

Classical MD and water models

MIT
10.637
Lab 3

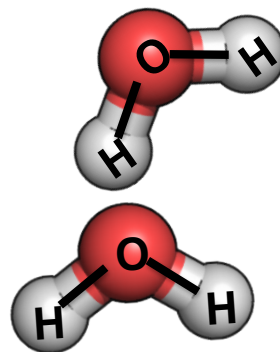
Professor Kulik
hjkulik@mit.edu

Water: simple models

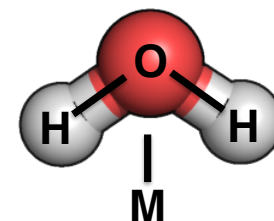
3-site rigid water models:

$$E_{ab} = \sum_i \sum_j \frac{k_C q_i q_j}{r_{ij}} + \frac{A}{r_{OO}^{12}} - \frac{B}{r_{OO}^6}$$

Parameters	TIP3P	SPC/E
$d(\text{O-H})$ (Å)	0.9572	1.0
H-O-H (°)	104.52	109.47
$q(\text{O}) = -2q(\text{H})$	-0.834	-0.8476
A (kcal Å ¹² /mol)	582000	629400
B (kcal Å ⁶ /mol)	595	625.5



Extension to 4 sites: charge is off-center



Parameters	TIP4P
$d(\text{O-H})$ (Å)	0.9572
H-O-H (°)	104.52
$d(\text{O-M})$ (Å)	0.15
$q(\text{M}) = -2q(\text{H})$	-1.04
A (kcal Å ¹² /mol)	600000
B (kcal Å ⁶ /mol)	610

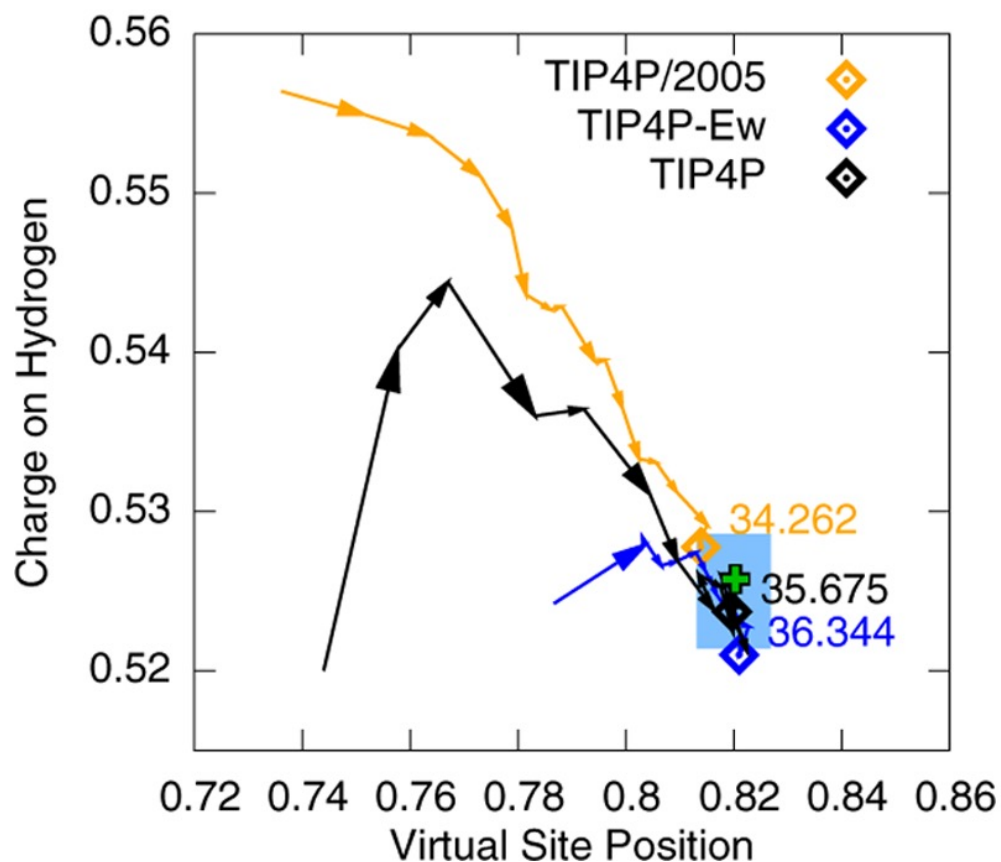
Flexible SPC (SPC/Fw): O-H stretch and bend added.

Note: NO such thing as flexible TIP3P, etc!

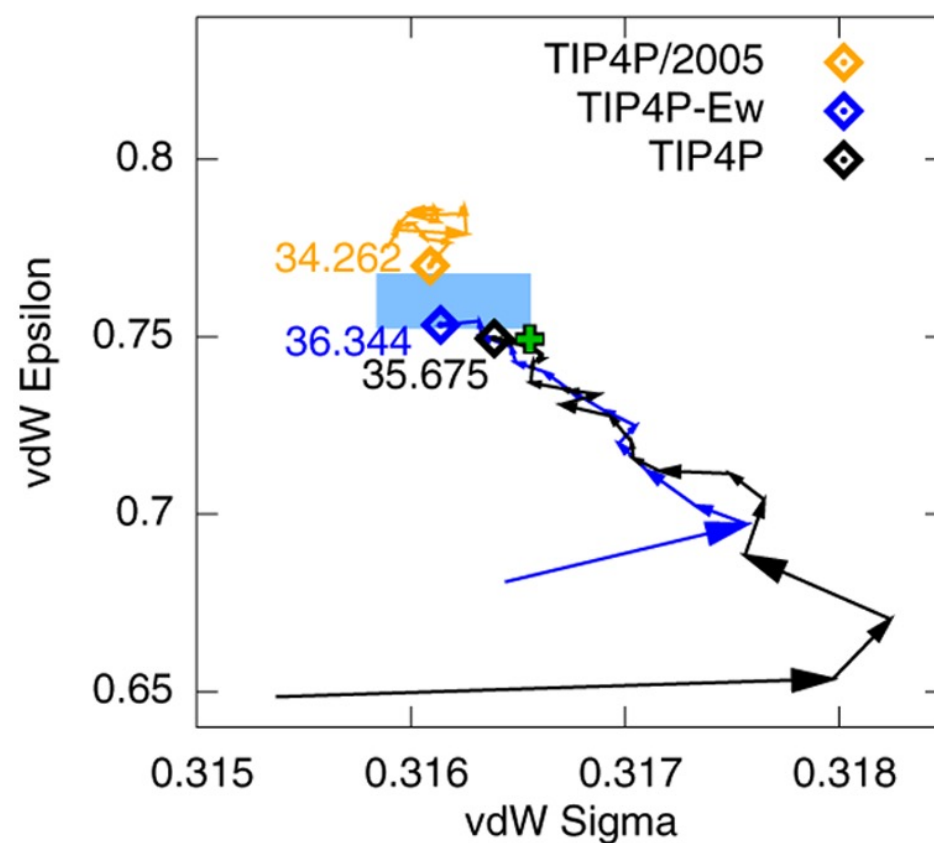
5- and 6-site water models: charges on lone pair sites + dummy atom.

TIP4PFB: an optimized TIP4P

Optimization History for Charge and V-Site



Optimization History for Lennard-Jones



Wang et al., JPC Lett (2014).

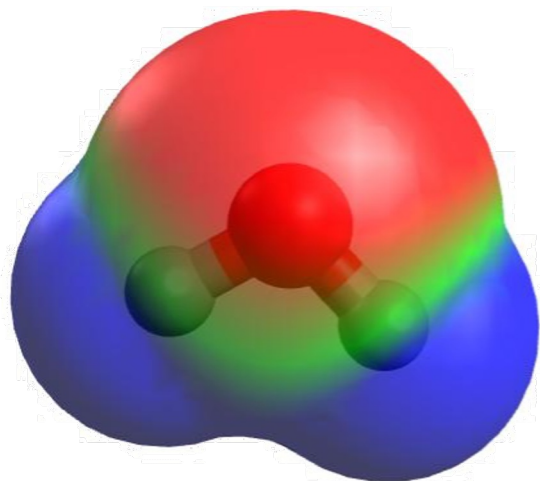
Electrostatic energy

Coulomb interactions between atoms A and B with partial charges:

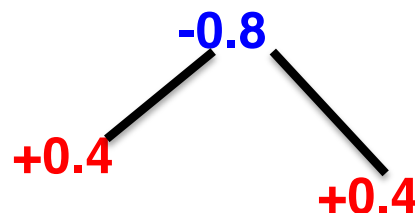
$$E_{\text{el}}(R^{AB}) = \frac{q_A q_B}{\epsilon_{AB} r_{AB}}$$

ϵ is Effective dielectric constant, typically 1 in vacuum.

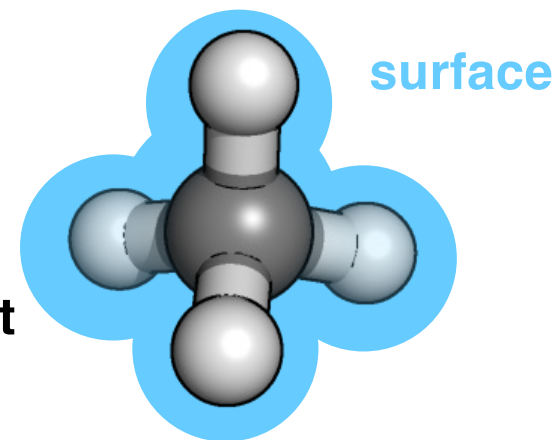
**Quantum-mechanical
electrostatic potential**



**Reproduce with point
charges**



**Uncertainty,
need for restraint:**

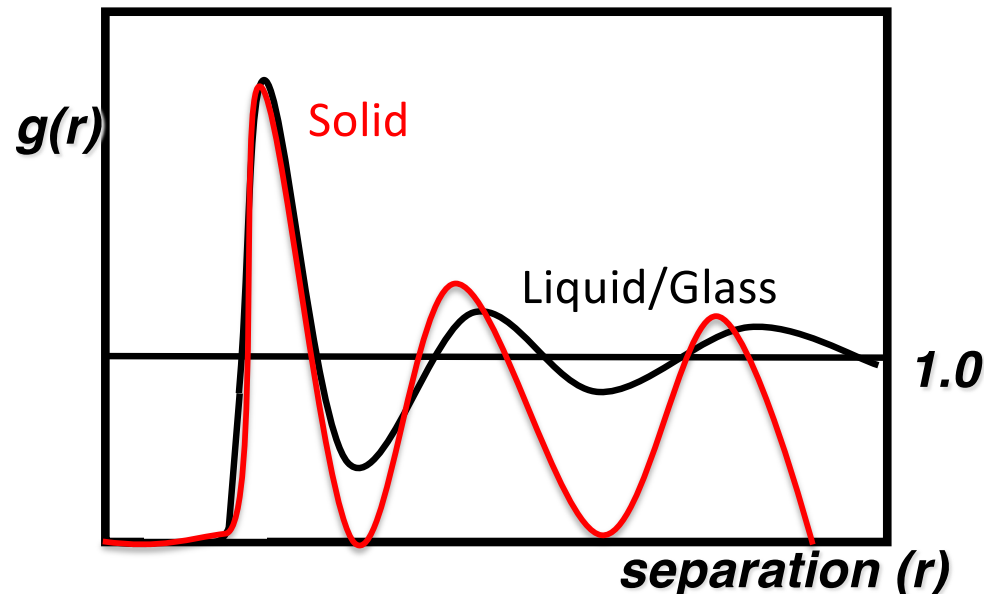
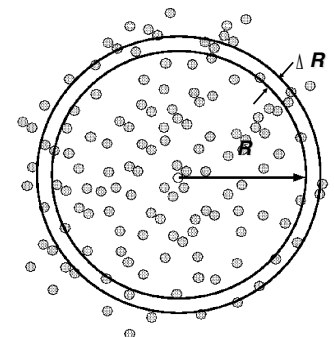


RESP: restrained
electrostatic potential
charges.

Properties from MD runs

Radial distribution function: $g(r) = \frac{1}{N} \frac{dn(r)}{4\pi r^2 dr \rho}$

Probability of finding two species near each other



What we learn:

- Integrate and get coordination number (e.g., Mg^{2+} is coordinated by 6 water molecules for the life of most MD simulation - microseconds).
- Structure: a solid will have zero probability between coordination shells, glass/liquid has diffusion.
- Can compare to experiment.

Properties from MD runs

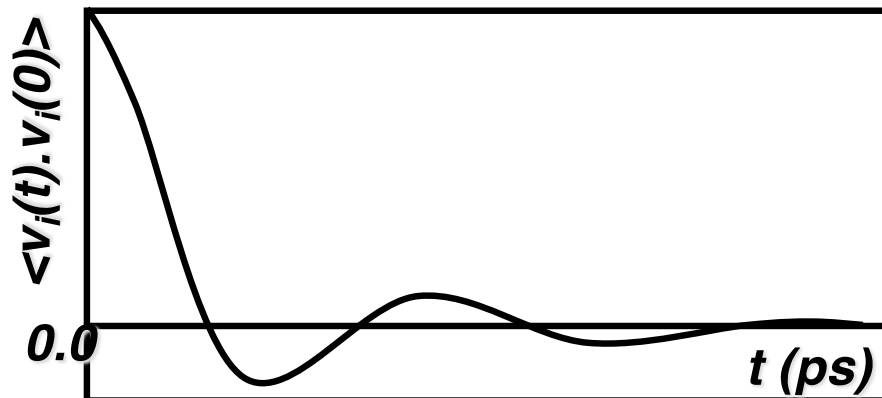
Autocorrelation functions:

$$C_{AA} = \frac{1}{N} \sum_{i=1}^N \frac{\langle A_i(t) A_i(0) \rangle - \langle A_i(0) \rangle^2}{\langle A_i(0) A_i(0) \rangle - \langle A_i(0) \rangle^2}$$

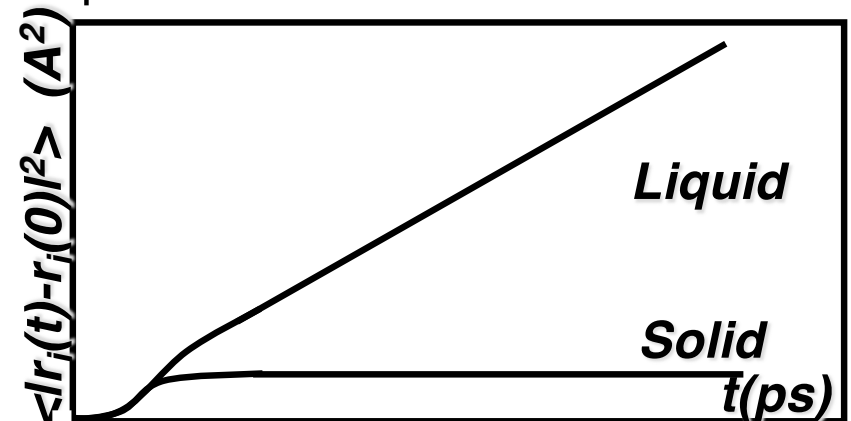
Autocorrelation functions (ACFs) can be defined and calculated for any particle quantity (e.g. v_i) or any system quantity (e.g. U , T , P , r). Starts at 1 and decays usually exponentially with time.

Diffusion coefficient:
$$D = \frac{1}{3N} \int_0^\infty \left\langle \sum_{i=0}^N \mathbf{v}_i(t) \mathbf{v}_i(0) \right\rangle dt$$

Velocity ACF integrated over time can give diffusion coefficient:



Mean squared displacements for phases of matter:



AMBER nuts and bolts

AMBER dat files + Online repository:
force field files, prepin files, to describe proteins, nucleotides, water, and ions.

antechamber/REDS server:
generates force field and prepin files for non-standard ligands and residues

We've done these first two steps for you for water, but you will do these for MgCl_2 salt solutions (time permitting)...

tleap: organize and load in force field properties, solvate, set box dimensions, neutralize, and generate topology/coordinate files

sander or pmemd.cuda: run minimizations and molecular dynamics using topology, coordinates, and input parameters.

cpptraj/ptraj: process and analyze output of trajectories.

Reminder to log on to XSEDE

Make sure you can do the following:

```
ssh <access username>@login.expense.sdsc.edu
```

If you have never set a password for Expanse, reset your password on the CILogon page on access-ci.org

Note: We are using a special GPU version of the main sander driver called “pmemd.cuda”. In job scripts, we have specified this executable and necessary dependent modules.

Expanse SLURM queue

- Comet uses SLURM as a queueing system.
- We write a *job script* and then *submit* it to the *queueing system* to run when resources are available.
- Example job script:

```
#!/bin/bash
#SBATCH -J watermd
#SBATCH -o equil.out
#SBATCH -partition=gpu-shared
#SBATCH -t 1:00:00
#SBATCH -gres=gpu:1
#SBATCH -ntasks-per-node=1
#SBATCH -A itm101

module load slurm
module load gpu/0.15.4 openmpi/4.0.4
module load amber/20

#Example run command
./exec < input > output
```

Expanse SLURM queue

- We issue commands to the SLURM queue to manage our jobs.
- Submit the job script:
`sbatch <jobscript>`
- Submit a job that depends on another one:
`sbatch -d <slurm job #> <jobscript>`
- Check the status of all of your jobs
`squeue -u `whoami``
- Check the status of everyone's jobs
`squeue`
- Check specifics of a running or queued job
`scontrol show job <slurm job #>`
- Cancel a job
`scancel <slurm job #>`

Parts of today's lab

Sections A and B require a lot of compute resources and should be prioritized:

A. Simulating boxes of water
(TIP3P/TIP4PFB)

B. Running MgCl_2 salt solutions

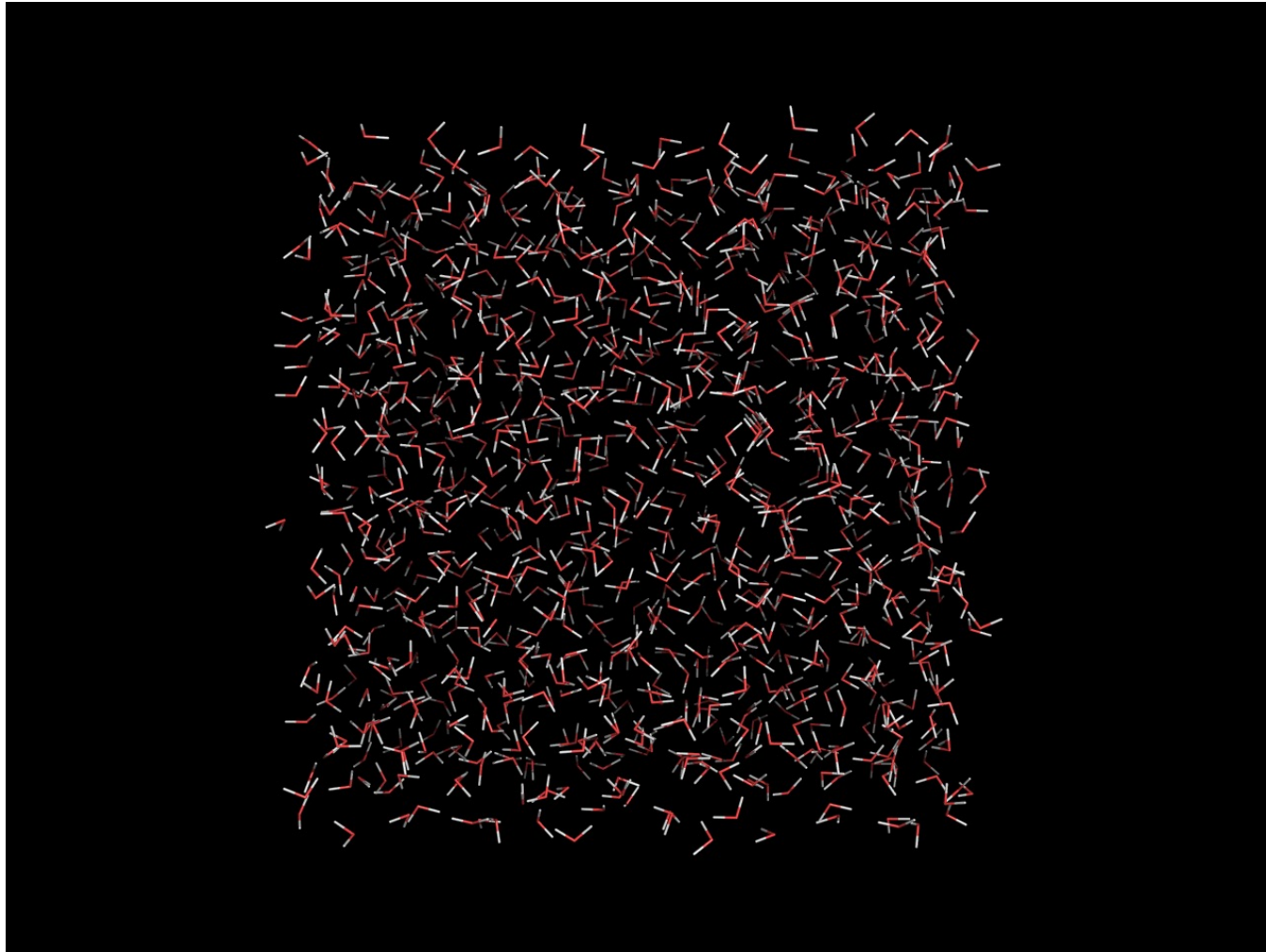
C. Typical charges in water and molecular simulations

D. Processing MD results with cpptraj (can be done after class)

Steps for water equilibration (TIP3P & TIP4PFB)

1. Edit “??”, in files as discussed in the worksheet, e.g.:
`nano constV_equil.in`
2. You will run the minimization, NVT, and NpT equilibration schemes.
3. Submit the queue script
`sbatch amber_equil.q`
4. Check on your job
`squeue -u <username>`
5. Come back to the script to check that everything worked – once you get a spot in the queue, it should take about **5** minutes.

The initial water box



The density will be too low! An important part of our equilibration scheme

What's needed for a calculation

Topology file (force field parameters/how atoms are connected)

tip3p.prmtop

Coordinate file (initial positions of atoms, could contain velocities)

tip3p.inpcrd

Equilibration queue script

amber_equil.q

Input files for each step of equilibration in AMBER

unrestrained_min.in

constV_equil.in

constP_equil.in

constP_run.in

Input file for production in AMBER

constP_production.in

Production queue script

amber_prod.q