

ECE 535: Introduction to Quantum Sensing

Uncertainties in measurements

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Fall 2025



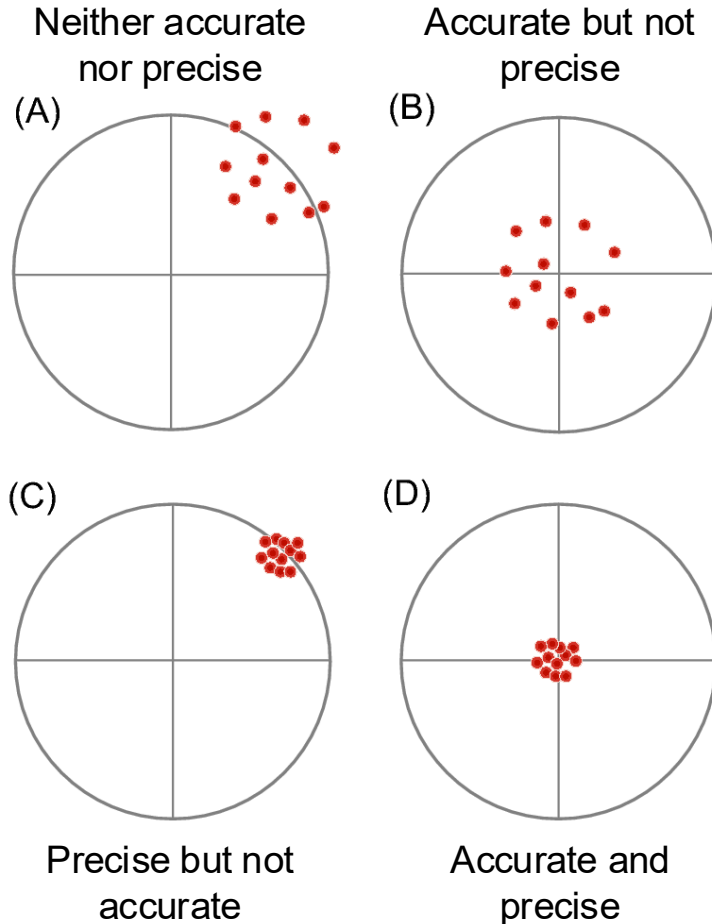
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Topics

- Error quantification in measurements
- Limits due to the uncertainty principle
 - Concept of squeezing in measurements
- Limits in measurements due to statistics (projection noise)
- Limits in measurements from discrete nature of light and matter (shot noise)
- Vacuum fluctuations

Error analysis in experiments

Recall: Accuracy vs precision



- Uncertainties in measurements can be categorized into errors of precision and accuracy
- Precision is related to the distribution of random errors, as a result of experimental conditions and/or the physical process being measured
 - Error of precision can be estimated by repeating measurements and decreases with increasing number of experiments
- Accuracy is related to systematic errors, which could be due to instrumentation
 - Error of accuracy can be inferred by referencing an experiment against a known standard or a measurement with better accuracy
- Usually one of these error sources will dominate

Natl Sci Rev, Volume 3, Issue 2, June 2016, Pages 189–200

References:

<https://reference.wolfram.com/applications/eda/ExperimentalErrorsAndErrorAnalysis.html>

<https://faraday.physics.utoronto.ca/PVB/Harrison/ErrorAnalysis/>

Is there such a thing called *over precision*?

Jen went to the store to buy a crinoid fossil. She was told by the owner that the fossil is 350 million years old. She took the fossil home and displayed it on a shelf.

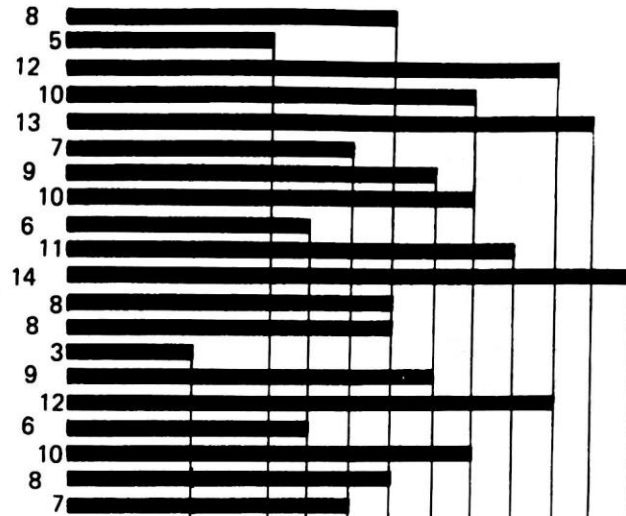
6 months later, some of her friends came over and asked her how old is the fossil.

She proudly proclaimed “350 million years and 6 months!”

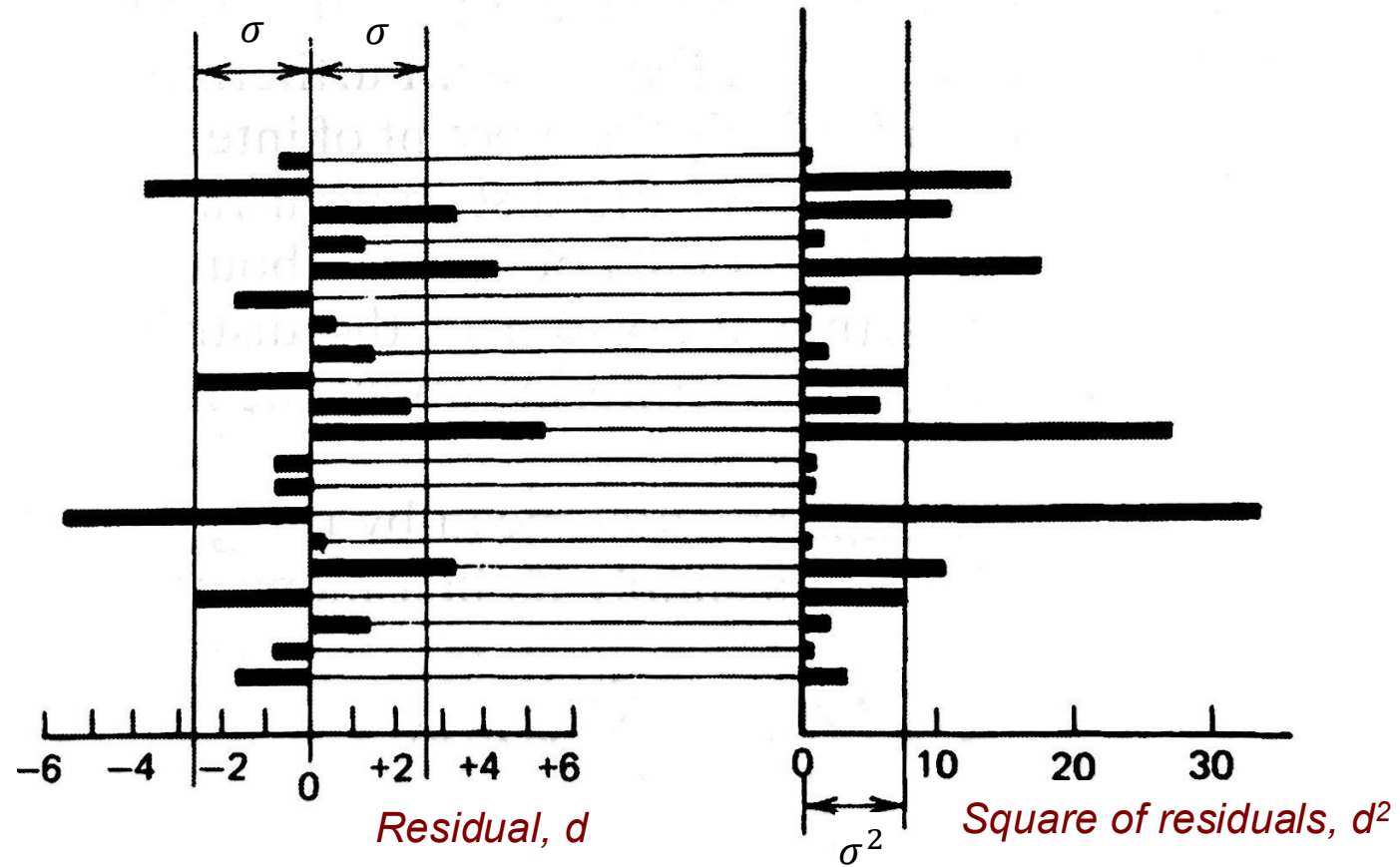
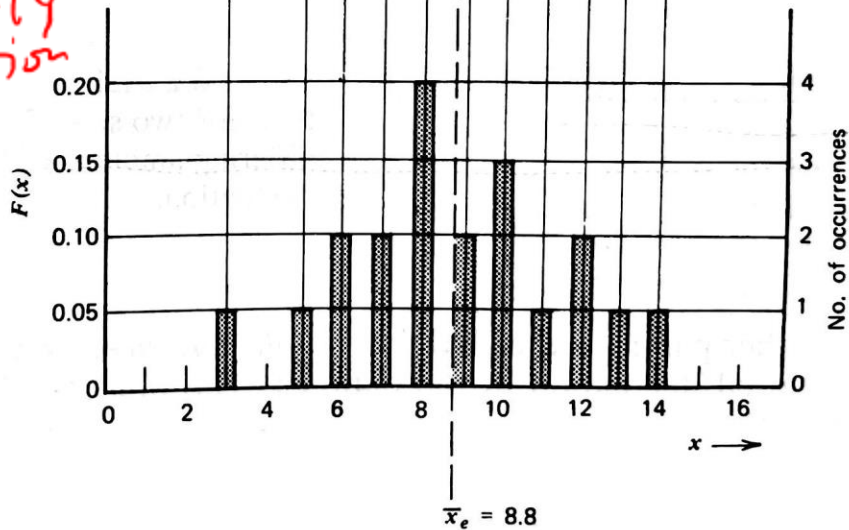


Determining the precision

Dataset



probability distribution



Scatter of data around mean is a measure for *randomness*

Residual

$$d_i \equiv x_i - \bar{x}_e$$

Variance

$$\sigma^2 \equiv \frac{1}{N-1} \sum_{i=1}^N (x_i - \bar{x}_e)^2$$

$$\bar{x}_e = \frac{\sum_{i=1}^N x_i}{N}$$

σ is the **standard deviation** and quantifies the amount of fluctuation in the data

Statistical models that describe the distribution of errors

- Can we predict statistical behavior under certain assumptions?

Consider a *binary process* with only a *true* or *false* result

Table 3.2 Examples of Binary Processes

Trial	Definition of Success	Probability of Success $\equiv p$	$P(x) = \frac{\text{no. occurrences of } x}{\text{no. measurements } (= N)}$
Tossing a coin	Heads	1/2	
Rolling a die	A six	1/6	
Observing a given radioactive nucleus for a time t	The nucleus decays during the observation	$1 - e^{-\lambda t}$	

- Common statistical models

Binomial Distribution

General model for *binary processes* with constant probability p for a certain outcome

Cumbersome to use with a large sample size (n)

Probability of getting x number of a certain outcome

$$P(x) = \frac{n!}{(n-x)!x!} p^x (1-p)^{n-x}$$

of permutations
Joint probability
of having success
 x times in a
sequence

Mean

$$\bar{x} = \sum_{x=0}^n x P(x) = pn$$

Variance

$$\sigma^2 \equiv \sum_{x=0}^n (x - \bar{x})^2 P(x) = np(1-p) = \bar{x}(1-p)$$

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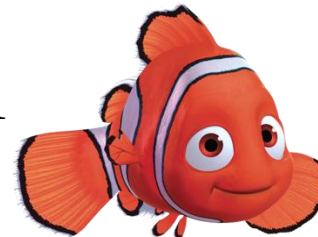
Poisson Distribution

Simplification for large n (less computationally intensive); assumes $p \ll 1$

$$P(x) = \frac{(pn)^x e^{-pn}}{x!} = \frac{(\bar{x})^x e^{-\bar{x}}}{x!}$$

$$\bar{x} = pn = \sigma^2$$

mean = variance



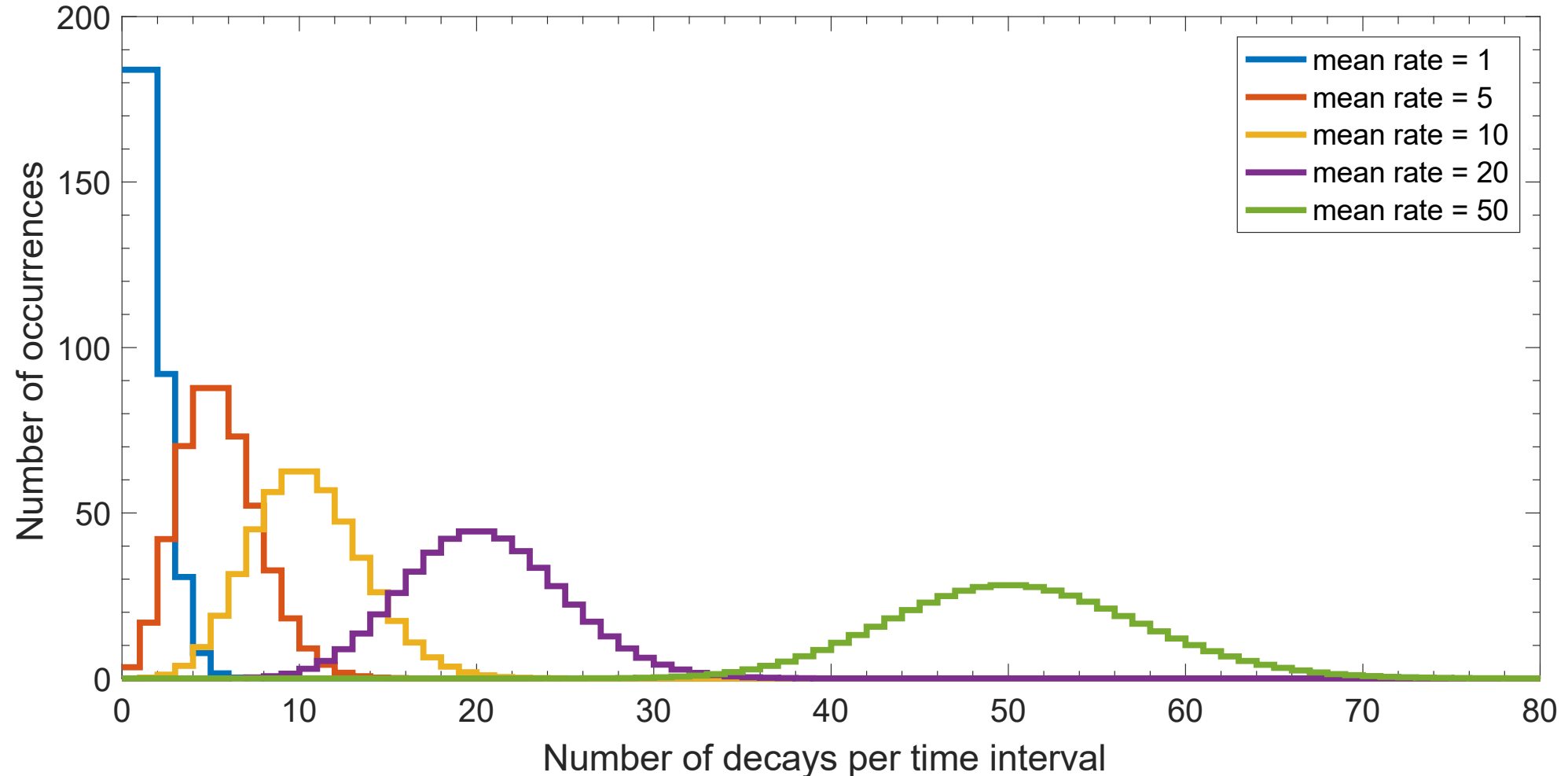
Example of Poisson distribution: radioactive decay

also fluorescence decay for many independent emitters

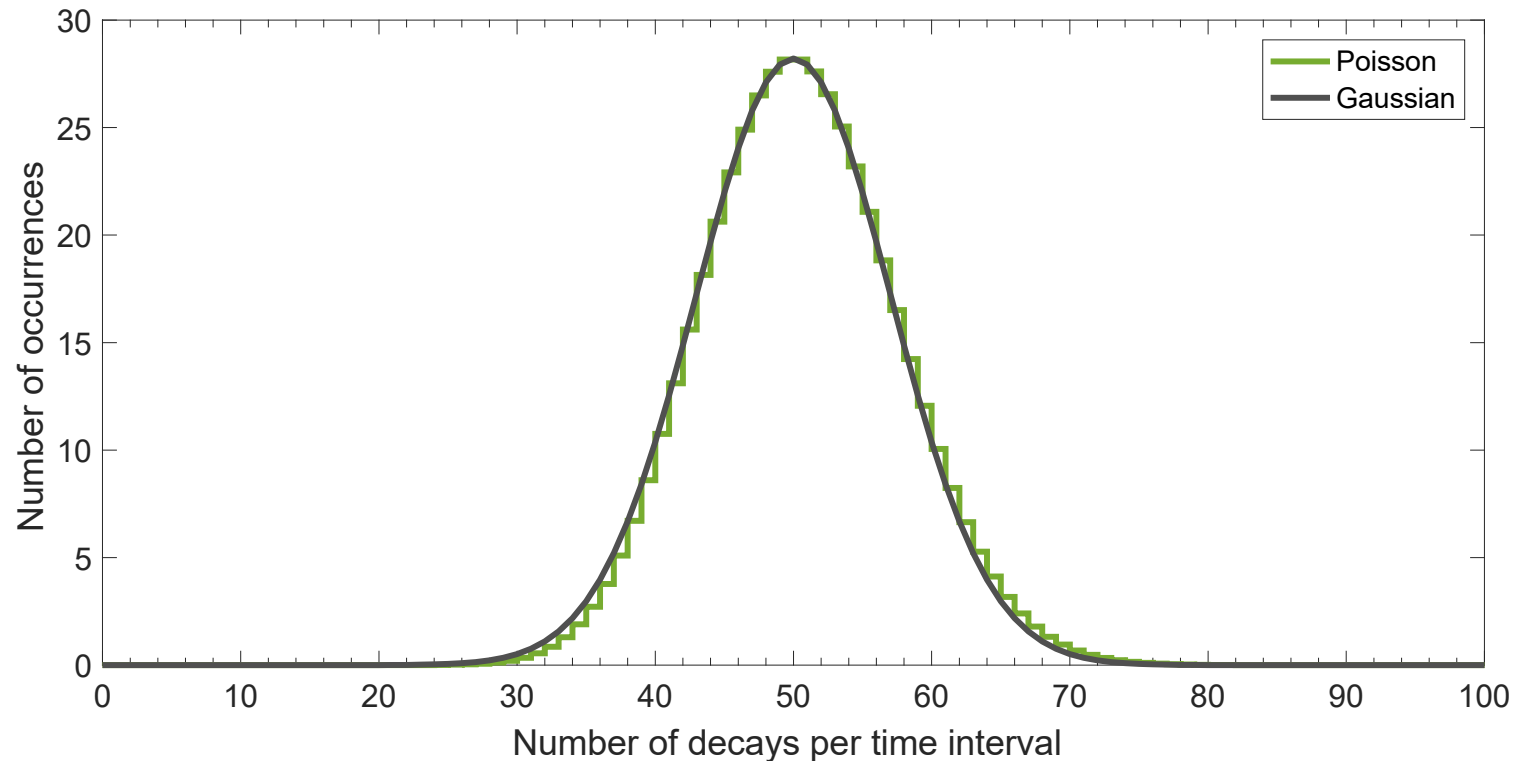
Probability of getting n decays in time t :

$$P(n, t) = \frac{(\lambda t)^n e^{-\lambda t}}{n!}$$

λ is the decay constant



Large mean \rightarrow distribution looks like Gaussian



Binomial Distribution

General model for *binary processes* with constant probability p

Cumbersome to use with a large sample size (n)



Poisson Distribution

Simplification for large n (less computationally intensive); assumes $p \ll 1$



Normal Distribution

Valid when $np \gg 1$

Also known as *Gaussian Distribution*

$$\int P(x) dx = 1$$

- **Probability function:** $P(x) = \frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{(x-\bar{x})^2}{2\sigma^2}}$
- **Mean:** $\bar{x} = pn$
- **Variance:** $\sigma^2 = \bar{x}$ for *Poisson statistics*
- We can use **discrete** or **continuous** versions

Statistics for photons

Discrete

Normalization: $\sum_{x=0}^{\infty} P(x) = 1$

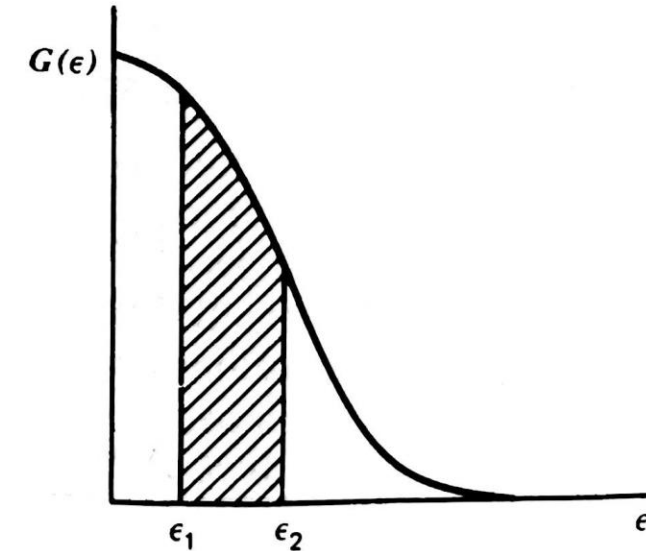
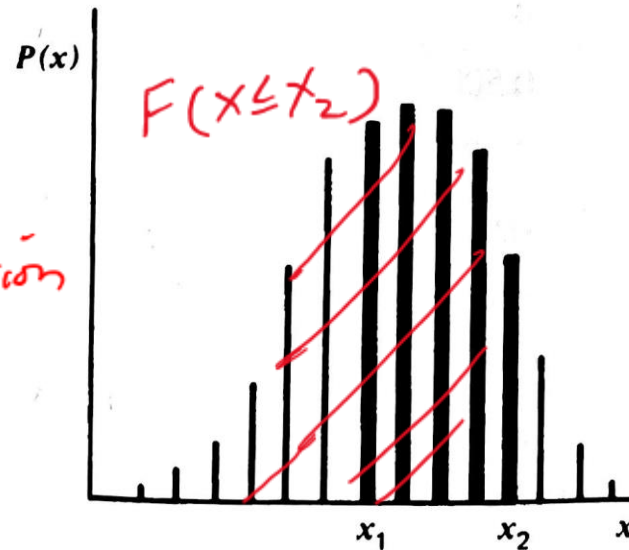
$\sum_{x=x_1}^{x_2} P(x) =$ Probability of observing a value of x between x_1 and x_2

Continuous

$\int_{\epsilon=0}^{\infty} G(\epsilon) d\epsilon = 1$

$\int_{\epsilon_1}^{\epsilon_2} G(\epsilon) d\epsilon =$ Probability of observing a value of ϵ between ϵ_1 and ϵ_2

Cumulative distribution function
 $F(x \leq x_2)$



Why are Gaussian distributions important in experiments?

Central limit theorem

Let X_1, X_2, \dots, X_n be a set of independent variables with the same distribution, with mean μ and standard deviation σ . For a large number of n , the mean of X ($\equiv \bar{X}_n$) is approximately a normal distribution with the same mean as X but variance σ^2/n :

$$\bar{X}_n \approx \mathcal{N}\left(\mu, \frac{\sigma^2}{n}\right)$$

$$P(x) = \frac{1}{\sqrt{2\pi}} \frac{1}{\sigma} \exp\left(-\frac{(x-\mu)^2}{2\sigma^2}\right)$$
$$\sigma = \sqrt{\frac{p(1-p)}{n}} \quad \text{projection noise}$$

<https://mathworld.wolfram.com/CentralLimitTheorem.html>

Example from political polling:

A two-candidate race has true support for Candidate A of $p = 0.48$ (and thus 0.52 for B). A polling firm takes a simple random sample of likely voters and reports the sample proportion \hat{p} favoring A. With $n = 500$, approximate the probability that the poll shows A *leading by more than 2 percentage points* (i.e., $\hat{p} > 0.51$).

\hat{p} will have a normal distribution with mean $p = 0.48$ and variance

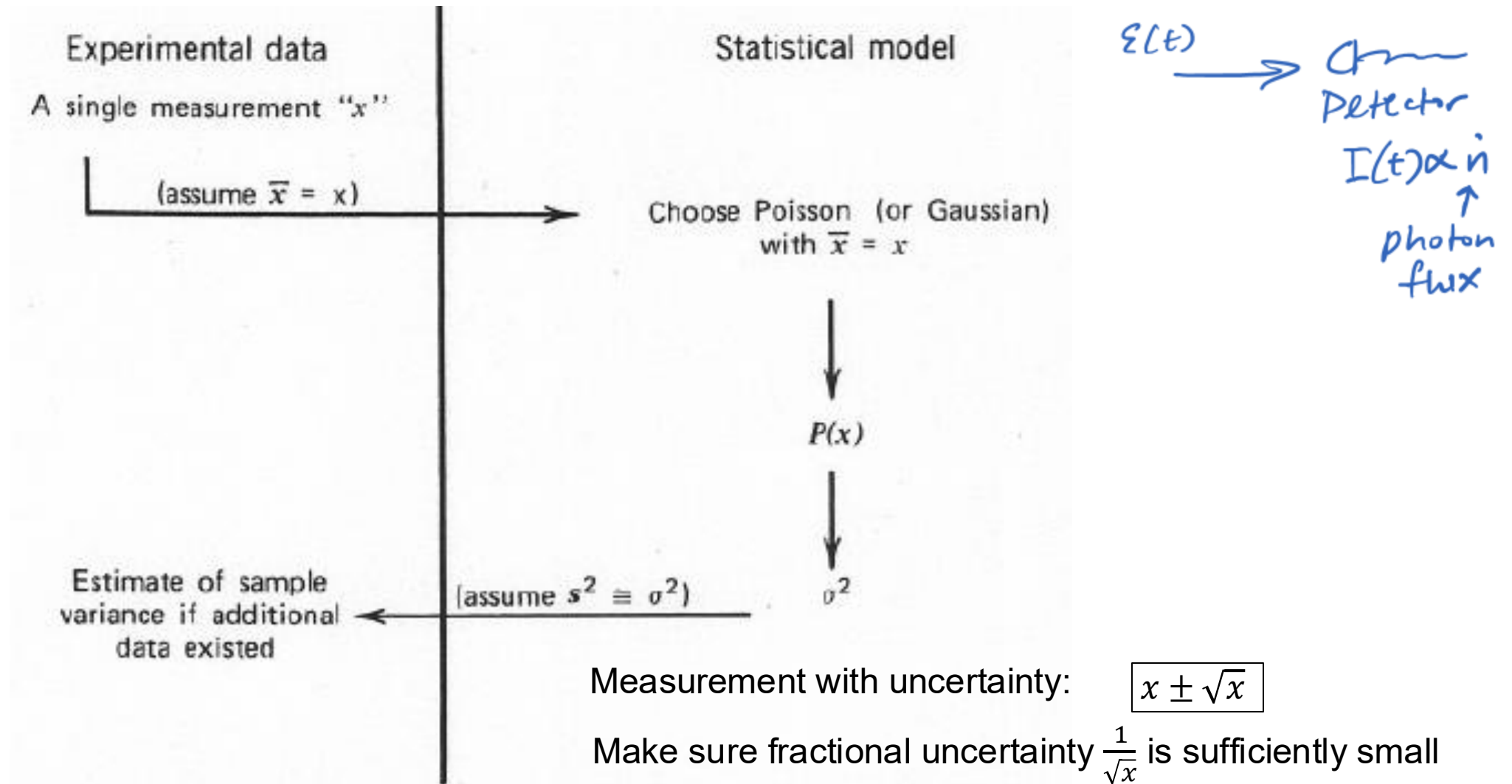
$$\frac{p(1-p)}{n} \rightarrow \sigma = \sqrt{\frac{0.48(1-0.48)}{500}}$$
$$= 0.022$$

$$P(\hat{p} > 0.51) = 1 - P(\hat{p} \leq 0.51)$$
$$= 8.9\%$$

See `sim_binomial_CLT.m`

Ascertaining uncertainty in a measurement

Single measurement of a random event (e.g., photon count rate on a detector)



Error propagation: how to estimate errors when we combine two or more measured quantities?

Gaussian error propagation law

Let $u(x,y,z)$ be a statistical quantity with x , y , and z as statistically independent variables. Each variable has the measurement uncertainty σ_i .

Then the combined uncertainty of σ_u of $u(x,y,z)$ is

$$\sigma_u^2 = \left(\frac{\partial u}{\partial x}\right)^2 \sigma_x^2 + \left(\frac{\partial u}{\partial y}\right)^2 \sigma_y^2 + \left(\frac{\partial u}{\partial z}\right)^2 \sigma_z^2 + \dots$$

$$u = u(x,y,z,\dots)$$

Error propagation: sums and differences

Simple sum or difference with a constant:

$$u = x \pm a$$

$$\frac{\partial u}{\partial x} = 1$$

$$\sigma_u = \sigma_x$$

Involving two or more independent variables:

$$u = x \pm y$$

$$\frac{\partial u}{\partial x} = 1, \frac{\partial u}{\partial y} = \pm 1$$

$$\sigma_u^2 = (1)^2 \sigma_x^2 + (\pm 1)^2 \sigma_y^2$$

$$\sigma_u = \sqrt{\sigma_x^2 + \sigma_y^2}$$

Weighted:

$$u = ax \pm by$$

$$\frac{\partial u}{\partial x} = a, \frac{\partial u}{\partial y} = \pm b$$

$$\sigma_u^2 = a^2 \sigma_x^2 + b^2 \sigma_y^2 + 2ab\sigma_{xy}^2$$

0 if x and y are
uncorrelated

Error propagation rules for other operations

Error propagation equation:

For $u = u(x, y, z, \dots)$, where all variables are uncorrelated and errors are small

$$\sigma_u^2 \approx \left(\frac{\partial u}{\partial x}\right)^2 \sigma_x^2 + \left(\frac{\partial u}{\partial y}\right)^2 \sigma_y^2 + \left(\frac{\partial u}{\partial z}\right)^2 \sigma_z^2 + \dots$$

Sums and Differences:

$$u = ax \pm by$$

$$\sigma_u = \sqrt{a^2 \sigma_x^2 + b^2 \sigma_y^2}$$

Fractional
uncertainty

Multiplication or division:

$$u = axy \text{ or } u = \frac{ax}{y}$$

$$\frac{\sigma_u}{u} = \sqrt{\left(\frac{\sigma_x}{x}\right)^2 + \left(\frac{\sigma_y}{y}\right)^2}$$

Powers:

$$u = ax^b$$

$$\frac{\sigma_u}{u} = b \frac{\sigma_x}{x}$$

Exponentials:

$$u = ae^{bx}$$

$$\frac{\sigma_u}{u} = b \sigma_x$$

(a, b are constants)

Mean of measurements

- To calculate the estimated mean in a measurement:

$$\bar{x}_{est} = \frac{\sum_{i=1}^N x_i}{N}$$

Each measurement has an uncertainty of σ_x .

- What is the error with N measurements?

$$\begin{aligned}\sigma_{\bar{x}_{est}} &= \frac{\sqrt{N} \sigma_x}{N} \\ &= \frac{\sigma_x}{\sqrt{N}}\end{aligned}$$

$$\begin{aligned}x_{sum} &= x_1 + x_2 + \dots \\ \sigma_{x_{sum}} &= \sqrt{\underbrace{\sigma_x^2 + \dots + \sigma_x^2}_N} = \sqrt{N \sigma_x^2}\end{aligned}$$

- How many times (in terms of N) do we need to do the measurement to reduce the error by half?

Do measurements $4 \times N$

Another approach: Monte Carlo error analysis

repeated random sampling

- Concept: For $u = u(x, y, z, \dots)$, randomly sample independent variables based on your knowledge of their mean, standard deviation, and distribution (e.g., Gaussian) and calculate u with each set of random values. The standard deviation in the computed u is then the estimated uncertainty.
- Need random number generator and a large number of samples. To be computationally efficient, should avoid using for loops.

See code simpleMC.m on Canvas

Example

We want to determine the resistance R of a cylindrical conductor based on the resistivity of the material ρ , the length of the cylinder L , and the radius r of the cylinder cross-section. The relationship between the resistance and cylinder properties is:

$$R = \frac{\rho L}{\pi r^2}$$

A cylinder made out of an alloy material has the following properties:

$$\rho = (4.41 \pm 0.03) \times 10^{-5} \Omega \cdot m$$

$$L = 0.1 \pm 0.0005 m$$

$$r = (20 \pm 0.01) \times 10^{-6} m$$

- Calculate R and its uncertainty using the standard error propagation approach.

$$R = \frac{\rho L}{y} \quad y = \pi r^2 \quad \frac{\sigma_y}{y} = \frac{2\pi r \sigma_r}{\pi r^2} = \frac{2\sigma_r}{r}$$
$$\left(\frac{\sigma_R}{R}\right)^2 = \left(\frac{\sigma_\rho}{\rho}\right)^2 + \left(\frac{\sigma_L}{L}\right)^2 + \underbrace{\left(\frac{\sigma_y}{y}\right)^2}_{4\left(\frac{\sigma_r}{r}\right)^2} = 29.8 \Omega$$

$$\boxed{R = 3509 \pm 29.8 \Omega}$$

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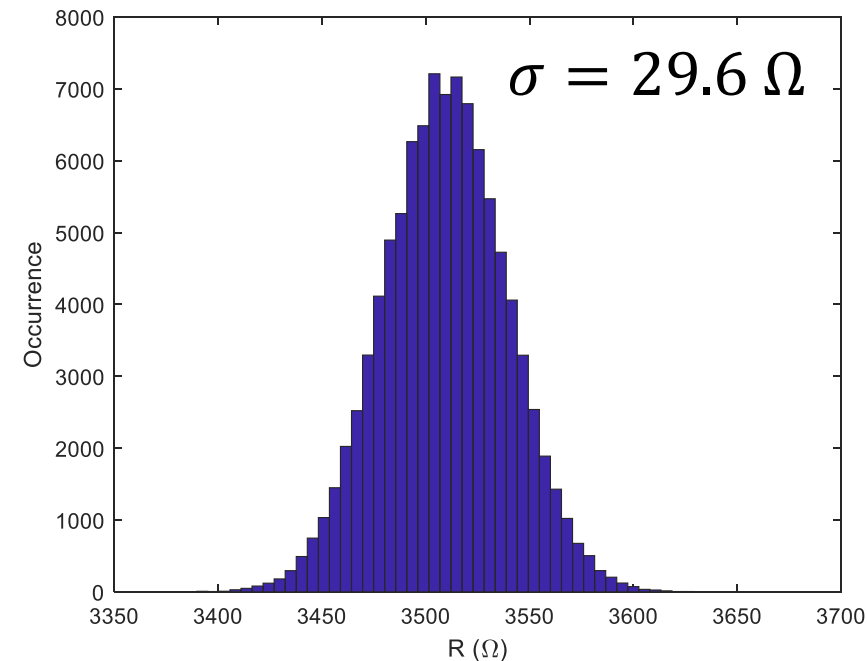
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$$r = (20 \pm 0.01) \times 10^{-6} m$$

- Calculate the uncertainty in R using the Monte Carlo approach.

See code `errPropExample_mc.m` on Canvas



Example

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Now consider the case of much larger errors for the measurements:

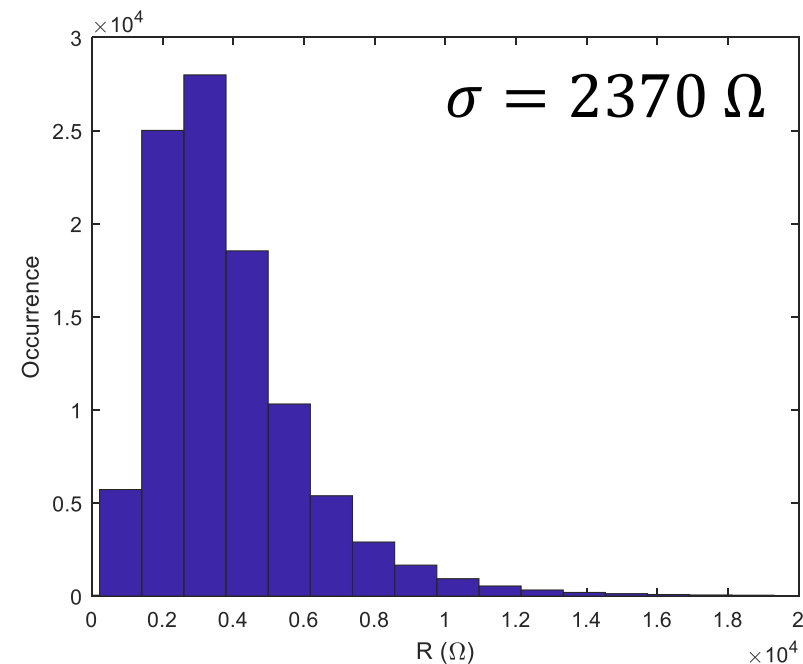
$$\rho = (4.41 \pm 1.2) \times 10^{-5} \Omega \cdot m$$

$$L = 0.1 \pm 0.025 m$$

$$r = (20 \pm 3.5) \times 10^{-6} m$$

Gaussian error propagation:

$$\sigma = 1786 \Omega$$

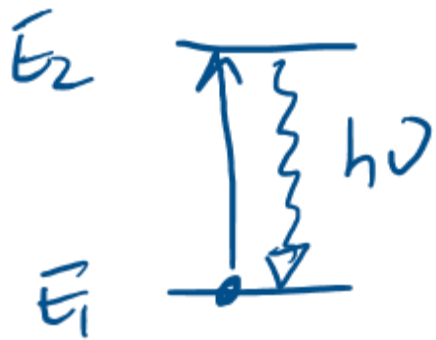


What are the fundamental limits to uncertainties in measurements?

- Uncertainty principle
- Statistical fluctuations due to probabilistic nature of quantum states (projection noise)
- Fluctuations from discreteness of light and matter (shot noise)
- Vacuum fluctuations

Bandwidth of a single photon

A photon is emitted by an atom at a decay rate of 1 ns. What is the bandwidth (frequency uncertainty) of the photon?



$$\underbrace{\Delta E}_{h\Delta\nu} \underbrace{\Delta t}_{1\text{ ns}} \geq \frac{\hbar}{2} = \frac{h}{4\pi}$$

$$\Delta\nu \Delta t \geq \frac{1}{4\pi}$$

$$\Delta\nu_{\min} \sim 8 \times 10^7 \text{ Hz} \sim 80 \text{ MHz}$$

Trading off position-momentum uncertainties to improve a measurement of electron energy (simple example of squeezing)

An electron is prepared such that its position is known to $\Delta x \approx 10^{-12}$ m.

- Estimate the uncertainty in the kinetic energy of the electron (ΔE_{kin}). Assume that the momentum uncertainty is on the order of the mean momentum of the electron.
- Say the electron position uncertainty can be changed by a factor of 100 to improve the energy uncertainty. Which way would you change the position uncertainty? What is the improved energy uncertainty?

$$\Delta x \Delta p \geq \frac{\hbar}{2} \rightarrow \Delta p \geq \frac{\hbar}{2\Delta x} \sim 5.3 \times 10^{-23} \text{ kg m/s}$$

$$\Delta E_{kin} = \Delta \left(\frac{p^2}{2m_e} \right) \approx \frac{\Delta p \cdot p}{m_e} \sim 1.5 \times 10^{-15} \text{ J}$$

Increase uncertainty in x : $\Delta x' \sim 10^{-10} \text{ m}$

Squeezing factor in x :

$$S_x = \frac{\Delta x'}{\Delta x} = 100$$

Squeezing factor in p

$$S_y = \frac{\Delta p'}{p} = \frac{1}{100}$$

ΔE_{kin} decreases

by factor of
 $10^2 - 10^4$

Quantum projection noise (QPN): measurement uncertainty from statistics

An electron has equal probability of being found in the left (L) or right (R) side of a region. What is the variance in the measurement of the electron position?

Electron wavefunction in which there is equal likelihood ($P = 1/2$) for the electron to be on the left and right region:

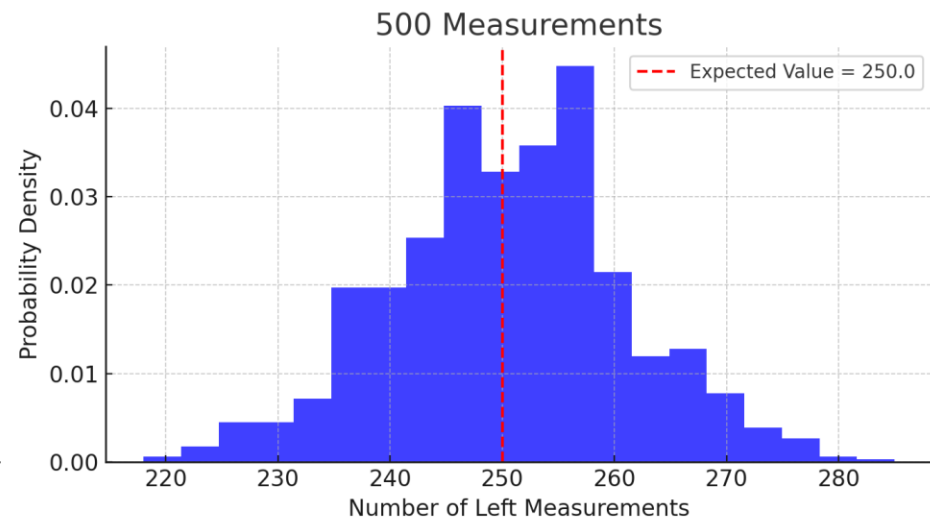
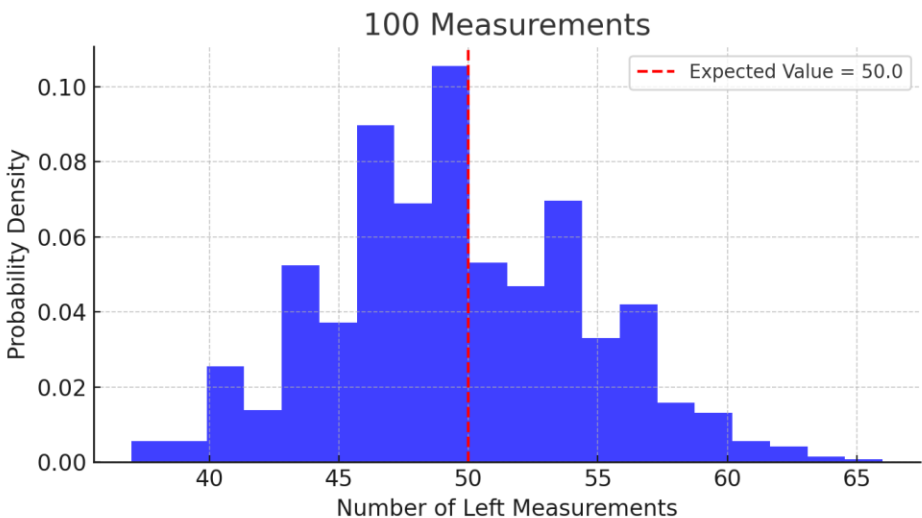
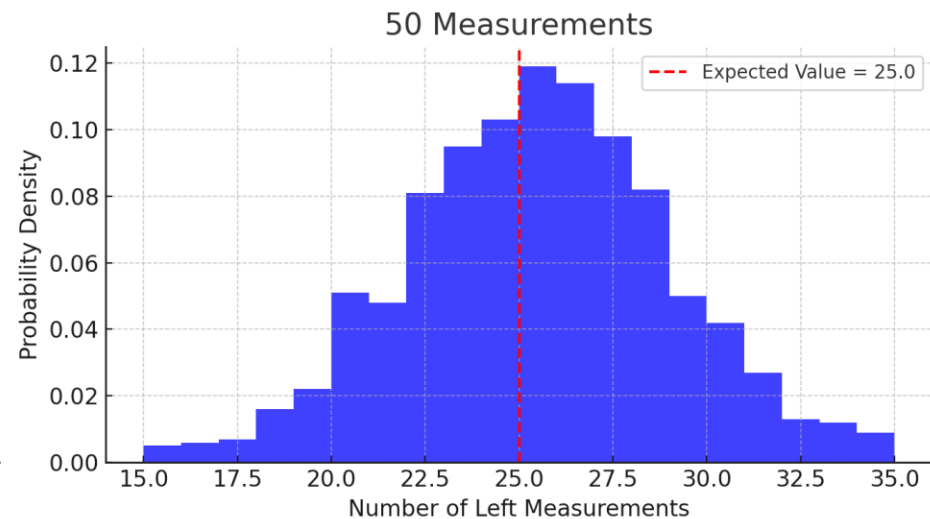
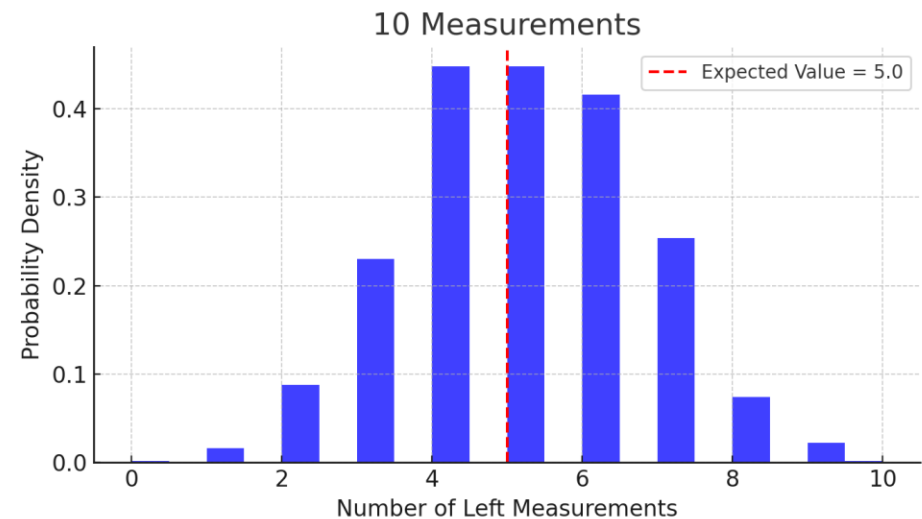
$$\psi(x) = \frac{1}{\sqrt{2}} (\psi_L(x) + \psi_R(x))$$

Variance of finding the electron on the left side after N measurements:

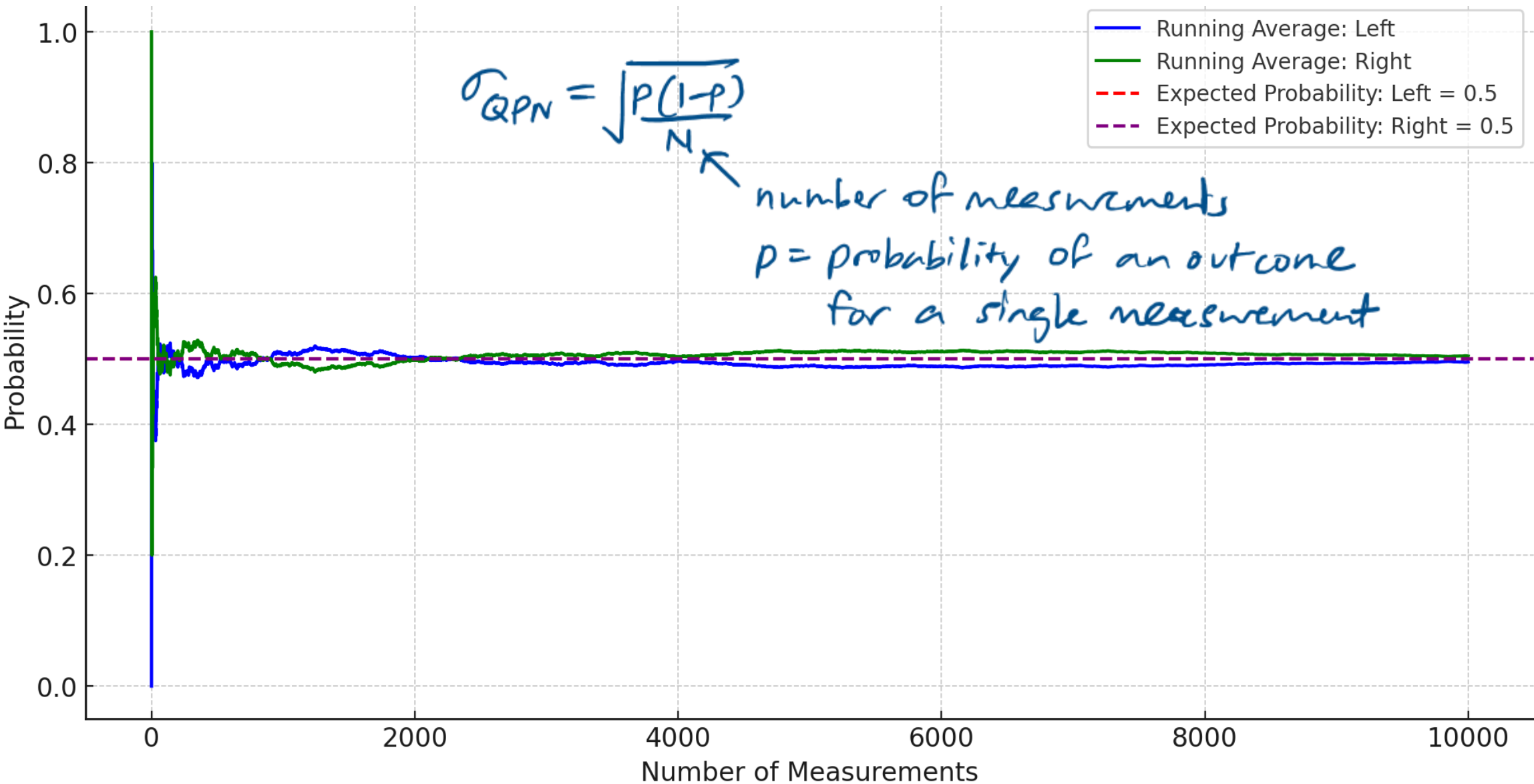
$$\text{Var}_L = \sigma_L^2 = \frac{1}{4N}$$

Same for the measuring the electron on the right:

$$\text{Var}_R = \sigma_R^2 = \frac{1}{4N}$$



Quantum projection noise (QPN): measurement uncertainty from statistics



Another example of QPN: observing electron diffraction

- Probability distribution of the electron's position on the screen

$$|\psi_s(x)|^2 \propto \frac{1}{2} \left(1 + \cos \left(\frac{2\pi s x}{\lambda z_0} \right) \right)$$

Handwritten annotations:
 - s : slit width
 - z_0 : source-detector distance
 - λ : de Broglie wavelength

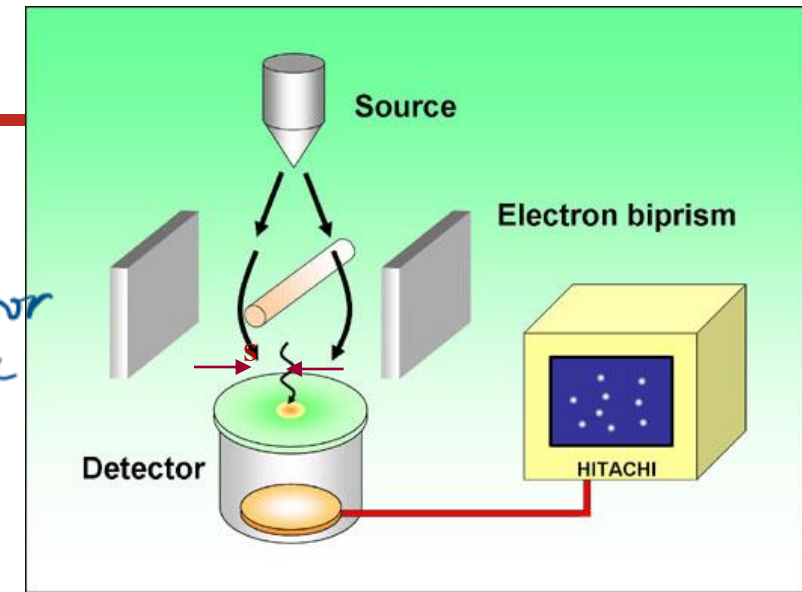
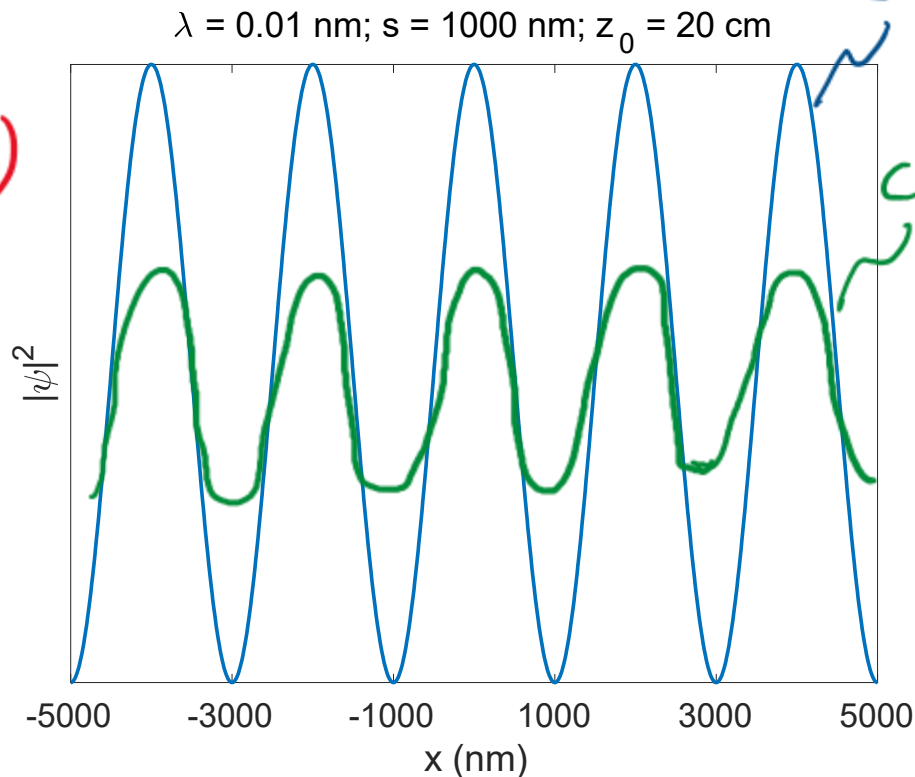
- What are relevant length scales here?

- $\lambda \sim 0.005 - 1 \text{ nm}$
- $s \sim 1 \text{ micron}$
- $z_0 \sim 10 \text{ cm}$

$$p(\phi) = \frac{1}{2} (1 + C \cos \phi)$$

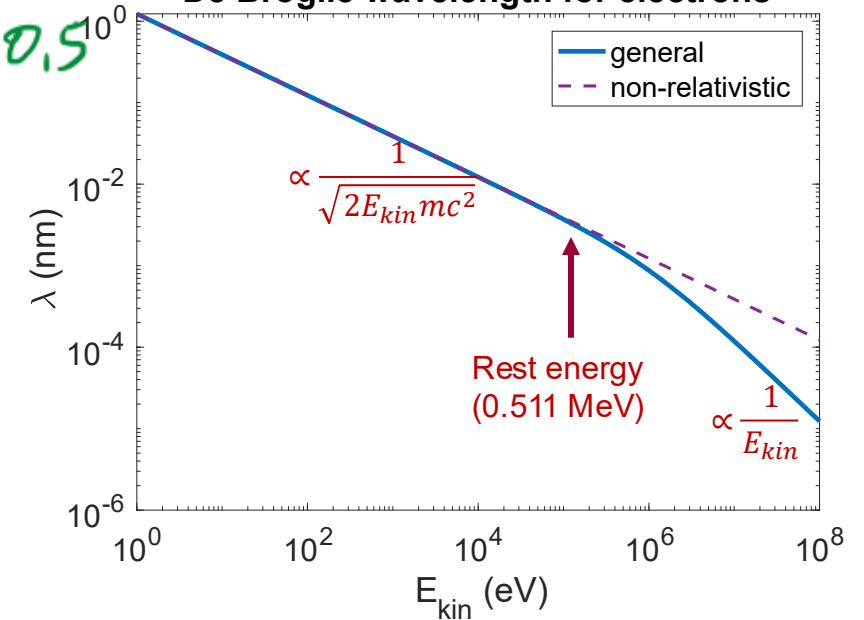
C is interferometer contrast
 $[0, 1]$

$$\phi = \frac{2\pi \cdot s}{\lambda \cdot z_0} x$$

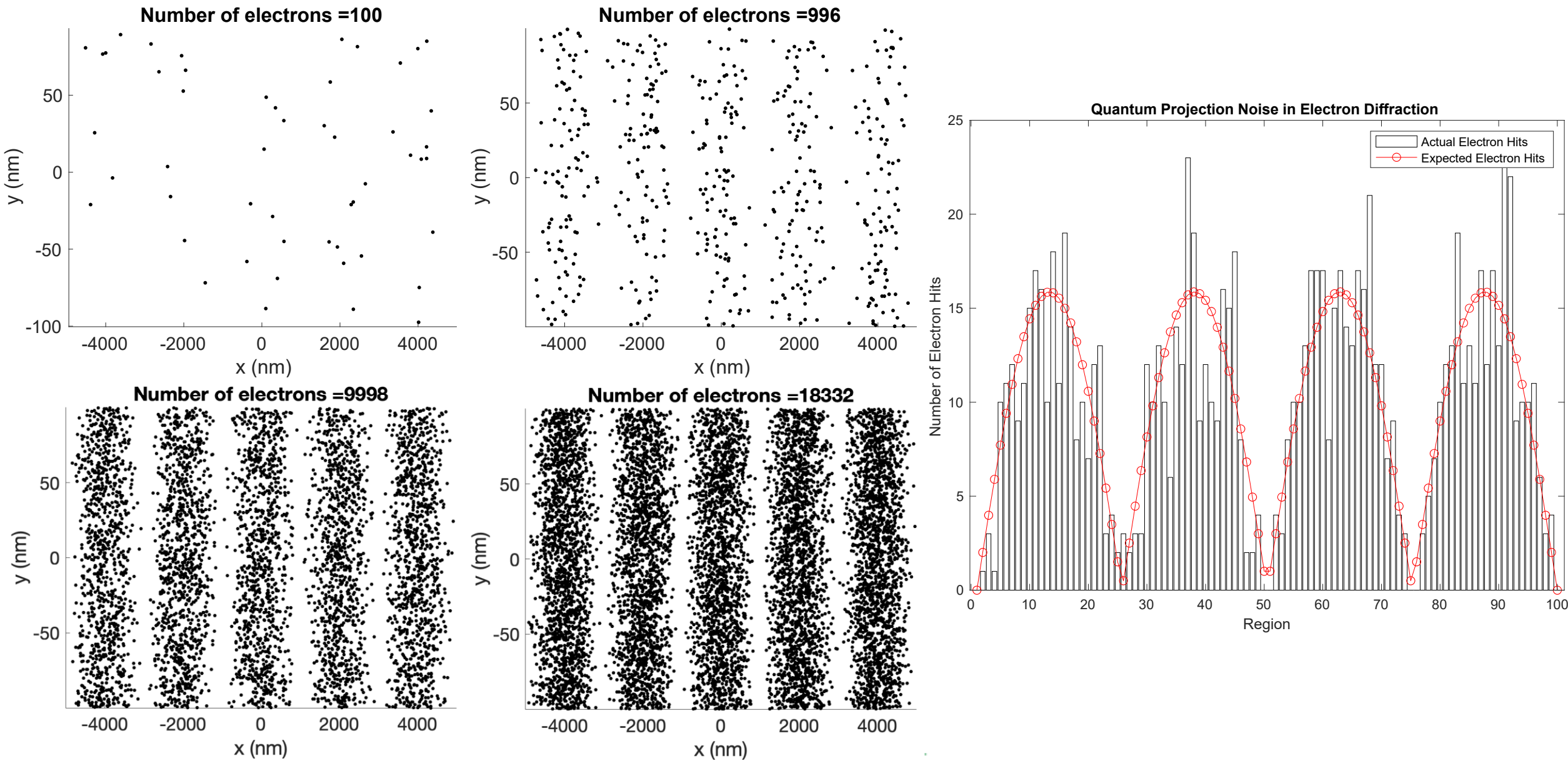


<https://www.hitachi.com/rd/research/materials/quantum/doubleslit/index.html>

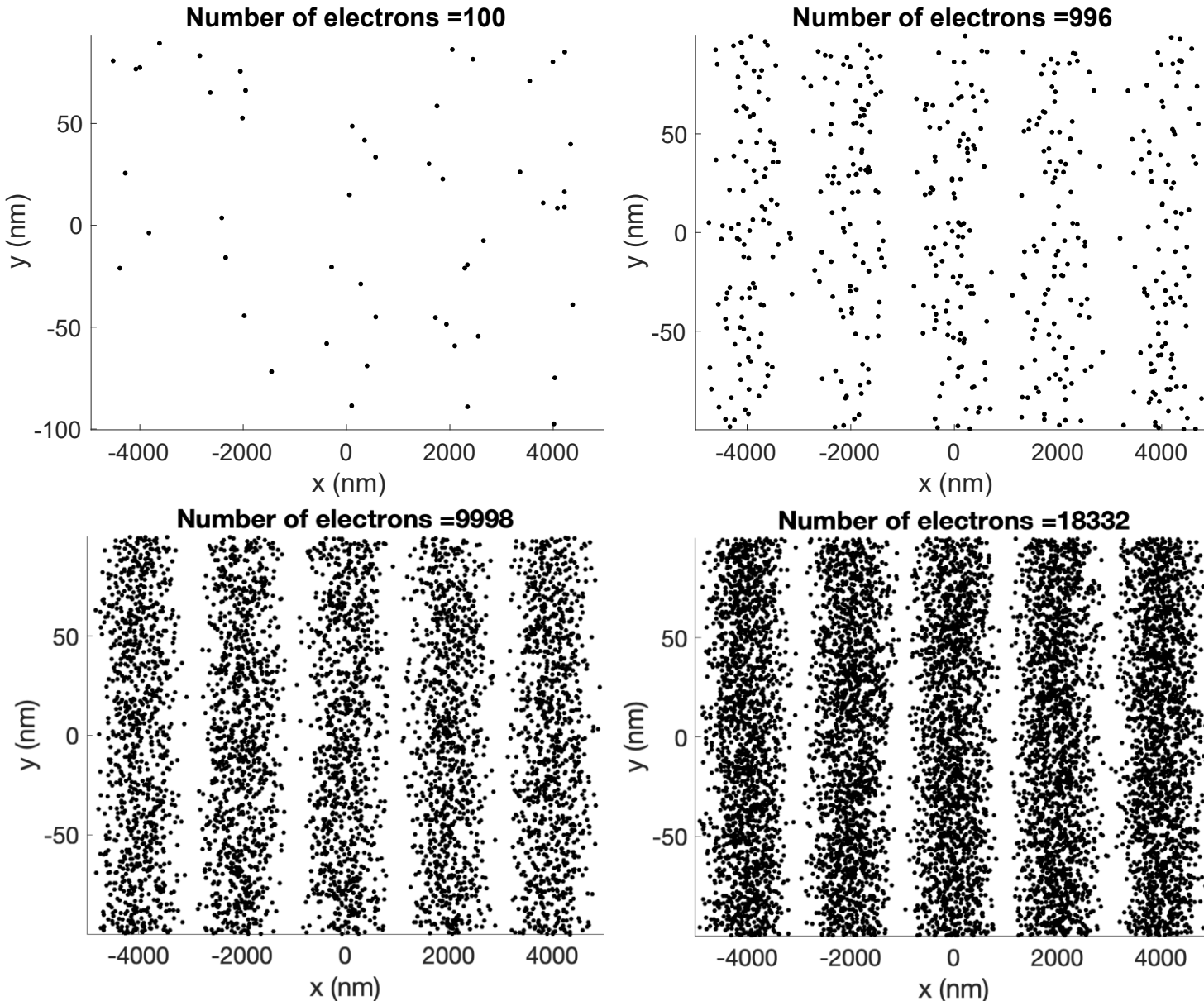
De Broglie wavelength for electrons



Another example of QPN: observing electron diffraction



Another example of QPN: observing electron diffraction



Electron detection at 1 pixel
 Number of detected e^- : Np
 total number of detected e^- : $\frac{1}{2}(1 + C \cos \phi)$

Variance in detected e^- : $\sigma_{QPN} = \sqrt{\frac{p(1-p)}{N}}$

Phase sensitivity: smallest detectable phase

$$\Delta\phi_{\min} \approx \frac{\sigma_{QPN}}{\left| \frac{\partial p}{\partial \phi} \right|} \quad \frac{\partial p}{\partial \phi} = -\frac{C \sin \phi}{2}$$

Want: C close to 1
 $|\sin^2 \phi| = 1$; $p = \frac{1}{2}$

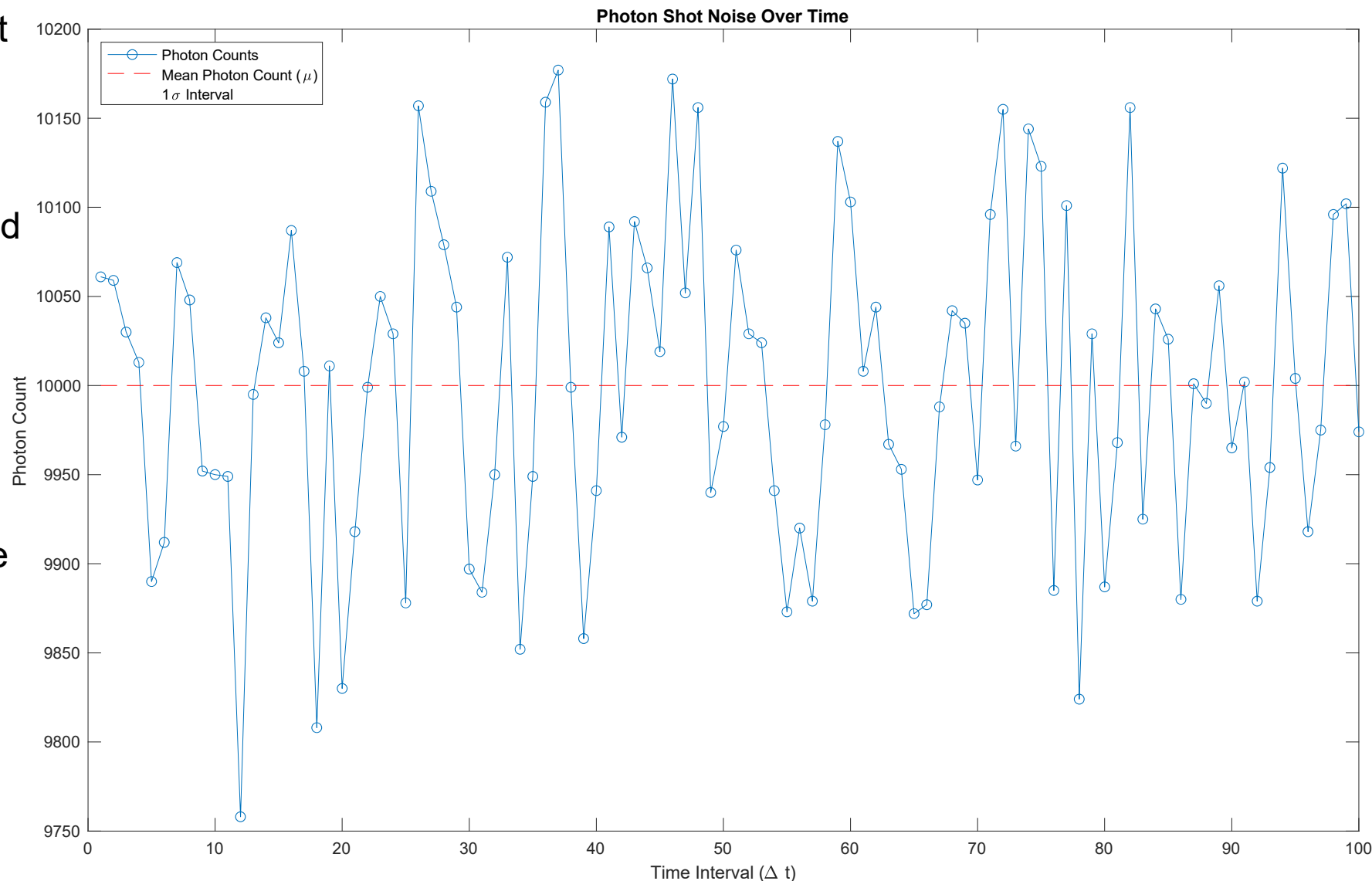
$$\Delta\phi_{\min} \approx \frac{1}{\sqrt{N}} \sqrt{\frac{1 - C^2 \cos^2 \phi}{C^2 \sin^2 \phi}}$$

Shot noise

- Coherent light (such as output of a laser) follows Poissonian photon statistics
- Discrete nature of photons and electrons leads to statistical fluctuations on short time-scales.
- Observed in the form of fluctuations in the number of particles observed over a time window

$$\sigma = \sqrt{\mu}$$

$$\frac{\sigma}{\mu} = \frac{1}{\sqrt{\mu}}$$



Shot noise numerical example

An attenuated beam from a 514-nm laser with a power of 0.1 pW is detected with a photon counter with efficiency 20%. The time interval of the counting system is set to 0.1 s. Calculate:

- Mean photon counts within a 0.1-s time window
- Standard deviation in the mean photon counts

$$1 \text{ pW} = 1 \times 10^{-12} \text{ W}$$
$$h = 6.626 \times 10^{-34} \text{ J} \cdot \text{s}$$

$$\text{Photon generation rate: } \dot{n} = \frac{10^{-13} \text{ W}}{hc/\lambda} = 2.6 \times 10^5 \text{ s}^{-1}$$

$$\text{Mean counts within measurement window: } \bar{N} = 0.2 \times (0.1 \text{ s}) \times (2.6 \times 10^5 \text{ s}^{-1}) = 5180$$

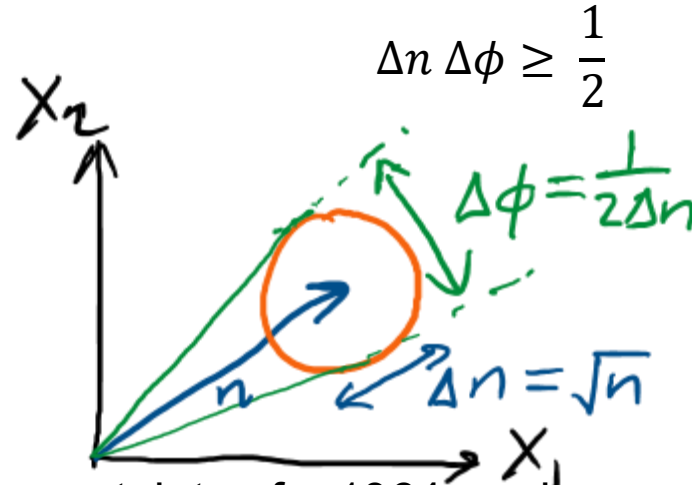
$$\text{Uncertainty in counts: } \sigma_{\bar{N}} = \sqrt{\bar{N}} = 72$$

Standard quantum limit

- A “coherent” electromagnetic field, say from a laser, can be represented by $\vec{\mathcal{E}} = \vec{\mathcal{E}}_0 e^{i(\vec{k} \cdot \vec{r} - \omega t)} = \vec{\mathcal{E}}_0 e^{i\phi(\vec{r}, t)}$

- The amplitude and phase of an EM field follow the uncertainty relationship:

- n is the photon number



phase

$$\sqrt{\frac{I}{c\epsilon_0}} \propto \sqrt{n}$$

- Example: What is the phase uncertainty of a 1064-nm laser with a power output of 300 W?

$$\dot{n} = \frac{300 \text{ W}}{hc/(1064 \text{ nm})} = 1.6 \times 10^{21} \text{ s}^{-1}$$

$$\Delta n = 4 \times 10^{10} = \sqrt{1.6 \times 10^{21}}$$

$$\Delta \phi = \frac{1}{2\Delta n} = 1.3 \times 10^{-12} \text{ rad}$$

in a second

$$\Delta \phi_{\text{SQL}} = \frac{1}{2\sqrt{n}} \propto \frac{1}{\sqrt{I}}$$

increase intensity to reduce phase noise

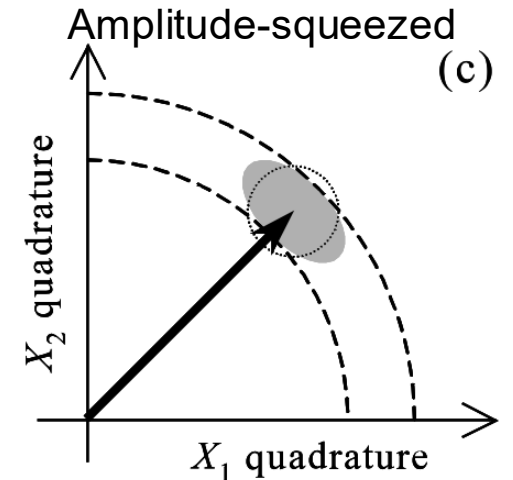
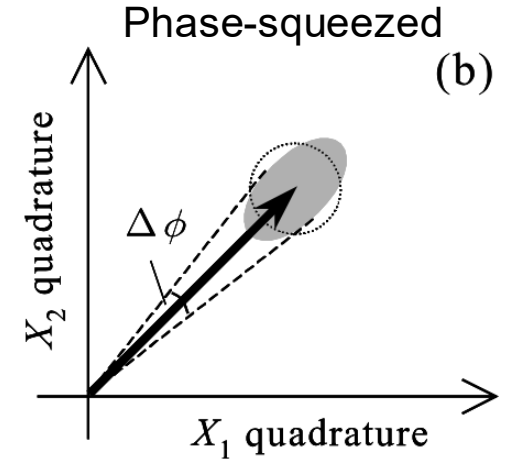
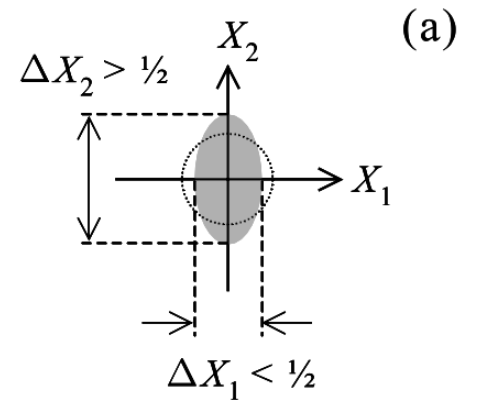
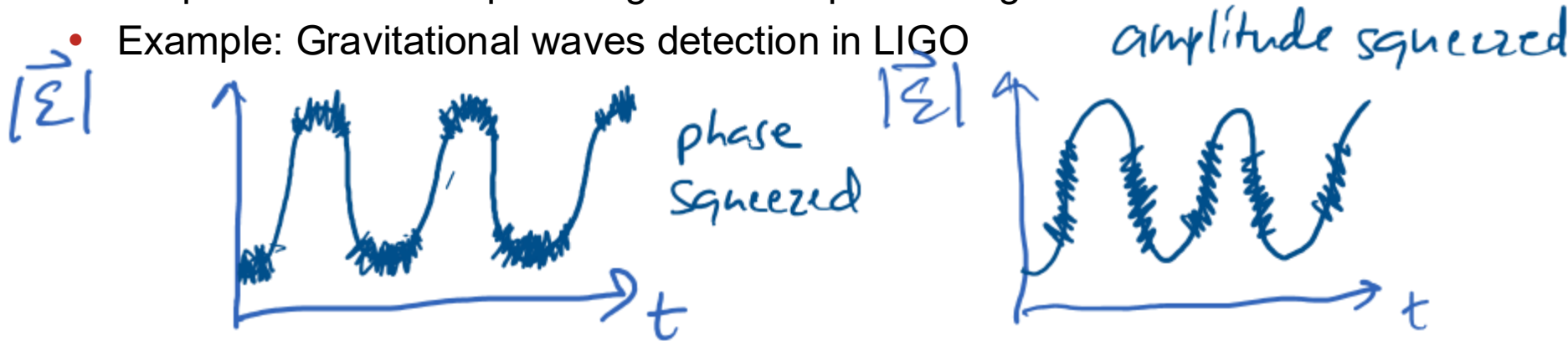
Heisenberg limit and squeezed light

- Can we beat the standard measurement limit?
 - What is the most uncertainty we can produce in n ?

$$\Delta n \approx n$$

$$\Delta \phi_{\text{Heisenberg limit}} = \frac{1}{2n} \propto \frac{1}{I}$$

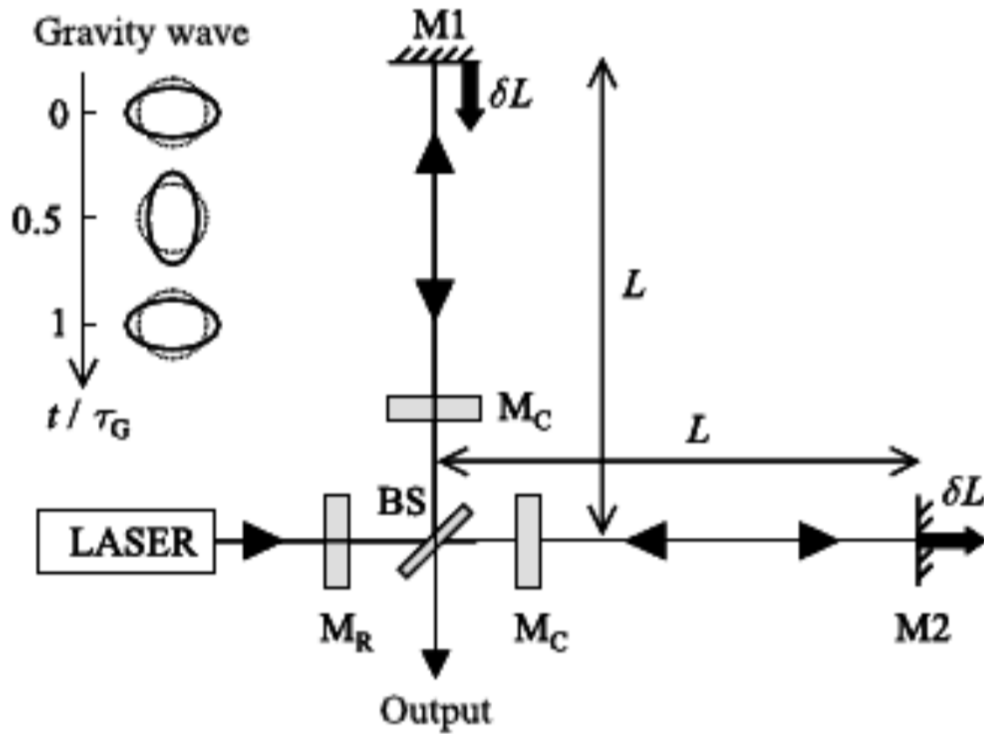
- Amplitude and phase noise can be traded off to improve precision in measurements
 - Requires nonlinear optics to generate squeezed light
 - Example: Gravitational waves detection in LIGO



Dowling, Jonathan P. "Quantum optical metrology—the lowdown on high-N00N states." *Contemporary physics* 49.2 (2008): 125-143. <https://arxiv.org/pdf/0904.0163>

Exam 1 covers up to this point

Laser Interferometer Gravitational-Wave Observatory (LIGO)



Hanford, Washington (H1)



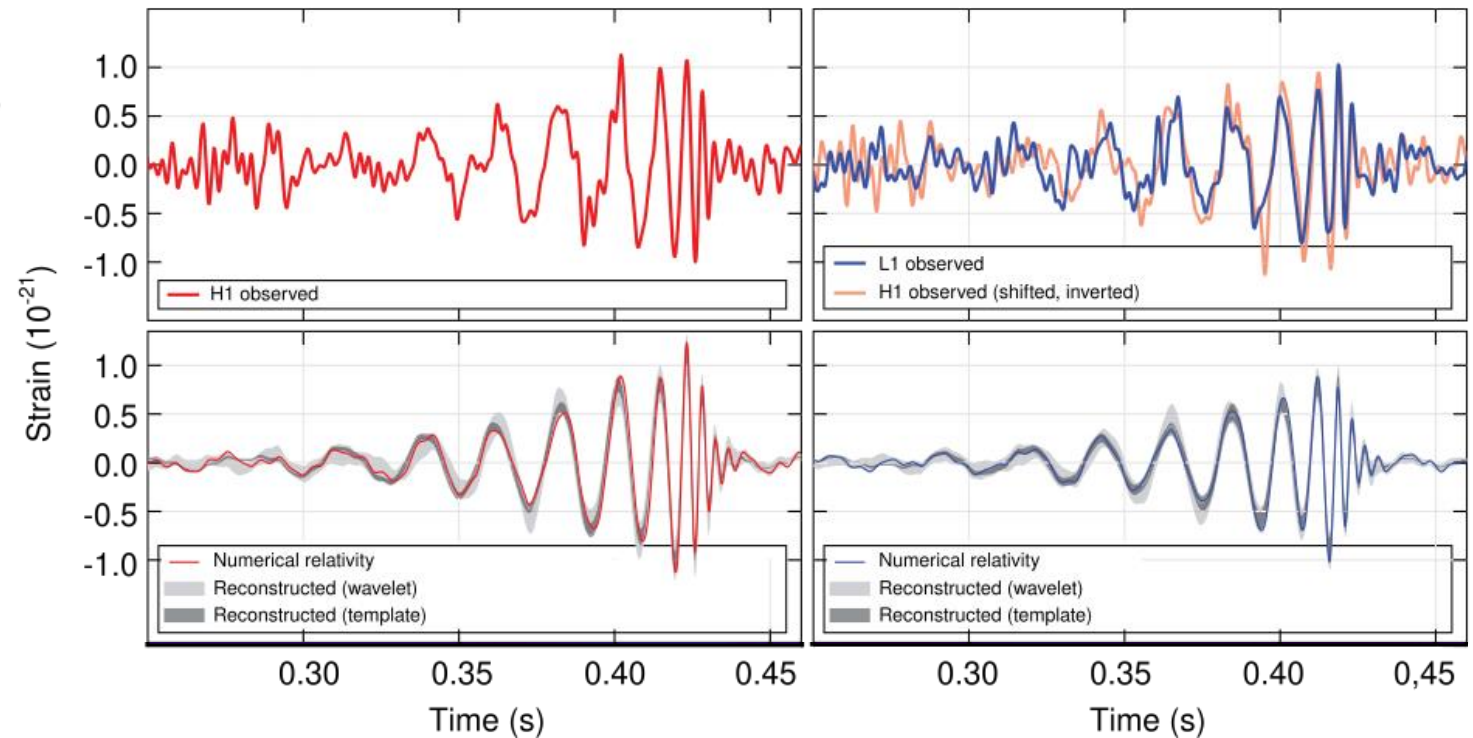
Livingston, Louisiana (L1)

Mirror displacement:

$$\frac{\delta L}{\lambda} > \frac{\delta \phi}{2\pi}$$

Strain:

$$h \approx \frac{\delta L}{L_{tot}}$$



Zero-point energy

- Explanation using Heisenberg uncertainty principle
 - Zero energy, corresponding to a particle sitting motionless, corresponds to precisely determined position and momentum → violates uncertainty principle
 - Distribution in momentum and position → particle energy must be greater than the minimum of the potential well

- Zero-point energy for harmonic oscillators (which approximate all potential wells near their minima):

$$E_0 = \frac{\hbar\omega}{2}, \text{ where } \omega \text{ is the angular frequency of the oscillation of the system}$$

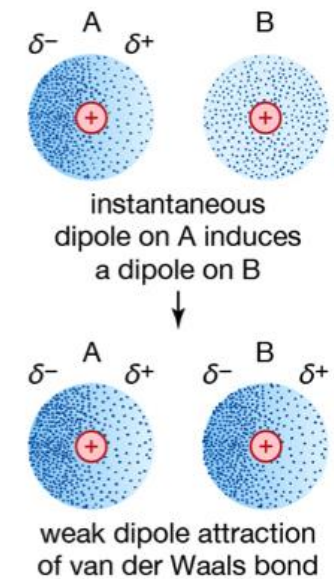
- Experimental observations:
 - Vacuum field in quantum optics
 - "Lamb shifts" in atomic spectra (later in the course)
 - Spontaneous emission (later in the course)

Connection of zero-point energy to vacuum field

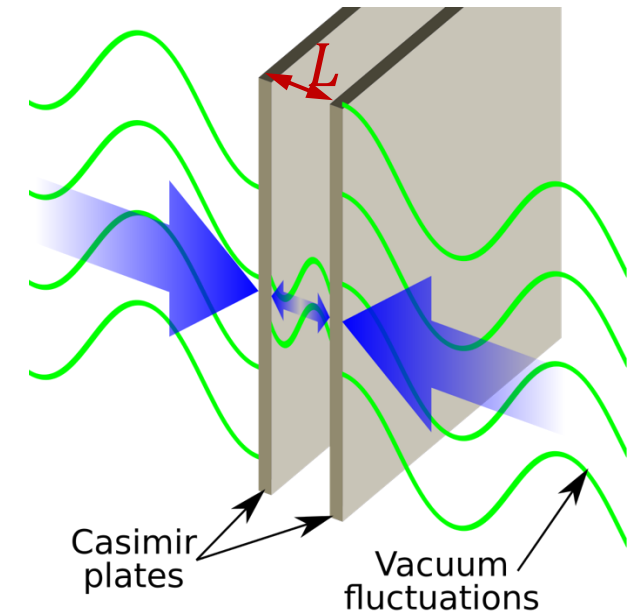
van der Waals and Casimir forces

- Zero-point energies of charges in objects (molecules, small particles, etc.) induce electromagnetic fluctuations that interact at small distances (few nm) → attractive van der Waals forces
 - Instantaneous interaction
 - At larger separations, interaction is non-instantaneous due to the finite speed of light (“Casimir force”)
- Casimir force between parallel conducting mirrors in vacuum of area A

$$F_{Casimir} = -\frac{\hbar c \pi^2}{240 L^4} A$$

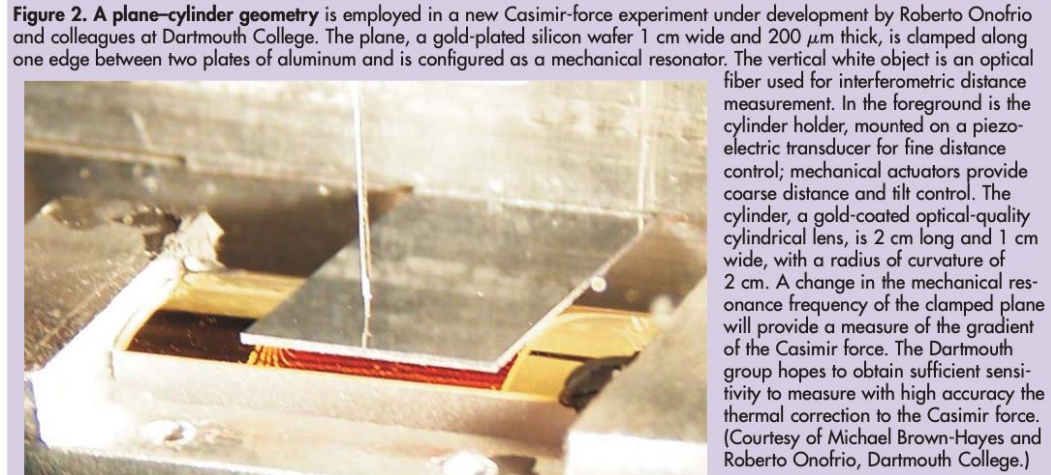


Source: Encyclopedia Britannica

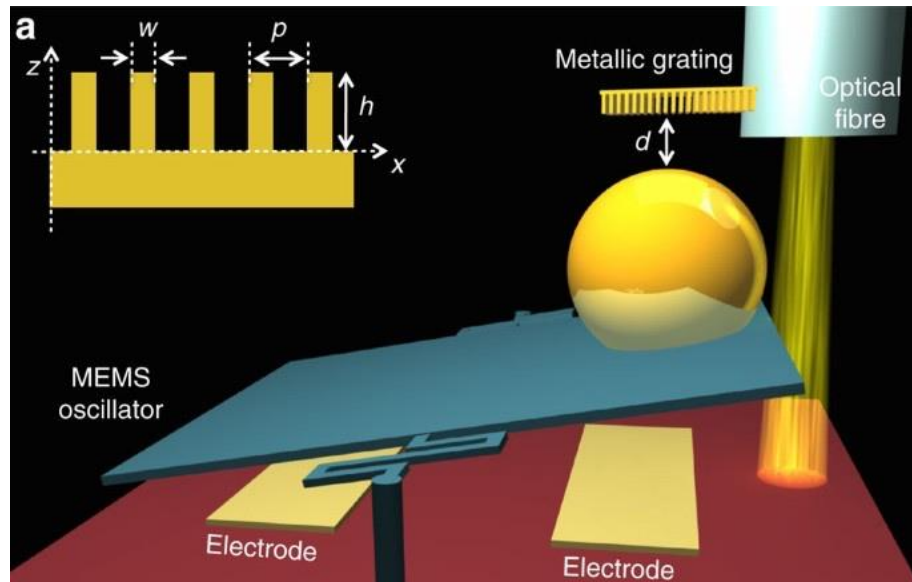


Source: Emok (Wikipedia), CC BY-SA 3.0

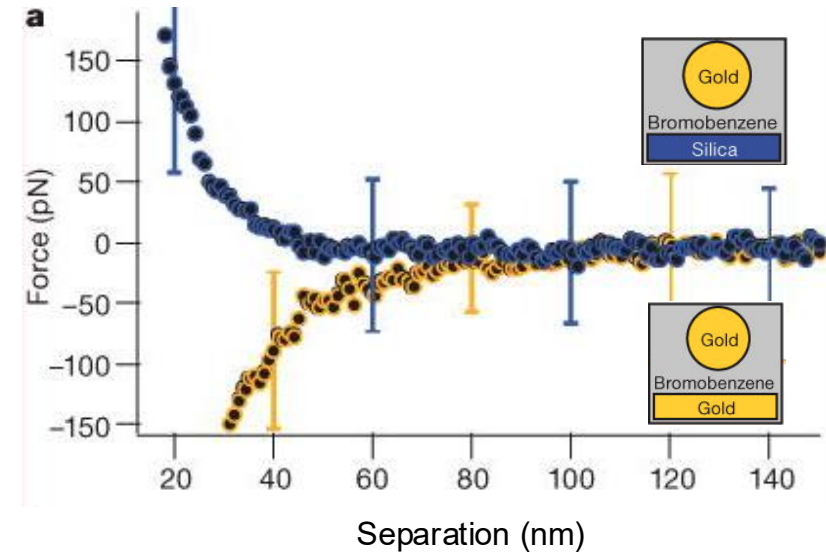
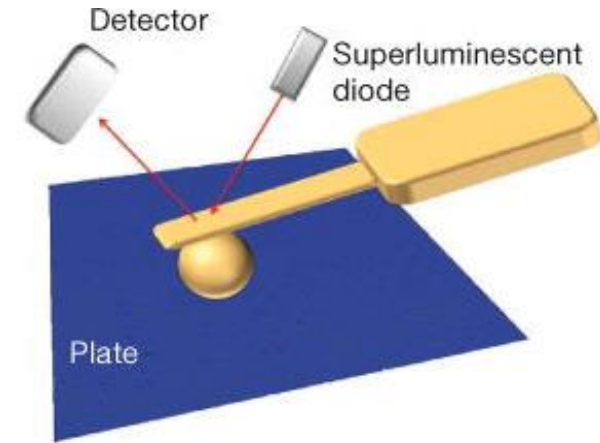
Measurement of Casimir forces



<https://physicstoday.scitation.org/doi/pdf/10.1063/1.2711635>



F. Intravaia et al Nature Communications 4 2515 (2013)



J. Munday et al Nature 457:170-173 (2009)