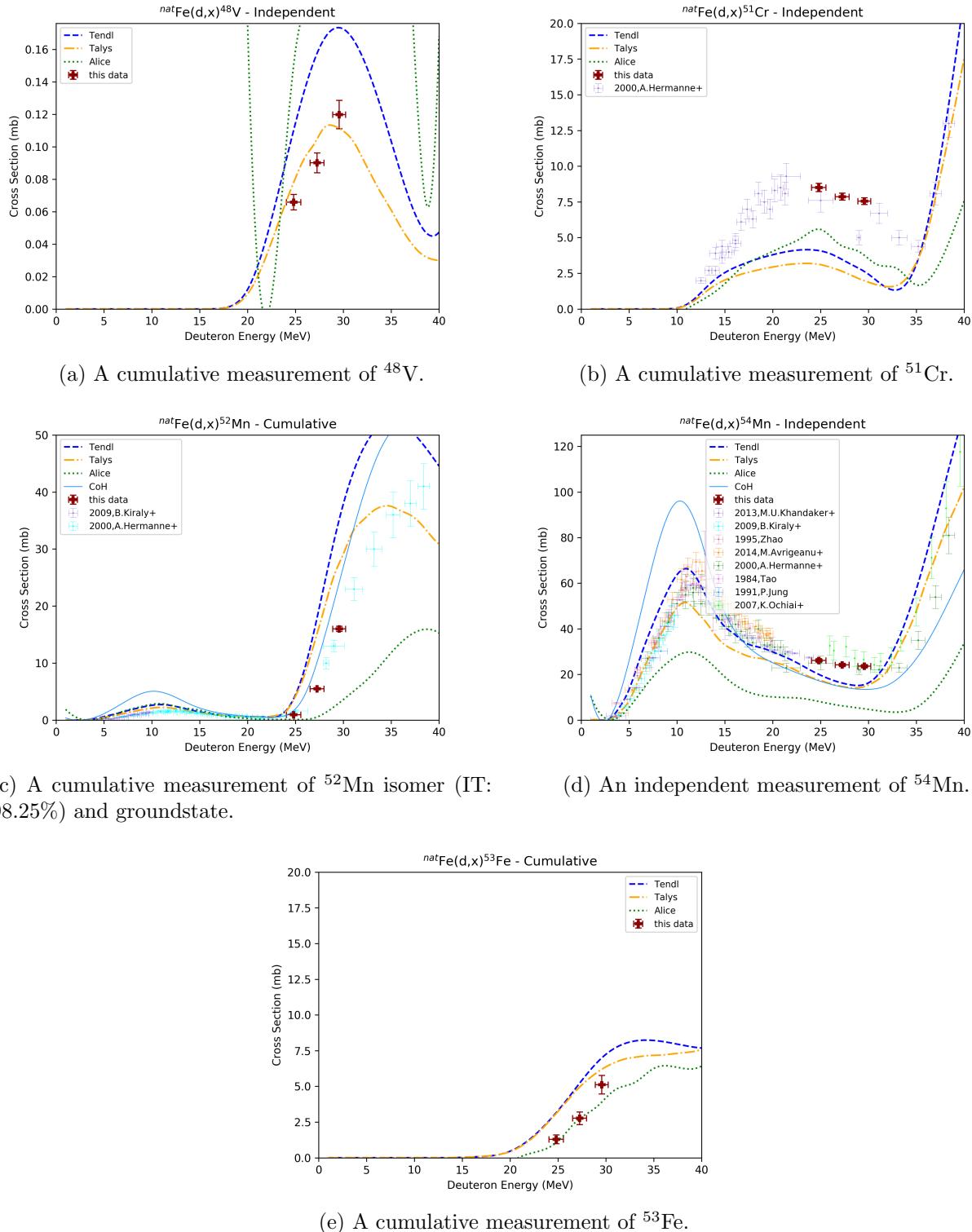


Figure 4.4: Excitation functions for ^{48}V , ^{51}Cr , ^{52}Mn , ^{54}Mn , ^{53}Fe produced from iron.

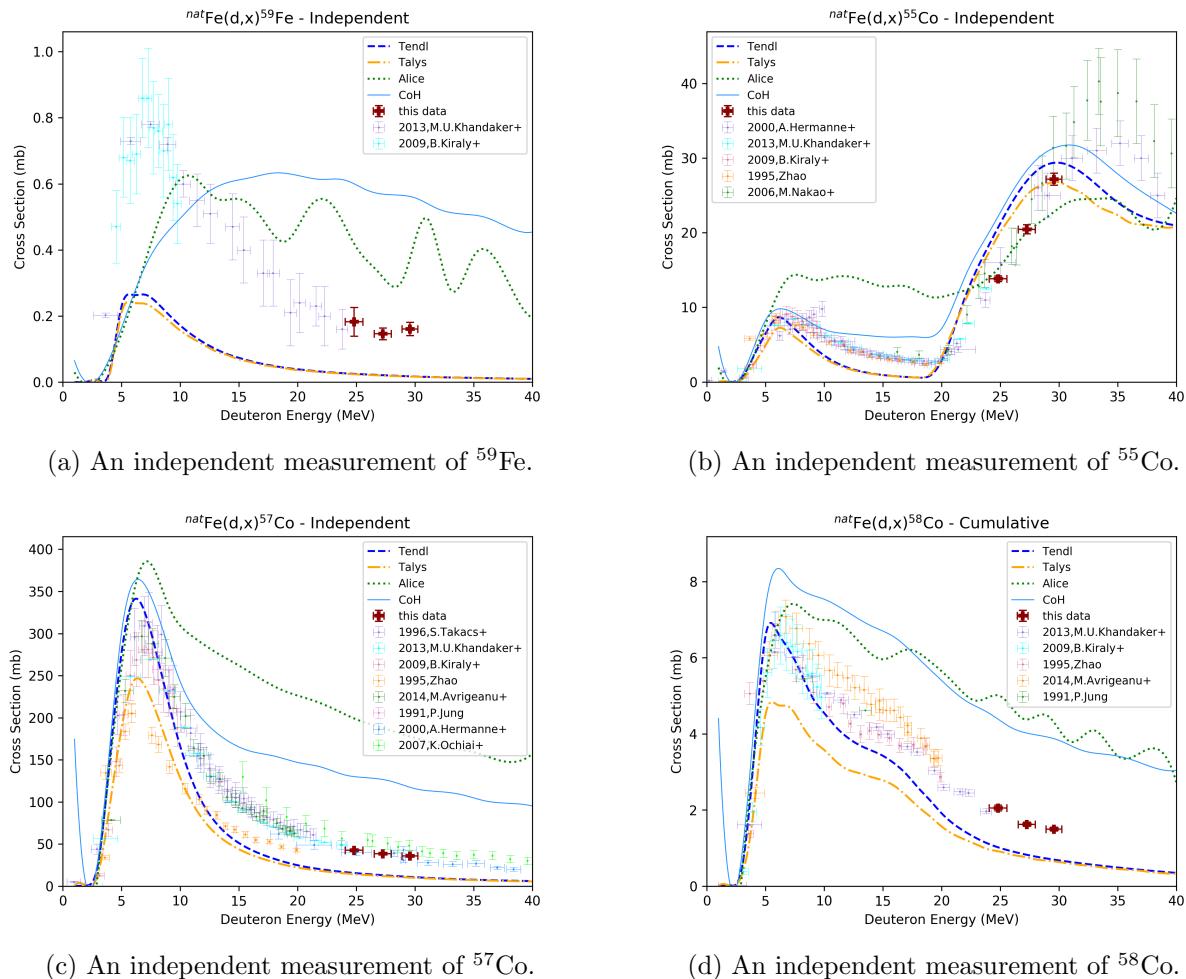


Figure 4.5: Excitation functions for ^{59}Fe , ^{55}Co , ^{57}Co , ^{58}Co produced from iron.

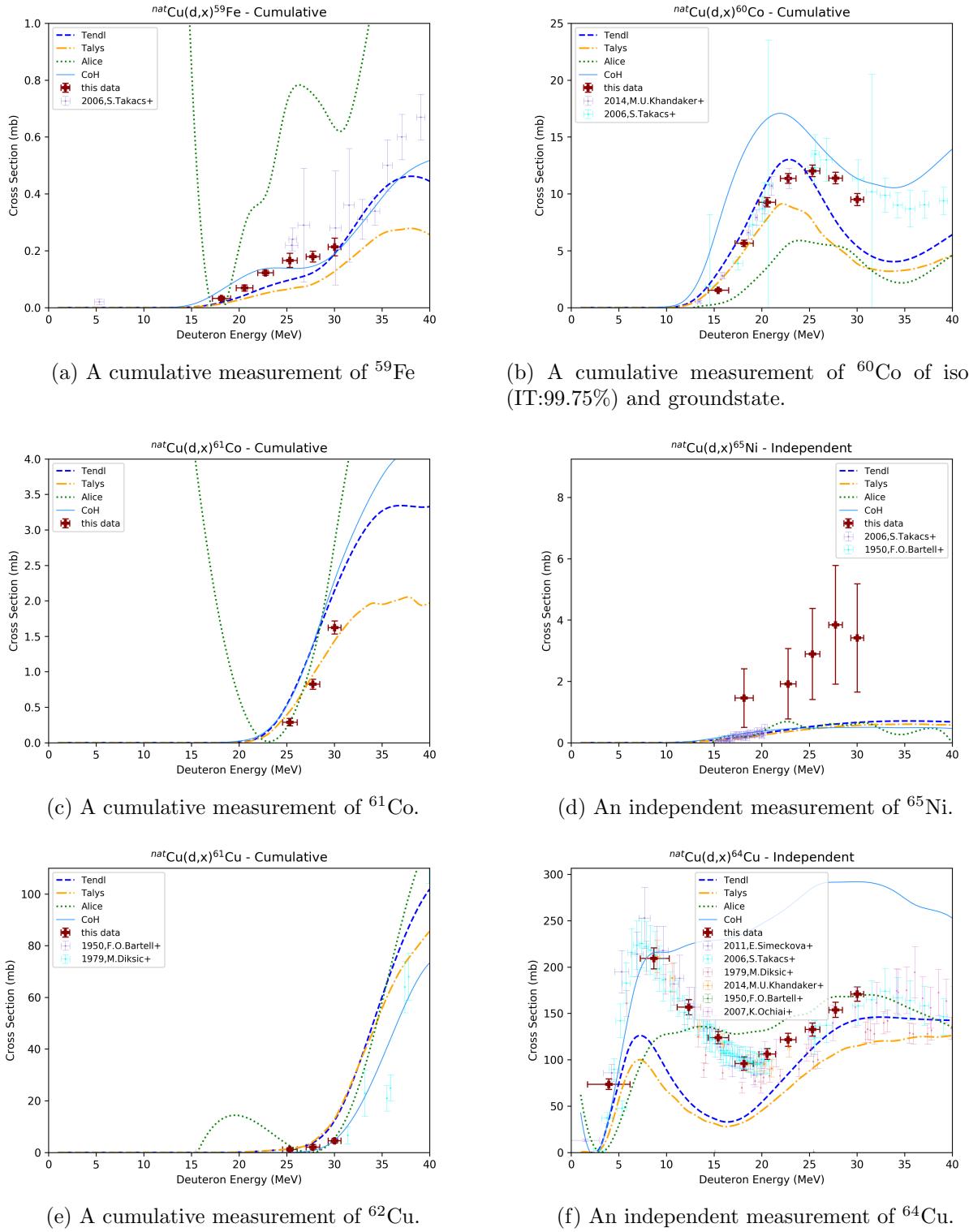


Figure 4.6: Excitation functions for ${}^{59}\text{Fe}$, ${}^{55}\text{Co}$, ${}^{57}\text{Co}$, ${}^{58}\text{Co}$ produced from iron.

Chapter 5

Discussion

5.1 Beam current

The variance minimization (figure 3.7) led to an overall more constant weighted average beam current after the variance minimization, where $\text{Ni}(\text{d},\text{x})^{56}\text{Co}$ and $\text{Cu}(\text{d},\text{x})^{62}\text{Zn}$ in particular have contributed. The other reactions have about the same values before and after variance minimization. With the exception of compartment 9, where the beam current is higher than before the variance minimization, the beam currents are more consistent (table 3.1). It is hard to tell whether it was a point to do this, but figure 3.7 clearly show an improvement of ^{56}Co and ^{62}Zn .

The estimated production cross sections for the monitor reactions, using the weighted average beam current (figure 3.8) in general shows good compliance with the recommended IAEA cross sections that were used in the calculation of the beam current. In general, the measured cross section values are within uncertainty of the recommended cross section. The high energy measurements are in general more precise, and it is clear that where it is a rapid change in the excitation function, the energy bins seems to be slightly wrong. However, the measured cross sections are within uncertainty in both energy and cross section, with the exception being the high energy exponential tail for ^{63}Zn .

5.2 Cross section products

In general, we got most of the results that we expected to see, based upon energy threshold, decay channel of compound nucleus which would increase the cross section and based upon the Coulomb barrier.

Generally, there is no evidence of decay by triton, deuteron or ^3He , although they are energetically possible. This is due to the higher spin, so that decay channels with lower spin is favoured instead. Alpha-decay happens for Iron, Nickel and Copper, but as we can see in table.... the Coulomb barrier (equation ...) is a lot higher. Thus we do not see products lower than one p down (platinum and iridium). For Iron ($Z=26$) Nickel ($Z=28$) and Copper ($Z=29$), the Coulomb barrier is almost half the

In general, can see on all cross sections that the energy uncertainty increases with decreasing beam energy.

5.2.1 Nickel products

In this work, cross sections for $^{52m+g}\text{Mn}$,
proton-barrier: ca 7 MeV, α -barrier: ca. 12 MeV

In general, only the lowest energy decay channel is listed. might be more opening.

^{52}Mn (cumulative)

A cumulative cross section for ^{52m}Mn ($t_{1/2}=21.1$ m) and ^{52g}Mn ($t_{1/2}=5.591$ d) is published in this work, along with independent measurements of both isomer and ground state. ^{52m}Mn decay by ϵ/β^+

(98.25%) to ^{52}Cr , and by isomeric transition (1.75%) to the ground state. ^{52g}Mn decay by ϵ/β^+ (100%) to ^{52}Cr . ^{52}Mn can be produced via $^{58}\text{Ni}(\text{d},2\alpha)$ with a reaction Q-value -1235.6 keV and energy threshold 1278.5 keV, or via $^{58}\text{Ni}(\text{d},2\text{n}2\text{p}\alpha)$ with reaction Q-value -29531.2 keV and energy threshold 30557.9 keV. Production via $^{60}\text{Ni}(\text{d},2\text{n}2\alpha)$ with a reaction Q-value -21622.6 keV and energy threshold 22349.2 keV and $^{61}\text{Ni}(\text{d},3\text{n}\alpha)$ is also possible with a reaction Q-value -29442.7 keV and energy threshold 30415.9 keV is also possible. Due to the Coulomb barrier, the energy demanded to overcome this barrier is ca 12 MeV, which is clear on the figure, as the excitation function starts to increase around where the alpha-channel opens. Since the energy threshold is so low for the $(\text{d},2\alpha)$ reaction, it is not a problem to emit to alpha-particles. The excitation function steadily increases as multiple reaction channels opens. The data from this work is in well compliance with other experimental data by A. Hermanne et. al. (2013), N. Amjad et. al. (2013) and A.r. Usman et. al. (2016), although the measured experimental point for ca. 29 and 31 MeV are slightly above the previous experimental data. Tendl and Talys suggests that the peak is slightly shifted towards the lower energy region with a maximum at ca. 26 MeV, while the experimental data suggests maximum at ca. 29 MeV. Alice follows the experimental data shape-wise, and is closer magnitude wise, but oscillates.

The cumulative cross section was estimated just using single step decay with using the gammas listed in the table, since first of all the branching from ^{52m}Mn is weak, along with short half life. The 1434.. keV gamma-line is the strongest and present in both isomer and ground state, however, the groundstate has multiple independent gamma-lines. After removing 1434 keV, a new end of beam activity was measured and a independent cross section of the ground state was made. Then, the independent measurement of the isomer was made by subtracting the cumulative cross section of the both isomer and ground state and the independent cross section of the ground state multiplied by the branching ratio.

^{54}Mn (cumulative)

^{54}Mn ($t_{1/2}=312.20$ d) is a radionuclide, which decay by ϵ/β^+ (100%) to stable ^{54}Cr . The cross section reported is an independent measurement. This radionuclide can be produced from $^{58}\text{Ni}(\text{d},2\text{p}\alpha)$ with a reaction Q-value -8538.3 keV and an energy threshold 8835.1 keV, frp, $^{60}\text{Ni}(\text{d},2\alpha)$ with a reaction Q-value -629.6 keV and energy threshold 650.8 keV, or via $^{60}\text{Ni}(\text{d},2\text{n}2\text{p}\alpha)$ with a reaction Q-value -28925.3 keV and energy threshold 29897.3 keV, from $^{61}\text{Ni}(\text{d},\text{n}2\alpha)$ with reaction Q-value -8449.7 keV and energy threshold 8729.0 keV, from $^{62}\text{Ni}(\text{d},2\text{n}2\alpha)$ with reaction Q-value -19045.4 keV and energy threshold 19664.8 keV. From the figure, we can clearly see that even though the reaction channel $(\text{d},2\alpha)$ with Q-value -629.6 keV, along with $(\text{d},2\text{p}\alpha)$ with Q-value -8538.3 keV and $(\text{d},\text{n}2\alpha)$ with Q-value -8449.7, and does not open before ca. 11 MeV, as the alpha-barrier is overcome, and the measured points after that are in the same cross section magnitude. When the reaction route via $(\text{d},2\text{n}2\alpha)$ with Q-value -19045.4 keV and $(\text{d},2\text{n}2\text{p}\alpha)$ with Q-value -28925.3, the cross section increases, and reaches local maximum at ca. 35 MeV. The measured cross section points in this work is in well compliance with the measured datapoints from A. Hermanne et. al. (2013), A.R. Usman et. al. (2016), N. Amjad et. al. (2013) and S. Takacs et. al. (2007).

5.2.2 Iridium Products

α -barrier: ca 25 MeV, proton-barrier: ca 13.7 MeV

For a 33 MeV deuteron beam on the target stack, cross sections for ^{188}Ir , ^{189}Ir , $^{190m1+g}\text{Ir}$, $^{190m2}\text{Ir}$, ^{192}Ir , $^{194m1+g}\text{Ir}$, $^{194m2}\text{Ir}$, ^{188}Pt , ^{189}Pt , ^{191}Pt and ^{193m}Pt was measured, which described below. There was however no evidence that anything which would emit more than one proton in the decay of the compound nucleus, like Os, Re, W or Ta was produced. Identifying the Iridium products was a difficult task. Firstly, as the Coulomb-barrier equation... for these reactions is higher than for Ni, Cu and Fe as target nuclei, since the number of protons is higher. Thus we known that reactions like (d,α) , or (d,xp) would not be heavily fed. However, most of the energetic thresholds are less than 33 MeV, so even though they are weakly fed, they will still be fed if the channel is possible. This we did not see any evidence of. Secondly, many of the nuclei have gammas that are so close in energy that

the germanium detectors would identify the peak as one. As the reaction routes producing Ir- and Pt-radionuclides is heavily favoured, the peaks with feeding from multiple nuclei would still mostly be caused by those nuclides. The question is then, was anything else than Ir and Pt radionuclides produced, or were they produced but could not be distinguished because of the peaks. As a summary, a few of the nuclides were excluded, but a few of the nuclides is questionable. It can also be mentioned that the compound nuclei ^{193}Pt and ^{195}Pt are nuclei with even Z and odd N, thus proton emission would be more likely, as discussed in section

Production routes via α emission is energetically possible. $^{191}\text{Ir}(\text{d},\alpha)^{189}\text{Os}$ (stable) with reaction Q-value=12988.8 keV could not be observed since it is a stable product, and $^{193}(\text{d},\alpha)^{191}\text{Os}$ (ground state $t_{1/2}=15.4$ d, isomer $t_{1/2}=13.10$ h) with reaction Q-value 12569.8 keV, was not observed. ^{191}Os was expected to be produced, as the cumulative cross sections within the deuteron energy region have been observed cite tarkanyi 2019. However measurements done in that work was small cross sections. The main hypothesis for ^{191}Os not to be observed is that the only gamma-line of the ground state which was intense enough to be observed was 129.431 keV (26.50%) which was too close to ^{191}Pt ($t_{1/2}=2.83$ d) in energy. Thus the way to measure the ^{191}Os cross section was to count long enough for ^{191}Pt to decay completely, which would be for ten half lives which wold be approximately 1500 hours. The isomer also has one single gamma-line at 74.38 keV (0.0729%), which was not observed due to low intensity, and multiple nuclei with similar gamma-line. $^{191}\text{Ir}(\text{d},2\alpha)^{185}\text{W}$ ($t_{1/2}=75.1$ d) with reaction Q-value 14964.9 keV was not observed, nor was $^{193}\text{Ir}(\text{d},2\alpha)^{187}\text{W}$ ($t_{1/2}=24.0$ h) with reaction Q-value 13653.6 keV. Thus it was concluded that W and below in mass number was produced at all.

Another reaction route which was expected to be observed was $^{193}\text{Ir}(\text{d},2\text{p})^{193}\text{Os}$ ($t_{1/2}=30.11$ h) with reaction Q-value -2584.0 keV. Gamma-lines matched well with the spectra, and unique gamma-lines 321.616 keV (1.245%) and 387.509 keV(1.226%) (which was not in the background) was used to estimate the end of beam activity, the cross sections did not look reasonable. Re-radionuclides had similar gamma-lines to Pt and Ir-radionuclides, so we could not distinguish. No observation of Re-radionuclides were made, but a question-mark remain on this behalf, especially since ^{190}Re and ^{192}Re is even (odd-odd).

^{188}Ir (Cumulative and independent)

^{188}Ir ($t_{1/2}=41.5$ h) is a radionuclide with one metastable isomer (with a half-life in the order of ms) along with the ground state, which decays by either isomeric transition or ϵ/β^+ , but the branching ratio is not stated. The reported cross sections for ^{188}Ir , regardless of whether it is reported as cumulative or independent is the cumulative cross section of the ground state and isomer. Since an independent measurement of ^{188}Pt ($t_{1/2}=10.2$ d), which feeds into ^{188}Ir (100%) was obtained, the independent measurement without feeding and the cumulative cross section with feeding is reported. The reaction threshold is $^{191}\text{Ir}(\text{d},2\text{nt})$ with a reaction Q-value of -16231.0 keV. In this work, we did not see any evidence of this decay-route being fed. Thus, ^{188}Ir can be produced via $^{191}\text{Ir}(\text{d},4\text{np})$ with a reaction Q-value -24802.0 keV and an energy threshold of 25064.0 keV. From ^{193}Ir as target nuclei, the threshold for 6np-particle emission is above energy threshold, so the only available route from this target nucleus is with emission of 4nt (Q-value is -30291.0 keV), which is not heavily fed either. It is clear that the excitation function first increases once the 4np-decay channel opens, but the measured points are at a low cross section value. The gamma-lines which were used are listed in table ..., The end of beam-activity was estimated using a single decay, where activity points ca. 200 hours after end of beam is slightly higher than the curve, because of the feeding. The end of beam activity however looks reasonable, and the cross section also looks reasonable. Comparing the measured cross section points to experimental data, F. Tarkanyi et. al. (2019), does not have any measured points below ca 35 MeV, but the points predicts that the excitation function increases in this region, where the measured cross sections in this work give reason to believe that the cross section increases from about 25 MeV. The independent measurement aligns well with TALYS and TENDL, but the large measurements 26 MeV is probably a little higher than then true value. reaction models.

^{189}Ir (Cumulative and independent)

This is weird. Expectedly, the cumulative cross section of ^{189}Ir should be larger than the independent measurement of ^{189}Pt , due to ϵ -decay (100%). This is however not the case so the independent measurement of ^{189}Ir is negative. This is also apparent in Tarkanyi et. al. (2019), so why is this? ^{189}Ir ($t_{1/2}=13.2$ d) is a radionuclide with two metastable states with half-lives in order of microseconds. Thus the total cross section in general is the two isomers and the ground state. However, ^{189}Pt ($t_{1/2}=10.87$ h) feeds into this radionuclide via β^+/ϵ (100%), so both an independent measurement of the three states of ^{189}Ir is reported along with the cumulative cross section with feeding from ^{189}Pt . The gamma-lines which were used are listed in table ..., and the end of beam activity was estimated using single-step decay, although ideally this could have been fitted using a two-step decay with feeding from ^{189}Pt , but due to overestimation in the end of beam activity, so it turned out large negative. Instead single-step decay was used where no peaks before 60 hours after end of beam was used, so that the ^{189}Pt had decayed out. ^{189}Ir can be produced via $^{191}\text{Ir}(\text{d},3\text{np})$ with a reaction Q-value -16626.0 keV and energy threshold 16802.0 keV, or via $^{193}\text{Ir}(\text{d},5\text{np})$ with reaction Q-value -30596.0 keV and energy threshold 30916.0 keV. This work provided four cross section measurements from 23-30 MeV, and it is clear that the cross section increases from ca. 20 MeV, comparing to experimental data from F. Tarkanyi et. al. (2006). The data provided by this work is in well compliance with the 2006 data, where the excitation function is at its highest between 30-35 MeV, when the other reaction channel has opened. However, 2019 data is shifted ca. 10 MeV higher in energy. Our data seems to agree with the 2006 version. The 2019 data seems to be agreeing more with when the excitation function starts to increase looking at ALICE, TENLD and TALYS, where at 20-25 MeV, the function is basically 0.

 $^{190m1+g}\text{Ir}$ (Cumulative and independent)

^{190}Ir ($t_{1/2}=11.78$ d) decays by ϵ/β^+ -decay to stable ^{190}Os (100%). This radionuclide can be produced via $^{191}\text{Ir}(\text{d},2\text{np})$ with a reaction Q-value -1769.3 keV and with an energy threshold 10359.2 keV or via $^{193}\text{Ir}(\text{d},4\text{np})$ with a reaction Q-value -24221.2 and with an energy threshold 24474.0 keV. For this radionuclide, a total of eight cross section measurements were made in the energy region 13-30 MeV, and as expected, we did not see evidence that this radionuclide was produced for deuterons below 10 MeV. The excitation function steadily increases from 13-20 MeV, and once the next reaction channel opens, the excitation function increases more steeply. Due to the independent measurement of $^{190m2}\text{Ir}$, an independent measurement of the ground state and the m1-isomer was made, which is given in figure ... The excitation function looks almost identical due to the weak branching from $^{190m2}\text{Ir}$ (8.60%), along with quite low measured cross sections of $^{190m2}\text{Ir}$ (table...). Comment on reaction models.

 $^{190m2}\text{Ir}$ (Independent)

$^{190m2}\text{Ir}$ ($t_{1/2}=3.087$ h) either decay by β^+/ϵ -decay (91.4%) to ^{190}Os or isomer transition (8.6%) to ^{190g}Ir . figure out Q-values and e thresholds

 ^{192g}Ir (Cumulative)

^{192g}Ir ($t_{1/2}=78.829$ d) either β^- decays to stable ^{192}Pt (95.24%) or β^+/ϵ -decays to stable ^{192}Os (4.76%). The radionuclide can be produced via $^{191}\text{Ir}(\text{d},\text{p})$ with a reaction Q-value of 3973.55 keV, or via $^{193}\text{Ir}(\text{d},2\text{np})$ with a reaction Q-value -9996.6 keV, and a energy threshold 10.1009 MeV¹. Thus the reaction channel is energetically possible for 0 MeV deuterons. However, since there are two possible ways to produce this radionuclide, an increasing cross section would appear as the probability of proton emission as decay channel increases. With a proton Coulomb-barrier at approximately 13 MeV for both stable targets², the first peak in the excitation function should be somewhere around there, which we can clearly see on the figure, in the experimental data in this work. Then the cross section decreases up to about 20 MeV, and increases up to about 35 MeV, where the experimental data suggests that the cross section flattens. This measurement is a cumulative cross section with

¹(<https://www.nndc.bnl.gov/qcalc/qcalcr.jsp>)

²(<http://hyperphysics.phy-astr.gsu.edu/hbase/NucEne/coubar.html>)

$^{192m^1}\text{Ir}$ ($t_{1/2}=1.45$ m) feeding in with isomer transition (99.98%), and $^{192m^2}\text{Ir}$ ($t_{1/2}=241$ y) feeding in with isomer transition (100%), where we did not make any independent measurements of either isomers due to too short and too long half life. We were able to produce and measure the cross section in each foil, and the cross section measurements done in this work is in well compliance with experimental data provided by Tarkanyi et. al. (2006, 2019), but the values are slightly higher for energies higher than 13 MeV. The gamma-lines which were used are listed in table ??, and all the gamma-lines are neither in the background or feeding from other gammas. The end of beam activity was estimated using a single-step decay, as there is nothing beta-feeding in, and the isomers are in the cumulative cross section. **The theoretical predictions are off. Make more comments.**

^{194g}Ir (Cumulative)

The reaction Q-value for $^{193}\text{Ir}(\text{d},\text{p})^{194g}\text{Ir}$ is 3842.22 keV, thus the energy threshold for the deuteron is 0 MeV. For this particular measurement, the cross section is reported as cumulative because of the $^{194m^1}\text{Ir}$ -isomer ($E=147.0785$ keV, $t_{1/2}=31.85$ ms) feeds into ^{194g}Ir by isomer transition (100%). ^{194g}Ir decays by β^- to stable ^{194}Pt . The gamma-lines which were used to calculate the activity in the foils were 293.541 (2.5%), 300.741 (0.35%), 589.179 (0.140%), 938.69 (0.60%), 1150.75 (0.60%), 1468.91 (0.19%). Most of the gamma-intensities are less than 1%, as more intense gamma-lines were either in the background or had gamma-lines which were overlapping with other nuclei. Thus the contribution to uncertainties can be of importance here. However, comparing the measured cross sections to other measured cross sections by Tarkanyi et. al. (2006, 2019), the uncertainty in cross section looks small. The end of beam activity was estimated using a single step decay, since the parent isomer $^{194m^2}\text{Ir}$ does not feed (decay by β^- (100%)), and there is no chance that we can measure the $^{194m^1}\text{Ir}$ isomer. The correspondence between the previous experimental data seems to be good. The measured points in this work seems to be a little on the high side on the tail, but yet within uncertainty. TENDL and TALYS does an ok job shaping the peak, but is both lower in magnitude and the maximum of the TENDL function shifted from about 13 to 10 MeV. The maximum of of the TALYS peak is in better compliance with the measured cross sections, and suggests a maximum at around 12 MeV. ALICE however is on **baertur**

$^{194m^2}\text{Ir}$ (Independent)

The reaction Q value for $^{193}\text{Ir}(\text{d},\text{p})^{194m^2}\text{Ir}$ ($t_{1/2}=171$ d) is 3652.22-X keV, due to the uncertain energy level of the isomer. Having a positive reaction Q-value, the deuteron energy threshold is 0 MeV. The cross section represented in this work is independent, as this isomer level is the highest one observed. $^{194m^2}\text{Ir}$ decays by β^- -decay to stable ^{194}Pt . In the estimation of cross section, one single line was used, which was 390.8 keV (35%). The more intense gamma-lines such as 482.6 keV (97%), 328.5 (93%) snf 600.5 (62%) were not used due to weak feeding from other nuclei, or being present in the background spectra. Thus, the total number of measured points are only five in the energy region 16-26 MeV, since stronger gamma-lines were excluded in the activity calculations. The end of beam activity was estimated using a one-step decay chain. The measurement of the reaction cross section for this nucleus is in well compliance with the results published by F. Tarkanyi et. al. (2006, 2019), magnitude wise (a little on the high side, but within uncertainty). There is however difficult to see if there is any clear curve shape of the excitation function, as especially the datapoints from F. Tarkanyi et. al. (2006) seems to have a high variation in the cross section values. The datapoints from F. Tarkanyi et. al. (2019) seems to be more systematic, along with smaller energy bins. The measured datapoints in this work suggests that the cross section increases with beam energy at least up to 26 MeV, which seems to be the general tendency for F. Tarkanyi et. al. (2019), without a few exceptions where the measured cross section decreases. Since this isomer only can be produced in one particular way, the expected shape of the excitation function is one global maximum, which should be shifted towards the low energy side due to the low energy threshold. There is no TALYS, TENDL or ALICE reaction code-data for this specific isomer (per april 2020), so it is difficult to determine the exact shape of the excitation function. With as long half-life as this isomer has, the foil should have been counted over a longer timeperiod after end of beam. Then multiple other gamma-lines could have

been used to measure the end of beam activity, by excluding the timepoints where multiple nuclei feeds the same gamma-energy.

^{188}Pt (Independent)

^{188}Pt ($t_{1/2}=10.2$ d), ϵ/β^+ -decay to ^{188}Ir (100%). $^{191}\text{Ir}(\text{d},5\text{n})$ with a reaction Q-value -26109.0 keV and an energy threshold 26384.0 keV. $^{193}\text{Ir}(\text{d},7\text{n})$ is not possible in this deuteron energy range under 40 MeV. Thus as expected, the measured cross sections are three points in the energy range 26-30 MeV, and the points are in the threshold region, which is in compliance with both talys and tendl. The experimental data done by T. Tarkanyi et. al. (2006, 2019) have not measured any cross sections below 32 MeV, but the measured points supports that the threshold should be close to where we measured the first points, and then the excitation function increases for this specific reaction route. The gamma-lines that were used are listed in table ..., and the end of beam activity was found via single-step decay. Around the threshold, ALICE, TALYS and TENDL agrees with the measured data.

^{189}Pt (Independent)

^{189}Pt ($t_{1/2}=10.8$ h) decays by ϵ/β^+ -decay to ^{189}Ir (100%). This radionuclide can be produced via $^{191}\text{Ir}(\text{d},4\text{n})$ with a reaction Q-value -19389.0 keV and energy threshold 19593.0 keV, or via $^{193}\text{Ir}(\text{d},6\text{n})$ with reaction Q-value -33359.0 keV and energy threshold 33707.0 keV. The measured cross sections suggest a energy threshold around 19 MeV, but the threshold of 19.5 MeV is within uncertainty in energy. Talys and Tendl suggests threshold 19.8 MeV. It seems like there is some disagreement on where the excitation function peak should be in the 30-35 MeV region, where in general, datapoints provided by this measurement suggests a higher cross sections shifted to slightly lower energies than experimental data by F. Tarkanyi et. al. (2006, 2019). It is also hard to say whether the next point would have been higher in cross section, if we could have measured that, but if not, TENDL and TALYS supports this curve shape. The 2006 data also suggests a lower but broader cross section peak, than the 2019 data. From TALYS and TENDL, it seems like the cross section is increasing from 37 MeV, where the 6n decay channel opens. The gamma-lines used are listed in table .., and to estimate the end of beam activity, the decay curve was estimated using a single-step decay.

^{191}Pt (Independent)

^{191}Pt ($t_{1/2}=2.83$ d) decays by ϵ/β^+ -decay to stable ^{191}Ir (100%). This radionuclide can be produced via $^{191}\text{Ir}(\text{d},2\text{n})$ with a reaction Q-value -4017.0 keV and an energy threshold 4060.0 keV, or via $^{193}\text{Ir}(\text{d},4\text{n})$ with a reaction Q-value -17988.0 keV and an energy threshold 18175.0 keV. In this work, ten independent measurements were made for this product nucleus, and the first one is right above energy-threshold at ca. 5.5 MeV, and the two peaks in the cross section matches the different reaction channels which opens at ca 4 MeV and 18 MeV. The measured data in this work matches the experimental data from F. Tarkanyi et. al (2016, 2019). The end of beam activity was estimated using a single step decay. **REMEMBER that when above threshold, all reactions can happen. And even though for instance d,n is not favoured at higher energies, it can still happen, and thus the second peak may look higher.**

^{193m}Pt (Independent)

^{193m}Pt ($t_{1/2}=4.33$ d) decays by isomeric transition to long-lived ^{193}Pt (100%). The isomer can be produced via $^{193}\text{Ir}(\text{d},2\text{n})$ with a reaction Q-value -3063.5 keV and energy threshold 3095.5 keV. Made 10 independent measurements for this reaction. Comparing to experimental data, the result look reasonable. Talys and tendl peak seems to be shifted toward higher energies, while ALICE seems reasonable position wise, suggesting a excitation function maximum at ca. 12 MeV, which matches the experimental data well. However, for higher energies ALICE looks strange. Due to the weak intensity of the gamma-line provided by the IT decay, we had to use X-rays to be sure that we got the right cross section. However due to many X-rays being fed only 66.831 keV was used. Since

66.831 keV is also an X-ray present in ^{192}Ir ($t_{1/2}=73.829$ d) (intensity is 4.44%), this X-ray was used to obtain the ^{193m}Pt activities. Single-step decay was used to obtain end of beam activity.

^{52}Mn (Cumulative)

^{52}Mn ($t_{1/2} = 5.591$ d) (figure ??) is reported as a cumulative cross section, because of the unidentified decay rate of ^{52m}Mn ($t_{1/2} = 21.1$ m). The isomer and ground state share multiple lines, and using the values after the isomer was decayed completely ($\simeq 5\text{ hours} = 10$ half lives after end of beam), the activities were fitted to a single step decay. Hence the cross section is cumulative. From table ??, the Q-value³ in which ^{52}Mn can be produced is -1235.6 keV, so the threshold is well within the energy range in which this work is operating in. As we can see in figure figure, the first observation is approximately at 14 MeV. The Coulomb barrier for the compound nucleus $^{58}\text{Ni}(d)^{58*}\text{Cu}$, the height is 12.7509 MeV, thus the chance of tunneling before that energy is relatively low. Once the reaction channel from ^{60}Ni (with a Q value of -21622.6 keV), the increase in cross section is visible. The gammaslines which were used are strong, and as a consequence, the uncertainties in the cross section is relatively small.

^{52}Mn is an odd-odd nucleus with 25 protons and 27 neutrons. Both proton and a neutron is unpaired, so this decay channel is not necessarily favoured as a d,pn, d,3p3n etc.

^{54}Mn (Independent)

^{54}Mn ($t_{1/2} = 312.20$ d) is an independent measurement of the cross section. There are multiple measurements at low cross section values. The (d,2 α) reaction channel is available at most energies, but the Coulomb barrier is which is at 12.7509 MeV, suggests that the probability of tunneling is relatively low below 10 MeV. The cross section first starts rising at around 18-20 MeV, where multiple reaction channels opens. In general, the measured cross sections done in this thesis matches the earlier experimental data within uncertainties, allthough a little on the high side in the low energy region.

^{59}Fe (Independent)

There is no experimental data in this reaction, however the measured cross section data points agree well with talys. Tendl apparently has a harder time to find a good cross section model. This is not a strongly fed channel, first of all this is an even-odd nucleus with 26 protons and 33 neutrons, leaving one unpaired neutron. If we interpret the cross section correctly, the cross section starts to increase around when overcoming the coulomb barrier, and keeps increasing while other reaction channel opens.

^{55}Co (Independent)

Agrees very well with other experimental data. Z=27, N=28. Cs not super high.

^{60}Co (Independent)

Tendl looks off, however talys follows the experimental data pretty well. As we can see the experimental data is spread in the high energy region, and it is hard to say if actually have made good values. However we are well above energy threshold.

^{56}Ni (Independent)

The only way ^{56}Ni can be produced is via ^{58}Ni . The threshold is

³Calculating the Q value for various reactions: <http://hyperphysics.phy-astr.gsu.edu/hbase/NucEne/coubar.html>

The cross sections for the iridium products are listed in table Give a brief discussion on the various products. And also, give a discussion of what we expected and what we did not see. THIS IS NOT A FINAL VERSION In this experiment, ^{188}Pt , ^{189}Pt , ^{191}Pt , ^{193m}Pt , ^{188}Ir , ^{189}Ir , ^{190}Ir , ^{192}Ir , ^{194}Ir , $^{194m2}\text{Ir}$, ^{188}Re , ^{189}Re and ^{190}Re was observed. The products we did not observe was ^{183}Ta ($t_{1/2}=5.1$ d), ^{186}Re ($t_{1/2}=3.7186$ d), ^{186}Ta ($t_{1/2}=10.5$ m), ^{18}W ($t_{1/2}=24.0$ h).

^{183}Ta has a Q-value equal to 0.0 MeV for $^{191}\text{Ir}(\text{d},\text{d}\alpha)^{183}\text{Ta}$ and 4.020 MeV for $^{193}\text{Ir}(\text{d},\text{nt}2\alpha)^{183}\text{Ta}$. The strongest gamma-line ($E_\gamma = 59.318$ keV, $I_\gamma = 42.1\%$) was not observed in any spectrum, and the second strongest gamma-line ($E_\gamma = 246.059$ keV, $I_\gamma = 27.2\%$) is within 1 keV from a gamma-line belonging to ^{189}Ir ($E_\gamma = 245.1$ keV, $I_\gamma = 6.0\%$). Hence, we conclude that ^{183}Ta was not observed.

^{186}Re has a Q-value equal to 0.0 MeV for $^{191}\text{Ir}(\text{d},\text{t}\alpha)^{186}\text{Re}$ and 13.1264 MeV for $^{193}\text{Ir}(\text{d},\text{2n}t\alpha)^{183}\text{Ta}$. There is no data in the EXFOR database, and the reaction modelled cross sections provided by TALYS predicts that the cross section is **very low or zero** in this energy range. The two strongest gamma-lines were observed, ($E_\gamma = 122.64$ keV, $I_\gamma = 0.603\%$, $E_\gamma = 137.157$ keV, $I_\gamma = 9.47\%$). However, the activity which was estimated was not looking natural, hence the cross section curve is difficult to interpret **Might need some more work**.

Appendix A

Statistics

Uncertainty in statistics refers to the standard deviation of the data, which gives a number of the spreading of the data from the mean value of the data citation. The variance is the standard deviation squared, which weights the variables to a higher degree.

$$std = \sqrt{\sigma^2} \quad (\text{A.1})$$

$$\sigma = \sqrt{\frac{1}{N} \sum_{i=1}^N (x_i - \bar{x})^2} \quad (\text{A.2})$$

where N is the number of measurements, x_i is a measurement and \bar{x} is the average over all measurements.

χ^2 is an estimation of the goodness of the fit, which includes the weight of the error

$$\chi^2 = \sum_i^n \left(\frac{y_i - \bar{y}}{\sigma_i} \right)^2 \quad (\text{A.3})$$

where \bar{y} is the mean value of y and σ_i is the error in y_i . The reduced χ^2 is defined as the χ^2 per degree of freedom

$$\chi_\nu^2 = \frac{\chi^2}{\nu} \quad (\text{A.4})$$

where ν is the degrees of freedom equal to the number of observations minus the number of fitted parameters. A value close to $\chi_\nu^2 = 1$ indicates that the observations and fit is in well accordance to the error, while $\chi_\nu^2 > 1$ indicates an underfitting and a $\chi_\nu^2 < 1$ indicates an overfitting¹.

A function f with input x and a set of variables $\vec{\beta} = \beta_1, \beta_2, \dots, \beta_n$ and output y can be written on the following form

$$y = f(x, \vec{\beta}) \quad (\text{A.5})$$

The uncertainty in y is dependent on the uncertainty in the different input variables $\vec{\beta}$. The matrix expression for error propagation is (Tellinghuisen, Joel, Statistical error propagation)²

$$\sigma_y^2 = \mathbf{J} \cdot \mathbf{V} \cdot \mathbf{J}^T \quad (\text{A.6})$$

where σ_y^2 is the variance in y, J is the Jacobian matrix

$$\mathbf{J} = \begin{bmatrix} \frac{\partial f}{\partial \beta_1} & \frac{\partial f}{\partial \beta_2} & \dots & \frac{\partial f}{\partial \beta_n} \end{bmatrix} \quad (\text{A.7})$$

and V is the variance-covariance matrix

¹https://en.wikipedia.org/wiki/Reduced_chi-squared_statistic

²A full derivation of the expression can be found in Uncertainty Propagation for Measurements with multiple output quantities, Dobbert, Schrijver

$$\mathbf{V} = \begin{bmatrix} \sigma_0^2 & \sigma_{0,1} & \cdots & \sigma_{0,n} \\ \sigma_{1,0} & \sigma_1^2 & \cdots & \sigma_{1,n} \\ \vdots & \vdots & \ddots & \vdots \\ \sigma_{n,0} & \sigma_{n,1} & \cdots & \sigma_n^2 \end{bmatrix} \quad (\text{A.8})$$

In the cases where the input parameters are uncorrelated, all non-diagonal elements in the variance-covariance matrix is equal to zero, and the expression for the variance is simplified to

$$\sigma_y^2 = \sum_{i=1}^n \left(\frac{\partial f}{\partial \beta_i} \right)^2 \sigma_{\beta_i}^2 \quad (\text{A.9})$$

Whenever the input parameters are correlated, which means that $\sigma_{\beta_i, \beta_j} \neq 0$, we have to apply equation A.6, otherwise, the simplification in equation A.9 will give wrong error propagation.

To evaluate the partial derivatives of f , the computational derivation is applicable

$$\frac{\partial f}{\partial \beta_i} \approx \frac{f(x, \beta_i + \frac{\Delta \beta_i}{2}) - f(x, \beta_i - \frac{\Delta \beta_i}{2})}{\Delta \beta_i} \quad (\text{A.10})$$

where $\Delta \beta_i$ is a small number, like $10^{-8} \beta_i$.

For a function $f = xy$, the variance can be expressed from equation A.6, where

$$\mathbf{J} = [y \ x]$$

and

$$\mathbf{V} = \begin{bmatrix} \sigma_x^2 & \sigma_{x,y} \\ \sigma_{y,x} & \sigma_y^2 \end{bmatrix}$$

$$\sigma_f^2 = x^2 \sigma_y^2 + y^2 \sigma_x^2 + 2xy \sigma_{x,y} \quad (\text{A.11})$$

If we multiply each term so that we can collect f^2 in the numerator, the variance in f can be expressed as

$$\sigma_f^2 = f^2 \left(\frac{\sigma_x^2}{x^2} + \frac{\sigma_y^2}{y^2} + \frac{2\sigma_{x,y}}{xy} \right) \quad (\text{A.12})$$

if the variables x and y are uncorrelated, the variance is further simplified, and more terms can be included easily. The simplified standard deviation of a function $f(\bar{\beta}) = \beta_1 \cdot \beta_2 \cdots \beta_n$ with uncorrelated variables is thus

$$\sigma_f = |f| \sqrt{\left(\frac{\sigma_{\beta_1}}{\beta_1} \right)^2 + \left(\frac{\sigma_{\beta_2}}{\beta_2} \right)^2 + \cdots + \left(\frac{\sigma_{\beta_n}}{\beta_n} \right)^2} \quad (\text{A.13})$$

Appendix B

Tables

For all tables, the assumption is that the main particle emission is due to alpha, proton or neutron emission. However, triton, ^3He and deuterons are still fed when above threshold, but from theory, the feeding is low. For triton, ^3He and deuteron respectively, subtract 8.5, 7.7 or 2.2 respectively from the Q value. Alphaparticles (due to the large binding energy and spin equal to 0) will be more fed. To calculate Q value for alpha emission subtract 28.3 MeV from Q value of 2p2n-reactions. Q values from ¹ are used. Q values below 40 MeV are included for comparison to experimental data.

B.1 Product nuclei, Q-values and gammarays

Table B.1: Products observed on Nickel foils. Nickel has five stable isotopes: ^{58}Ni (68.077%), ^{60}Ni (26.223 %), ^{61}Ni (1.1399%), ^{62}Ni (3.6346%) and ^{64}Ni (0.9255%). If the nucleus has provided energy level, the nucleus is an isomer, if nothing then ground state. The table is inspired by Tarkanyi et al 2019 (in paper)

Nuclide level (keV)	Half life	Decay mode	Reaction route	Q value (keV)	E_γ (keV)	I_γ (%)
^{52}Mn (0.0)	5.591 d 21.1 m	$\epsilon : 100\%$	$^{58}\text{Ni}(\text{d},2\alpha)$	-1235.6	744.233	90.0
			$^{60}\text{Ni}(\text{d},2\text{n}2\alpha)$	-21622.6	935.544	94.5
			$^{61}\text{Ni}(\text{d},3\text{n}2\alpha)$	-29442.7	1246.278	4.21
					1434.092	100.0
^{54}Mn (0.0)	312.20 d	$\epsilon : 100\%$	$^{58}\text{Ni}(\text{d},2\text{p}\alpha)$	-8538.3	834.848	99.9760
			$^{60}\text{Ni}(\text{d},2\alpha)$	-629.6		
			$^{61}\text{Ni}(\text{d},\text{n}2\alpha)$	-8449.7		
			$^{62}\text{Ni}(\text{d},2\text{n}2\alpha)$	-19045.4		
^{59}Fe (0.0)	44.490 d	$\beta^- : 100\%$	$^{60}\text{Ni}(\text{d},3\text{p})$	-12539.5	1291.590	43.2
			$^{61}\text{Ni}(\text{d},\text{n}3\text{p})$	-20359.6		
			$^{62}\text{Ni}(\text{d},\text{p}\alpha)$	-2659.7		
			$^{64}\text{Ni}(\text{d},2\text{n}\text{p}\alpha)$	-19154.9		

¹<https://www.nndc.bnl.gov/qcalc/>

APPENDIX B. TABLES

^{55}Co (0.0)	17.53 h	$\epsilon : 100\%$	$^{58}\text{Ni}(\text{d},\text{n}\alpha)$ $^{60}\text{Ni}(\text{d},3\text{n}\alpha)$ $^{61}\text{Ni}(\text{d},4\text{n}\alpha)$	-3559.4 -23946.4 -31766.5	91.9 385.4 477.2 520.0 803.7 827.0 931.1 984.6 1212.8 1316.6 1370.0 1408.5 1792.1 2177.6 2872.4	1.16 0.54 20.2 0.83 1.87 0.21 75 0.52 0.26 7.1 2.9 16.9 0.082 0.29 0.118
^{56}Co (0.0)	77.236 d	$\epsilon : 100\%$	$^{58}\text{Ni}(\text{d},\alpha)$ $^{61}\text{Ni}(\text{d},2\text{n}\alpha)$ $^{61}\text{Ni}(\text{d},3\text{n}\alpha)$ $^{62}\text{Ni}(\text{d},4\text{n}\alpha)$	6522.5 -13864.5 -21684.6 -32280.4	787.743 846.770 977.372 1175.101 1963.741 2015.215 2034.791	0.3111 99.9399 1.421 2.252 0.707 3.016 7.77
^{58}Co (0.0)	70.86 d	$\epsilon : 100\%$	$^{58}\text{Ni}(\text{d},2\text{n})$ $^{60}\text{Ni}(\text{d},\alpha)$ $^{61}\text{Ni}(\text{d},\text{n}\alpha)$ $^{62}\text{Ni}(\text{d},2\text{n}\alpha)$ $^{64}\text{Ni}(\text{d},4\text{n}\alpha)$	-1823.8 6084.9 -1735.3 -12331.0 -28826.2	810.7593 863.951 1674.725	99.450 0.686 0.517
^{60}Co (0.0)	1925.28 d	$\beta^- : 100\%$	$^{60}\text{Ni}(\text{d},2\text{p})$ $^{61}\text{Ni}(\text{d},\text{n}2\text{p})$ $^{62}\text{Ni}(\text{d},\alpha)$ $^{64}\text{Ni}(\text{d},2\text{n}\alpha)$	-4265.0 -12085.1 5614.8 -10880.4	1173.228 1332.492	99.85 99.9826
^{56}Ni (0.0)	6.075 d	$\epsilon : 100\%$	$^{58}\text{Ni}(\text{d},3\text{np})$	-24688.4	158.38 480.44 749.95 811.85 1561.80	98.8 36.5 49.5 86.0 14.0
^{57}Ni (0.0)	35.60 h	$\beta^+ : 100\%$	$^{58}\text{Ni}(\text{d},2\text{np})$ $^{60}\text{Ni}(\text{d},4\text{np})$	-14440.8 -34827.8	379.94 673.44 906.98 1046.68 1224.00 1377.63 1730.44 1757.55 1897.42 1919.52 2133.04 2804.20	0.0670 0.0491 0.0613 0.134 0.063 81.7 0.052 5.75 0.028 12.3 0.0286 0.098
^{65}Ni (0.0)	2.51719 h	$\beta^- : 100\%$	$^{64}\text{Ni}(\text{d},\text{p})$	3873.51	366.27 1481.84 1623.42 1724.92	4.81 23.59 0.498 0.399

^{60}Cu (0.0)	23.7 m	$\epsilon : 100\%$	$^{60}\text{Ni}(\text{d},2\text{n})$ $^{61}\text{Ni}(2,\text{3n})$ $^{62}\text{Ni}(\text{d},4\text{n})$	-9134.9 -16955.0 -27550.7	467.3 497.9 643.2 952.4 1035.2 1110.5 1293.7 1791.6 1861.6 1936.9 2061.0 2158.9 2403.3 2687.9 2746.1	3.52 1.67 0.97 2.73 3.70 1.06 1.85 45.4 4.8 2.20 0.79 3.34 0.77 0.44 1.06
^{61}Cu	3.339 h	$\epsilon : 100\%$	$^{60}\text{Ni}(\text{d},\text{n})$ $^{61}\text{Ni}(\text{d},2\text{n})$ $^{62}\text{Ni}(\text{d},3\text{n})$ $^{64}\text{Ni}(\text{d},5\text{n})$	2575.3 -5244.8 -15840.5 -32335.7	282.956 373.050 529.169 588.605 625.605 656.008 816.692 841.211 902.294 1032.162 1073.465 1132.351 1185.234 1446.492	12.2 2.1 0.38 1.17 0.040 10.8 0.31 0.21 0.083 0.043 0.033 0.090 3.7 0.045
^{64}Cu	12.701 h	$\epsilon : 100\%$ $\beta^- : 38.5\%$	$^{64}\text{Ni}(\text{d},2\text{n})$	-4681.3	1345.77	0.475

APPENDIX B. TABLES

Table B.2: Products observed on Iron foils. Iron has five stable isotopes: ^{54}Fe (5.845%), ^{56}Fe (91.754 %), ^{57}Fe (2.119%) and ^{58}Fe (0.282%). If the nucleus has provided energy level, the nucleus is an isomer, if nothing then ground state. **The table is inspired by Tarkanyi et al 2019 (ir paper)**

Nuclide level (keV)	Half life	Decay mode	Reaction route	Q value (keV)	E_γ (keV)	I_γ (%)
^{48}V (0.0)	15.9735 d	$\epsilon : 100\%$	$^{54}\text{Fe}(\text{d},2\alpha)$	-3490.9	944.130	7.870
			$^{56}\text{Fe}(\text{d},2\text{n}2\alpha)$	-23986.1	983.525	99.98
			$^{57}\text{Fe}(\text{d},3\text{n}2\alpha)$		1312.106	98.2
^{51}Cr (0.0)	27.704 d	$\epsilon : 100\%$	$^{54}\text{Fe}(\text{d},\text{p}\alpha)$	-1381.3	320.0824	9.910
			$^{56}\text{Fe}(\text{d},2\text{n}\alpha)$	-21876.5		
			$^{57}\text{Fe}(\text{d},3\text{n}\alpha)$	-29522.6		
			$^{58}\text{Fe}(\text{d},4\text{n}\alpha)$	-39567.2		
^{52}Mn (0.0)	5.591 d d	$\epsilon : 100\%$	$^{54}\text{Fe}(\text{d},\alpha)$	5163.6	346.02	0.980
			$^{56}\text{Fe}(\text{d},2\text{n}\alpha)$	-15331.6	744.233	90.0
			$^{57}\text{Fe}(\text{d},3\text{n}\alpha)$	-22977.7	848.18	3.32
					935.544	94.5
					1246.278	4.21
					1333.649	5.07
					1434.092	100.0
^{54}Mn (0.0)	312.20 d	$\epsilon : 100\%$	$^{54}\text{Fe}(\text{d},2\text{p})$	-2139.1	834.8480	99.9760
			$^{56}\text{Fe}(\text{d},\alpha)$	5661.4		
			$^{57}\text{Fe}(\text{d},\text{n}\alpha)$	-1984.7		
			$^{58}\text{Fe}(\text{d},2\text{n}\alpha)$	-12029.3		
^{53}Fe (0.0)	8.51 m ????	$\epsilon : 100\%$	$^{54}\text{Fe}(\text{d},2\text{np})$	-15602.9	377.9	42%
			$^{56}\text{Fe}(\text{d},4\text{np})$	-36098.1		
^{59}Fe (0.0)	44.490 d	$\beta^- : 100\%$	$^{58}\text{Fe}(\text{d},\text{p})$	4356.44	1099.245	56.5
					1291.590	43.2
^{55}Co (0.0)	17.53 h	$\epsilon : 100\%$	$^{54}\text{Fe}(\text{d},\text{n})$	2839.8	91.9	1.16
			$^{56}\text{Fe}(\text{d},3\text{n})$	-17655.4	477.2	20.2
			$^{57}\text{Fe}(\text{d},4\text{n})$	-25301.5	803.7	1.87
					827.0	0.21
					931.1	75
					1316.6	7.1
					1370.0	2.9
					1408.5	16.9
					2177.6	0.29
					2872.4	0.118
					2938.9	0.057

^{56}Co	77.236 d	$\epsilon : 100\%$	$^{56}\text{Fe}(\text{d},2\text{n})$	-7573	263.434	0.0220
(0.0)			$^{57}\text{Fe}(\text{d},3\text{n})$	-15219.7	486.55	0.0540
			$^{58}\text{Fe}(\text{d},4\text{n})$	-25264.3	733.514	0.191
					787.743	0.311
					846.770	99.9399
					852.732	0.049
					896.510	0.073
					977.372	1.421
					996.948	0.111
					1037.843	14.05
					1140.368	0.132
					1159.944	0.094
					1175.101	2.252
					1198.888	0.049
					1238.288	66.46
					1335.40	0.1224
					1360.212	4.283
					1771.357	15.41
					1963.741	0.707
					2015.215	3.016
					2034.791	7.77
					2212.944	0.388
					2276.131	0.118
					2598.500	16.97
^{57}Co	271.74 d	$\epsilon : 100\%$	$^{56}\text{Fe}(\text{d},\text{n})$	3802.9	122.06065	85.60
(0.0)			$^{57}\text{Fe}(\text{d},2\text{n})$	-3843.2	136.47356	10.68
			$^{58}\text{Fe}(\text{d},3\text{n})$	-13887.8		
^{58}Co	70.86	$\epsilon : 100\%$	$^{57}\text{Fe}(\text{d},\text{n})$	4729.7	810.7593	99.450
(0.0)			$^{58}\text{Fe}(\text{d},2\text{n})$	-5314.9		

APPENDIX B. TABLES

Table B.3: Products observed on Copper foils. Copper has two stable isotopes: ^{63}Cu (69.15%) and ^{65}Cu (30.85 %). If the nucleus has provided energy level, the nucleus is an isomer, if nothing then ground state. **The table is inspired by Tarkanyi et al 2019 (in paper)**

Nuclide level (keV)	Half life	Decay mode	Reaction route	Q value (keV)	E_{γ} (keV)	I_{γ} (%)
^{59}Fe (0.0)	44.490 d	β^- : 100%	$^{63}\text{Cu}(\text{d},2\text{p}\alpha)$	-8782.1	1099.245	56.5
			$^{65}\text{Cu}(\text{d},2\alpha)$	1687.0	1291.590	43.2
^{60}Co (0.0)	1925.28 d	β^- : 100%	$^{63}\text{Cu}(\text{d},\text{p}\alpha)$	-507.6	1173.228	99.85
			$^{65}\text{Cu}(\text{d},\text{t}\alpha)$	-9852.4	1332.492	99.9826
^{61}Co (0.0)	1.649 h	β^- : 100%	$^{63}\text{Cu}(\text{d},\text{n}3\text{p})$	-19484.2	67.412	84.7
			$^{65}\text{Cu}(\text{d},\text{np}\alpha)$	-9015.1		
^{65}Ni (0.0)	2.51719 h	β^- : 100%	$^{65}\text{Cu}(\text{d},2\text{p})$	-3580.2	1481.84	23.59
^{61}Cu (0.0)	3.339 h	ϵ : 100%	$^{63}\text{Cu}(\text{d},3\text{np})$	-21962.9	282.956	12.2
			$^{65}\text{Cu}(\text{d},5\text{np})$	-39789.4	656.008	10.8
					1185.234	3.7
^{64}Cu (0.0)	12.701 h	ϵ : 61.5% β^- : 38.5	$^{63}\text{Cu}(\text{d},\text{p})$	5691.54	1345.77	0.475
			$^{65}\text{Cu}(\text{d},2\text{np})$	-12135.0		
^{62}Zn (0.0)	9.193 h	ϵ : 100%	$^{63}\text{Zn}(\text{d},3\text{n})$	-15490.0	40.85	25.5
			$^{65}\text{Cu}(\text{d},5\text{n})$	-33316.6	243.36	2.52
					246.95	1.90
					260.43	1.35
					304.88	0.29
					394.03	2.24
					548.35	15.3
					596.56	26.0
					637.41	0.25
^{63}Zn (0.0)	38.47 m	ϵ : 100%	$^{63}\text{Cu}(\text{d},2\text{n})$	-6373.3	449.93	0.236
			$^{65}\text{Cu}(\text{d},4\text{n})$	-24199.8	669.62	8.2
					962.06	6.5
^{65}Zn (0.0)	243.93 d	ϵ : 100%	$^{65}\text{Cu}(\text{d},2\text{n})$	-4358.6	1115.539	50.04

Table B.4: Products observed in Iridium foils. Iridium has two stable isotopes: ^{191}Ir (37.3%) and ^{93}Ir (62.7 %). If the nucleus has provided energy level, the nucleus is an isomer, if nothing then ground state. **The table is inspired by Tarkanyi et al 2019 (ir paper)**

Nuclide level (keV)	Half life	Decay mode	Reaction route	Q value (keV)	E_{γ} (keV)	I_{γ} (%)
^{188}Ir (0.0)	41.5 h	ϵ : 100%	$^{191}\text{Ir}(\text{d},4\text{np})$	-24802.0	1209.80	6.9
					1715.67	6.2
					2059.65	7.0
^{189}Ir (0.0)	13.2 d	ϵ : 100%	$^{191}\text{Ir}(\text{d},4\text{np})$	-16626.0	95.23	0.38
					216.7	0.52
			$^{193}\text{Ir}(\text{d},5\text{np})$	-30596.0	233.5	0.30
					245.1	6.0
^{190}Ir (0.0)	11.78 d	ϵ : 100%	$^{191}\text{Ir}(\text{d},2\text{np})$	-10251.1	294.75	6.6
					380.03	2.03
			$^{193}\text{Ir}(\text{d},4\text{np})$	-24221.2	1036.05	2.42
$^{190m^2}\text{Ir}$ (376.4)	3.087 h	IT:8.6% ϵ : 91.4%	361.2	86.72
					502.5	89.38
					616.5	90.14
^{192}Ir (0.0)	73.829 d	ϵ : 4.76% β^- : 95.24%	$^{191}\text{Ir}(\text{d},\text{p})$	3973.55	201.3112	0.471
					295.95650	28.71
			$^{193}\text{Ir}(\text{d},2\text{np})$	-9996.6	374.4852	0.727
					416.4688	0.670
					468.06885	47.84
					489.06	0.438
					612.46215	5.34
					1061.49	0.0531
			$^{194}\text{Ir}(\text{d},\text{p})$	3842.22	293.541	2.5
					300.741	0.35
					589.179	0.140
					938.69	0.60
					1150.75	0.60
					1468.91	0.19
$^{194m^2}\text{Ir}$ (190+X)	171 d	β^- : 100%	338.8	55
					482.6	97
					562.4	35
					687.8	3.6
^{188}Pt (0.0)	10.16 d	ϵ : 99.999974% α : $2.6E - 5\%$	$^{191}\text{Pt}(\text{d},2\text{n})$	-26109.0	195.05	18.4
					381.43	7.4
^{189}Pt (0.0)	10.87 h	ϵ : 100%	$^{191}\text{Ir}(\text{d},4\text{n})$	-19389.0	94.34	6.5
					113.82	2.5
					243.50	5.9
					317.65	2.8
					721.38	7.9
^{191}Pt (0.0)	2.802 d	ϵ : 100%	$^{191}\text{Ir}(\text{d},2\text{n})$	-4017.0	178.96	12.5
					351.17	42.6
			$^{193}\text{Ir}(\text{d},4\text{n})$	-17988.0	409.44	100
					456.47	42
					538.87	181
					624.06	18.5
^{193m}Pt (149.783)	4.33 d	IT:100%	$^{193}\text{Ir}(\text{d},2\text{n})$	-3063.5	66.831	7.21
					135.5	0.1145475

Table B.5: Iridium production cross sections produced from Iridium

E_d (MeV)	Production cross section (mb) for iridium radionuclides				
	$^{188m1+g}\text{Ir}_{\text{cum}}$	$^{188m1+g}\text{Ir}_{\text{ind}}$	$^{189}\text{Ir}_{\text{cum}}$	$^{190m1+g}\text{Ir}_{\text{cum}}$	$^{190m1+g}\text{Ir}_{\text{ind}}$
30.65 $^{+0.76}_{-0.75}$	1.37 \pm 0.10	0.42 \pm 0.03	332.49 \pm 24.20	86.65 \pm 2.89	85.88 \pm 2.86
28.40 $^{+0.80}_{-0.79}$	0.45 \pm 0.07	0.17 \pm 0.02	237.84 \pm 17.44	62.80 \pm 2.14	62.36 \pm 2.13
26.03 $^{+0.82}_{-0.82}$	0.34 \pm 0.08	0.17 \pm 0.03	91.49 \pm 5.47	44.26 \pm 1.47	44.01 \pm 1.46
23.54 $^{+0.88}_{-0.87}$	-	-	19.23 \pm 2.65	27.29 \pm 1.02	27.19 \pm 1.02
21.38 $^{+0.94}_{-0.92}$	-	-	-	18.73 \pm 0.71	18.69 \pm 0.70
19.03 $^{+1.00}_{-0.99}$	-	-	-	14.02 \pm 0.55	14.00 \pm 0.55
16.43 $^{+1.11}_{-1.08}$	-	-	-	12.40 \pm 0.51	12.39 \pm 0.51
13.51 $^{+1.28}_{-1.22}$	-	-	-	8.26 \pm 0.43	8.25 \pm 0.42
10.09 $^{+1.55}_{-1.41}$	-	-	-	-	-
5.63 $^{+2.21}_{-1.83}$	-	-	-	-	-

Table B.6: Iridium production cross sections produced from Iridium

E_d (MeV)	Production cross section (mb) for iridium radionuclides			
	$^{190m2}\text{Ir}_{\text{ind}}$	$^{192}\text{Ir}_{\text{cum}}$	$^{194g}\text{Ir}_{\text{cum}}$	$^{194m2}\text{Ir}_{\text{ind}}$
30.65 $^{+0.76}_{-0.75}$	8.87 \pm 0.25	188.43 \pm 5.27	50.92 \pm 2.18	-
28.40 $^{+0.80}_{-0.79}$	5.03 \pm 0.15	152.55 \pm 4.39	51.39 \pm 2.89	-
26.03 $^{+0.82}_{-0.82}$	2.92 \pm 0.08	124.33 \pm 3.42	61.37 \pm 2.39	0.74 \pm 0.17
23.54 $^{+0.88}_{-0.87}$	1.16 \pm 0.04	100.03 \pm 3.14	69.68 \pm 2.76	0.68 \pm 0.26
21.38 $^{+0.94}_{-0.92}$	0.45 \pm 0.01	90.41 \pm 2.80	86.38 \pm 3.18	0.65 \pm 0.13
19.03 $^{+1.00}_{-0.99}$	0.16 \pm 0.01	90.65 \pm 3.01	97.79 \pm 3.99	0.60 \pm 0.14
16.43 $^{+1.11}_{-1.08}$	0.06 0.00	99.61 \pm 3.14	121.54 \pm 4.54	0.50 \pm 0.09
13.51 $^{+1.28}_{-1.22}$	0.03 \pm 0.00	107.41 \pm 3.48	143.27 \pm 5.52	-
10.09 $^{+1.55}_{-1.41}$	0.02 \pm 0.00	64.27 \pm 2.56	92.78 \pm 4.21	-
5.63 $^{+2.21}_{-1.83}$	0.02 \pm 0.00	6.67 \pm 0.37	6.32 \pm 0.42	-

B.2 Production cross sections

B.2.1 $^{\text{nat}}\text{Ir}(\text{d},\text{x})$

Table B.7: Platinum production cross sections produced from Iridium

E_d (MeV)	Production cross section (mb) for platinum radionuclides			
	$^{188}\text{Pt}_{\text{ind}}$	$^{189}\text{Pt}_{\text{ind}}$	$^{191}\text{Pt}_{\text{ind}}$	$^{193m}\text{Pt}_{\text{ind}}$
$30.65^{+0.76}_{-0.75}$	0.94 ± 0.13	486.47 ± 21.86	597.10 ± 16.55	48.11 ± 6.33
$28.40^{+0.80}_{-0.79}$	0.30 ± 0.09	341.24 ± 16.64	483.60 ± 13.79	46.78 ± 2.19
$26.03^{+0.82}_{-0.82}$	0.17 ± 0.05	172.11 ± 8.03	353.99 ± 9.67	55.68 ± 2.17
$23.54^{+0.88}_{-0.87}$	-	30.72 ± 1.48	165.12 ± 5.15	51.79 ± 2.12
$21.38^{+0.94}_{-0.92}$	-	1.04 ± 0.07	71.05 ± 2.19	58.31 ± 1.96
$19.03^{+1.00}_{-0.99}$	-	0.09 ± 0.02	77.53 ± 2.57	77.98 ± 2.89
$16.43^{+1.11}_{-1.08}$	-	-	128.24 ± 4.03	115.33 ± 4.09
$13.51^{+1.28}_{-1.22}$	-	-	137.37 ± 4.42	148.98 ± 5.54
$10.09^{+1.55}_{-1.41}$	-	-	53.45 ± 2.12	56.18 ± 2.85
$5.63^{+2.21}_{-1.83}$	-	-	1.05 ± 0.06	1.56 ± 0.12

Table B.8:

E_d (MeV)	Production cross section (mb) for ...				
	$^{48}\text{V}_{\text{cum}}$	$^{51}\text{Cr}_{\text{cum}}$	$^{52}\text{Mn}_{\text{cum}}$	$^{54}\text{Mn}_{\text{ind}}$	$^{53}\text{Fe}_{\text{cum}}$
$29.57^{+0.68}_{-0.68}$	0.12 ± 0.01	7.54 ± 0.23	16.00 ± 0.36	23.85 ± 0.70	5.12 ± 0.65
$27.26^{+0.73}_{-0.72}$	0.09 ± 0.01	7.86 ± 0.25	5.48 ± 0.16	24.18 ± 0.72	2.77 ± 0.44
$24.80^{+0.77}_{-0.76}$	0.06 ± 0.00	8.51 ± 0.29	0.91 ± 0.03	26.12 ± 0.79	1.29 ± 0.30

Table B.9:

E_d (MeV)	Production cross section (mb) for ...			
	$^{59}\text{Fe}_{\text{ind}}$	$^{55}\text{Co}_{\text{ind}}$	$^{57}\text{Co}_{\text{ind}}$	$^{58}\text{Co}_{\text{ind}}$
0.16 ± 0.02	27.15 ± 0.80	35.91 ± 1.06	1.50 ± 0.05	$29.57^{+0.68}_{-0.68}$
$27.26^{+0.73}_{-0.72}$	0.15 ± 0.02	20.44 ± 0.60	38.37 ± 1.13	1.62 ± 0.05
$24.80^{+0.77}_{-0.76}$	0.18 ± 0.04	13.82 ± 0.40	42.63 ± 1.27	2.05 ± 0.07

Table B.10:

E_d (MeV)	Production cross section (mb) for ...					
	$^{59}\text{Fe}_{\text{cum}}$	$^{60}\text{Co}_{\text{cum}}$	$^{61}\text{Co}_{\text{cum}}$	$^{65}\text{Ni}_{\text{ind}}$	$^{61}\text{Cu}_{\text{cum}}$	$^{64}\text{Cu}_{\text{ind}}$
$30.03_{-0.67}^{+0.67}$	0.21 ± 0.03	9.49 ± 0.52	1.62 ± 0.09	3.42 ± 1.76	4.54 ± 0.87	170.76 ± 7.76
$27.74_{-0.71}^{+0.72}$	0.18 ± 0.02	11.38 ± 0.51	0.82 ± 0.07	3.84 ± 1.94	2.07 ± 0.95	153.78 ± 8.20
$25.32_{-0.76}^{+0.77}$	0.17 ± 0.02	12.02 ± 0.51	0.29 ± 0.05	2.89 ± 1.48	$1.18 \text{ pm } 0.72$	132.56 ± 6.93
$22.77_{-0.81}^{+0.83}$	0.12 ± 0.01	11.36 ± 0.43	-	1.92 ± 1.14	-	121.54 ± 7.12
$20.57_{-0.87}^{+0.89}$	0.07 ± 0.01	9.27 ± 0.41	-	-	-	106.07 ± 5.81
$18.14_{-0.94}^{+0.97}$	0.03 ± 0.01	5.65 ± 0.26	-	1.46 ± 0.95	-	95.92 ± 7.14
$15.43_{-1.04}^{+1.08}$	-	1.53 ± 0.12	-	-	-	123.79 ± 6.62
$12.34_{-1.20}^{+1.27}$	-	-	-	-	-	156.65 ± 8.20
$8.68_{-1.43}^{+1.62}$	-	-	-	-	-	209.38 ± 11.27
$3.94_{-2.22}^{+2.25}$	-	-	-	-	-	73.54 ± 5.70