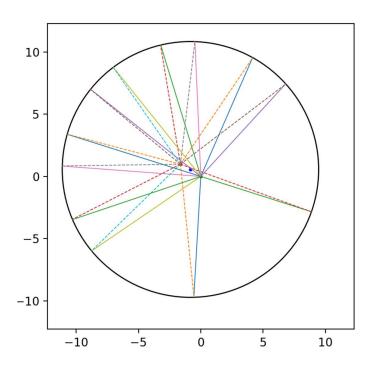
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M.Tech Artificial Intelligence SR No. 17862

Results



Equant: Red Sun: Green Center: Blue

Dashed line: Equant Longitudes Straight line: Heliocentric Longitudes

| Parameters | Optimum Values |
|--------------------------------|--|
| Center angle (c) | 145.887° |
| Equant distance (e1) | 1.906 * sun-center distance |
| Equant angle wrt equant-0 (e2) | 93.05° |
| Equant-0 (z) | 55.87° |
| Orbit radius (r) | 10.28 * sun-center distance |
| Angular velocity (s) | 360/686.918 °/days |
| Errors | [-0.077, 0.0791, 0.0619, -0.0534, -0.0243, 0.0799, -0.0098, -0.0791, 0.0402, 0.0618, -0.0558, -0.0799] |
| Maximum error | 0.0799° |

Initial guess for Orbit radius

Semi-major axis(a) = 1.52400 Semi-minor axis(b) = 1.51740 Eccentricity(e) = 0.093

Data is taken from here: https://en.wikipedia.org/wiki/Semi-major and semi-minor axes#Semi-major and semi-minor axes#Semi-major and semi-minor axes of the planets' orbits

Average radius = (a+b)/2 = 1.5207Focal distance(f) = a*e = 0.141732 (Assuming Sun at Foci) Scaling radius to unit of focal distance = Average radius / f = **10.6**

Learning & Conclusion

- Using datetime module is important to get Equant longitudes correctly.
- Use conversion between polar to cartesian or vice versa as few as possible. These conversions introduce floating-point errors a lot.