

- (a) (b) as graphed above. In order to simulate gravity,

$$A = \omega^2 R = g \Rightarrow \omega = \sqrt{\frac{g}{R}} = 0.5 \text{ rad/s} \quad (1)$$

- (c) effective gravity is

$$g_{\text{eff}} = \omega^2 R \propto R$$

$$\Rightarrow \frac{g_{\text{head}} - g_{\text{feet}}}{g_{\text{feet}}} = \frac{R_{\text{head}} - R_{\text{feet}}}{R_{\text{feet}}} = \frac{-2}{40} = -5\% \quad (2)$$

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- (a) recall EOM for a particle in rotating frame with angular velocity Ω ,

$$m\ddot{\vec{r}} = -m\Omega^2\rho + \overbrace{2m\dot{\vec{r}} \times \vec{\Omega}}^{F_{\text{cor}}} + \overbrace{m\vec{\Omega} \times (\vec{r} \times \vec{\Omega})}^{F_{\text{centripetal}}} + F_{\text{buoyancy}} \quad (3)$$

noticing that the droplet is relatively stationary with respect to the rotation frame, $F_{\text{cor}} = 0$, $F_{\text{centripetal}} \neq 0$

- (b) (c) The direction of $F_g + F_{\text{cent}}$ is downward, tangential to the water surface.

The water level is a equipotential surface of the combined potential of gravity and centrifugal force. So for a given height z ,

$$\begin{aligned} U &= mgz + U_{\text{cent}} \\ \text{where } U_{\text{cent}} &= - \int F_{\text{cent}} d\rho = - \int m\Omega^2\rho d\rho = -\frac{1}{2}m\Omega^2\rho^2 \\ \Rightarrow U &= mgz - \frac{1}{2}m\Omega^2\rho^2 = \text{constant} \\ \Rightarrow z(\rho) &= \frac{\Omega^2\rho^2}{2g} + \text{const} \end{aligned} \quad (4)$$

It is obvious that $z(\rho)$, i.e. water surface, is a parabola.

we choose the frame to be the following: $\{\hat{x}, \hat{y}, \hat{z}\}$, where \hat{x} points eastward, \hat{y} points northward, and \hat{z} points upward. Let particle above earth surface has position \vec{r} , whose angle with earth's rotation axis is θ .

EOM of particle

$$\begin{cases} \dot{\vec{v}} = -g\hat{z} + 2\vec{v} \times \vec{\Omega} \\ \vec{\Omega} = \Omega \sin \theta \hat{y} + \Omega \cos \theta \hat{z} \\ \vec{v} = v_x \hat{x} + v_y \hat{y} + v_z \hat{z} \end{cases} \Rightarrow \dot{\vec{v}} = -g\hat{z} + 2\Omega[(v_y \cos \theta - v_z \sin \theta)\hat{x} - v_x \cos \theta \hat{y} + v_x \sin \theta \hat{z}] \quad (5)$$

let $\vec{v}(t) = \vec{v}_1 + \vec{v}_2$, with the first term being velocity neglecting rotation, and 2nd term is the correction with rotation.

for $\vec{v}_1, \Omega = 0$.

$$\Rightarrow \begin{cases} \dot{v}_x = \dot{v}_y = 0 \\ \dot{v}_z = -g \end{cases} \Rightarrow \begin{cases} \vec{v}_1(t) = \vec{v}_0 - gt\hat{z} \\ \vec{r}(t) = \vec{r}_0 + \vec{v}_0 t - \frac{1}{2}gt^2\hat{z} \end{cases} \quad (6)$$

Therefore, when recalling the coriolis force, $\dot{\vec{v}} = -g\hat{z} + 2\vec{v} \times \vec{\Omega}$, we have

$$\begin{aligned} \dot{\vec{v}}_1 + \dot{\vec{v}}_2 &= -g\hat{z} + 2\vec{v}_1 \times \vec{\Omega} + 2\vec{v}_2 \times \vec{\Omega} \\ &\text{recognizing } \Omega = \text{small}, v_2 = \text{small} \end{aligned} \quad (7)$$

$$\vec{v}_2 \approx 2(\vec{v}_0 - gt\hat{z}) \times \vec{\Omega}$$

Now consider:

1. free fall from height h , with $v_0 = 0$, over a time lapse of $t = \sqrt{\frac{2h}{g}}$

$$\dot{\vec{v}} = 2gt\Omega \sin \theta \hat{x} \Rightarrow \vec{v} = \int \dot{\vec{v}} dt = gt^2\Omega \sin \theta \hat{x} \quad (8)$$

$$\vec{v} = -gt\hat{z} + gt^2\Omega \sin \theta \hat{x}$$

$$\vec{r} = \int \vec{v} dt = \begin{pmatrix} \frac{1}{3}gt^3\Omega \sin \theta \\ 0 \\ -\frac{1}{2}gt^2 \end{pmatrix} \quad (9)$$

Noticing that $t = \sqrt{\frac{2h}{g}}$, displacement on the x-y plane is

$$x_f = \frac{1}{3}g\Omega \left(\sqrt{\frac{2h}{g}} \right)^3 \sin \theta \quad (10)$$

2. compare with particle thrown from ground, with $\vec{v}_0 = v_0\hat{z}$, $t = 2\sqrt{\frac{2h}{g}} = \frac{2v_0}{g}$

$$\begin{aligned} \dot{\vec{v}}_2 &= 2(v_0 - gt)\hat{z} \times \vec{\Omega} \\ &= 2(v_0 - gt)\hat{z} \times (\Omega \sin \theta \hat{y} + \Omega \cos \theta \hat{z}) \\ &= \hat{x}[-2(v_0 - gt)\Omega \sin \theta] \\ &= (-2v_0\Omega \sin \theta + 2gt\Omega \sin \theta)\hat{x} \end{aligned} \quad (11)$$

$$\begin{aligned}
\vec{v}_2 &= \int \vec{v}_2 \, dt = (-2v_0\Omega \sin \theta t + g\Omega \sin \theta t^2) \hat{x} \\
\Rightarrow \vec{v} &= \vec{v}_1 + \vec{v}_2 = \begin{pmatrix} -2v_0\Omega \sin \theta t + g\Omega \sin \theta t^2 \\ 0 \\ v_0 - gt \end{pmatrix}
\end{aligned} \tag{12}$$

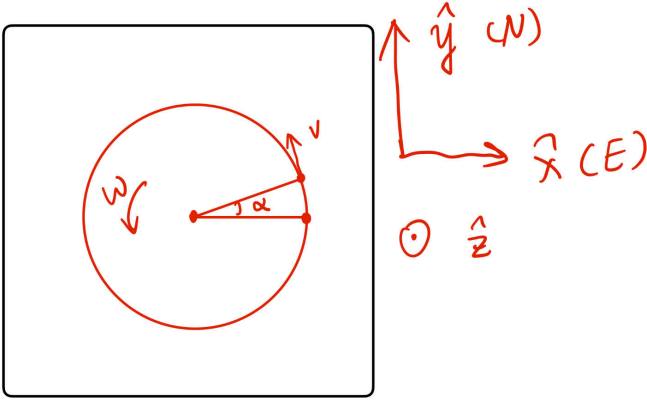
Considering $t = 2\sqrt{\frac{2h}{g}} = \frac{2v_0}{g}$

$$\begin{aligned}
\Rightarrow x_t &= \int v_x \, dt = -v_0\Omega \sin \theta t^2 + \frac{1}{3}g\Omega \sin \theta t^3 \\
&= -v_0\Omega \sin \theta \left(2\frac{v_0}{g}\right)^2 + \frac{1}{3}g\Omega \sin \theta \left(2\frac{v_0}{g}\right)^3 \\
&= -\frac{4}{3}g\Omega \left(\sqrt{\frac{2h}{g}}\right)^3 \sin \theta
\end{aligned} \tag{13}$$

Comparing Equation 10 and Equation 13, we have $(x_t)/(x_f) = -4$, thus showing that the deflection is indeed opposite and 4 times larger when the particle is thrown from the ground.

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We propose the following frame: $\{\hat{x}, \hat{y}, \hat{z}\}$, where \hat{x} points eastward, \hat{y} points northward, and \hat{z} points upward. Picture a hoop lying on ground with the following illustration:



Consider the rotation of an infinitesimal element of the hoop over an infinitesimal time interval, across small angle α . Find coriolis force

$$dF_{\text{cor}} = 2 \, dm (\vec{v} \times \vec{\Omega}) \quad (14)$$

where

$$\vec{v} = \omega r \begin{pmatrix} -\sin \alpha \\ \cos \alpha \\ 0 \end{pmatrix}, \quad \vec{\Omega} = \Omega \begin{pmatrix} 0 \\ \sin \theta \\ \cos \theta \end{pmatrix}, \quad dm = m \frac{d\alpha}{2\pi} \quad (15)$$

Torque on said infinitesimal element can be found by

$$d\tau = \vec{r} \times dF_{\text{cor}}, \quad \vec{r} = r \begin{pmatrix} \cos \alpha \\ \sin \alpha \\ 0 \end{pmatrix} \quad (16)$$

$$\Rightarrow d\tau = 2 \, dm (\vec{v} (\vec{r} \cdot \vec{\Omega} - \vec{\Omega} (\vec{r} \cdot \vec{v}))) = 2\omega r^2 \Omega \sin \theta \, dm \begin{pmatrix} -\sin^2 \alpha \\ \sin \alpha \cos \alpha \\ 0 \end{pmatrix}$$

noticing $dm = m \frac{d\alpha}{2\pi}$, integrating the above w.r.t. α from 0 to 2π ,

$$\tau = -(m\omega\Omega r^2 \sin \theta) \hat{x} \quad (17)$$

The above shows that the torque due to coriolis force is westward with magnitude $m\omega\Omega r^2 \sin \theta$

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According to the hints in problem statement of Taylor 9.34, we write the puck's position vector, relative to the earth's center O, as $\vec{R} + \vec{r}$, where \vec{R} is the position of thi point P and $\vec{r} = (x, y, 0)$ is the puck's position relative to P.

Ignoring centrifugal force, we can write the EOM as the following

$$\ddot{\vec{r}} = g_0(r) + 2\dot{\vec{r}} \times \vec{\Omega} \quad (18)$$

$$\text{where } g_0(r) = -GM \frac{\vec{R} + \vec{r}}{\|(\vec{R} + \vec{r})\|^3} = -GM \frac{\vec{R} + \vec{r}}{R^3} \left(1 + \frac{r^2}{R^2}\right)^{-\frac{3}{2}} \quad (19)$$

recognizing that $r \ll R$, expanding the above function gives an approximation:

$$g(r) = -GM \frac{\vec{R} + \vec{r}}{R^3} = \vec{g}(0) + g(0) \frac{\vec{r}}{R} \quad (20)$$

putting the above into Equation 18, we have

$$\begin{aligned} \ddot{\vec{r}} &= \vec{g}(0) + g(0) \frac{\vec{r}}{R} + 2\dot{\vec{r}} \times \vec{\Omega} \\ &= g(0) \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} - \frac{g}{R} \begin{pmatrix} x \\ y \\ 0 \end{pmatrix} + 2 \begin{pmatrix} \dot{y}\Omega \cos \theta - \dot{z}\Omega \sin \theta \\ -\dot{x}\Omega \cos \theta \\ \dot{x}\Omega \sin \theta \end{pmatrix} \end{aligned} \quad (21)$$

The above is set of 2nd order ODEs:

$$\begin{aligned} \ddot{x} &= -g \frac{x}{R} + 2\dot{y}\Omega \cos \theta \\ \ddot{y} &= -g \frac{y}{R} - 2\dot{x}\Omega \cos \theta \end{aligned} \quad (22)$$

This is the same as Foucault pendulum equation with length of pendulum being R .

The oscillation frequency is $\omega_0 = \sqrt{\frac{g}{R}} \approx 1.24e(-3) \text{ s}^{-1}$, frequency of Foucault precession, $\Omega_z = \Omega \cos \theta$