

Confidence Intervals and a taste of ML

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TUM - winter term 2020/21

Large Sample Theory: Likelihood Ratio Testing

Reminder

given two hypotheses $H0: \theta = \theta_0$ and $H1: \theta \neq \theta_0$ the likelihood ratio test-statistic $\lambda(x)$ is defined as

$$\lambda(x) = -2\log\Lambda(x) = -2\log\left\{\frac{\sup_{\nu}L(\theta_0, \nu|x)}{\sup_{\nu, \theta}L(\theta, \nu|x)}\right\}$$
(1)

to perform the hypothesis test, we also need to know the sampling distribution of this test-statistic:

$$\lambda \sim f_{\lambda}(\lambda; \boldsymbol{\theta}, \boldsymbol{\nu})$$
 (2)

Often, this is non-trivial and one needs extensive Monte-Carlo computations (see example 3)

Luckily, as the sample size increases, the distribution is known to **converge**! (beware of conditions!)

Large Sample Theory: Likelihood Ratio Testing

Wilk's Theorem

As the sample size increases, the distribution of the likelihood ratio test-statistic (eq. 10) converges to a χ^2 distribution with number of degrees of freedom k equal to the difference in number of free parameters specified by each hypothesis. In our notation $k = \dim \theta$.

$$f_{\lambda}(\lambda; \boldsymbol{\theta}_0) \underset{n \to \infty}{\longrightarrow} \chi^2(k)$$
 (3)

Large Sample Theory: Likelihood Ratio Testing

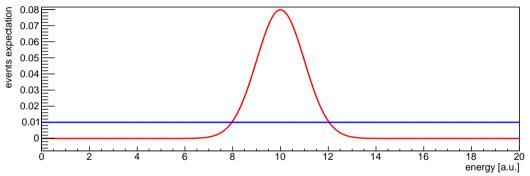
Wilk's Theorem (cont'd)

Unfortunately there are strict regularity conditions. Here are the two most important ones

- $heta_0$ needs to be an interior point of Θ
- ullet nuisance parameters u that are only present under H1 are another issue
- ... several minor ones (typically not important)

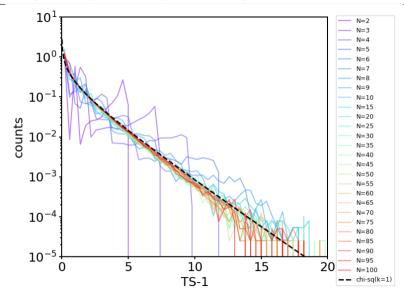
Some extensions exists that might be useful (see Chernoff 1954, Gross, Vitells 2010) in such situations.

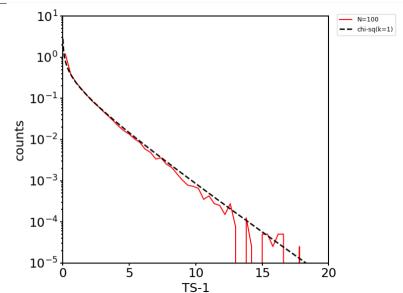
Application to our standard toy problem (with 2 parameters: p_s , μ_s)



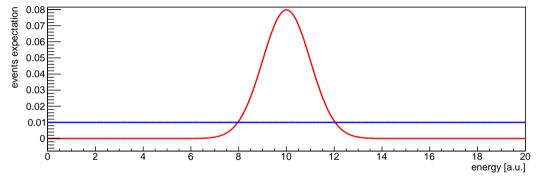
Two different hypothesis tests satisfying Wilk's theorem

Case 1: $H0: p_s = 0.2 \text{ and } H1: p_s \neq 0.2 \text{ (k=1) } (\mu_s \text{ is nuisance!})$





Application to our standard toy problem (with 2 parameters: p_s , μ_s)



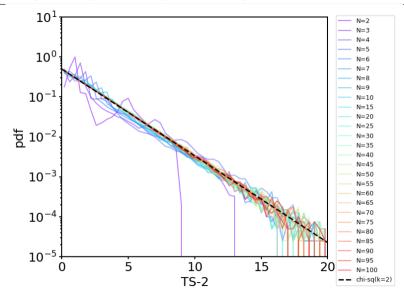
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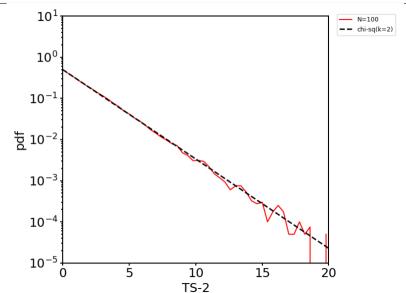
Case 1:
$$H0: p_s = 0.2$$
 and $H1: p_s \neq 0.2$ (k=1)

Case 2:
$$H0: p_s = 0.2, \ \mu_s = 10.0 \ \text{and} \ H1: p_s \neq 0.2, \ \mu_s \neq 10.0 \ (k=2)$$

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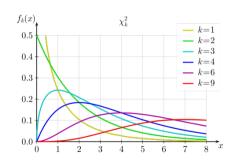


Critical Values from Chi-squared Distribution

for different number of degrees of freedom ($n = \Delta \mathrm{dim} \theta$) and various choices of common levels α .

(here: critical value $c \equiv Q_{lpha}$))

$1-\alpha$	Q_{lpha}				
	n = 1	n = 2	n = 3	n=4	n = 5
0.683	1.00	2.30	3.53	4.72	5.89
0.90	2.71	4.61	6.25	7.78	9.24
0.95	3.84	5.99	7.82	9.49	11.1
0.99	6.63	9.21	11.3	13.3	15.1



Questions?

Confidence Intervals

Goal: calculate some range/region that has some probability to contain the true (unknown) parameter/s.

- Probability does not refer to the parameter (the true parameter is a fixed constant, not a random variable.) but to the region/interval that we obtain from the data.
- Generally speaking: different data results in a different region/interval (albeit construction is the same).

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Mathematically, from data X we calculate function values L(X) and U(X) which are random variables.

$$[L(X), U(X)] \quad (two - sided)$$

$$(-\infty, U(X)] \quad or \quad [L(X), \infty) \quad (one - sided)$$

$$(5)$$

in physics: two-sided intervals often called "uncertainties", one-sided intervals often called "limits". (sometimes gets mixed ... e.g. hard to tell difference on bounded parameter spaces. always check how the construction was done.)

Confidence Intervals: Coverage

coverage := probability that the random interval [L(X), U(X)] (or limit) happens to overlap with the unknown, true parameter value.

$$P_{\theta}\left(\theta\in\left[L(\mathsf{X}),\ U(\mathsf{X})\right]\right)\tag{6}$$

confidence coefficient of an interval (denoted by $1-\alpha$) defined by

$$\inf_{\theta} P_{\theta} \left(\theta \in [L(X), U(X)] \right) = 1 - \alpha$$
 (7)

Can not always guarantee exact coverage (hello nuisance parameters!) - strive to guarantee confidence coefficient (i.e. minimum coverage!). That is usually possible.

Confidence Intervals: Coverage in the normal mean problem

Consider the problem of constructing a confidence interval for the unknown mean μ of a normal distribution (variance σ^2 known) from n observations $(X = \{X_1, ..., X_n\})$. This can be done using a **pivot** (a function of the parameter and observations, that has a distribution which is independent of the parameter).

$$Q(\mu, \mathsf{X}) = \frac{\bar{\mathsf{X}} - \mu}{\sigma / \sqrt{n}}$$

$$Q \sim N(0, 1)$$
(8)

i.e. here Q is a standard normal random variable. Thus can solve

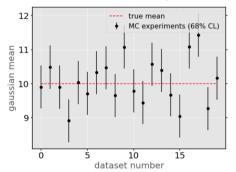
$$P_{\mu}\left(-a \le Q \le a\right) = 1 - \alpha \tag{10}$$

which corresponds to the following confidence set

$$\left\{ \mu : \, \bar{X} - a \frac{\sigma}{\sqrt{n}} \le \mu \le \bar{x} + a \frac{\sigma}{\sqrt{n}} \right\} \tag{11}$$

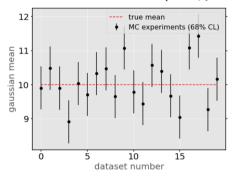
Confidence Intervals: Coverage in the normal mean problem

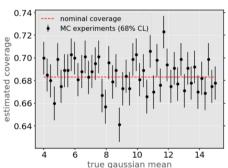
Check with Monte-Carlo (see ipython notebook)



Confidence Intervals: Coverage in the normal mean problem

Check with Monte-Carlo (see ipython notebook)





Confidence Intervals from inversion of hypothesis tests

If you can construct a level α hypothesis test for the unknown parameter/s specified by H_0 it is always possible to use this test to construct a confidence interval with guaranteed confidence coefficient $1 - \alpha$.

This is called **inverting a hypothesis test**. Whether you get two-sided or one-sided intervals depends on the alternative hypothesis

- $H_0: \theta = \theta_0$ and $H_1: \theta \neq \theta_0$ produces two-sided intervals
- $H_0: \theta = \theta_0$ and $H_1: \theta < \theta_0$ produces one-sided intervals (upper-limit)
- $H_0: \theta = \theta_0$ and $H_1: \theta > \theta_0$ produces one-sided intervals (lower-limit)

Confidence Intervals from inversion of hypothesis tests

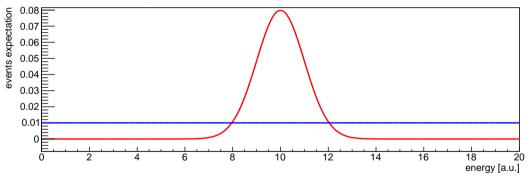
Why does it work?

- perform the test on every possible point in parameter space
- if the test rejects the point, simply discard it
- if the point is accepted, add the point to your confidence set
- Whats the coverage of this strategy? (probability that the random set contains true parameter)
- Probability to rejected a parameter if it is true is $\leq \alpha$ by definition (size of test)
- Thus, probability for true parameter to contribute to set is $\geq 1 \alpha$.
- ullet Hence, probability for set to cover true parameter is $\geq 1-lpha$ by construction

Questions?

Confidence Intervals from inversion of LRT

We have learned how to construct likelihood ratio tests. Let's invert a likelihood ratio test to obtain a confidence set on the signal fraction p_s in our toy model. (see ipython notebooks)



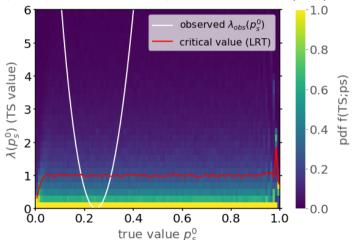
$$f_X(x;\mu,\sigma,p_s) = \left[p_s \cdot rac{1}{\sqrt{2\pi\sigma^2}} e^{rac{-(x-\mu)^2}{2\sigma^2}} + (1-p_s) \cdot rac{1}{20}
ight]$$

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(12)

Confidence Intervals from inversion of LRT in toy problem.

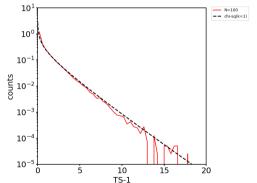
 $H_0: p_s = p_s^0$ and $H_1: p_s \neq p_s^0$, sample size n=100 endpoints of interval: intersection points of obs. TS value (white) with critical value (red)



Reminder: Asymptotic Chi-Squared.

For large n:

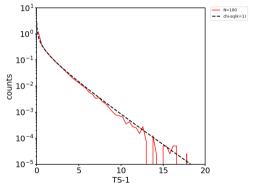
The distribution of λ is independent from (true) p_s^0 with $\lambda \sim \chi^2(k=1)$ We just have to find the points, where λ changes by 1 (for $1-\alpha=0.68$). (because $p(\lambda \geq c) = \alpha$ yields c=1 for k=1 and $1-\alpha=0.68$.

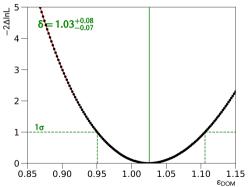


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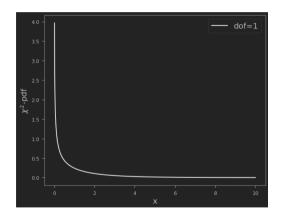
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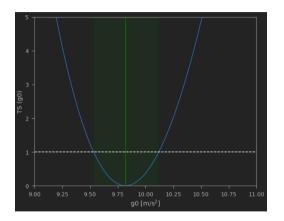
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Exercise 4, Problem 3



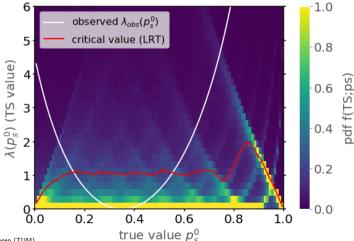


Questions?

Confidence Intervals from inversion of LRT in toy problem.

The problem is much harder, if Wilk's theorem does not apply.

 $H_0: p_s = p_s^0$ and $H_1: p_s \neq p_s^0$, sample size n=10 endpoints of interval: intersection points of obs. TS value (white) with critical value (red)



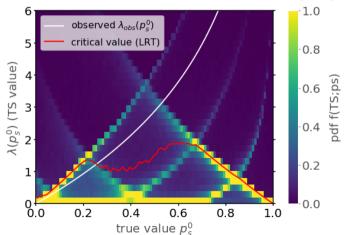
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Confidence Intervals from inversion of LRT in toy problem.

The problem is much harder, if Wilk's theorem does not apply.

 $H_0: p_s = p_s^0$ and $H_1: p_s \neq p_s^0$, sample size n=3 endpoints of interval: intersection points of obs. TS value (white) with critical value (red)



Summary: Confidence Intervals from inversion of LRT

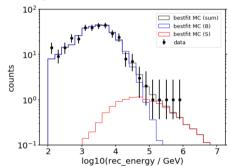
- very simple if your measurement is in the asymptotic regime (lots of data!)
- • obtain the critical value (red curve) from wilk's theorem (i.e. appropriate $\chi^2\text{-pdf})$
- generalizes well to high dimensions, if analysis remains asymptotic
- if asymptotics don't apply, you will run out of CPU quickly as the dimensionality increases (since you need to costruct the TS distributions for each point in parameter space)
- always check a few representative parameter combinations (and also a few extreme ones) first

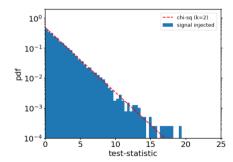
Example: Inversion of LRT in the IceCube diffuse flux measurement

To construct a joint confidence interval for the normalization and spectral index of the astrophysical neutrino flux, we need to invert a LRT:

 $H_0: (\Phi, \gamma) = (\Phi_0, \gamma_0)$ and $H_1: (\Phi, \gamma) \neq (\Phi_0, \gamma_0)$

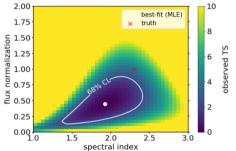
The asymptotic expectation for the TS distribution would be χ^2 with 2 dof.





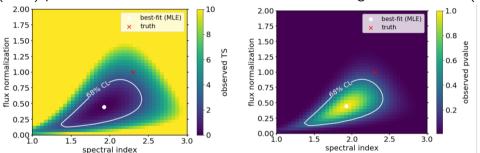
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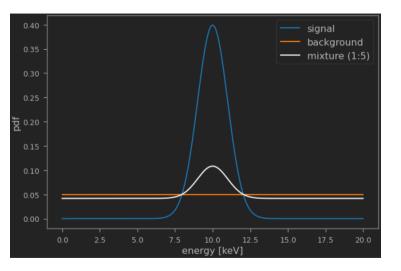


$$p(x_{obs}) = \sup_{\theta \in \Theta_0} P_{\theta} (TS(X) \ge TS(x_{obs}))$$
 (13)

Questions?

A taste of ML: Classification

Let's come back to our toy example.



$$f_X(x;\mu,\sigma,p_s) = \left[p_s \cdot \frac{1}{\sqrt{2\pi\sigma^2}} e^{\frac{-(x-\mu)^2}{2\sigma^2}} + (1-p_s) \cdot \frac{1}{20}\right]; p_s \frac{\lambda_s}{\lambda_s + \lambda_b}$$
(14)

To classify events into signal and background based on the observed energy, we need to setup some notation.

- introduce binary random variable c for the class-membership of each event (c=1 for signal, c=0 for background).
- need to find posterior probability $p(c \mid E)$.
- one possible criterion: consider event as signal if p(c = 1 | E) > 0.5.

We need the following ingredients: p(c), $f(E \mid c)$ and f(E).

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$$p(c \mid E) = \frac{f(E \mid c) \cdot p(c)}{f(E)}$$
(15)

where f(E) is the usual energy pdf (mixture).

The marginal pmf p(c) is easy.

•
$$p(c=1) = \frac{\lambda_s}{\lambda_s + \lambda_b}$$

•
$$p(c = 0) = 1 - p(c = 1) = \frac{\lambda_b}{\lambda_s + \lambda_b}$$

```
def class prob(c, pars):
    lambda s, lambda b = (pars['lambda s'], pars['lambda b'])
    lambda tot = lambda s + lambda b
    if c == 0:
        return lambda b / lambda tot
        return lambda s / lambda tot
        raise ValueError(f"ValueError: c can only be 0 or 1. but value c={c} gi
pars = {'lambda s':200, 'lambda b':1000}
print(class prob(0, pars))
print(class prob(1, pars))
```

And so is the conditional pdf $f(E \mid c)$.

For c=1: this is just the normal $N(\mu,\sigma)$ distribution For c=0: we have the uniform $u(E_{min},E_{max})$ distribution

```
def conditional energy pdf(energy, c, pars):
    if c == 0:
        return uniform.pdf(energy, pars['emin'], pars['emax'])
    elif c == 1:
        return norm.pdf(energy, pars['mu'], pars['sigma'])
        raise ValueError(f"ValueError: c can only be 0 or 1. but value c=\{c\} given.")
pars.update({'emin':0.. 'emax':20.. 'mu':10.. 'sigma':1.})
print(conditional energy pdf(10, 0, pars))
print(conditional energy pdf(10, 1, pars))
 0.3989422804014327
```

Putting it all together ...

```
def joint_pdf(energy , c, pars):
    if c == 0 or c == 1:
        return conditional_energy_pdf(energy, c, pars) * class_prob(c, pars)
    else:
        raise ValueError(f'ValueError: c can only be 0 or 1. but value c={c} given.*)
    print(joint_pdf(10, 0, pars))
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        a.sissexexexexexex
        0.04649310000000544
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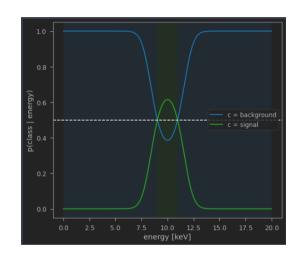
0.01206806806606544
```

Voila! The result.

```
energies = np.linspace(0.0, 20.0, 1000)
background probs = class prob given energy(0, energies, pars)
boundary high = energies[boundary idx[1]]
plt.plot(energies, background probs, label='c = background', color='tab:blue', linewidth=2)
plt.plot(energies, signal_probs, label='c = signal', color='tab:green', linewidth=2)
plt.axvspan(pars['emin'], boundary low, color='tab:blue', alpha=0.1)
plt.axvspan(boundary high, pars['emax'], color='tab:blue', alpha=0.1)
plt.axyspan(boundary low, boundary high, color='tab:green', alpha=0.1)
plt.axhline(0.5, color='white', linestyle='dashed')
plt.xlabel("energy [keV]")
plt.ylabel("p(class | energy)")
plt.legend(fontsize=12)
```

Voila! The result.

```
energies = np.linspace(\theta.0. 20.0. 1000)
background probs = class prob given energy(0, energies, pars)
plt.plot(energies, background probs, label='c = background', color='tab:blue', linewidth=2)
plt.axhline(0.5, color='white', linestyle='dashed')
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```

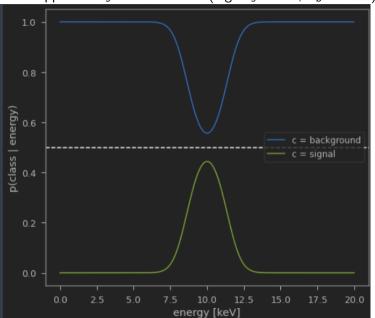


The green region shows where $f(c=1 \mid E) > 0.5$, i.e. energies for which the conditional signal probability is larger than the background probability.

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What happens if λ_s becomes small (e.g. $\lambda_s=100$, $\lambda_b=1000$)?

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Questions?

What if we do not know the right model?

- true relationship $\mu(x) = g(x; \vec{\theta})$ e.g. in regression problems (like $s(t) = \frac{1}{2}gt^2$ in HW ex. 4, problem 2)
- correct sampling pdfs $f(x; \vec{\theta})$ e.g. for classification (like the mixture model in our toy-problem)

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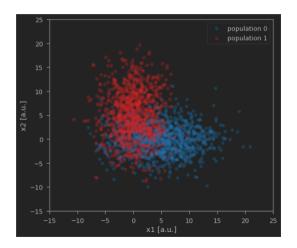
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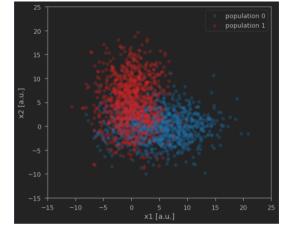
One can still perform inference and learn from the data!

- use flexible, non-parametric models and fit!
- re-formulate the problem in a clever way.

Example: consider data generated from a **bi-variate** gaussian mixture with 2 components.

Pretend we do not know the true data-generating process (i.e. the gaussians!)





Each data point comes with two observables: $\vec{x} = (x_1, x_2)$.

Goal: Classify data points depending on \vec{c} into class 0 (blue) or class 1 (red) but without using the correct pdfs.

Goal: Calculate an estimate of $p_1(\vec{x}) \equiv p(c=1; \vec{x})$.

Solution: Logistic regression - or more complex neural networks (or decision tree forests etc...).

Reminder: The multi-variate normal distribution

$$f\left(\vec{x}\right) = \left(\frac{1}{2\pi}\right)^{d/2} |\Sigma|^{-1/2} \exp\left(-\frac{1}{2} \left(\vec{x} - \vec{\mu}\right) \Sigma^{-1} \left(\vec{x} - \vec{\mu}\right)^{T}\right)$$

for d=2 the covariance matrix Σ can be expressed as

$$\Sigma = egin{bmatrix} \sigma_1^2 &
ho\sigma_1\sigma_2 \
ho\sigma_1\sigma_2 & \sigma_2^2 \end{bmatrix}$$

and $0 < \rho < 1$ measured the amount of correlation between the components x_1 and x_2 .

(and now let's forget about it again ...)

Assume we have a **training dataset**, which consists of samples $\vec{x_i}$ with known/observed labels c_i ($c_i = 0$ or $c_i = 1$).

One reasonable statistical model would be the *tossing a biased coin*, which can be described with the **bernoulli pmf** for the random variable *c*.

$$c_i \sim \text{bernoulli}(p_1(\vec{x_i}))$$
 (16)

$$p(c) = p_1^c (1 - p_1)^{1 - c}, \qquad c \in \{0, 1\}$$
(17)

where $0 \le p_1(\vec{x}) \le 1$ is the probability for class 1 (success) with some dependence on the values for $\vec{x} = (x_1, x_2)$.

Now we need a parameterization of $p_1(\vec{x})$.

A simple linear relationship $p_1(\vec{x}) \equiv g(\vec{x}) = \vec{\omega} \cdot \vec{x} + b$ won't work, since $g(\vec{x})$ ranges (in principle) from $-\infty$ to $+\infty$. But probabilities are bounded by 0 and 1.

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sigmoid function to the rescue! It maps from $(-\infty, +\infty)$ to (0, 1)

$$\sigma(z) = \frac{1}{1 + e^{-z}}$$

$$p_{1}(\vec{x}) = \sigma(g(\vec{x}))$$

$$= \sigma(\vec{\omega} \cdot \vec{x} + b)$$

$$(18)$$

$$= \sigma(\vec{\omega} \cdot \vec{x} + b)$$

$$(20)$$

What does this mean?

$$\vec{\omega} \cdot \vec{x} + b = \log\left(\frac{p_1}{1 - p_1}\right)$$

$$= \log\left(\frac{p_1(\vec{x})}{p_0(\vec{x})}\right)$$
(21)

Implicit Assumption: log-odds are linear in \vec{x} . This could easily be relaxed (this will yield more complex ML models).

What does this mean?

$$\vec{\omega} \cdot \vec{x} + b = \log\left(\frac{p_1}{1 - p_1}\right)$$

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Implicit Assumption: log-odds are linear in \vec{x} . This could easily be relaxed (this will yield more complex ML models).

Staying with the linear log-odds (logistic regression) gives 3 free parameters: slopes $\vec{\omega} = (\omega_1, \omega_2)$ and intercept b.

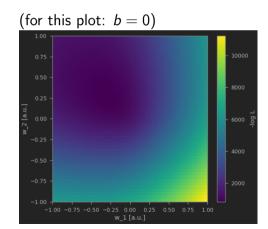
Can fit parameters using **maximum likelihood**!

$$\log L(\vec{\omega}, b | \{\vec{x}_i, c_i\}) = \sum_{i=0}^{N} \{c_i \log[\sigma(g(\vec{x}, \vec{\omega}, b))] + (1 - c_i) \log[1 - \sigma(g(\vec{x}, \vec{\omega}, b))]\}$$
(23)

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(23)

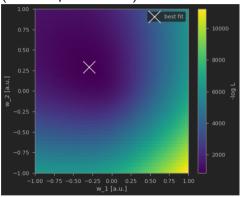
```
def neg logl(pars, samples, labels):
   weights = pars[:2]
    b = pars[2]
    zs = np.dot(weights, samples.T) + b
   ps = 1./ (1 + np.exp(-zs))
    ll = np.sum(bernoulli.logpmf(labels, ps))
```



numerical minimization in (ω_1, ω_2, b) using scipy (real problems: gradient descent)

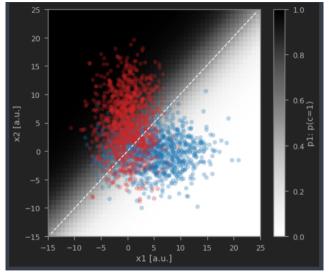
```
from scipy.optimize import minimize
starting slopes = np.array([-1., 1])
starting intercepts = np.array([0.0])
pars starting point = np.concatenate([starting slopes, starting intercepts])
obi = lambda x: neg logl(x, combined samples, combined labels)
          [ 0.00000000e+00. 1.0000000e+00. 0.0000000e+00].
```





Let's **predict**! (assuming the best fit $\hat{\vec{\omega}}$. \hat{b}

```
def get p(x, pars):
edges = np.linspace(-15, 25, 51)
centers = 0.5*(edges[:-1] + edges[1:])
xv, vv = np.meshgrid(centers, centers, sparse=False, indexing='xv')
pos = np.stack([xv.flatten(), yv.flatten()], axis=1)
pls = get p(pos, best pars)
fig, ax = plt.subplots()
h = ax.hist2d(xv.flatten(), vv.flatten(), bins=[edges]*2, weights = pls.flatten(), norm=Normalize
ax.set xlabel("x1 [a.u.]")
ax.set ylabel("x2 [a.u.]")
cbar = fig.colorbar(h[3], ax=ax)
cbar.set label("p1: p(c=1)")
add scatter points(c0 samples, label="population 0", color="tab:blue")
add scatter points(cl samples, label="population 1", color="tab:red")
ax.contour(centers, centers, pls.reshape(len(centers), len(centers)), [0.5], colors=['white'], l
```



white dashed line: estimated that $p_0 = p_1 = 0.5$.

Summary

Achieved decent estimates of the **class-membersip probabilities** without knowledge of the data generating process (underlying process)!

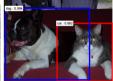
For more complex problems: increase flexibility of $g(\vec{x})$ - introducing more parameters.

Neural networks (and other techniques) are natural extensions of this example. Sometimes involving several million parameters!

Eventually it turns into this:











Hans Niederhausen (TUM)

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