

# Installing grtrans

- `git clone https://github.com/jadexter/grtrans`
- Instructions in `grtrans_tutorial.pdf` file, briefly here
- Install `cfitsio`
- Install `pyfits`
- `cp Makefile.top.sample Makefile.top` and edit `Makefile.top`
- `make`

# Install cfitsio

- `wget http://heasarc.gsfc.nasa.gov/FTP/software/fitsio/c/cfitsio3390.tar.gz`
- `cd cfitsio`
- `./configure; make -j4; make install`
- `cd ..`

# Compile grtrans

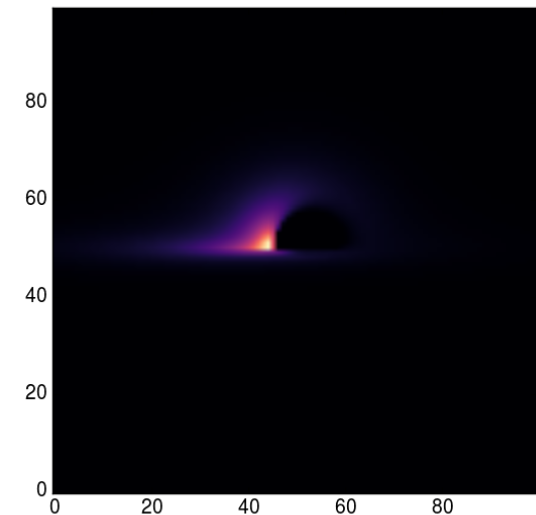
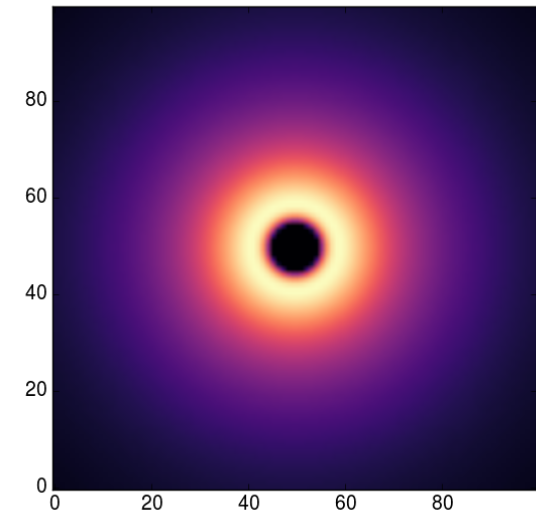
- `cp Makefile.top.sample Makefile.top`
- Edit `Makefile.top` to include paths to `grtrans` and `cfitsio` directories
- `make -j4`

# grtrans via python (pgrtrans)

- Best supported / easiest way is to use python 2.7 to run grtrans
- Alternative: inputs.in input file, run grtrans from command line, write to FITS format output
- For python: install pyFITS, pip install pyfits
- Also requires numpy, matplotlib

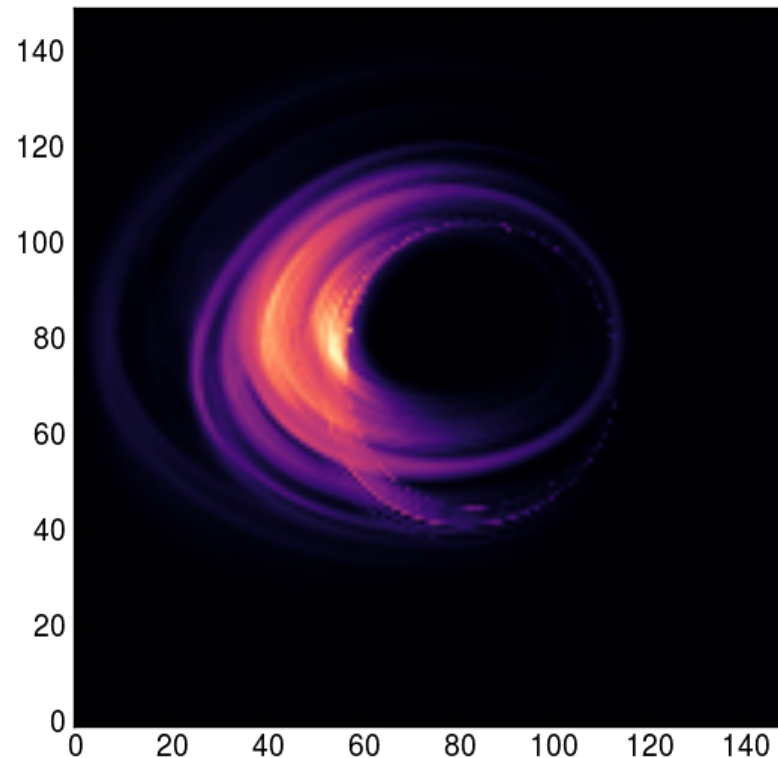
# Running pgrtrans

- in grtrans directory, ipython
- import grtrans\_batch as gr
- x = gr.grtrans()
- x.run\_pgrtrans()
- x.disp\_pgrtrans\_image(-1)
- x.disp\_pgrtrans\_image(0)
- Default: thin disk total intensity images vs. inclination



# A HARM image

- `x.run_pgrtrans(standard=1,nvals=1,fname='HARM',nfreq=1,fmin=2.3e11,fmax=2.3e11,ename='POLSYNCHTH',spin=0.9375,nn=[150,150,400],uout=0.04,mbh=4e6,mdotmin=4e15,mdotmax=4e15,nmdot=1,nmu=1,mumin=.6428,mumax=.6428,gridvals=[-13.,13.,-13.,13.],fhfile='dump040',fdfile='dump',findf=40,fnt=1,muval=1./4.)`
- `x.disp_pgrtrans_image(0)`

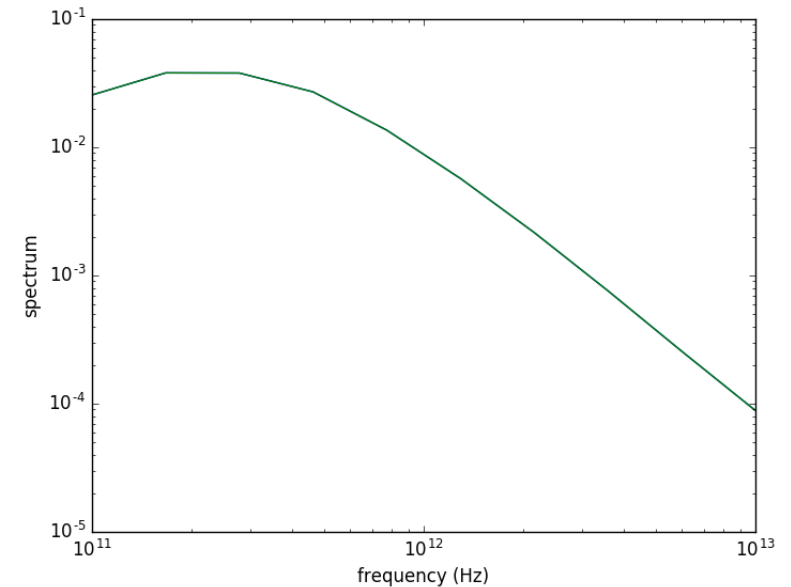


# A HARM image

- inputs:
  - standard: 1 is normal ray tracing, 2 is equatorial plane only
  - nvals: 1 for Stokes I, 4 for full polarization
  - fname: fluid model, analytic or simulation
  - fmin,fmax,nfreq: observed  $\nu$  (Hz) for nfreq frequencies
  - ename: emissivity (mostly various synchrotron emissivities)
  - spin:  $a/M$  in Kerr
  - mbh: black hole mass in  $M_{\text{sun}}$
  - nn: camera resolution # of pixels (x,y), # of points along each ray
  - uout:  $1/r_{\text{out}}$ , the maximum radius used in rad. trans. calculation
  - mumin, mumax, nmua:  $\cos(\text{inclination})$  for nmua angles
  - mdotmin,mdotmax,nmdot: same for mdot
  - gridvals: camera size in units of  $M$ , [xmin, xmax, ymin, ymax]
  - fhfile, fdfile, findf: fluid model specific inputs (e.g. dump file name, header file name)
  - fnt: = 1 “fast light”, > 1 load fnt files than one dump file for slow light
  - muval: constant fraction of internal energy in  $e^-$ ,  $1 / (T_p/T_e + 1)$
- (good / bad): Many others!

# Calculating spectra

- re-run,  $\text{nfreq} = 10$ ,  $\text{fmin} = 5\text{e}10$ ,  $\text{fmax} = 1\text{e}13$
- `x.calc_spec_pgrtrans(x.nx)`
- `import matplotlib.pyplot as plt`
- `plt.loglog(x.nu,x.spec)`





# HARM polarization maps

- same as before, but change  
nvals = 4
- x.disp\_pol\_map(0)
- Now set Faraday  
coefficients = 0: add input  
emiscoefindx=[1,1,1,1,0,0,0]
- x.disp\_pol\_map(0)