



# First signals

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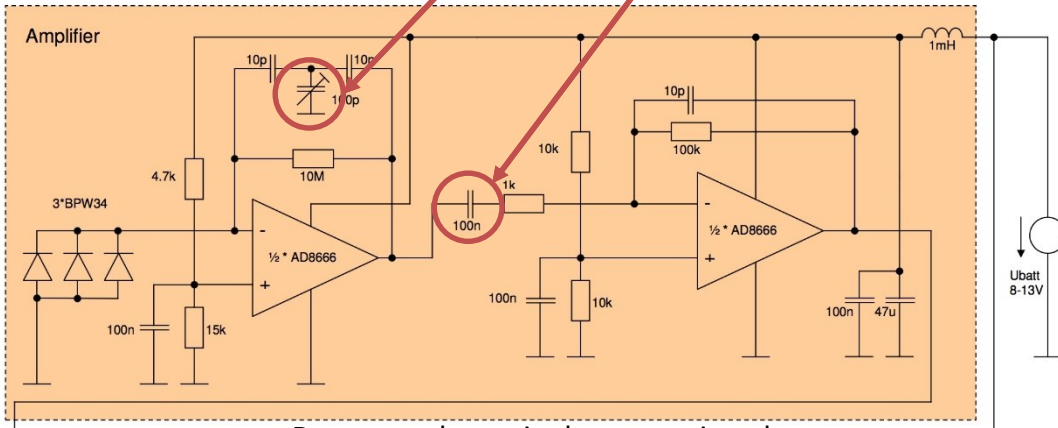
# Better understanding the opengeiger circuit

## Trimming capacitor

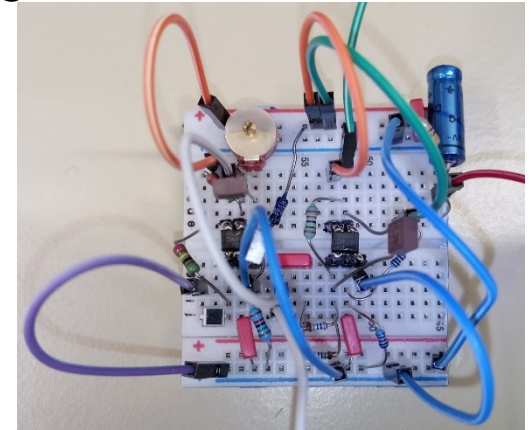
- Smooths the amplification towards ground
- Recommendation by Texas Instruments

## Capacitor between op-amps

- Cuts the DC component of the signal and differentiates the signal

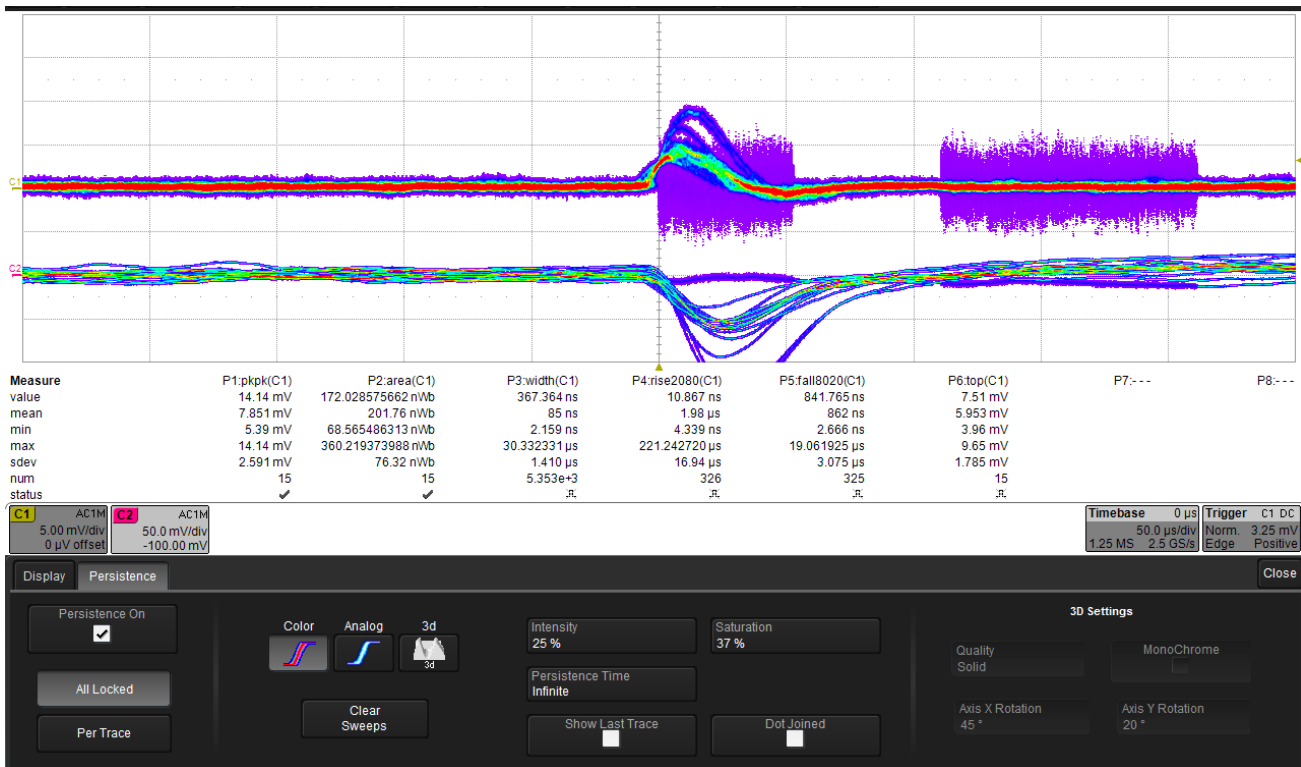


Detector schematics by opengeiger.de



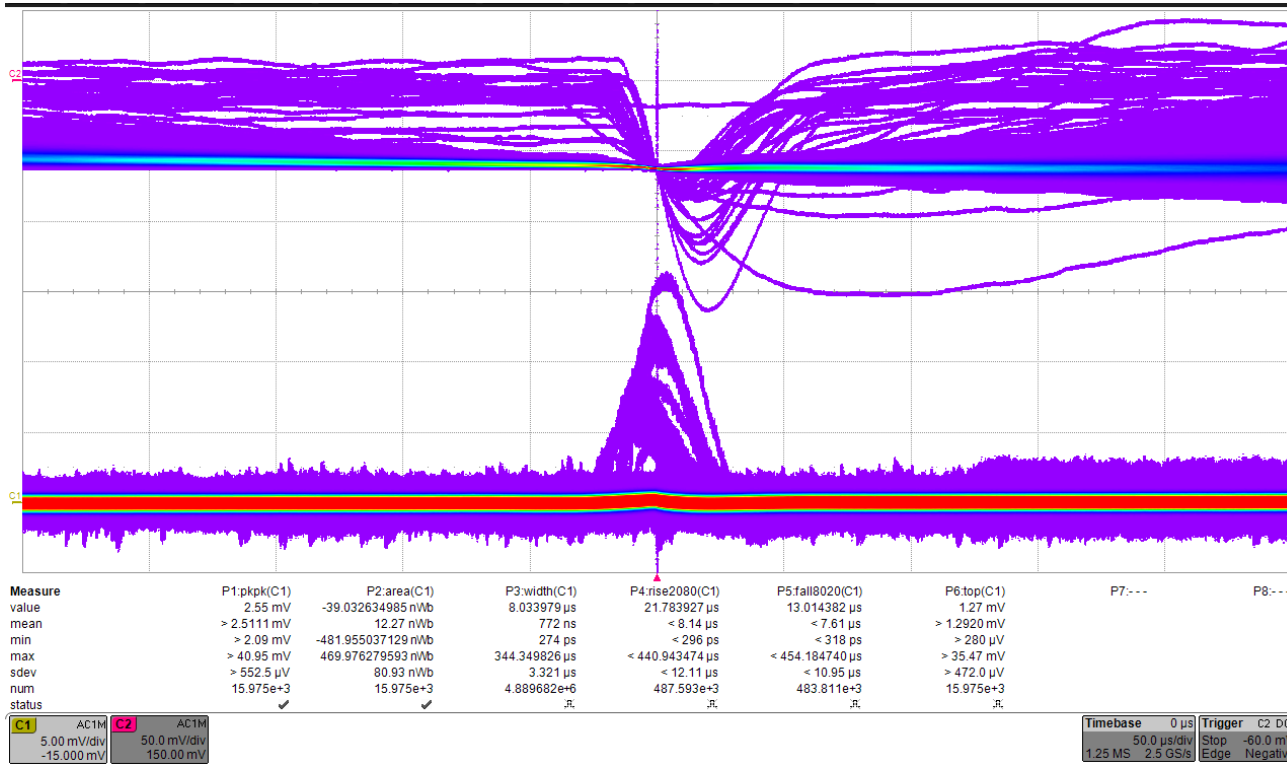
Breadboard circuit

# Results with the opengeiger circuit



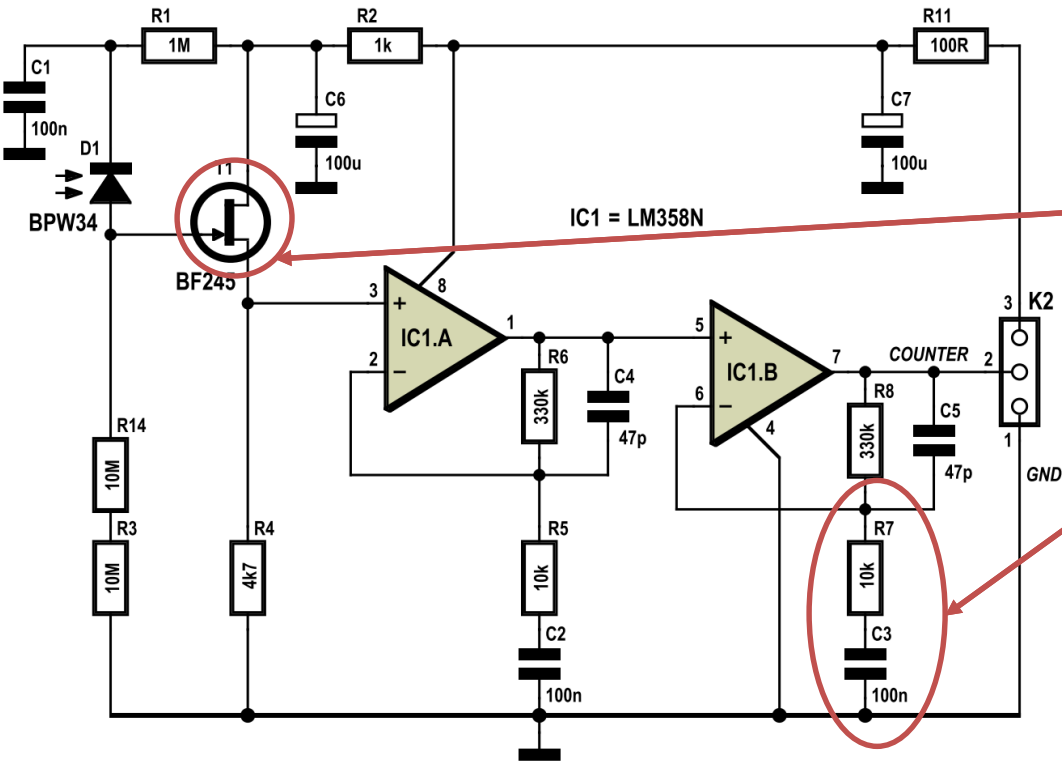
- Signals
  - C1: First op-amp
  - C2: Second op-amp
- Pulse characteristics
  - Height: 50~100 mV
  - Length: ~50  $\mu$ s
- Second op-amp filters HF noise
- Shows coincidences with the CosMo detector

# Results with the opengeiger circuit



- Signals
  - C1: First op-amp
  - C2: Second op-amp
- Second channel shows significant low-frequency noise
- Noise improvements should be possible with the trimming capacitor
- Osram and Vishay diodes show similar characteristics

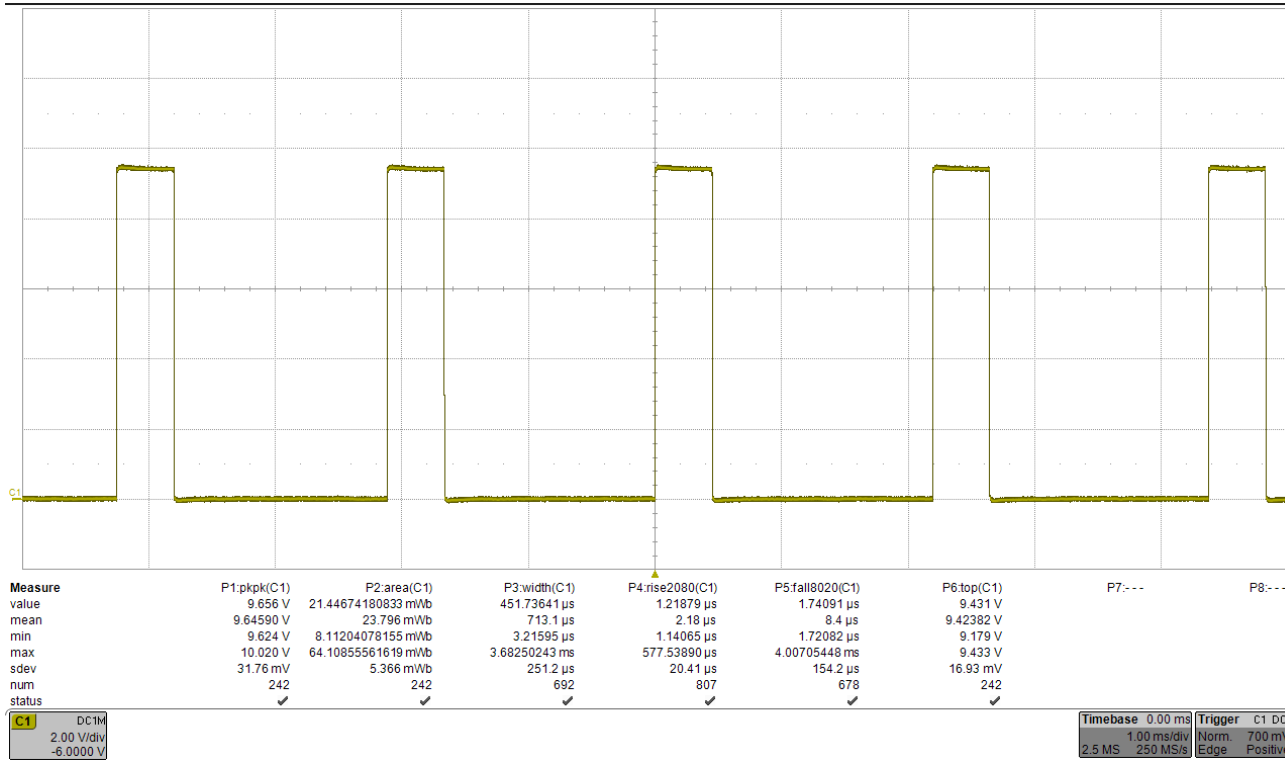
# Better understanding Burkhard Kainkas circuit



Detector schematics by Burkhard Kainka

- JFET for pre-amplification
- Low-end op-amps for further amplification
- Coupling towards ground for noise reduction
- No signal inversion
- Same circuit is on the  $\mu$ Telescope
- **Circuit is inherently unstable** (see Michaels calculations)

# Results with Kainkas circuit



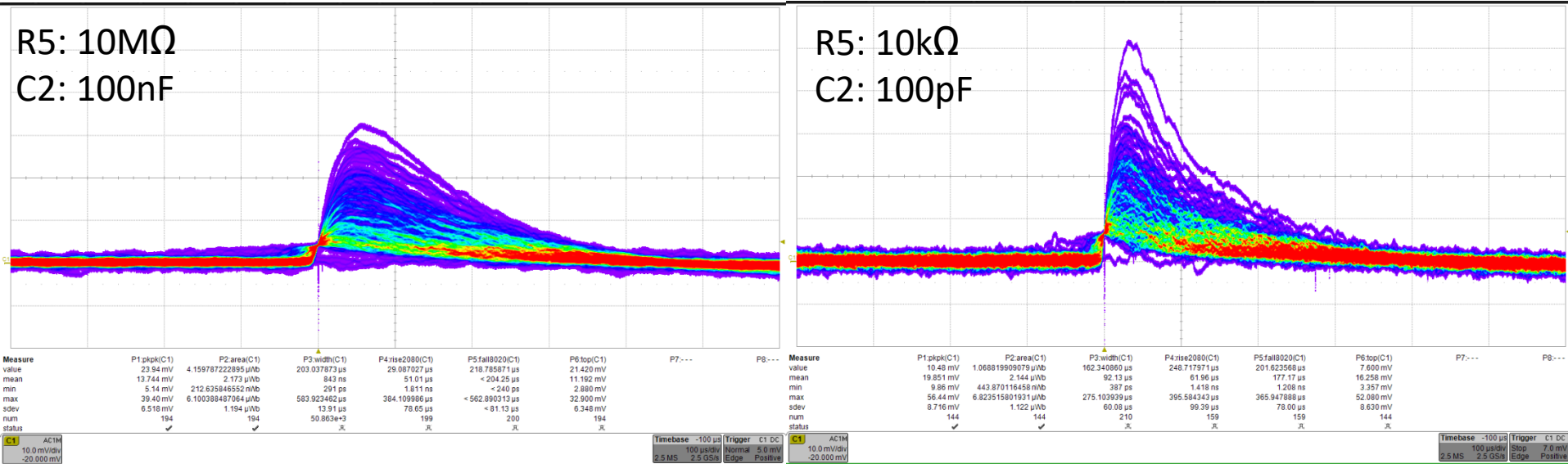
- Signals
  - C1: Second op-amp
- Internal oscillations visible
- Start as soon as light is removed
- Induced via the coupling towards ground and picked up at the gate of the JFET
- Can be suppressed by changing capacitor and resistor values at the noise reduction circuits (R5, R7, C2, C3)

# Results with Kainkas circuit



- Signals
  - C1: JFET output
  - C2: Second op-amp
- Stopping the internal oscillation:
  - Change C2:  $\geq 1\text{mF}$  or  $\leq 100\text{pF}$
  - Change R5:  $\geq 10\text{M}\Omega$
- Pulse characteristics
  - Height: 10~20 mV
  - Length:  $\sim 150\text{ }\mu\text{s}$

# Results with Kainkas circuit



- Very high resistance
- Strength of the ground coupling is reduced
- Visible reduction of noise and signal

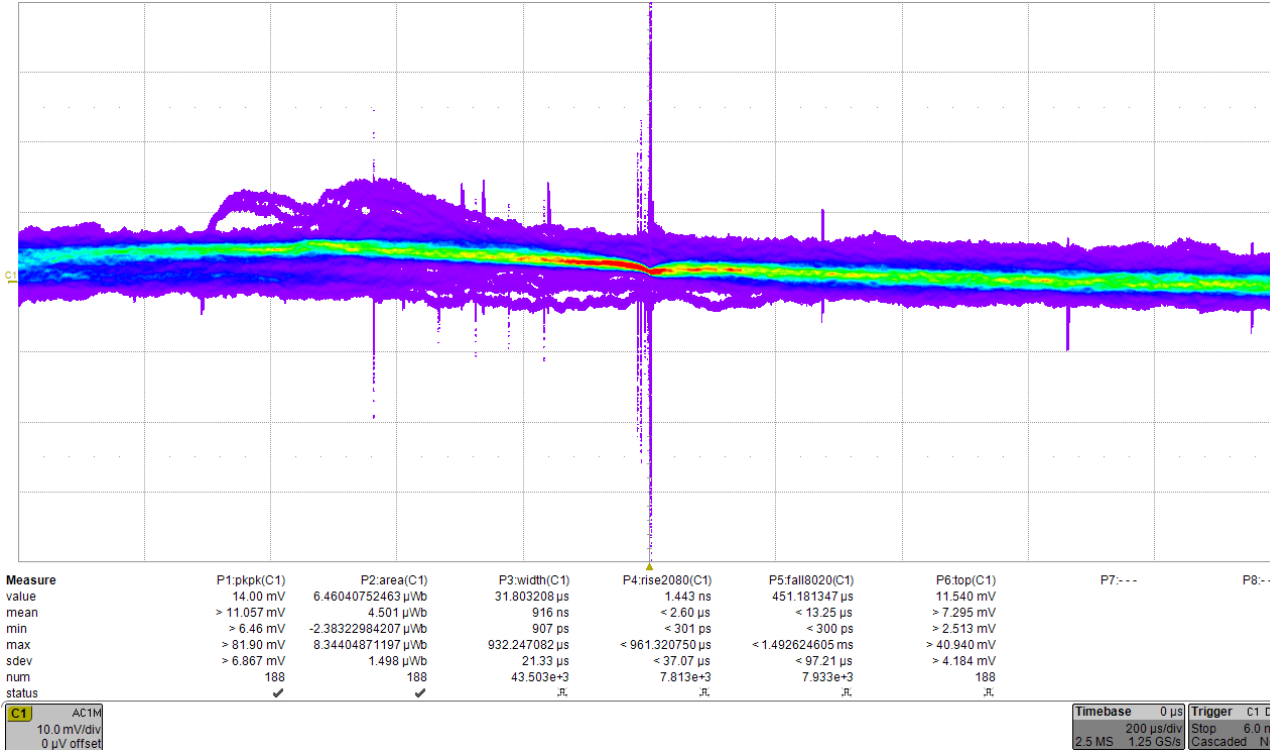
- Very low capacitance
- Close to no effect from the noise reduction



# Optimizing Kainkas circuit

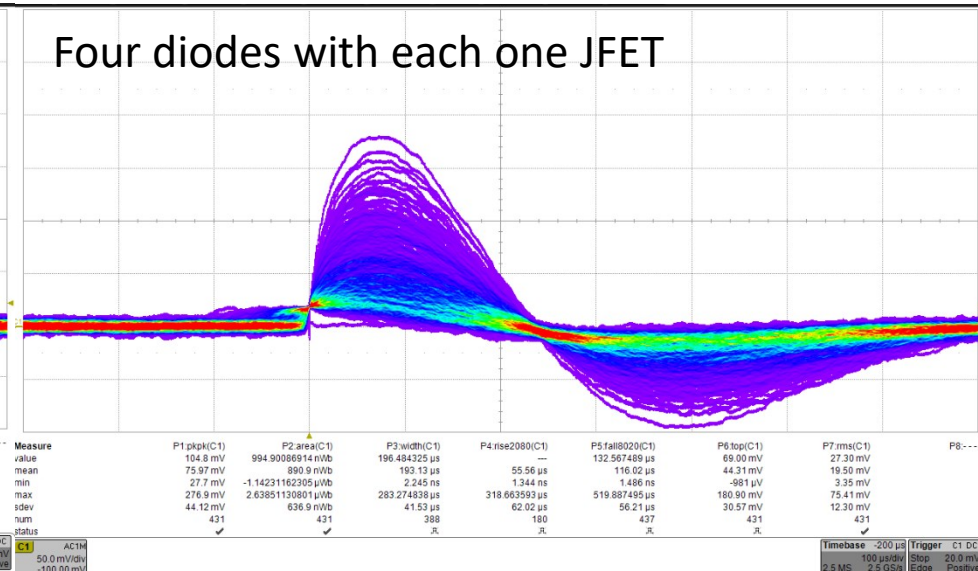
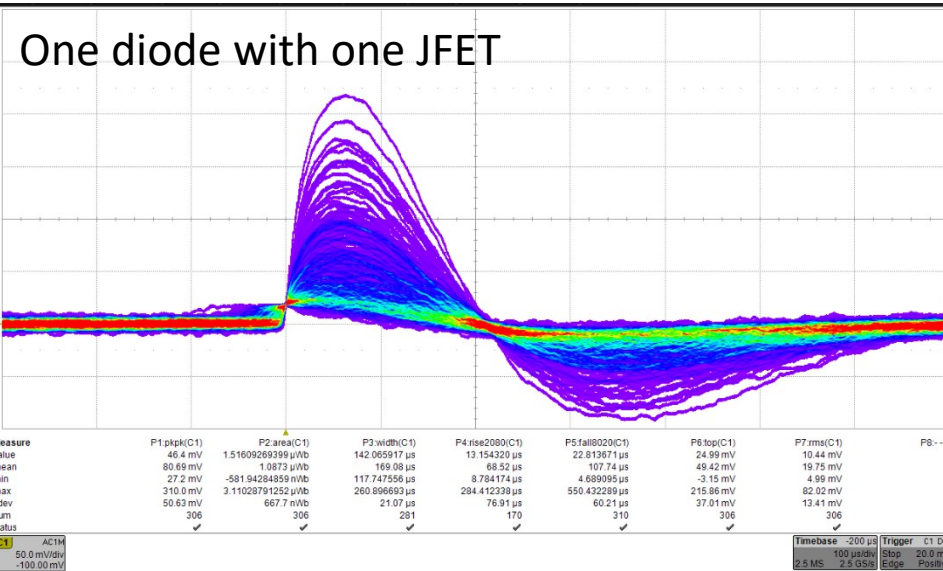
- Tuning the oscillation vs signal strength
  - Tunable parameters: R5, R7, C2, C3
  - Resistors are easily tunable via potentiometers
- Tuning characteristics
  - Non linear signal reaction to changes in R5 or R7
  - Reducing the resistance from R5=10k $\Omega$  and R7=10M $\Omega$  increases the signal strength, as well as signal undershoot
  - Minimum for R7 = 200k $\Omega$  (and R5 at default)
  - Minimum for R5 = 1k $\Omega$  (and R7 at default and 500k $\Omega$ )
- Good results were found at: R5 = 1k $\Omega$  and R7 = 500k $\Omega$ 
  - Pulse height: ~50 mV

# Increasing the detector area



- Signals
  - C1: Second op-amp
- Two diodes in parallel
- Almost halved peak height
- Noise stays the same
- Signal close to vanishes in the background
- On  $\mu$ Telescope: Four diodes in parallel are likely to have a very small signal

# Increasing the detector area



- Very clear signal
- Strong undershoot
- Noise:  $\pm 15$  mV
- Events/min:  $4.08 \pm 0.23$  (SR90 source)

- Slightly lower signal
- Noise:  $\pm 15$  mV
- Events/min:  $3.91 \pm 0.19$  (SR90 source)

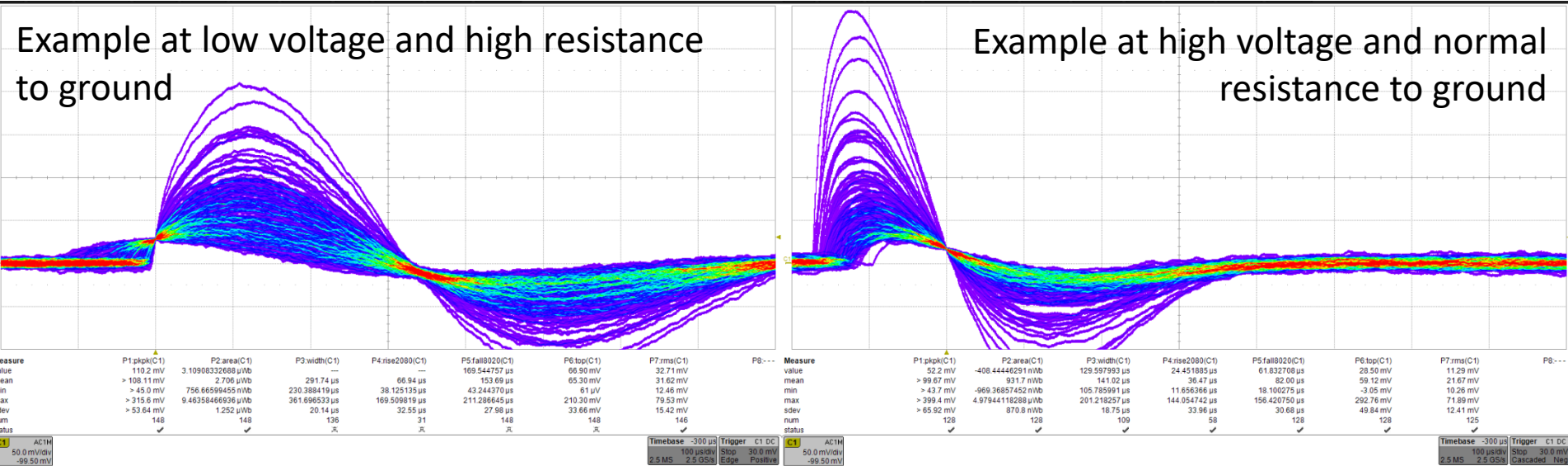
# Conclusions on increasing the detector area

## Tested configurations

- Multiple diodes in parallel
  - Heavily degrades the signal
  - Makes decent triggering impossible
- Multiple diodes with each one JFET in parallel
  - Signal doesn't degrade
  - Can't provide the expected count rate
- All used diodes and JFETS were tested for their performance
  - No problems were found and all have similar performance

Because the whole circuit is very cheap it would still be possible to have one amplification circuit per diode.

# Measuring the capacitance of the PIN-diodes



- Reverse bias = 6.4V
- Resistor to ground R3 = 30MΩ
- Diode capacitance: ~19 pF

- Reverse bias = 26V
- Resistor to ground R3 = 10MΩ
- Diode capacitance: ~14 pF
- Slightly higher signal

# Next steps

- Tomorrow 10:30: Partial PCBA (Printed Circuit Board Assembly) of two channels on the  $\mu$ Telescope
  - Channel 1: LM358
  - Channel 2: AD8666
- Test the maximum possible number of diodes on the PCB
- Simulate modifications to the circuit
- Design and build new PCB