ECE235 HW1, Harry Luo

3.17

By Stefan-Boltzmann Law, set total power $P=\kappa R$ and initial tempreture T_0 , we have

$$R = \sigma T^4 \Rightarrow P = \kappa \sigma T^4$$

$$\frac{P'}{P} = \frac{T'^4}{T^4} = \frac{(2T_0)^4}{T_0^4} = 16$$
(1)

Power increases by a factor of 16.

3.19

• (a)

Let initial tempreture be T_0 and the new tempreture be T^\prime . By Wien's Law, we have

$$\lambda T_0 = 2.898e^{-3} \ m \cdot K \quad \Rightarrow T_0 = \frac{2.898e^{-3}}{27e^{-6}} K = 107.33K.$$
 (2)

Using Stefan-Boltzmann Law to find the new tempreture.

$$\frac{P'}{P} = \frac{(T')^4}{T_0^4} = 2 \quad \Rightarrow T' = \sqrt[4]{2T_0^4} = \sqrt[4]{2 \times 107.33^4} = \boxed{127.63K}$$

• (b) By Wien's law,

$$\lambda = \frac{2.898e-3}{T'} = \frac{2.898e-3}{127.63}m = \boxed{22.7\mu m}$$
(4)

3.24

• (a)

Energy quantization shows:

$$E = hf = \frac{hc}{\lambda}. (5)$$

Given $\lambda \in (380, 750)$ nm,

$$\frac{hc}{750\text{nm}} < E < \frac{hc}{380\text{nm}} \quad \Rightarrow \boxed{E \in (1.653, 3.542)\text{eV}}$$
 (6)

• (b)

$$E = hf = 4.136 \times 10^{-15} \times 100 \times 10^6 \text{ eV} = \boxed{4.136 \times 10^{-7} \text{ eV}}$$
 (7)

3.25

• (a)

By the photoelectric effect equation, at therashold wavelength, we have

$$\Phi = hf_t = h\frac{c}{\lambda_t} \quad \Rightarrow \lambda_t = \frac{hc}{\Phi} = \frac{1.24e\text{-}6}{4.87}m = \boxed{2.546e\text{-}7m}$$
 (8)

• (b)

As suggested on Piazza, we use mid-point approximation to approximate the integrated energy density of sunlight from 0nm to 254.6nm by using the intensity at 254.6/2 = 127.3 nm as constant density:

$$u(127.3\text{nm}) \times (254.6 \text{ nm}) = \frac{8\pi hc (127.3e-9m)^{-5}}{e^{hc/(k\times 5800K\times 127.3e-9m)} - 1} \times (254.6e-9m) \approx 1.23e-4 \quad J/m^3$$
 (9)

Energy density is thus approximately

$$R' = \frac{1}{c}(1.23e-4) \quad J/m^3 \tag{10}$$

Total energy is given by

$$R = \sigma T^4 = \sigma \times 5800 K^4 \approx 6.42e7 \quad W/m^2$$
 (11)

Thus the maximal fractional power is

$$\frac{R'}{R} \approx 1.4e-4 \tag{12}$$

3.26

• (a)

Using the photoelectric equation, we can find threshold freq and wavelength, f_t, λ_t as follows,

$$\Phi = h f_t = \frac{hc}{\lambda_t}$$

$$\Rightarrow f_t = \frac{\Phi}{h} = \frac{1.9eV}{4.136e - 15eV \cdot s} = \boxed{4.59e4 \text{ Hz}},$$

$$\lambda_t = \frac{hc}{\Phi} = \frac{1240 \text{ eV} \cdot \text{nm}}{1.9 \text{ eV}} = \boxed{653 \text{ nm}}$$

• (b,c,d) The stopping potential can be found as follows,

$$eV_0 = \frac{hc}{\lambda} - \Phi \quad \Rightarrow V_0 = \frac{hc}{\lambda e} - \frac{\Phi}{e}. \tag{14}$$

For $\lambda = 300$ nm:

$$V_0 = \frac{1240 \text{ eV} \cdot \text{nm}}{e \times 300e - 9m} - \frac{1.9 \text{ eV}}{e} = \boxed{2.23V}$$
 (15)

For $\lambda = 400$ nm,

$$V_0 = \frac{1}{e} \frac{1240 \text{eV} \cdot \text{nm}}{400 \text{nm}} - 1.9 \text{eV} = \boxed{1.20V}$$
 (16)

3.28

• (a)

$$f_t = \frac{\Phi}{h} = \frac{4.22 \text{ eV}}{4.14e - 15 \text{ eV} \cdot s} = \boxed{1.02 \times 10^{15} \text{ Hz}}$$
 (17)

• (b)

$$f = \frac{c}{\lambda} = \frac{3e8}{560e_{-9}} \text{ Hz} = \boxed{5.36 \times 10^{14} \text{ Hz} < f_t}$$
 (18)

Frequency is less than the threshold frequency, so **no** photoelectrons are emitted.

Consider the photoelectric effect equation for n = 60 photons,

$$E = n \frac{hc}{\lambda} = \frac{60 \times 6.63e - 34 \times 3e8}{550e - 9} J = 2.17e - 17J$$
 (19)

3.32

• (a)

$$\Phi = \frac{hc}{\lambda} = \frac{1240}{653} \text{ eV} = \boxed{1.9 \text{ eV}}$$
(20)

• (b)

$$E_k = \frac{hc}{\lambda} - \Phi = \frac{1240}{300} \text{ eV} - 1.9 \text{ eV} = \boxed{2.23 \text{ eV}}$$
 (21)

3.42

Consider the stopping potential function for both cases, we have

$$eV = \frac{hc}{\lambda} - \Phi$$

$$\Rightarrow \begin{cases} V_1 = \frac{1}{e} \frac{hc}{\lambda_1} - \Phi \\ V_2 = \frac{1}{e} \frac{hc}{\lambda_2} - \Phi \end{cases}$$
(22)

Where $V_1=0.52V,\, \lambda_1=450$ nm; $V_2=1.9V, \lambda_2=300$ nm.

Solving Equation 22 for h and Φ :

$$\begin{cases} 0.52V = \frac{1}{e} \frac{hc}{450 \text{nm}} - \Phi \\ 1.9V = \frac{1}{e} \frac{hc}{300 \text{nm}} - \Phi \end{cases} \Rightarrow \begin{cases} h = 6.6376e - 34J \cdot s & \text{(a good approximation!)} \\ \Phi = 2.24 \text{ eV} \end{cases}$$
 (23)