Raytracer

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aytracing accurately computes a 2D representation of a 3D scene. It models the physics behind light interaction with surfaces by tracing rays of lights from a light source to the camera. However, this is not as efficient as rasterizing and hence not commonly used in real-time applications.

1. Compiling and running the code

The project includes a Makefile, which compiles the code. It contains the -03 flag in CC_Opts so that the code is compiled in the most optimised way. The -1X11 flag is added to the linker options LL_Opts to allow communication with the X11 Windows Manager. The flag -fopenmp is added to the CC flags. The last two flags allow for parallelism in the code, which will be elaborated on in Section 3.

The project is compiled and run using the command: $make \&\& .\Build\$ skeleton

2. Parts 1 and 2

The scene is stored as a collection of triangles in a global list. Rays are cast for every pixel on the screen, with their directions given by the offset of their pixels from the center of the screen along with the focal length.

For every ray, the program finds the intersection, if any, with each of the triangles in the scene. These intersections are sorted by distance, such that the closest intersection can be assumed to be with the surface that should be visible to the camera at the pixel's location. The colour of the closest intersected surface's triangle is then used as the pixel's colour.

In order to allow navigation of the scene, the Update method called on every frame was expanded to listen for keyboard input. If the up, down, left or right keyboard keys are pressed the camera's position is translated. This was further expanded with the ability to also rotate the camera, such that the left and right keys rotate the camera in either direction, using a rotation matrix.

A light source with colour and position properties was added to the scene to begin illuminating it with direct light. The direct light reaching any given ray's intersection point and triangle is calculated in the <code>DirectLight</code> method. It calculates the power of the light source as a sphere around its position, with a radius given by the distance of a surface from the light source. Whether a surface is facing the light source is determined using the dot product of the unit normal vector of the surface's triangle with the unit radius of the

sphere. If this value is negative, the surface receives no direct light. This method is then called with every pixel's closest intersection in order to calculate the pixel's direct light. Multiplying the colour of the pixel's closest intersected triangle by the direct light gives its illumination-adjusted colour value.

Further keyboard input functionality was added to the Update method to to enable movement of the light source. The light source can be translated using the following keyboard keys: W, S, A, D, Q and E.

To draw shadows, additional rays are casted in the DirectLight method from the surface intersection point to the light source. By finding the distance of the closest intersected point and comparing it to the distance to the light source, it is possible to determine whether a surface is in the way of the light. If this is the case, the pixel is given zero direct illumination.

Our final implementation for parts 1 and 2 was indirect illumination, to simulate light 'bouncing' between surfaces. This was done using a global constant for indirect illumination and factoring it into each pixel's colour value alongside direct light.

3. Extensions

Cramer's rule for inverting matrices

Using Cramer's rule instead of glm::inverse to invert matrices sped up the render time significantly. The rule is described below:

$$A = \begin{bmatrix} a_{00} & a_{01} & a_{02} \\ a_{10} & a_{11} & a_{12} \\ a_{20} & a_{21} & a_{22} \end{bmatrix} \quad A^{-1} = \frac{1}{\det(A)} \begin{bmatrix} A_{00} & A_{01} & A_{02} \\ A_{10} & A_{11} & A_{12} \\ A_{20} & A_{21} & A_{22} \end{bmatrix}$$

$$A_{00} = \begin{vmatrix} a_{11} & a_{12} \\ a_{21} & a_{22} \end{vmatrix} \quad A_{01} = - \begin{vmatrix} a_{10} & a_{12} \\ a_{20} & a_{22} \end{vmatrix} \quad A_{02} = \begin{vmatrix} a_{10} & a_{11} \\ a_{20} & a_{21} \end{vmatrix}$$

$$A_{10} = - \begin{vmatrix} a_{01} & a_{02} \\ a_{21} & a_{22} \end{vmatrix} \quad A_{11} = \begin{vmatrix} a_{00} & a_{02} \\ a_{20} & a_{22} \end{vmatrix} \quad A_{12} = - \begin{vmatrix} a_{00} & a_{01} \\ a_{20} & a_{21} \end{vmatrix}$$

$$A_{20} = \begin{vmatrix} a_{01} & a_{02} \\ a_{11} & a_{12} \end{vmatrix} \quad A_{21} = - \begin{vmatrix} a_{00} & a_{02} \\ a_{10} & a_{12} \end{vmatrix} \quad A_{22} = \begin{vmatrix} a_{00} & a_{01} \\ a_{10} & a_{11} \end{vmatrix}$$

A check for det(A) = 0 is also added, in which case the inversion step is skipped.

Multiple light sources

In addition to Cramer's rule, the render time was improved by parallelising loops with no data dependencies. This was done with means of OpenMP – pragma omp parallel for was used above all but one of the outermost for loops in the Draw() function.

As mentioned above, the -lX11 is necessary to allow communication with the X11 Windows Manager. The flag -fopenmp and the header #include <X11/Xlib.h> were also needed to allow usage of OpenMP.

The current render time allows for the camera and light movement to be responsive even with a 500×500 screen.

The final touch was adding multiple light sources - of different colours and positions. The structure used is: struct Light {vec3 pos; vec3 colour;}; the main light following it is now named ceilingLight. The new lights are placed on the side walls and simulate fairy lights. While a simple alteration, it is interesting to observe how the lights' colours alter the walls they are positioned on.

References

1. http://www.math.cornell.edu/andreim/Lec17.pdf

Anti-aliasing

In order for an anti-aliasing effect to be achieved multiple rays – instead of just one – are shot for each pixel. The colours for each ray are averaged, and the result is the current pixel's colour.

As expected, the more rays per pixel, the slower the render time gets. After experimenting with different values, 9 ray per pixel seemed to give the best results in terms of gain vs. time efficiency.

Depth of field and Robert's Operator for edge detection

The next extension implemented is depth of field. A global variable focus was created. For each pixel in the 2D scene, the 3D depth of its respective intersection was preserved in the 2D array float intersections [SCREEN_WIDTH] [SCREEN_HEIGHT]. Once the whole image is rendered, the intersection different from focus±slack are blurred. In the submitted version focus = 0.004f and slack = 0.001f. The blurring is achieved via a 5×5 Gaussian filter. If the camera is moved, the pixel in focus change respectively.

As the objects in the scene lack texture, the effect is not always easily noticeable. To ensure it is functioning, Robert's Operator for edge detection was implemented. The resulting image is produced side by side with the original and also responds to camera movements. The scene is first represented in grayscale. The Operator is then applied as follows: $\nabla f = |f(x,y) - f(x+1,y+1)| + |f(x,y+1) - f(x+1,y)|$, which is derived from:

$$h_1 = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix} \quad h_2 = \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix}$$