Physics 415 - Lecture 23: Classical Limit, Equipartition Theorem

March 14, 2025

Summary

- Canonical Ensemble (CE): Fixed T, V, N. $P_r = e^{-\beta E_r}/Z$, $Z = \sum_r e^{-\beta E_r}$ ($\beta = 1/T$). $F = -T \ln Z$.
- Classical Partition Function (N identical particles):

$$Z = \frac{1}{N!} \int \prod_{i=1}^{N} \frac{d^3 q_i d^3 p_i}{(2\pi\hbar)^3} e^{-\beta E(q,p)}$$

$$Z = \frac{1}{N!} \xi^{N} \int \prod_{i=1}^{N} \frac{d^{3}q_{i}}{V} e^{-\beta U(q)}$$

where $\xi = V/\lambda_{th}^3$ is related to the single-particle partition function, and $\lambda_{th} = h/\sqrt{2\pi mT}$ is the thermal de Broglie wavelength. U(q) is potential energy (for ideal gas U=0).

Validity of the Classical Approach

We have seen that a sensible partition function in classical statistical mechanics requires ingredients from Quantum Mechanics (QM): the phase space cell volume $(2\pi\hbar)^S$ and the N! factor for indistinguishable particles. Question: When is the classical approach valid? When must we resort to a full quantum treatment?

We can use the Heisenberg uncertainty relation $\Delta x \Delta p \sim \hbar$ to make some estimates of the "classical regime". (Note: This mixes classical and quantum reasoning, but gives the correct estimate. A more careful calculation uses a full QM treatment). Estimate the typical momentum spread Δp : Average kinetic energy per DOF is $\sim T/2$, so $p^2/(2m) \sim T/2 \implies p^2 \sim mT$. Assume the spread is roughly the typical momentum magnitude: $\Delta p \sim \sqrt{p^2} \sim \sqrt{mT}$. This momentum uncertainty determines a minimum "size" or position uncertainty of the particle via QM:

$$\Delta x \sim \frac{\hbar}{\Delta p} \sim \frac{\hbar}{\sqrt{mT}}$$

This leads to a characteristic length scale:

$$\lambda_{th} \equiv \sqrt{\frac{2\pi\hbar^2}{mT}} = \frac{h}{\sqrt{2\pi mT}}$$

This is the **thermal de Broglie wavelength**, roughly the quantum "size" of a particle at temperature T. Note that λ_{th} already appeared in the classical partition function:

$$\xi = V \left(\frac{mT}{2\pi\hbar^2}\right)^{3/2} = V \left(\frac{h^2}{2\pi m T (2\pi\hbar)^2}\right)^{3/2}$$
?

Let's re-check $\xi = V(m/(2\pi\hbar^2\beta))^{3/2} = V(mT/(2\pi\hbar^2))^{3/2}$. $\lambda_{th}^2 = h^2/(2\pi mT) = (2\pi\hbar)^2/(2\pi mT) = 2\pi\hbar^2/(mT)$. $mT/(2\pi\hbar^2) = 1/\lambda_{th}^2$. So $\xi = V(1/\lambda_{th}^2)^{3/2} = V/\lambda_{th}^3$. Correct.

$$Z = \frac{1}{N!} \left(\frac{V}{\lambda_{th}^3} \right)^N \int \prod \frac{d^3 q_i}{V} e^{-\beta U(q)} \quad \text{(Ideal gas } U = 0 \implies \text{integral} = V^N \text{)}$$

It is reasonable to expect the classical treatment to be appropriate when the average distance 'a' between particles is much larger than their thermal wavelength λ_{th} :

$$a \gg \lambda_{th}$$
 (Classical Regime)

When $a \sim \lambda_{th}$, quantum mechanical effects (wave nature, interference, identity) become important. We can rephrase this condition in terms of the particle density n = N/V. The average volume per particle is 1/n, so $a \sim (1/n)^{1/3}$. The condition becomes: $n^{-1/3} \gg \lambda_{th}$, or $1 \gg n \lambda_{th}^3$.

$$n\lambda_{th}^3 \ll 1$$
 (Classical Limit)

The classical limit is reached if the density n is sufficiently low or the temperature T is sufficiently high $(\lambda_{th} \propto 1/\sqrt{T})$.

Example 1: He gas at T=300 K and p=1 atm. $n\approx 3\times 10^{19}$ cm⁻³. Average distance $a=n^{-1/3}\approx 3\times 10^{-7}$ cm. Mass $m_{He}\approx 4\times 1.67\times 10^{-24}$ g. $\lambda_{th}=h/\sqrt{2\pi mT}\approx 0.5\times 10^{-8}$ cm. Here $a\gg \lambda_{th}$. Classical treatment is valid.

Example 2: Electrons (e^-) in a typical metal. $m_e \ll m_{He}$. Density $n \sim 10^{22} - 10^{23}$ cm⁻³. $a \sim \text{few} \times 10^{-8}$ cm (lattice spacing). $\lambda_{th} \sim 50 \times 10^{-8}$ cm at T = 300 K. Here $\lambda_{th} \gtrsim a$. Electrons in a metal form a highly quantum system even at room temperature.

Equipartition Theorem

This is an important general result in classical statistical mechanics. In many situations, the energy E contains terms that are quadratic in momenta and/or coordinates.

- Ideal gas kinetic energy: $K = \sum_{i=1}^{N} \frac{1}{2m} (p_{ix}^2 + p_{iy}^2 + p_{iz}^2)$. (Quadratic in momenta).
- Harmonic oscillator: $E = \frac{p^2}{2m} + \frac{1}{2}kq^2$. (Quadratic in p and q).

Theorem Statement: In classical statistical mechanics, for a system in thermal equilibrium at temperature T, each degree of freedom (coordinate or momentum) that enters the Hamiltonian quadratically contributes, on average, $\frac{1}{2}T$ to the internal energy \overline{E} . (Using T in energy units; otherwise $\frac{1}{2}k_BT$). It also contributes $\frac{1}{2}$ (or $\frac{1}{2}k_B$) to the heat capacity.

Derivation: Suppose the energy \tilde{E} contains a term $\epsilon = \frac{1}{2}\kappa u^2$, where u is some coordinate or momentum variable, and $E = \epsilon(u) + \tilde{E}(\text{other vars})$, where \tilde{E} is independent of u. The classical partition function Z (ignoring N! for simplicity here, as it won't affect averages) is:

$$Z = \int \frac{dq dp}{(2\pi\hbar)^S} e^{-\beta E} = \left(\int' \frac{dq' dp'}{(2\pi\hbar)^{S-1}} e^{-\beta \tilde{E}} \right) \left(\int_{-\infty}^{\infty} \frac{du}{2\pi\hbar \text{ or } 1?} e^{-\beta\kappa u^2/2} \right)$$
$$Z = \tilde{Z} \times \xi_u$$

where \tilde{Z} involves integration over all variables except u, and ξ_u is the factor from integrating over u:

$$\xi_u = C' \int_{-\infty}^{\infty} du \, e^{-\beta \kappa u^2/2} = C' \sqrt{\frac{2\pi}{\beta \kappa}}$$

The average energy associated with the variable u is $\bar{\epsilon} = \frac{1}{2}\kappa u^2$. We can calculate this using the formula $\bar{E} = -\partial(\ln Z)/\partial\beta$. The total average energy $\bar{E} = \bar{\tilde{E}} + \bar{\epsilon}$. $\ln Z = \ln \tilde{Z} + \ln \xi_u$.

 $\overline{E} = -\frac{\partial}{\partial\beta}(\ln\tilde{Z}) - \frac{\partial}{\partial\beta}(\ln\xi_u)$. The first term is the average energy from all other degrees of freedom. The second term is the contribution from u:

$$\bar{\epsilon} = -\frac{\partial}{\partial \beta} (\ln \xi_u) = -\frac{\partial}{\partial \beta} (\ln C' + \ln \sqrt{2\pi/(\beta \kappa)}) = -\frac{\partial}{\partial \beta} (\text{const} - \frac{1}{2} \ln \beta)$$
$$\bar{\epsilon} = -(-\frac{1}{2} \frac{1}{\beta}) = \frac{1}{2\beta} = \frac{1}{2}T$$

So, each quadratic term $\frac{1}{2}\kappa u^2$ contributes $\frac{1}{2}T$ to the average energy. The contribution to the heat capacity is $C_u = \partial \overline{\epsilon}/\partial T = \partial (\frac{1}{2}T)/\partial T = 1/2$. (Or $k_B/2$).

Note: This theorem is only true in classical statistical mechanics.

Simple Applications

- Single molecule (or atom) in gas at T: Translational kinetic energy $K = \frac{1}{2m}(p_x^2 + p_y^2 + p_z^2)$. Three quadratic terms. Contribution to average energy $= 3 \times (\frac{1}{2}T) = \frac{3}{2}T$.
- Monatomic ideal gas (N particles): $E = \sum_{i=1}^{N} K_i$. Total 3N quadratic terms. $\overline{E} = 3N \times (\frac{1}{2}T) = \frac{3}{2}NT$.
- Classical 1D Harmonic Oscillator: $E = \frac{p^2}{2m} + \frac{1}{2}kq^2$. Two quadratic terms. $\overline{E} = 2 \times (\frac{1}{2}T) = T$. Note that the average kinetic energy $\overline{K} = p^2/(2m) = T/2$ and average potential energy $\overline{U} = \overline{kq^2/2} = T/2$. Thus $\overline{K} = \overline{U} = \overline{E}/2$.

Quantum Harmonic Oscillator Revisited

It is interesting to see how the harmonic oscillator result is modified in QM. Energy levels: $E_n=(n+1/2)\hbar\omega,\, n=0,1,2,\ldots$ Partition function: $Z=\sum_{n=0}^\infty e^{-\beta E_n}=\sum_{n=0}^\infty e^{-\beta(n+1/2)\hbar\omega}=e^{-\beta\hbar\omega/2}\sum_{n=0}^\infty (e^{-\beta\hbar\omega})^n$. This is a geometric series $\sum_{n=0}^\infty x^n=1/(1-x)$ with $x=e^{-\beta\hbar\omega}$.

$$Z = \frac{e^{-\beta\hbar\omega/2}}{1 - e^{-\beta\hbar\omega}}$$

Average energy: $\overline{E} = -\frac{\partial}{\partial \beta} (\ln Z)$. $\ln Z = -\frac{\beta \hbar \omega}{2} - \ln(1 - e^{-\beta \hbar \omega})$.

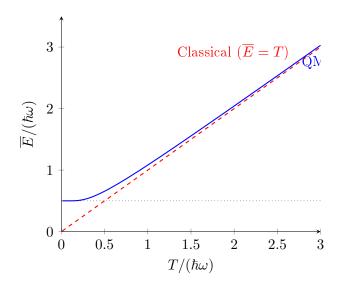
$$\begin{split} \overline{E} &= -\left[-\frac{\hbar\omega}{2} - \frac{1}{1 - e^{-\beta\hbar\omega}} (-e^{-\beta\hbar\omega}) (-\hbar\omega) \right] \\ \overline{E} &= \frac{\hbar\omega}{2} + \frac{\hbar\omega e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}} = \frac{\hbar\omega}{2} + \frac{\hbar\omega}{e^{\beta\hbar\omega} - 1} \\ \overline{E} &= \hbar\omega \left(\frac{1}{2} + \frac{1}{e^{\hbar\omega/T} - 1} \right) \end{split}$$

(Includes zero-point energy $\hbar\omega/2$).

Limits:

- High T limit $(T \gg \hbar\omega \implies \beta\hbar\omega \ll 1)$: Let $x = \beta\hbar\omega = \hbar\omega/T$. $e^x \approx 1 + x$. $\overline{E} \approx \hbar\omega(\frac{1}{2} + \frac{1}{(1+x)-1}) = \hbar\omega(\frac{1}{2} + \frac{1}{x}) = \hbar\omega(\frac{1}{2} + \frac{T}{\hbar\omega}) = \frac{\hbar\omega}{2} + T$. Excluding the zero-point energy, $\overline{E} E_0 \approx T$. This recovers the classical equipartition result $(\overline{E}_{classical} = T)$.
- Low T limit $(T \ll \hbar\omega \implies \beta\hbar\omega \gg 1)$: $e^{\beta\hbar\omega} \gg 1$. $\overline{E} \approx \hbar\omega(\frac{1}{2} + e^{-\beta\hbar\omega}) = \frac{\hbar\omega}{2} + \hbar\omega e^{-\hbar\omega/T}$. As $T \to 0$, $\overline{E} \to \hbar\omega/2$ (ground state energy).

Sketch of \overline{E} vs T:



Heat Capacity $C = \partial \overline{E} / \partial T$.

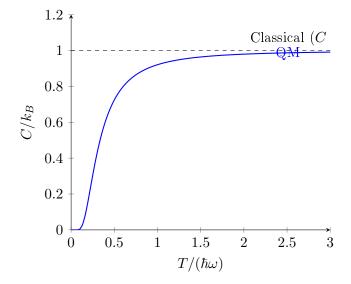
$$C = \frac{\partial}{\partial T} \left[\hbar \omega \left(\frac{1}{2} + \frac{1}{e^{\hbar \omega/T} - 1} \right) \right] = \hbar \omega \frac{-1}{(e^{\hbar \omega/T} - 1)^2} \left(e^{\hbar \omega/T} \right) \left(-\frac{\hbar \omega}{T^2} \right)$$
$$C = \left(\frac{\hbar \omega}{T} \right)^2 \frac{e^{\hbar \omega/T}}{(e^{\hbar \omega/T} - 1)^2}$$

(Using T in energy units, so C is dimensionless, equivalent to C/k_B).

Limits:

- High T limit $(x = \hbar \omega/T \ll 1)$: $e^x \approx 1 + x$. $C \approx x^2 \frac{1+x}{((1+x)-1)^2} = x^2 \frac{1+x}{x^2} \approx 1$. Agrees with classical C = 1 (or k_B).
- Low T limit $(x = \hbar \omega/T \gg 1)$: $e^x \gg 1$. $C \approx x^2 \frac{e^x}{(e^x)^2} = x^2 e^{-x} = (\frac{\hbar \omega}{T})^2 e^{-\hbar \omega/T}$. Goes to 0 exponentially as $T \to 0$.

Sketch of C vs T:



Quantized energy levels lead to a dramatic reduction ("freezing out") of the heat capacity at low temperatures $(T \ll \hbar\omega)$.