

CHAPTER 37

SPECIAL RELATIVITY (SR)

Einstein's Postulates
Relativity of simultaneity
Relativity of time intervals - time dilation
Relativity of length - length contraction
Lorentz transformations
(Doppler effect)
Relativistic momentum and energy

TIMELINE

Classical Mechanics = study of macroscopic world at speeds $\ll c$

Classical Electro-magn.

1642-1727	Newton	light is not a wave; it is made out of particles
1801	Young	double-slit experiment
1861-2	Maxwell	wave theory of el.mg. radiation/light

Quantum Physics = study of microscopic world - not restricted to a type of phenomenon

quantum - quanta (pl)- elementary amount of quantized physical quantities

1800s(late)		black-body radiation
1886-87	Hertz:	discovers photo-electric effect (later Von Lenard does more experiments) experimental proof of el.mg. waves

1893 Thomson *electron*

1895 Roentgen X-rays

1900 Planck mathematical model for BB radiation

$$E = n h f$$

1905 Einstein 03/18 radiation/light is quantized in photons

$$E = n h f$$

(annus mirabilis) & light absorption and emission occur in atoms

05/11 Brownian motion - kinetic theory of gases

06/30 special relativity

09 27 mass-energy equivalence

$$E_0 = m c^2$$

1909 Taylor double-slit experiment in single-photon version

1911 Rutherford *nucleus*

1916 Einstein quanta of light have linear momentum

$$p = \frac{E}{c} = \frac{h f}{c} = \frac{h}{\lambda}$$

1917 Rutherford *proton*

1923 Schrodinger main QM equation

Compton single- λ X-rays on carbon target scattered

$$\Delta\lambda = \frac{h}{mc} (1 - \cos\varphi)$$

1924 De Broglie wavelength

$$\lambda = \frac{h}{p}$$

1927 Heisenberg uncertainty principle = basic limitation on our experimental capabilities

Davisson-Germer/Thomson experiment

CH.37 SPECIAL RELATIVITY

INTRODUCTION

Relativity

concepts of space **and time** and their transformation between different reference frames

Special vs. general relativity

Special - inertial reference frames (IRF) - not undergoing acceleration

General - reference frames can have gravitational acceleration

Low vs. high speeds

Classical/slow relativity is wrong at high speed

SR - is an extension to all physically possible speeds

- includes classical/slow relativity as a limit for $v \ll c$

Space and time

Are entangled (time between events depends on how far apart they are)

and entanglement is \neq for \neq observers

Time does not pass at a fixed rate

Easy and difficult

Easy - math

Difficult - who/what/how/when

POSTULATES (1905)

1. Relativity postulate (principle of relativity)

Laws of physics are the same in all inertial reference frames IRF

Note: laws NOT measured values

Consequences - alter the definitions of energy/momentum

- energy/mass equivalence

1887 Michelson-Morley experiment

measured same speed of light in two directions - using interference pattern

2. Speed of light postulate

a. The speed of light in vacuum has the same value c in all directions and in all reference frames, and it is independent of the motion of IRF

OR

b. There is an ultimate speed c , no entity can exceed it, and no particle with mass can reach it.

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = 299\,792\,458 \frac{m}{s} \cong 3 \cdot 10^8 \frac{m}{s}$$

1964 experiment with accelerated electrons

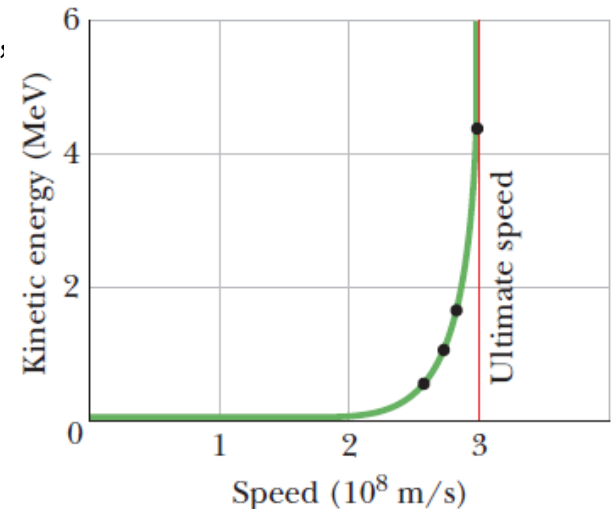
$$v_e = 0.999\,999\,999\,95\,c$$

green - special relativity prediction

1964 CERN experiment with $\pi^0 \rightarrow 2\gamma$

$$v_\pi = 0.999\,75\,c$$

$$v_\gamma = c \text{ (same as for } v_\pi = 0 \text{)}$$



RELATIVITY OF SIMULTANEITY

Thesis: Simultaneity is not an absolute concept, but a relative one, depending on the motion of the observer

OR

Two observers in relative motion might not agree that two events are simultaneous

Example:

RF = spaceships or trains

Start at same point

Sally has relative speed v

Observers at midpoint

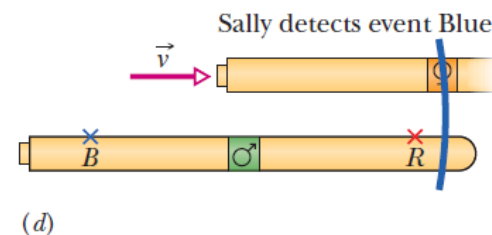
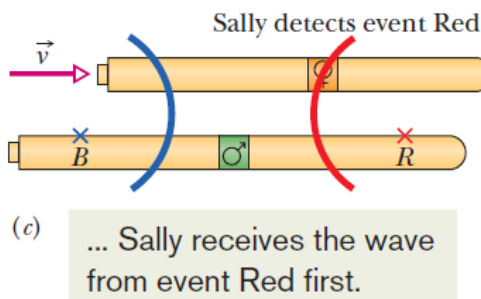
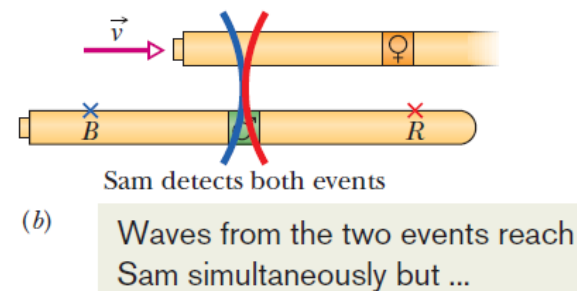
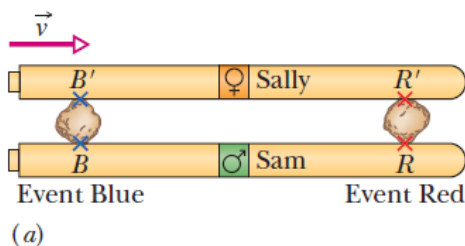
Events = 2 meteorites (R/B) strike
and emit red/blue light

Result:

Sam - simultaneous events

Sally - not simultaneous

Both are right!



CH.37 SPECIAL RELATIVITY

RELATIVITY OF TIME



Thesis: The time interval between two events depends on how far apart they occur in both space and time (entangled)

Example: observer on moving train (v) with light source, mirror, clock

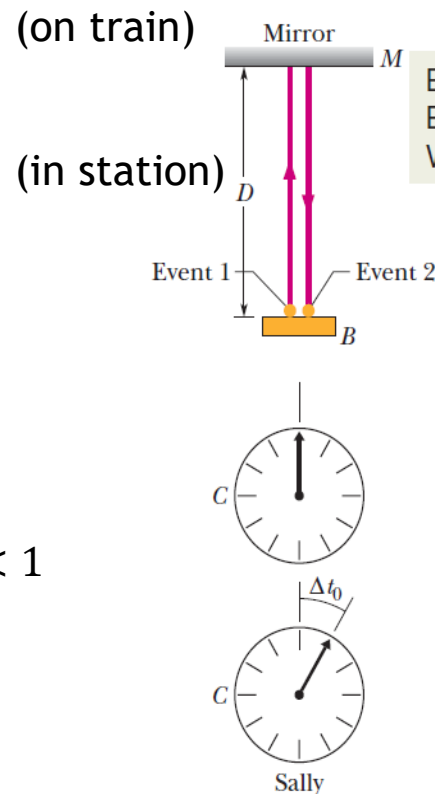
$$\Delta t_0 = \frac{2D}{c}$$

$$\Delta t = \frac{2L}{c} = \frac{2}{c} \sqrt{D^2 + \left(\frac{v\Delta t}{2}\right)^2}$$

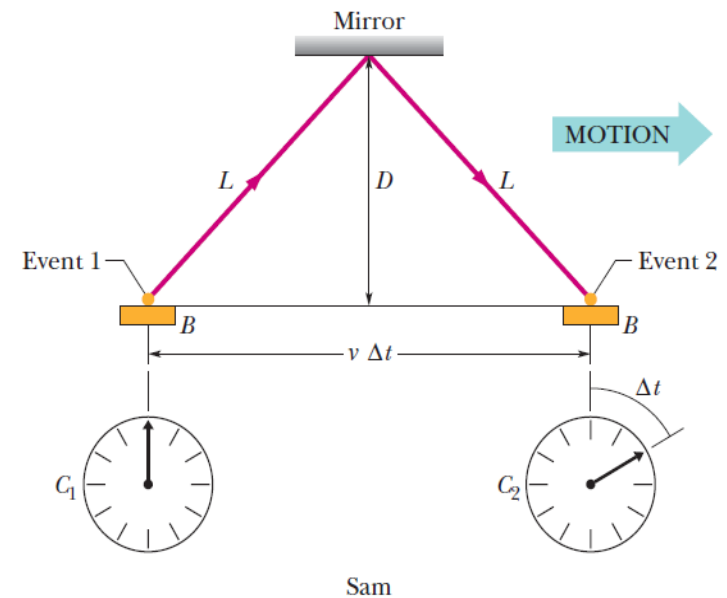
$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - (v/c)^2}} > \Delta t_0$$

→ shorter time in own RF

since $\sqrt{1 - \frac{v^2}{c^2}} < 1$



Event 1 is the emission of light.
Event 2 is the return of the light.
We want the time between them.



RELATIVITY OF TIME

Previous result

$$\Delta t = \frac{\Delta t_0}{\sqrt{1-(v/c)^2}}$$

Time dilation

$$\Delta t = \gamma \Delta t_0 > \Delta t_0$$

Proper time (interval)

Δt_0 time between two events occurring at the same location in an inertial RF
shortest possible measured time interval

Speed parameter

$$\beta = \frac{v}{c} \leq 1$$

Lorentz factor

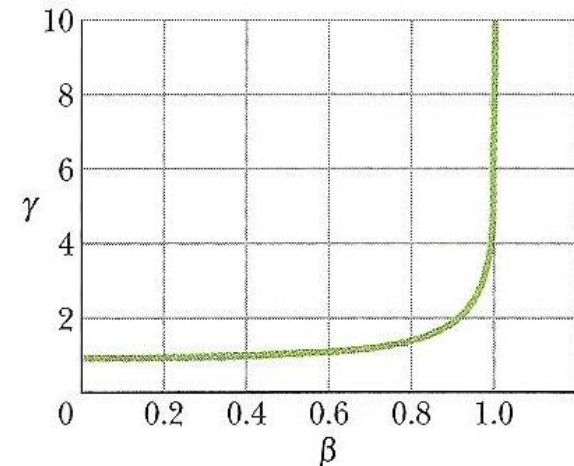
$$\gamma = \frac{1}{\sqrt{1-(\frac{v}{c})^2}} = \frac{1}{\sqrt{1-\beta^2}} > 1 \quad \text{for } v \neq 0$$

Low-speed approx.

$$\gamma = 1 + \frac{1}{2} \beta^2 \cong 1 \quad \text{for } v < 0.1c$$

NR limit

$$\gamma = 1.02 \quad \text{for } v = 0.2c$$



TIME DILATION TESTS

1. Microscopic clocks -

muon (μ meson) - 1936 cosmic rays and 1937 cloud chamber

$$m = 105.66 \frac{\text{MeV}}{c^2}; \quad q = -e; \quad s = \frac{1}{2}$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

at rest $\tau_0 = 2.2 \mu\text{s}$

at $v = 0.9994c$ $\tau = \gamma\tau_0 = 63.51 \mu\text{s}$

Rest frame = stationary RF at the muon itself

2. Macroscopic clocks -

1971 Hafele & Keating

flew 4 atomic clocks around the world once in each direction

results within 10% of special relativity

~1975 UMD - 15 h flight w/ atomic clock over the Chesapeake Bay

results within 1%

RELATIVITY OF LENGTH

Thesis: The distance between two points depends on RF and is related to relativity of simultaneity, therefore, is relative

Proper (rest) length L_0 -

distance between 2 points measured at the same time by an observer
at rest relative to both points

length measured in the RF of the object - where the object is at rest

Length contraction $L = \frac{L_0}{\gamma} < L_0$

Notes:

1. “length” can be the distance between to points/objects
2. length contraction - occurs only along the direction of relative motion
 - is a direct consequence of time dilation

RELATIVITY OF LENGTH

Proper length (rest length) L_0

distance between 2 points measured at the same time by an observer
at rest relative to both points

length measured in the RF of the object - where the object is at rest

Example:

1. Measuring the length of the station platform



Sam (at station): L_0 proper length (at rest)
 Δt time Sally moves through this length (not proper time - at \neq places)

$$L_0 = v \Delta t$$

Sally (in train): sees the platform moving past her
 Δt_0 time it passes through the station (proper time)

$L = v \Delta t_0$ not proper length / not in RF of the station

Use time dilation

$\Delta t = \gamma \Delta t_0$ to obtain length contraction

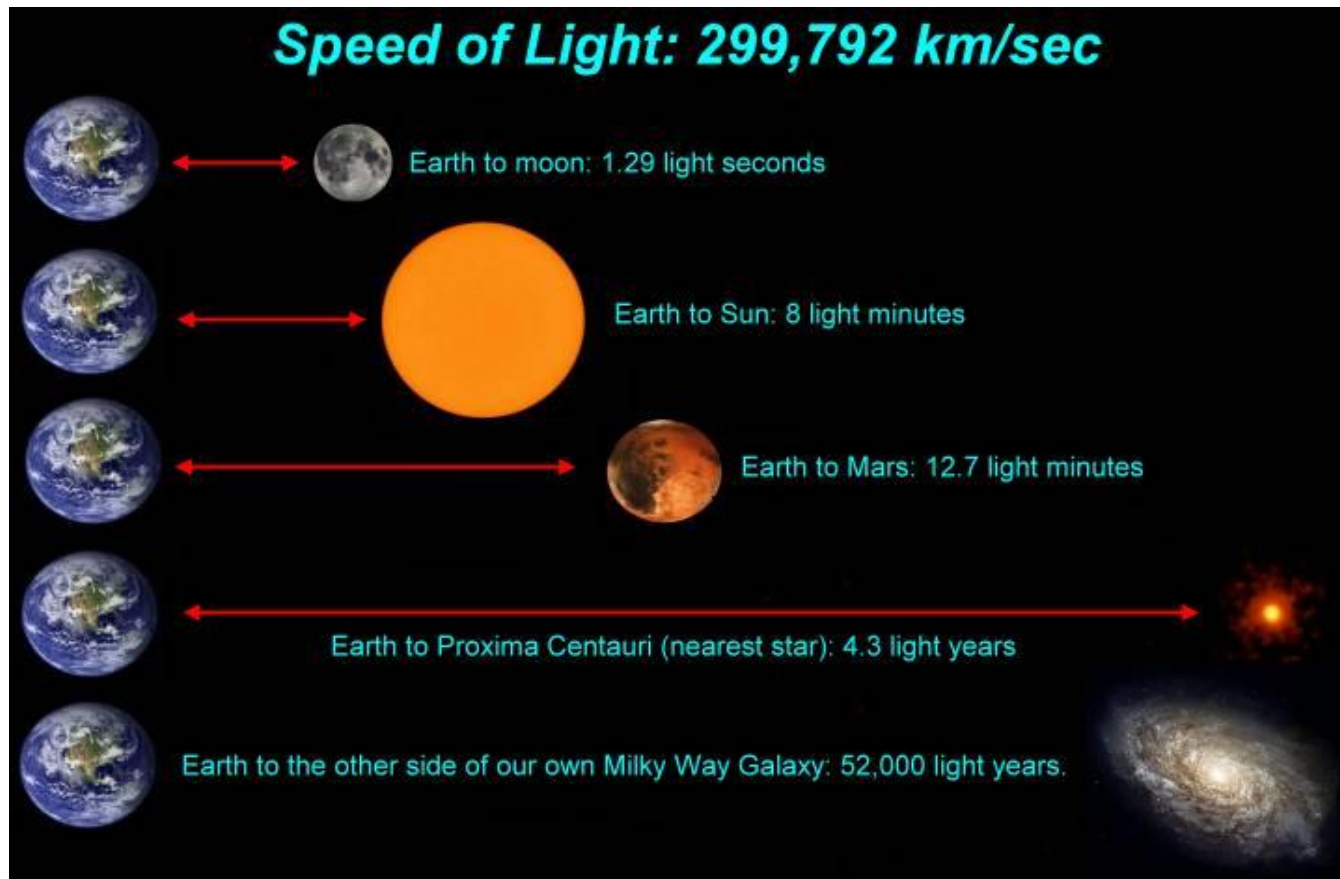
$$\frac{L}{L_0} = \frac{\Delta t_0}{\Delta t} = \frac{1}{\gamma} \rightarrow \boxed{L = \frac{L_0}{\gamma} < L_0}$$

2. Ruler w/light source and mirror on train - YF p.1233

RELATIVITY OF LENGTH

Light-year = distance traveled by light in one year

$$1 \text{ ly} = 9.46 \cdot 10^{15} \text{ m}$$



LORENTZ TRANSFORMATIONS

Galilean transformations (NR speed)

$$x = x' + u t$$

$$x' = x - u t$$

$$v = v' + u$$

$$v' = v - u$$

$$t = t'$$

$$t' = t$$

time passes at the same rate

Lorentz transformation (\forall speed) - Lorentz factor

$$\gamma = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

$$x = \gamma(x' + u t')$$

$$x' = \gamma(x - u t)$$

$$y = y'$$

$$y' = y$$

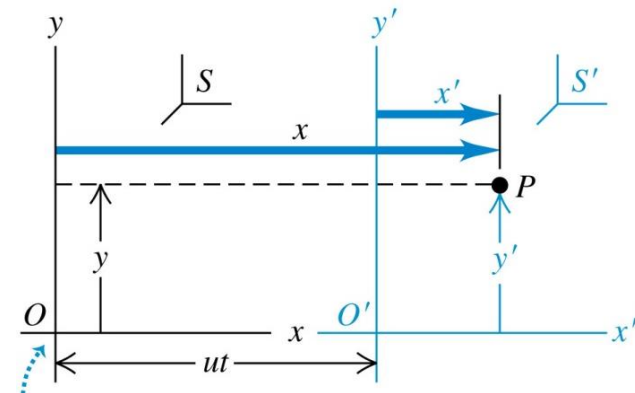
$$z = z'$$

$$z' = z$$

$$t = \gamma\left(t' + \frac{u x'}{c^2}\right)$$

$$t' = \gamma\left(t - \frac{u x}{c^2}\right)$$

proper length



Lorentz velocity transformation

$$v = \frac{v' + u}{1 + \frac{v' u}{c^2}}$$

$$v' = \frac{v - u}{1 - \frac{v u}{c^2}}$$

Notes:

Can be derived from the two relativity postulates

Time dilation and length contraction are ensured / confirmed

Classical limit $\gamma \rightarrow 1$ leads to Galilean transformation

Simultaneity, time dilation, and length contraction are consequences

RELATIVISTIC MOMENTUM + ENERGY

Energy and momentum re-defined so that conservation laws hold in all RFs

1905 Einstein - mass is a form of energy = consequence of theory of relativity

Rest energy

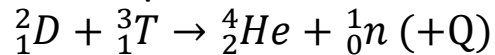
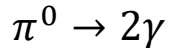
$$E_0 = m c^2$$

energy associated with the mass of an object

Chemistry - energy and mass are conserved separately ($\Delta m/m$ too small)

Special relativity - unifies conservation of mass and energy

Nuclear reactions - large ΔE & can measure Δm



where

fusion

$\gamma = \text{photon}$ (not Lorentz factor)

$$E_i = E_f + Q$$

where $Q > 0$

Examples

1 penny $\sim 78 \text{ GW h}$

few 100 kg matter \sim annual US energy production

MOMENTUM AND ENERGY

Total energy

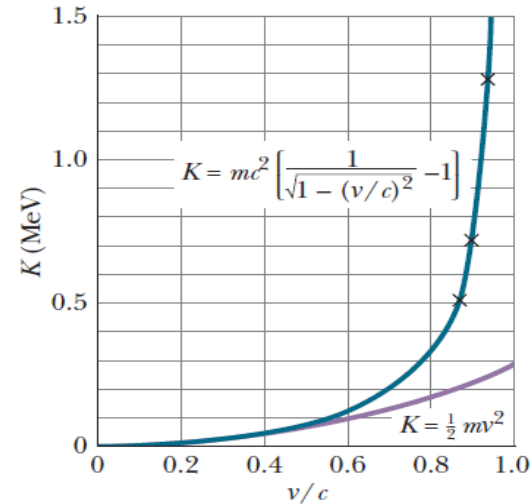
$$E = \gamma m c^2 \quad (\text{w/out proof})$$

$$E = E_0 + K = mc^2 + \gamma mc^2$$

Kinetic energy

$$K = (\gamma - 1) mc^2$$

$$K = (\gamma - 1) E_0$$

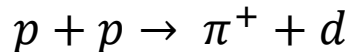


Conservation total energy for Isolated systems

$$E_i = E_f$$

$$\Sigma(E_0 + K)_i = \Sigma(E_0 + K)_f$$

Example: Find the minimum kinetic energy of the protons in a head-on collision to produce a pion and a deuteron (proton+neutron). Assume all masses are known.



MOMENTUM AND ENERGY

Rest energy $E_0 = m c^2$

Total energy $E = \gamma m c^2$

Kinetic energy $K = (\gamma - 1) m c^2 = (\gamma - 1) E_0$

Relativistic momentum $\vec{p} = \gamma m \vec{v}$

Note: $\vec{p} = m \vec{v}$

NR limit $v \ll c$ i.e. $\gamma \rightarrow 1$

Classical

$$p = mv$$

$$K = \frac{1}{2} m v^2$$

eliminate $v \rightarrow K = \frac{p^2}{2m}$

MOMENTUM AND ENERGY

Relativistic

$$p = \gamma m v$$

$$E = \gamma m c^2$$

eliminate γ, v^2

Momentum-energy relation

$$E^2 = (pc)^2 + (mc^2)^2$$

$$E^2 = (pc)^2 + (E_0)^2$$

Notes

1. NR /at rest

$$E \rightarrow E_0$$

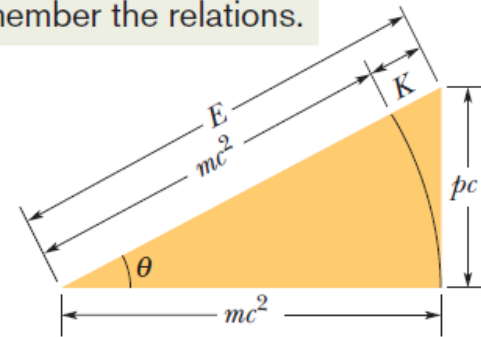
2. Highly relativistic

$$E \rightarrow pc$$

3. $m = 0$

$$E = pc$$

This might help you to remember the relations.



MOMENTUM AND ENERGY

Momentum-energy relation

$$E^2 = (pc)^2 + (mc^2)^2$$

Units

$$[E] = eV$$

$$1eV = 1e \ 1V = 1.6 \ 10^{-19} J$$

$$1 \ keV = 10^3 eV \quad (\text{atomic scale})$$

$$1 \ MeV = 10^6 eV \quad (\text{nuclear scale})$$

$$[m] = eV/c^2$$

OR

$$[m] = u = 1.66 \ 10^{-27} kg \sim \text{mass of nucleon}$$

$$[p] = \frac{[E]}{c} = \frac{eV}{c}$$

This might help you to remember the relations.

