

# Summary

02/07/24

①

• Cyclic coord. 's:

$$p_i = \frac{\partial L}{\partial \dot{q}_i}$$

$$F_i = \frac{\partial L}{\partial q_i}$$

$$\Rightarrow \dot{p}_i = F_i$$

if  $F_k = \frac{\partial L}{\partial q_k} = 0$ , then  $p_k = \text{conserved}$  ( $\dot{p}_k = 0$ )

$q_k = \text{"cyclic coord."}$

• Symmetry & conservation laws.

(1) homog. of time  $\Rightarrow E = \sum_i \dot{q}_i \frac{\partial L}{\partial \dot{q}_i} - L$  conserved.

$$= T + U$$

(2) homog. of space  $\Rightarrow \vec{P} = \sum_a m_a \vec{v}_a$  " "

(3) isotropy of space  $\Rightarrow \vec{L} = \sum_a \vec{r}_a \times \vec{p}_a$  " "

$$\left( \begin{array}{l} \vec{r}_a \rightarrow \vec{r}_a + \epsilon \hat{n} \times \vec{r}_a \\ \vec{v}_a \rightarrow \vec{v}_a + \epsilon \hat{n} \times \vec{v}_a \end{array} \right)$$

Wrap up from last time. Transformation properties of  $\vec{L}$ .

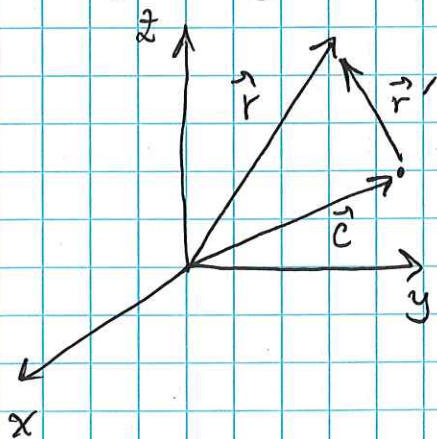
• choice of origin:  $\vec{r}_a = \vec{r}_a' + \vec{c} \Rightarrow$

$$\vec{L} = \sum_a \vec{r}_a \times \vec{p}_a$$

$$= \sum_a \vec{r}_a' \times \vec{p}_a + \vec{c} \times \sum_a \vec{p}_a$$

$$= \vec{L}' + \vec{c} \times \vec{P}$$

$\Rightarrow \vec{L}$  indep. of origin when  $\vec{P} = 0$  (COM frame).

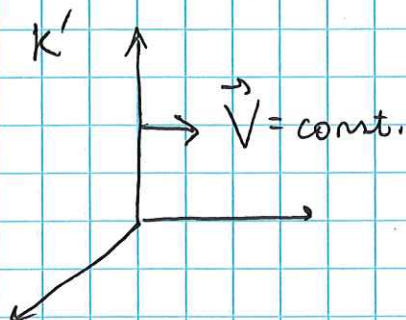
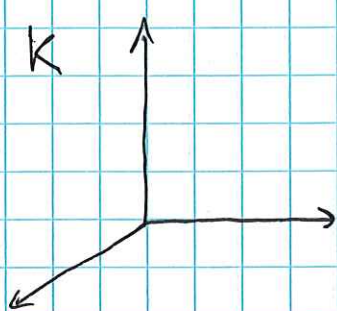




transform. 'n btwn. inertial frames

62/07/24

(2)



$$\vec{v}_a = \vec{v}_a' + \vec{V}$$

(suppose origins of  $K$  &  $K'$  coincide at instant considered)

$$\vec{L} = \sum_a \vec{r}_a \times \vec{p}_a = \sum_a \vec{r}_a \times m_a \vec{v}_a = \sum_a \vec{r}_a \times m_a (\vec{v}_a' + \vec{V})$$

$$\Rightarrow \vec{L} = \cancel{\sum_a \vec{r}_a \times m_a \vec{v}_a'} + \left( \sum_a m_a \vec{r}_a \right) \times \vec{V}$$

$$\vec{R} = \frac{1}{M} \sum_a m_a \vec{r}_a$$

$$= \vec{L}' + M \vec{R} \times \vec{V}$$

$$M = \sum_a m_a$$

↑  
ang. mom. in frame  $K'$ .

def  $K'$  = COM frame ~~then~~, where  $\vec{P}' = 0$ , then  $\vec{P} = M \vec{V}$

$$\Rightarrow \vec{L} = \vec{L}' + \vec{R} \times \vec{P}$$

↑  
"intrinsic  
ang. mom."

↑  
motion "as a whole".

Conservation of  $\vec{L}$  in ext. fields:

• motion in central potential  $U(r)$ ,  $r = |\vec{r}|$ .

under rotation:  $\vec{r} \rightarrow \vec{r}' = R \cdot \vec{r}$ ,  $U(r) \rightarrow U(r') = U(r)$

$\Rightarrow$  Lagrangian invariant &  $\vec{L}$  = conserved.

more generally, components of  $\vec{L}$  conserved when there is an axis of symmetry.

Ex:  $U = U(z) \Rightarrow$  inv. under rot. about  $\hat{z} \Rightarrow L_z = \hat{z} \cdot \vec{L}$  conserved.



Explicitly, if  $U = U(z)$ , we use cylindrical coord.'s & write:

$$L = T - U = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2 + \dot{z}^2) - U(z).$$

E-L eqn.'s:  $\phi: \frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}} = \frac{\partial L}{\partial \phi} = 0$

$$\Rightarrow \frac{\partial L}{\partial \dot{\phi}} = m r^2 \dot{\phi} = \text{conserved}.$$

$$\& \quad L_z = \hat{z} \cdot \vec{L} = m r^2 \dot{\phi}$$

check:  $L_z = x p_y - y p_x$   $\left\{ \begin{array}{l} x = r \cos \phi \\ y = r \sin \phi \end{array} \right. \quad \left\{ \begin{array}{l} \dot{x} = \dot{r} \cos \phi - r \dot{\phi} \sin \phi \\ \dot{y} = \dot{r} \sin \phi + r \dot{\phi} \cos \phi \end{array} \right.$

~~$E = m r^2 \dot{\phi}^2$~~

$$\begin{aligned} &= m (r \dot{r} \cos \phi \sin \phi + r^2 \dot{\phi} \cos^2 \phi \\ &\quad - r \dot{r} \cos \phi \sin \phi + r^2 \dot{\phi} \sin^2 \phi) \\ &= m r^2 \dot{\phi} \quad \checkmark \end{aligned}$$

### Mechanical ~~System~~ Similarity

(This is a "bonus" topic. you will not be tested on this material).

From symmetries of the Lagrangian  $L$  we have deduced a number of conservation laws. We'll now exploit one more transformation property of  $L$  to obtain some non-trivial consequences regarding the motion:

• Rescaling  $L \rightarrow L' = \alpha L$  does not affect EOM.

$$\left( \frac{d}{dt} \frac{\partial L'}{\partial \dot{q}} = \alpha \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = \alpha \frac{\partial L}{\partial q} = \frac{\partial L'}{\partial q} \right)$$



- Keeping this rescaling property of  $L$  in mind, consider (02/07/24)  
a rescaling of the coord.'s:

$$\vec{r} \rightarrow \vec{r}' = \alpha \vec{r}$$

$$\begin{aligned} \Rightarrow T = \sum_a \frac{1}{2} m_a \dot{\vec{r}}_a^2 &\rightarrow T' = \sum_a \frac{1}{2} m_a (\dot{\vec{r}}'_a)^2 \\ &= \sum_a \frac{1}{2} m_a \alpha^2 \dot{\vec{r}}_a^2 \\ &= \alpha^2 T \end{aligned}$$

$$\Rightarrow T' = \alpha^2 T.$$

we say  $T$  is a "homogeneous" fn., which means the value of the fn. is simply rescaled upon rescaling its arguments.

- Now suppose the potential is also a homogeneous fn:

$$\vec{r} \rightarrow \vec{r}' = \alpha \vec{r}$$

$$\begin{aligned} \Rightarrow U \rightarrow U' &= U(\alpha \vec{r}_1, \dots, \alpha \vec{r}_n) = \alpha^k U(\vec{r}_1, \dots, \vec{r}_n) \\ U &= \text{"homog. fn. of degree } k\text{"} \\ (T &= \text{homog. fn. of degree } 2). \end{aligned}$$

Ex: Uniform force:  $U \sim z \Rightarrow k=1$

Harmonic force:  $U \sim r^2 \Rightarrow k=2$

Newton / Coulomb:  $U \sim 1/r \Rightarrow k=-1$

- So, if  $U = \text{homog. fn. of deg. } k$ , then upon rescaling:

$$L \rightarrow L' = T' - U' = \alpha^2 T - \alpha^k U.$$



• now also make a rescaling of time:

$$t \rightarrow t' = \beta t$$

$$\Rightarrow \vec{v} = \frac{d\vec{r}}{dt} \rightarrow \vec{v}' = \frac{d\vec{r}'}{dt'} = \frac{\alpha}{\beta} \frac{d\vec{r}}{dt} = \frac{\alpha}{\beta} \vec{v}$$

$$\Rightarrow T \rightarrow T' = \frac{\alpha^2}{\beta^2} T.$$

$$\& L \rightarrow L' = \frac{\alpha^2}{\beta^2} T - \alpha^k U.$$

• Now we choose  $\beta^2 = \alpha^{\frac{2-k}{k}}$ ; then we get:

$$\begin{aligned} L' &= \frac{\alpha^2}{\alpha^{2-k}} T - \alpha^k U \\ &= \alpha^k T - \alpha^k U \\ &= \alpha^k L \end{aligned}$$

• Thus we conclude that, under the transform<sup>n</sup>:

$$\vec{r}(t) \rightarrow \vec{r}'(t') = \alpha \vec{r}(t)$$

$\uparrow$   
 $= \beta t$

$L$  is simply rescaled & EOM unchanged.

• written a slightly different way:

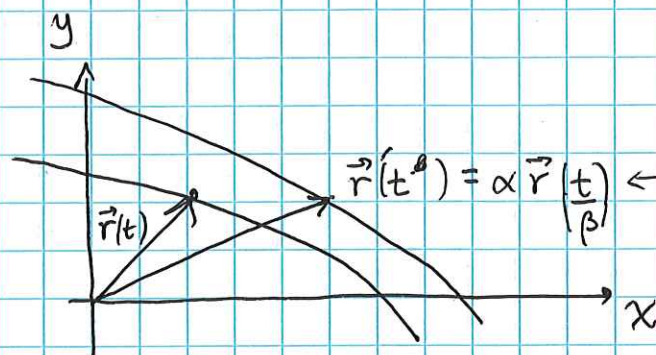
$$\vec{r}'(\beta t) = \alpha \vec{r}(t) \quad \downarrow \text{ set } t \rightarrow t/\beta$$

$$\begin{aligned} \vec{r}'(t) &= \alpha \vec{r}(t/\beta) \\ &= \alpha \vec{r}(\alpha^{\frac{k}{2}-1} t) \end{aligned}$$



(6)

So, if  $U = \text{homog. fn. of deg. } k$ , we can generate entire families of geometrically similar paths ~~from~~ (solutions to EOM) from one solution  $\vec{r}(t)$ .

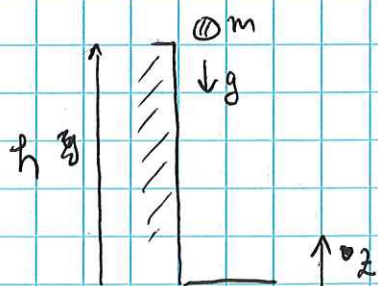


new solution to  
EOM: rescaling of  
old soln. ( $\vec{r}(t)$ ) w/  
a diff. time of motion  
through trajectory.

Some consequences of this:

Q: Suppose we drop a ball, initially at rest, from height  $h$  & it takes time  $T$  to reach the ground. How long will it take to reach ground if dropped from height  $2h$ ?

• straight-forward approach:



$$m\ddot{z} = -mg$$

$$\Rightarrow z(t) = h - \frac{1}{2}gt^2$$

$$0 = z(T) = h - \frac{1}{2}gT^2 \Rightarrow T = \sqrt{2h/g}$$

Thus, if  $h \rightarrow 2h$ , then  $T \rightarrow \sqrt{2} \times T$ .

Note: to get this result, we had to solve EOM



Alternatively, use the fact that  $U = mgz$  is a homog. fn. of deg.  $k=1$

$\Rightarrow$  if there is a solution in which particle travels distance  $h$  in time  $T$ , then there is another solution in which particle travels  $h' = \alpha h$  in time  $T' = \beta T = \alpha^{1-k/2} T = \sqrt{\alpha} \times T$

So, if  $h' = 2h$ , then  $T' = \sqrt{2} T$  ✓

$\rightarrow$  note we didn't have to solve EOM!

More generally, suppose  $U = \text{homog. fn. of deg. } k$  & particle travels distance  $l$  in time  $t$ . Then there is another trajectory (sol.'n of EOM) in which particle travels  $l'$  in time  $t'$  w/  $l' = \alpha l$ ,  $t' = \beta t = \alpha^{1-k/2} t$ .

$$\Rightarrow \left(\frac{t'}{t}\right) = \alpha^{1-k/2} = \left(\frac{l'}{l}\right)^{1-k/2}$$

Ex: • uniform force,  $k=1$   $\frac{t'}{t} = \sqrt{\frac{l'}{l}} \Rightarrow$  time of fall

$\sim$  square root of initial height.

• harmonic force,  $k=2$   $\frac{t'}{t} = \alpha^0 = \text{const.} \Rightarrow$  period of oscillation indep. of amplitude



• Newton / Coulomb,  $k = -1$

$$\frac{t'}{t} = \left(\frac{L'}{L}\right)^{3/2}$$

02/07/24

8

$\Rightarrow$  for orbit of size  $L$  & period  $T$ , we have

$T^2 \sim L^3$ . That is, orbital period varies in proportion

to the cube of the size of orbit

$\rightarrow$  Kepler's third law of planetary motion.