

# **Physical conditions in intermediate-energy heavy-ion collisions: Estimates from JAM**

**Hidetoshi Taya  
( Keio U. )**

Based on: [\[HT\]](#), Nishimura, Ohnishi, 2402.17136]

[\[HT\]](#), Jinno, Kitazawa, Nara, 2409.07685]

[\[HT\]](#), 2501.18171]

# Plan of the talk

Messages:

- (1) What is the **BEST** energy range to study high baryon-density physics ?  $\Rightarrow \sqrt{s_{NN}} = 3 \sim 5 \text{ GeV}$
- (2) Such an intermediate-energy regime is also **EXTREMELY** interesting for studying **physics of strong electromagnetic field**

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2. Best energy range for high density physics
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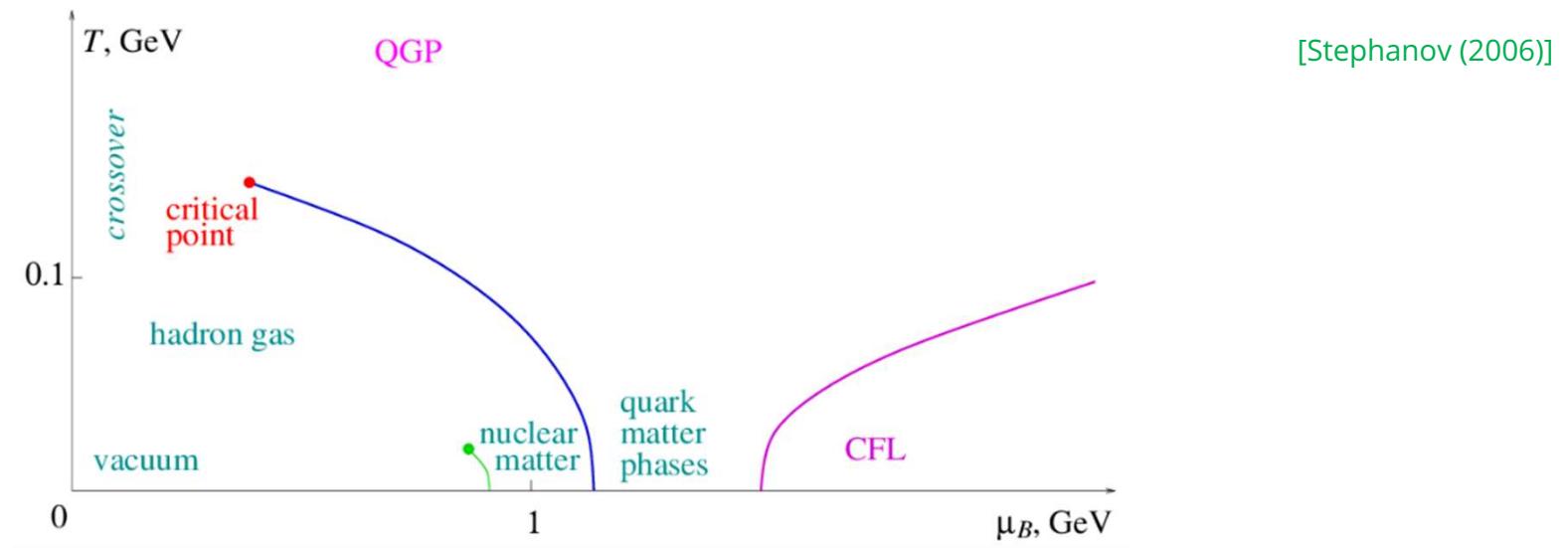
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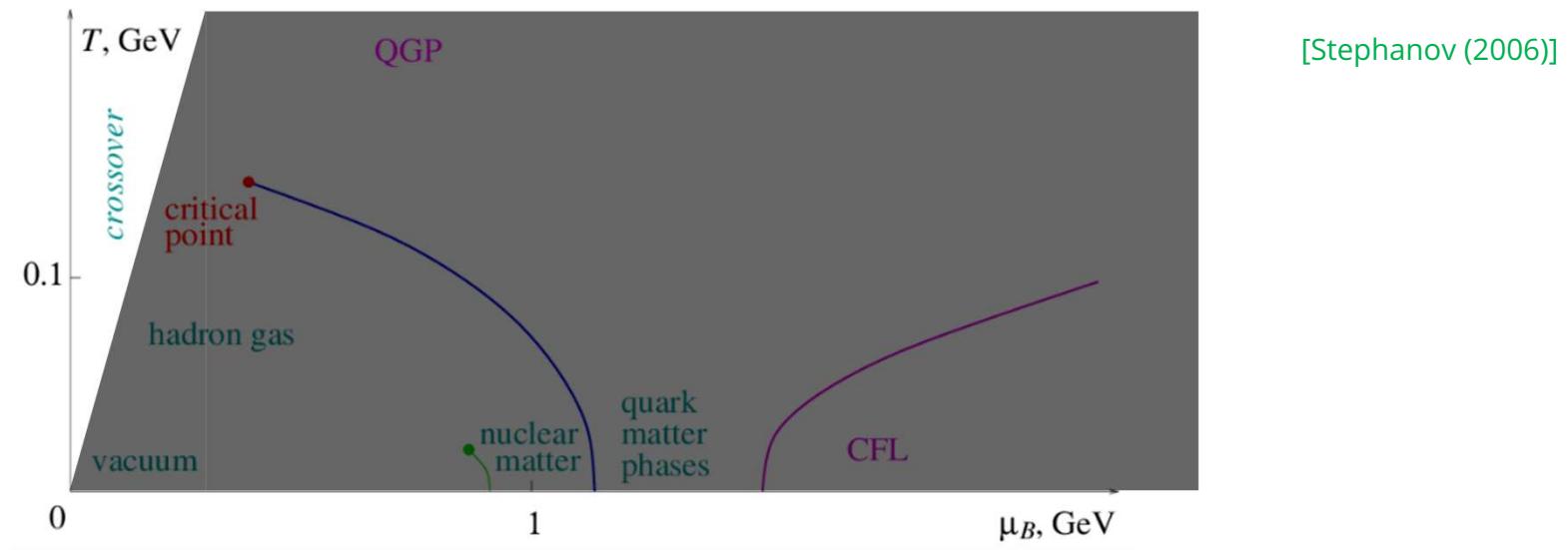
# Q: Matter under extreme condition

- ✓ Conditions: Temperature, density, ...
- ✓ Current expectation: QCD phase diagram



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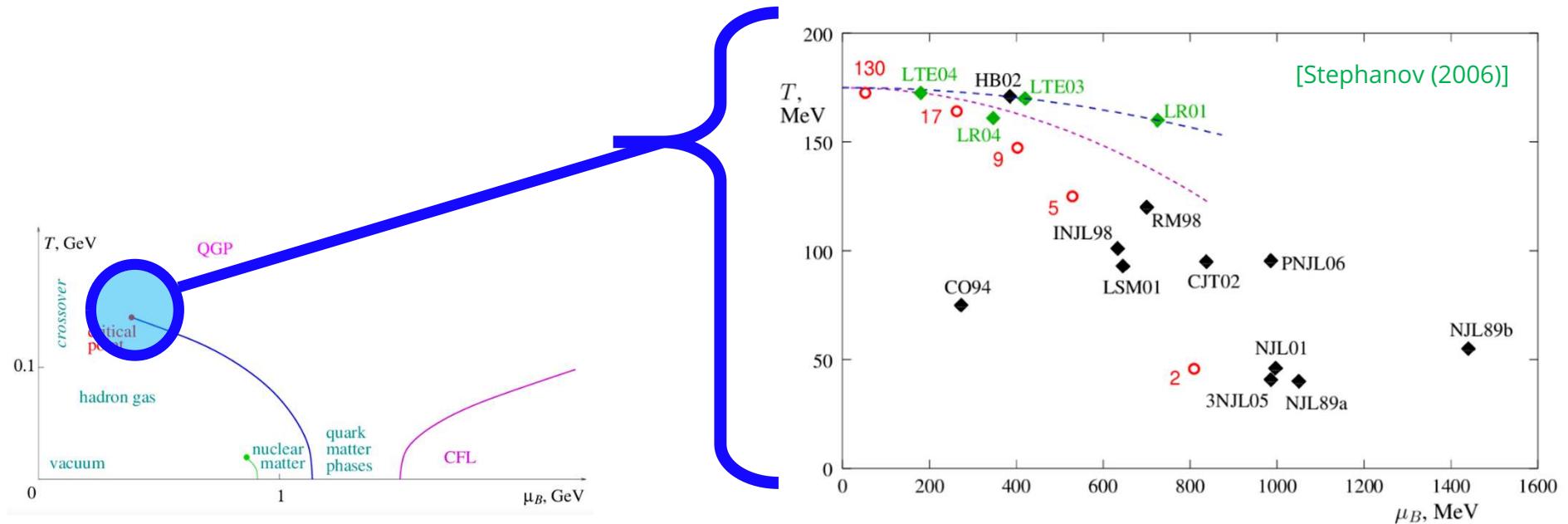
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- Certain for the left (low density region)  
     $\Leftarrow$  theory: lattice QCD exp.: high-energy heavy-ion coll.
- Huge uncertainties in the right (high density region)     [See also Yamamoto's talk]  
     $\Rightarrow$  need exp. Inputs: **intermediate-energy heavy-ion collisions**

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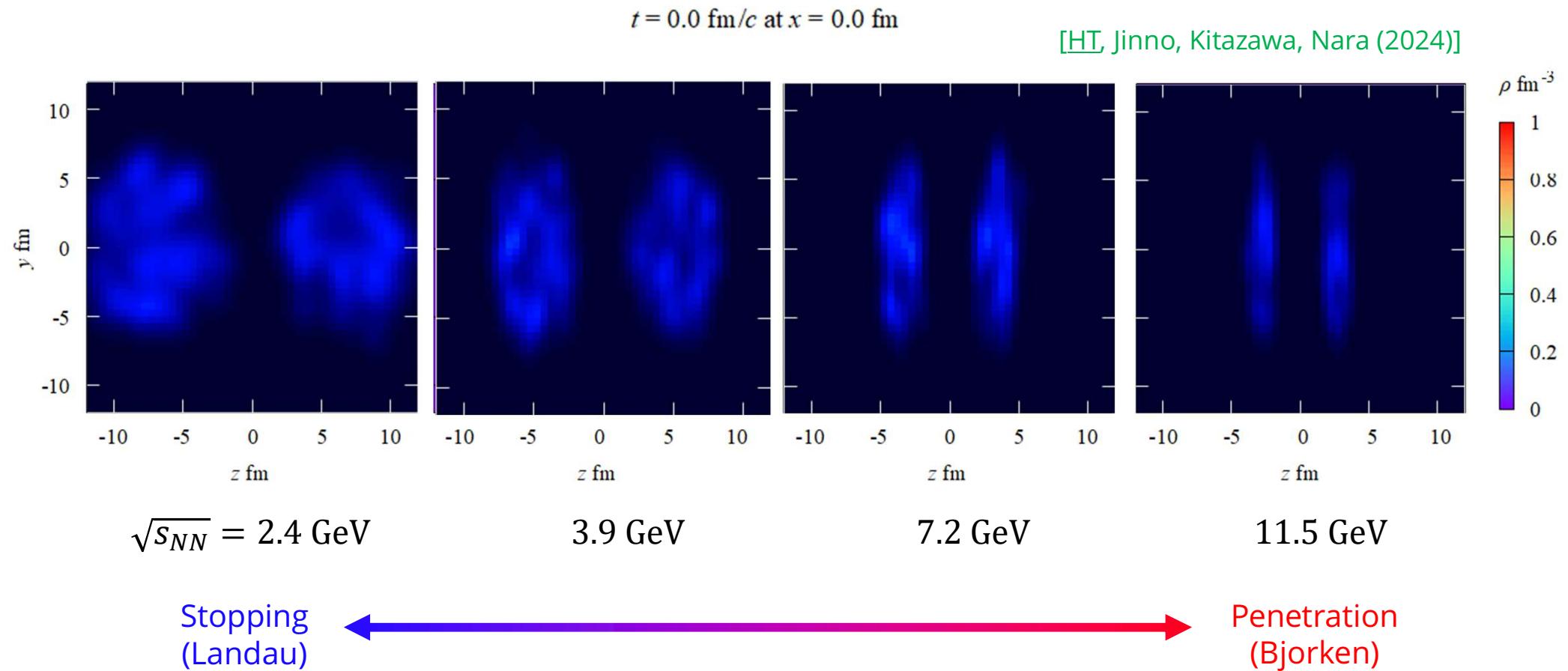
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# Intermediate-energy HIC

✓ Heavy-ion collisions at  $\sqrt{s_{NN}} = O(2 - 10 \text{ GeV})$  [AGS, SPS, RHIC BES, FAIR, NICA, HIAF, J-PARC-HI, ...]  
expected to be the only means to create “high density matter” in lab.



Idea: baryon stopping at lower energies  $\Rightarrow$  dense matter [See also Wolf's talk]

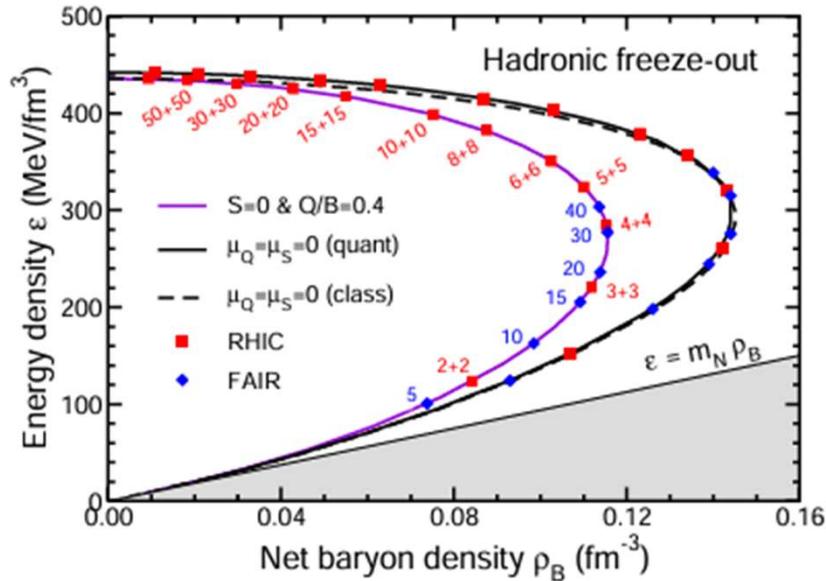
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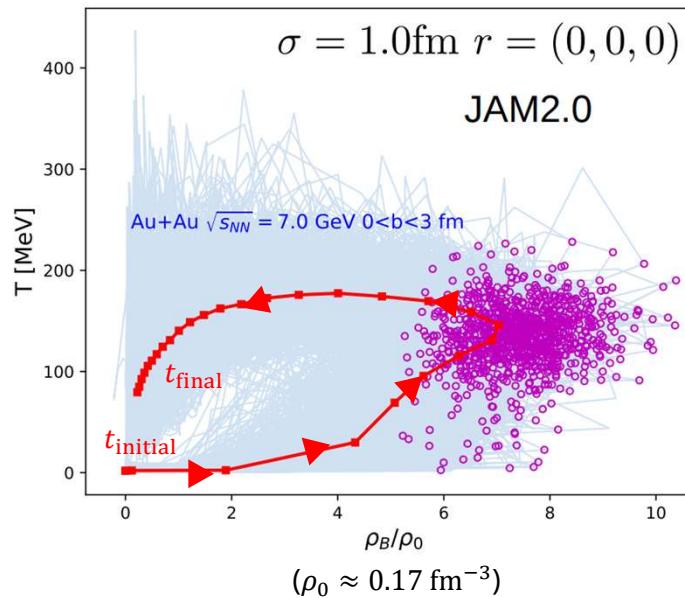
# Preceding thy. work on intermediate-energy HIC

✓ There exist many but highlight two:

- Statistical model: [Randrup, Cleymans (2006)]  
⇒ estimate from exp. particle yields  
at the freeze out



- Transport model: "Ohnishi plot" by A. Ohnishi (~2010)  
⇒ time evolution of density  
(at the center cell)

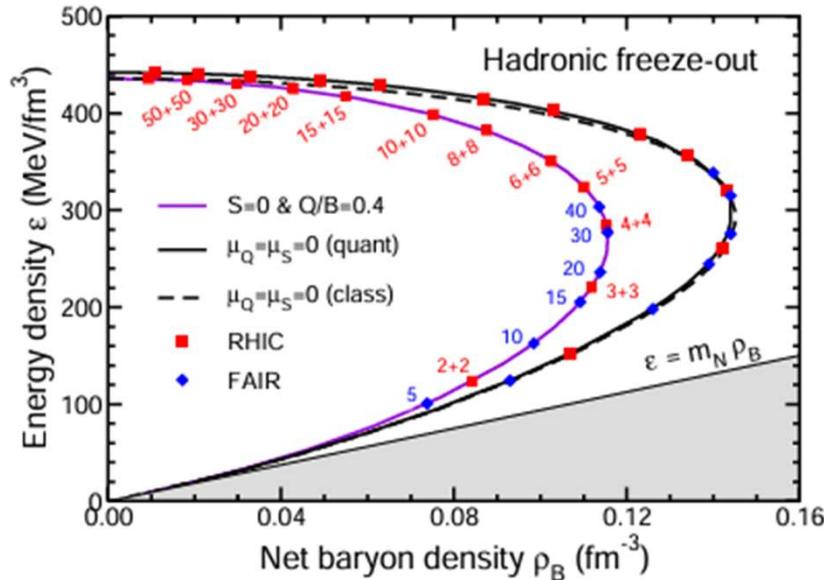


- ⇒ • the best energy range seems to be  $\sqrt{s_{NN}} = 6 \sim 10 \text{ GeV}$   
• high density  $\rho = (6 \sim 10) \times \rho_0$  can be realized

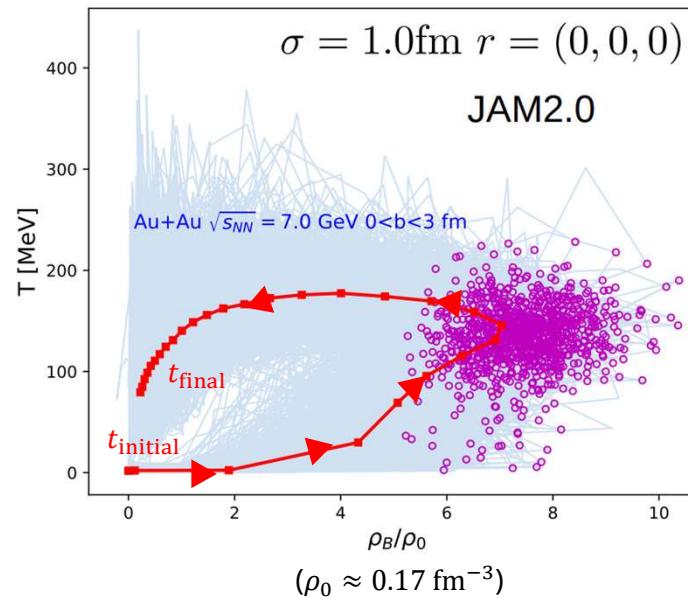
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✓ Problem: Volume of the dense region ?

⇒ Purpose:

- can the dense region have sufficiently large volume?
- what is the best energy to simultaneously maximize density & volume?

# Method

## ✓ Microscopic transport model: JAM (Jet AA Microscopic transport model)

[Nara, Otsuka, Ohnishi, Nitta, Chiba (2000)]

- A successful model to simulate the realtime dynamics of heavy-ion collisions, reproducing various data ( $v_1$ , yields, ...)
- Basic idea: superposition of collisions of individual hadrons  
(incl. inelastic ones such as resonance, string breaking, mini-jet)
- Anyway, the phase-space of each hadron ( $x_i^\mu, p_i^\mu$ ) can be obtained  
 $\Rightarrow$  integration of it gives physical observables

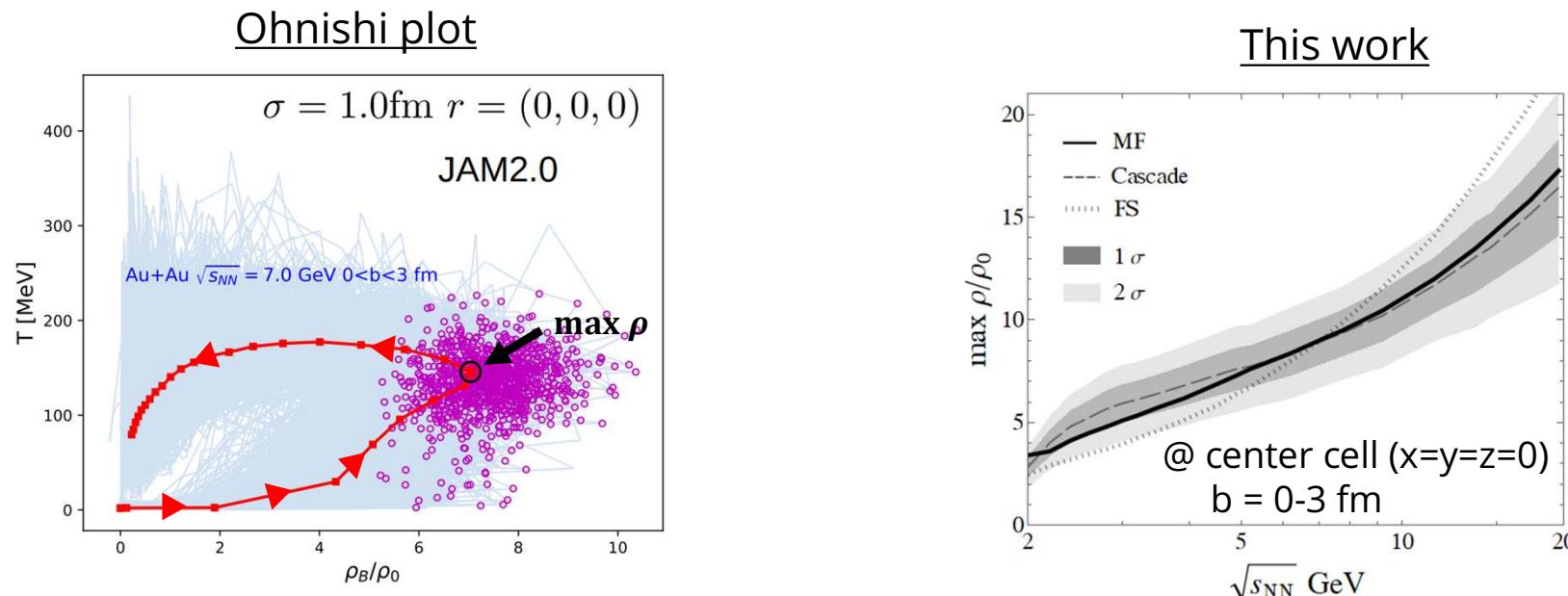
e.g.) baryon current in the local rest frame ( $J_{\text{LRF}}^0 =: \rho$  gives the baryon density):

$$J_{\text{LRF}}^\mu = \Lambda^\mu_\nu J_{\text{lab}}^\nu = \Lambda^\mu_\nu \sum_{i: \text{all hadrons}} \rho(x_i) Q_i \frac{p_i^\mu}{p_i^0}$$

Boost matrix (move to Eckart frame)      Gaussian smearing ( $\sigma=1\text{fm}$ )

# Result (1/3): maximum density @ center cell ( $x = 0$ )

✓ Consistency check w/ Ohnishi plot: calculate  $\max \rho := \max_t J_{\text{LRF}}^0$

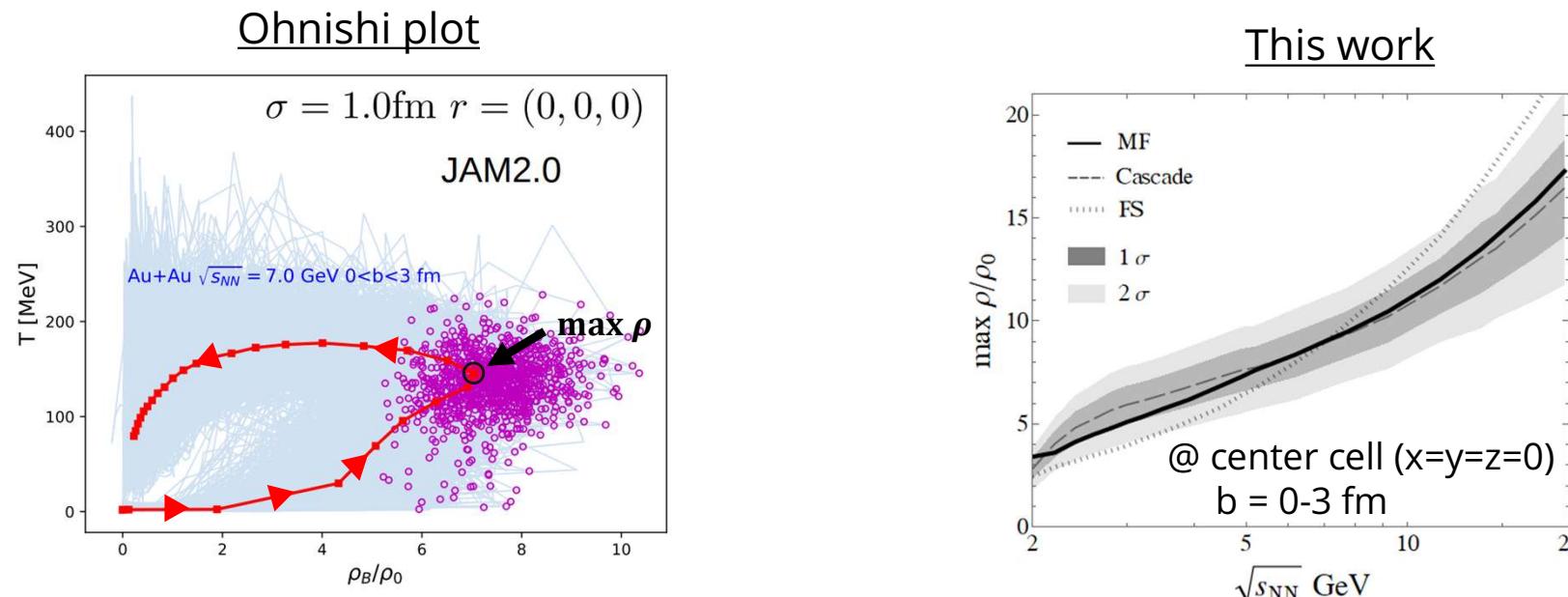


⇒ At  $\sqrt{s_{NN}} = 7 \text{ GeV}$ ,  $\max \rho \approx 7\rho_0$  and is fluctuating by ~30%

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(the time evo. is also consistent)

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- **Basic: Classical physics = Overlapping of Lorentz contracted “uniform” nuclei**

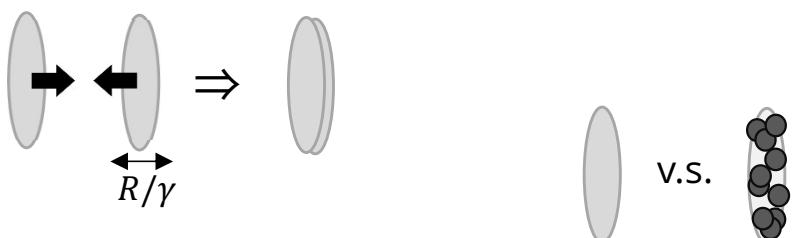
max density is at the maximally overlap

$$\Rightarrow \text{roughly: } \max \rho = 2 \times \rho_0 \gamma \approx \sqrt{s_{NN}} [\text{GeV}] \times \rho_0$$

- **Reality: A nucleus is not uniform but is fluctuating**

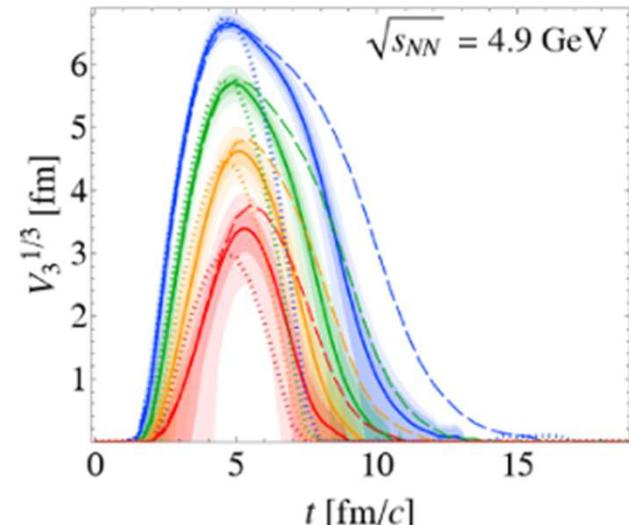
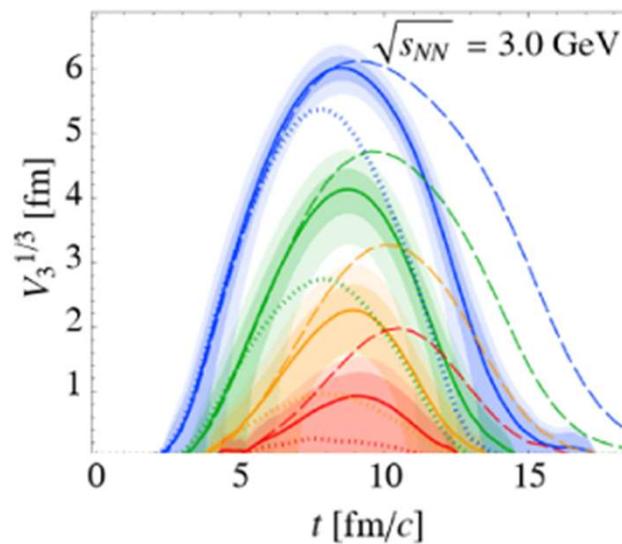
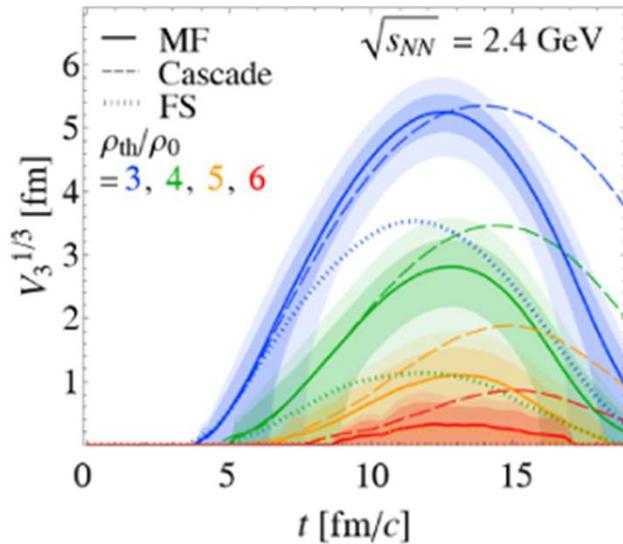
happen to overlap more/less ⇒ denser/diluter ⇒ ~30% fluctuation

- fluctuation is important for creating a dense region (e.g., top 10% of 7.7 GeV reaches  $10 \times \rho_0$ )
- the dense region should be “local” = cannot extend over the whole overlapping region



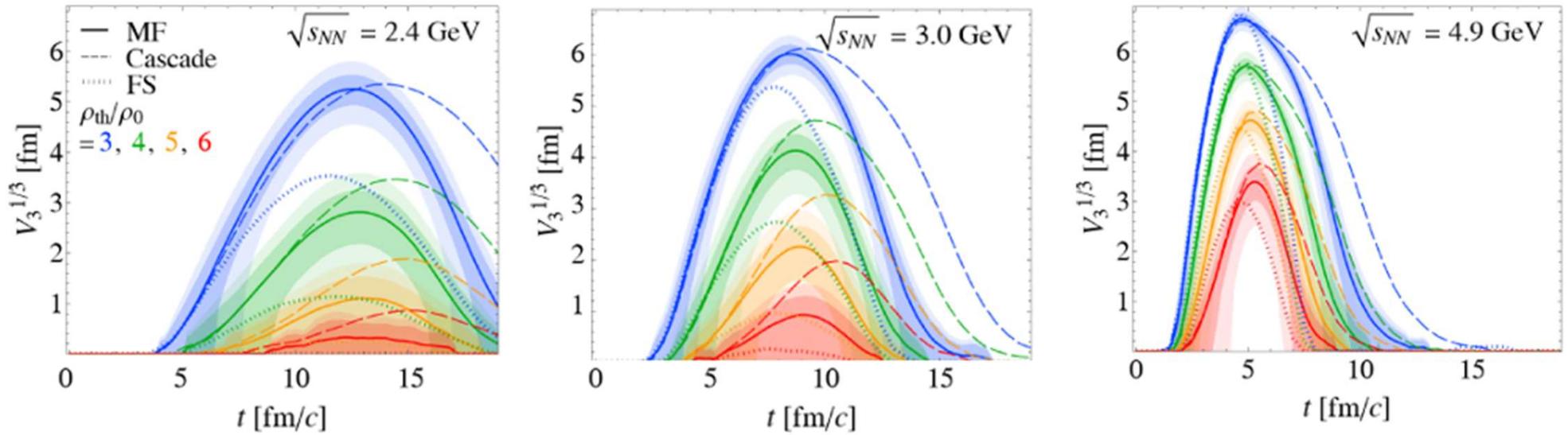
# Result (2/3): spatial volume $V_3$

✓ Spatial volume in the local rest frame:  $V_3(t) := \int_{\rho(t,x) > \rho_{\text{th}}} d^3x \gamma(t, x)$



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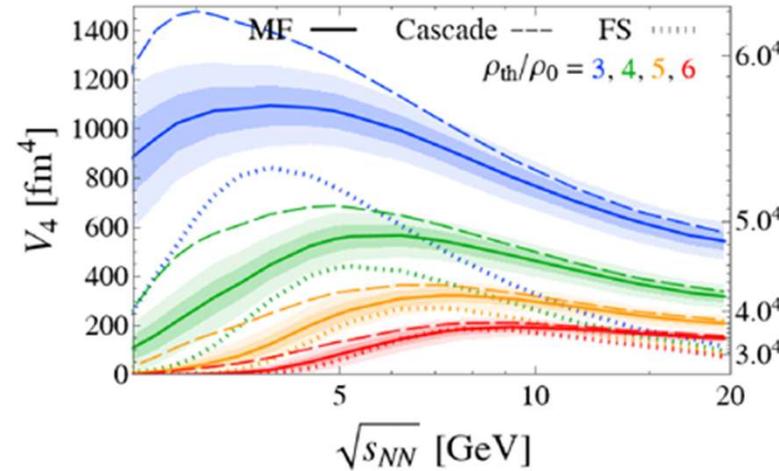
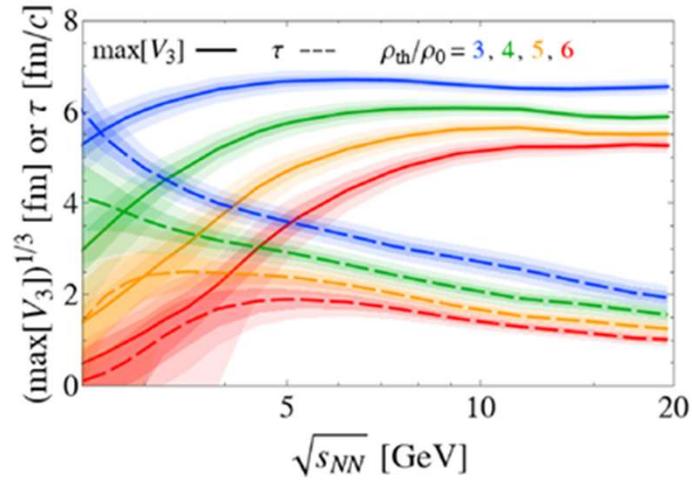
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- Super-dense region originates from nucleons fluct.  $\Rightarrow$  volume  $\sim$  nucleon size  $V_3 = O((1 \text{ fm})^3)$ 
  - event selection enables us to go to higher densities, but it's not matter (i.e., not macroscopically large)
  - even if such a super-dense local region  $V_3 = O((1 \text{ fm})^3)$  were giving a nontrivial signal, it's buried in the other signals by other regions  $V_3 = O((5 \text{ fm})^3)$
- “Relatively” dense region  $\rho \approx \sqrt{s_{NN}} [\text{GeV}] \times \rho_0$  that is created by the naïve overlapping of two uniform nuclei can have sufficiently large volume  $\sim$  nucleus size  $V_3 = O((6 \text{ fm})^3)$ 
  - under the requirement of large  $V_3 = O((6 \text{ fm})^3)$ , the achievable density is limited to  $\rho \approx \sqrt{s_{NN}} [\text{GeV}] \times \rho_0$
  - Q: Larger  $\sqrt{s_{NN}}$  is better for creating a high dense matter?  
A: No because the lifetime becomes shorter  
 $\Rightarrow$  Q: what is the best energy  $\sqrt{s_{NN}}$  that maximize the volume and lifetime simultaneously, with keeping a high density ?

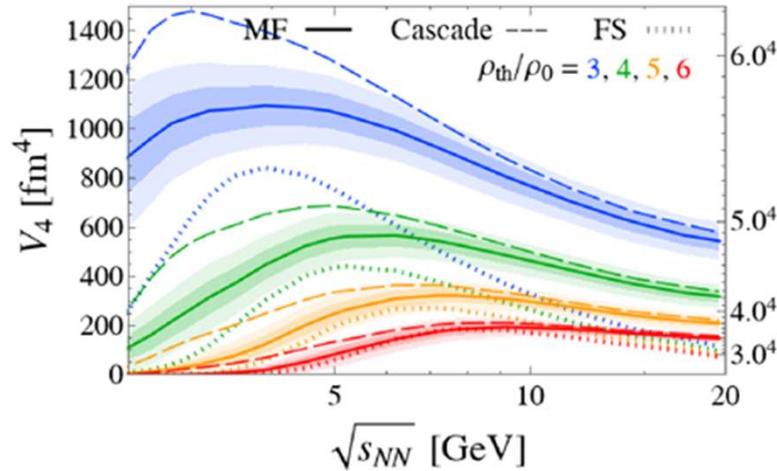
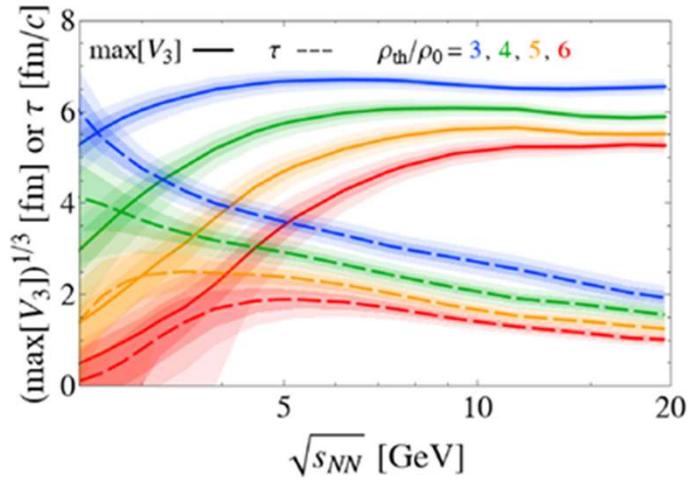
# Result (3/3): lifetime and four volume

✓ Four volume and lifetime:  $V_4 := \int_{\rho(t,x) > \rho_{\text{th}}} dt d^3x$  &  $\tau := V_4 / \max_t V_3$



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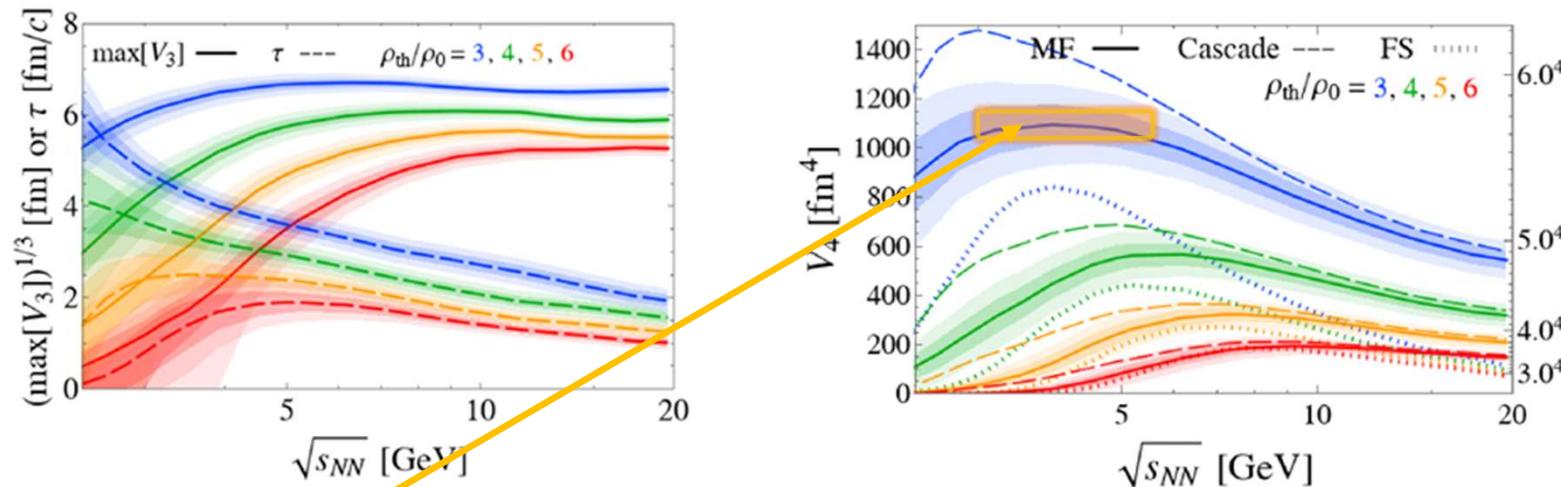
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- $\sqrt{s_{NN}} \nearrow \Rightarrow V_3 \nearrow \& \tau \searrow$
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- $\sqrt{s_{NN}} \uparrow \Rightarrow V_3 \nearrow \& \tau \searrow$
- $V_4$  has a **plateau** where lifetime and volume are simultaneously “maximized”  
⇒ interpreted as “the best energy range for creating high dense matter i.t.o. spacetime vol.”
- for  $\rho > 3\rho_0$  it's  $\sqrt{s_{NN}} = 3 \sim 5$  GeV
- the range can also be identified for even higher densities like  $\rho > 5\rho_0$ ,  
for which however  $V_4$  is no longer large (e.g.,  $V_4 \sim ((4 \text{ fm})^4)$  for  $\rho > 5\rho_0$ )

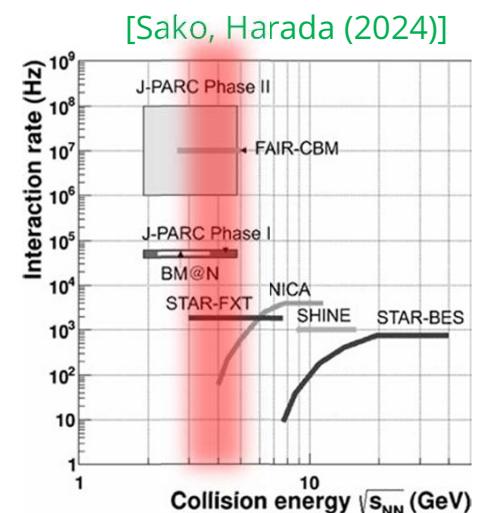
∴ The best energy range is  $\sqrt{s_{NN}} = 3 \sim 5$  GeV where you can explore  $\rho > 3\rho_0$

Comment 1: should be contrasted w/ the existing result [Randrup, Cleymans (2006)]

$$\sqrt{s_{NN}} = 6 \sim 10 \text{ GeV} \text{ and } \rho = (6 \sim 10) \times \rho_0$$

[Ohnishi plot (~2010)]

Comment 2: the future experiments like FAIR-CBM & J-PARC-HI will exactly cover this regime with great statistics !



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# Motivation (1/2): Strong-field physics

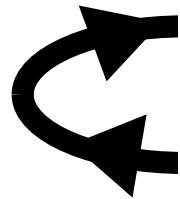
## ✓ Matter under extreme conditions: temperature, density, EM field, ...

- Not only at the QCD scale but also even at the QED scale, many nontrivial non-perturbative phenomena expected (=: strong-field physics)

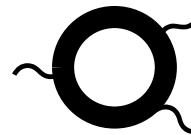
### Novel QED processes ( $eE, eB \gtrsim m_e^2$ )

Review: [Fedotov, Ilderton, Karbstein, King, Seipt, HT, Torgrimsson (2022)]  
[Hattori, Itakura, Ozaki (2023)]

e.g.) Schwinger effect



Photon splitting

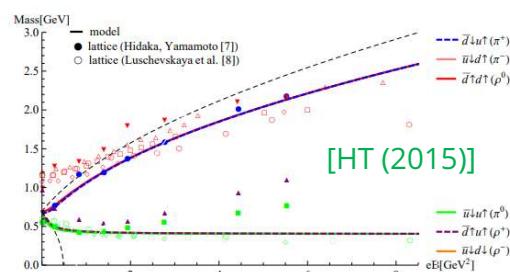


Vacuum birefringence

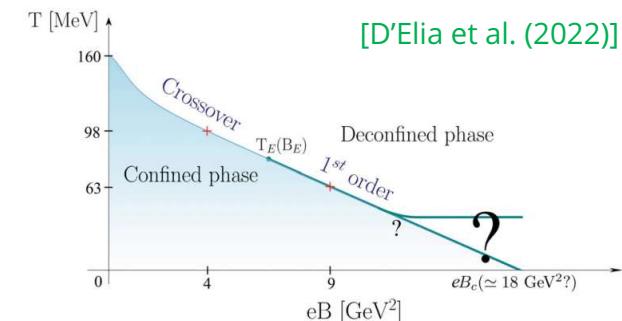


### Impacts on QCD/hadron physics ( $eE, eB \gtrsim \Lambda_{\text{QCD}}^2$ )

- Hadron properties  
(e.g., mass, charge density, decay mode, ...)



- QCD phase diagram  
(e.g., new phase, magnetic catalysis, ...)

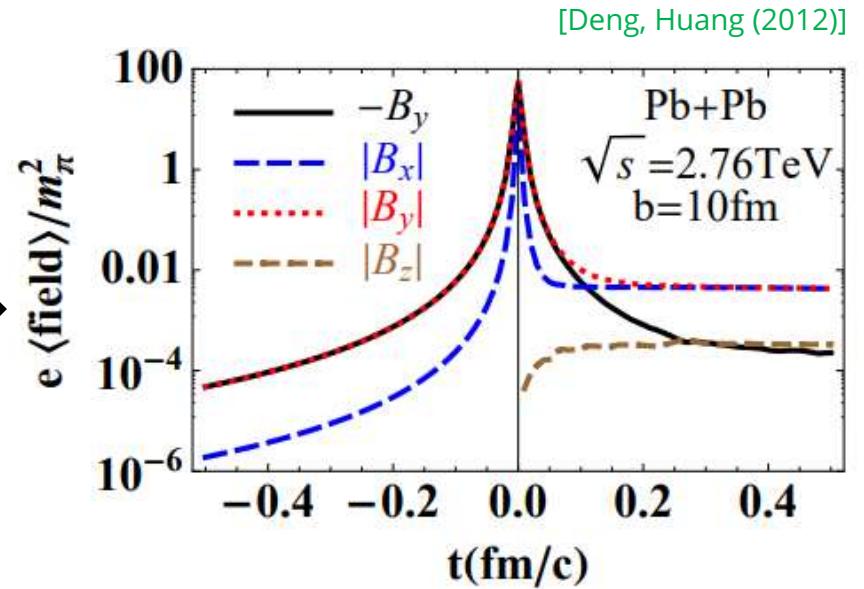
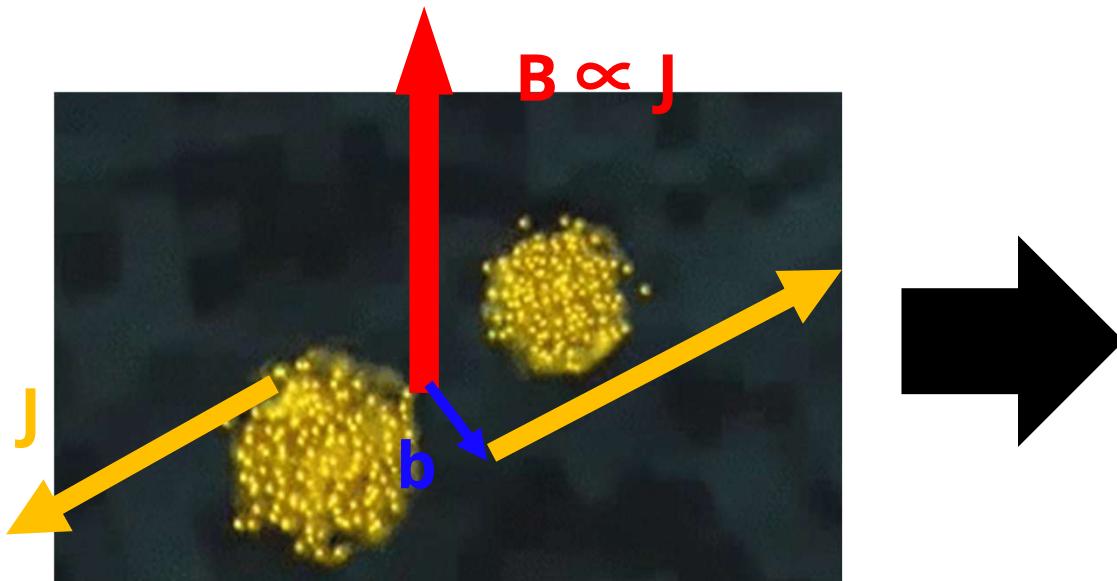


- However, NO experimental verification

∴ no exp. to realize strong field ⇒ New idea/approach strongly needed !

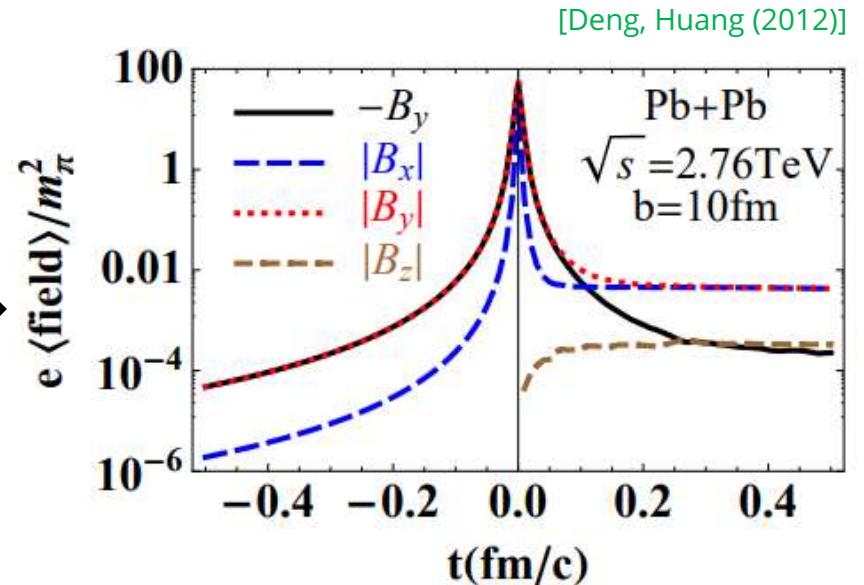
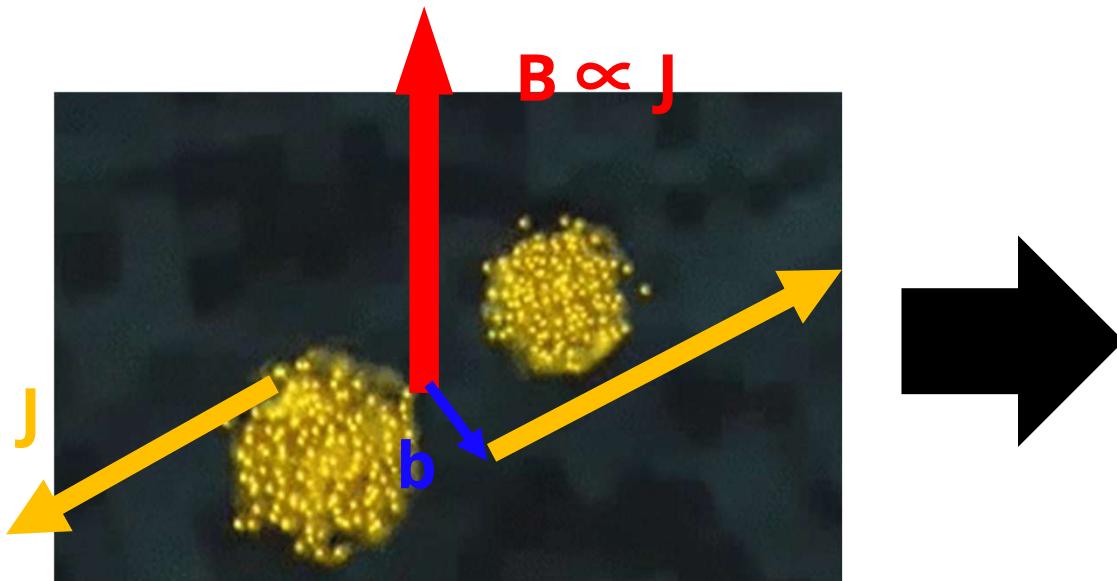
# Motivation (2/2): Use of HIC

- Such a possibility has been discussed at high energies  $\sqrt{s_{NN}} = O(100 - 1000 \text{ GeV})$  over the past decade

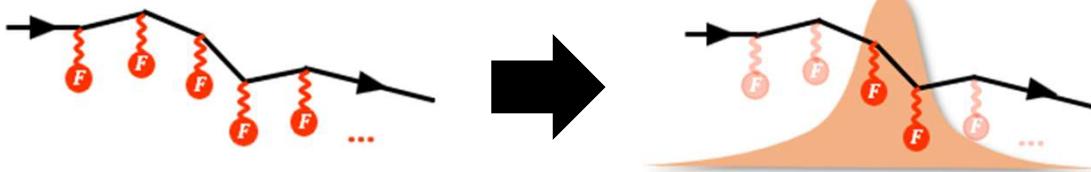


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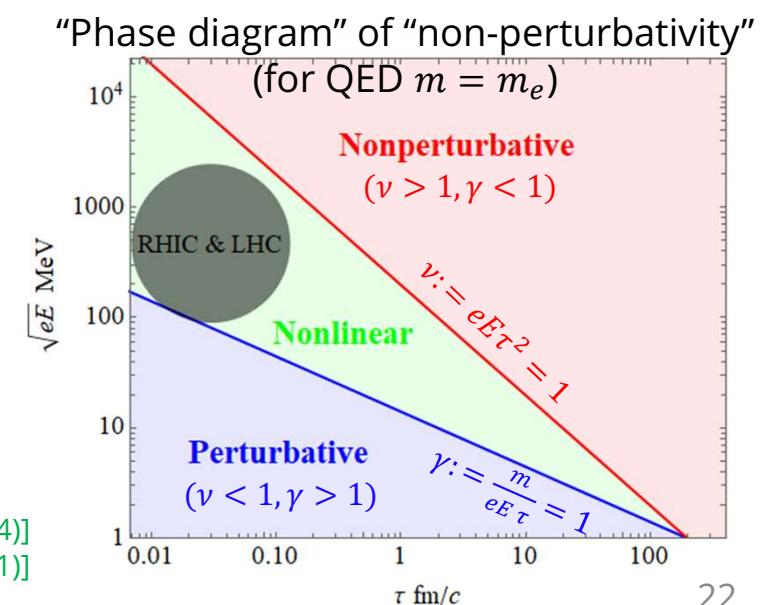
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- But, extremely short-lived, so not interesting for non-perturbative strong-field physics



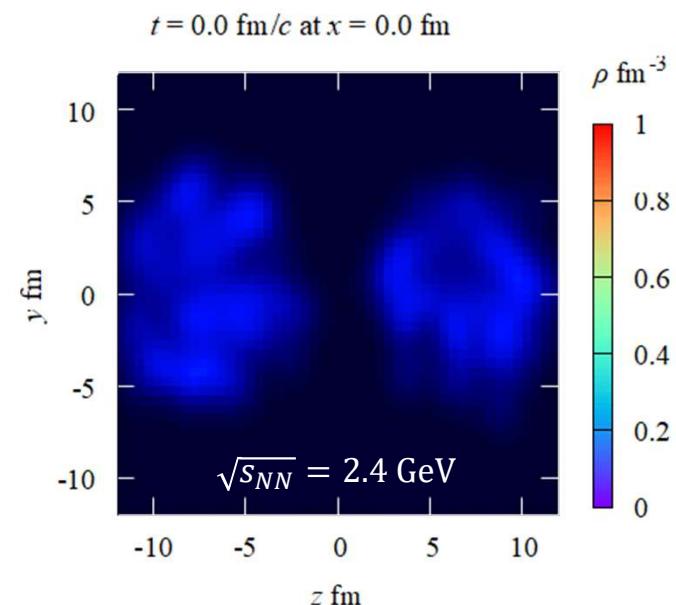
[HT, Fujiiii, Itakura (2014)]  
[HT, Fujimori, Misumi, Nitta, Sakai (2021)]



# Idea: Intermediate energy should be useful

## ✓ Baryon stopping at intermediate energies

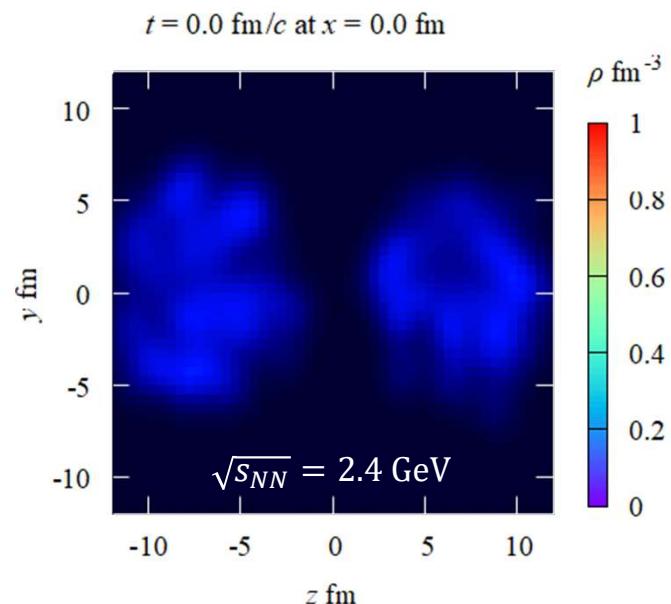
- high density is realized = large “atomic” number  $Z \sim 200$ 
  - ⇒ strong Coulomb field  $eE \sim \frac{Z\alpha}{r^2} \sim \Lambda_{\text{QCD}}^2 \sim (100 \text{ MeV})^2$
  - ⇒ much stronger than the QED scale and is still comparable to QCD
- sticks together for a long time  $O(10 \text{ fm}/c)$ 
  - ⇒ created Coulomb field should also be long-lived  $\tau \sim 10 \text{ fm}/c$  (more than 100 times longer than high energy)
  - ⇒ may resolve the issue of lifetime of high-energy HIC



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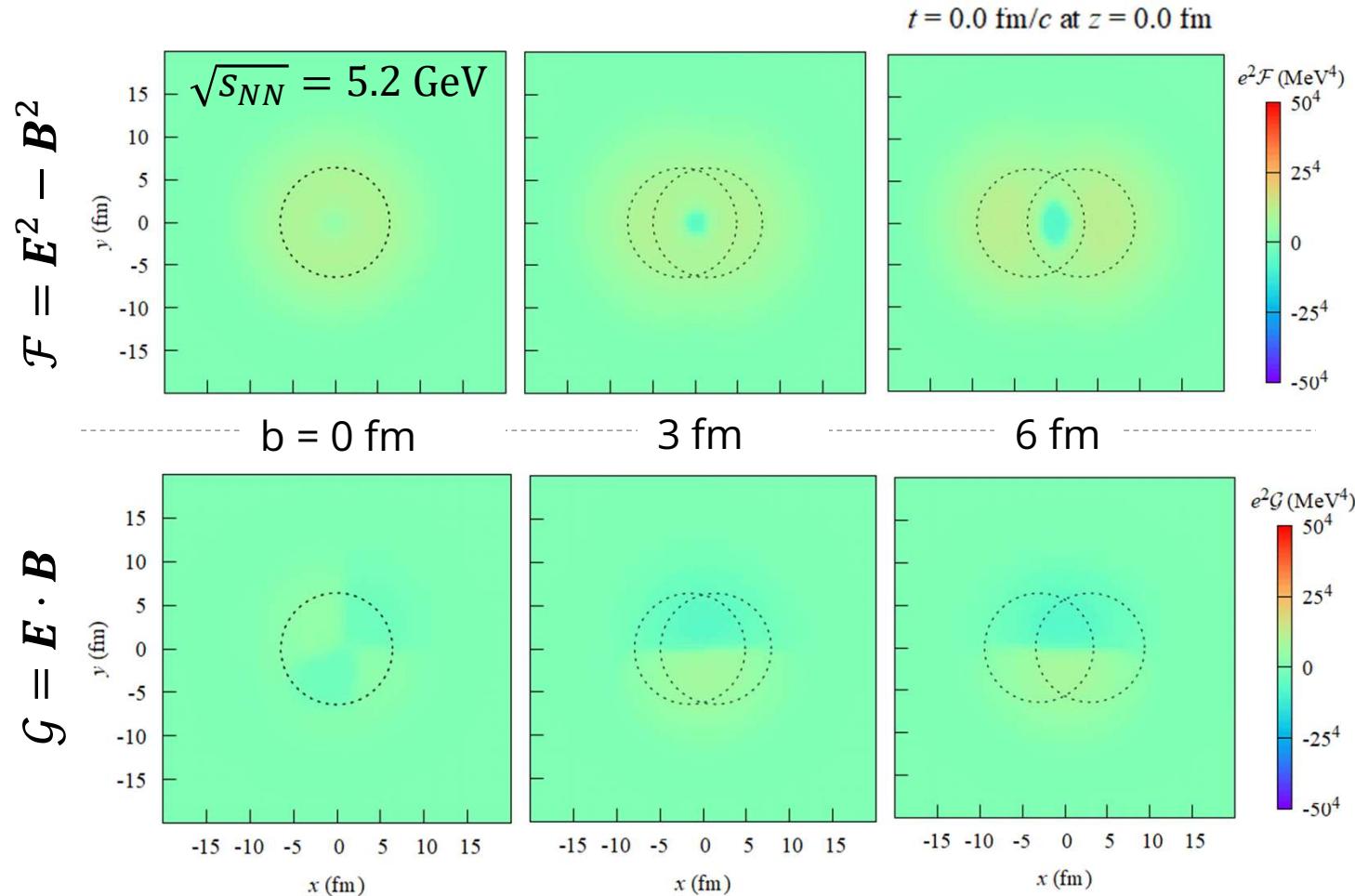
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∴ can be a new tool to study strong-field physics; should also affect QCD

⇒ Purpose: Let's check if this is correct/wrong by using JAM

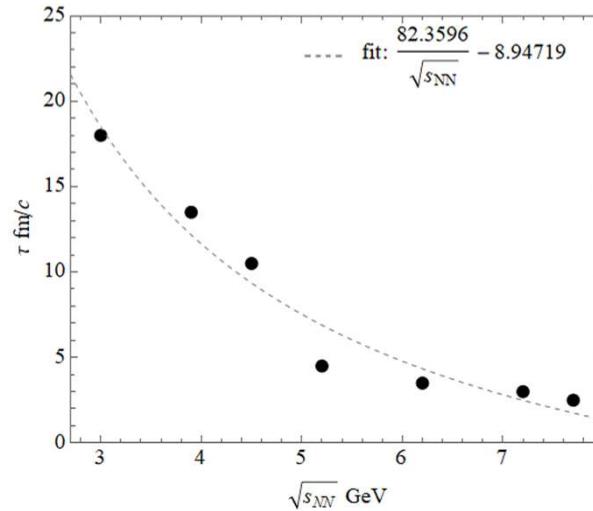
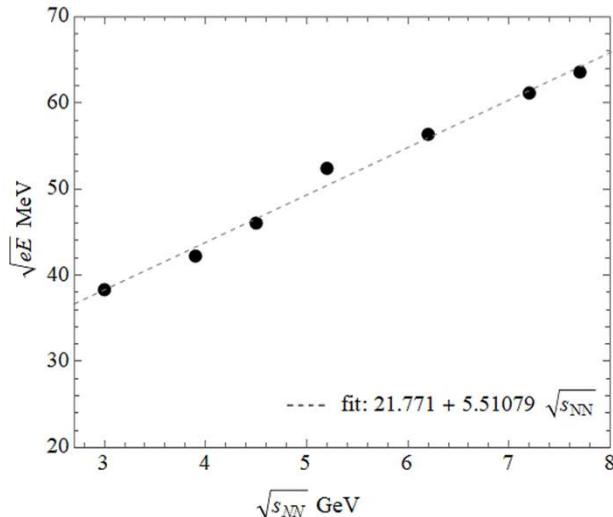
# Result (1/2): Spacetime profile of EM field



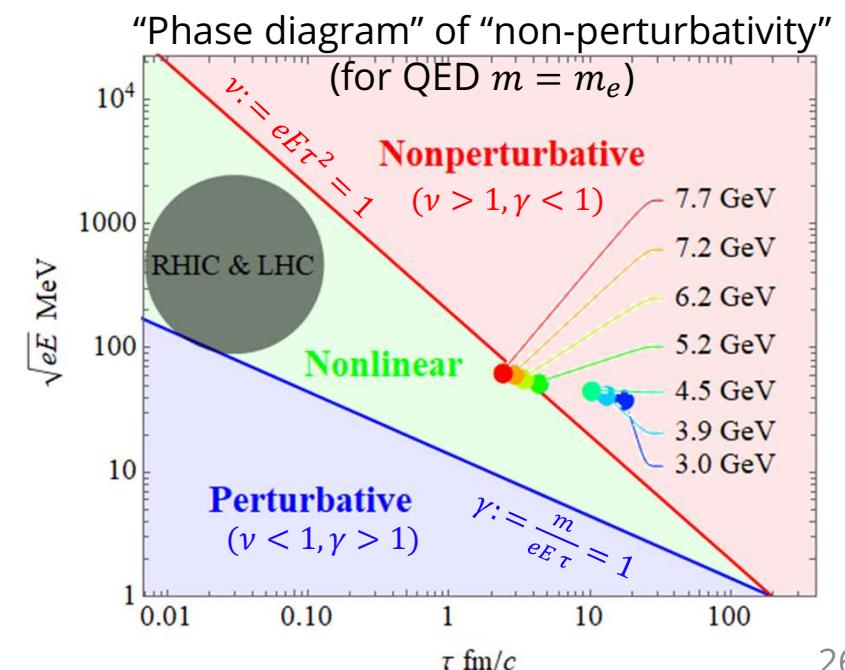
- B field appears w/ going to more peripheral but E field is always larger in space  
⇒ E field would be more important than B field in intermediate energies
- “topological” EM field configuration such that  $G = E \cdot B \neq 0$   
⇒ can be a source of chiral physics  $\partial_\mu J_5^\mu \propto E \cdot B$

# Result (2/2): More quantitative analysis

## ✓ Peak E-field strength and lifetime (FWHM) (at central coll. $b=0$ )



- “strong” O(50 MeV) and long-lived O(10 fm/c) realized
- can cover non-pert. regime
- ∴ Intermediate-energy HIC must be useful as a new tool to study strong-field physics
- non-negligible to QCD/hadron scale  
⇒ should affect, e.g., EM probe



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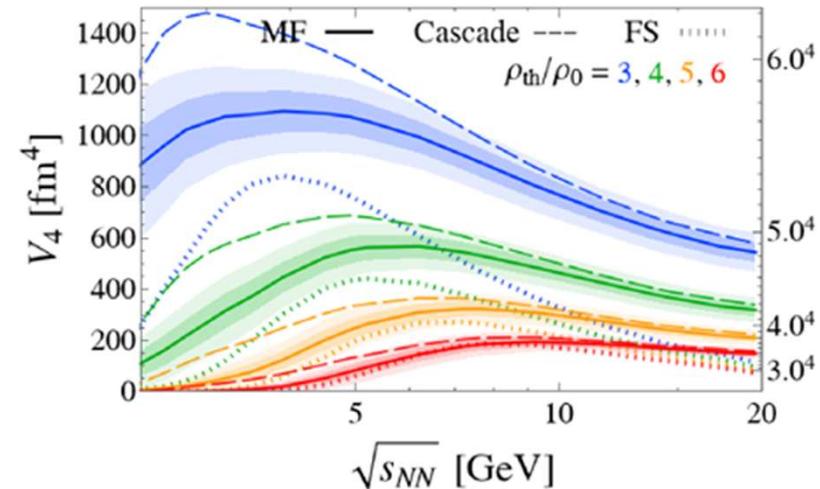
# Summary

✓ I discussed physical conditions realized in intermediate-energy heavy-ion collisions by using a hadron transport model JAM

✓ Two take-home messages:

- (1) What is the BEST energy range to study high baryon-density physics ?  $\Rightarrow \sqrt{s_{NN}} = 3 \sim 5 \text{ GeV}$

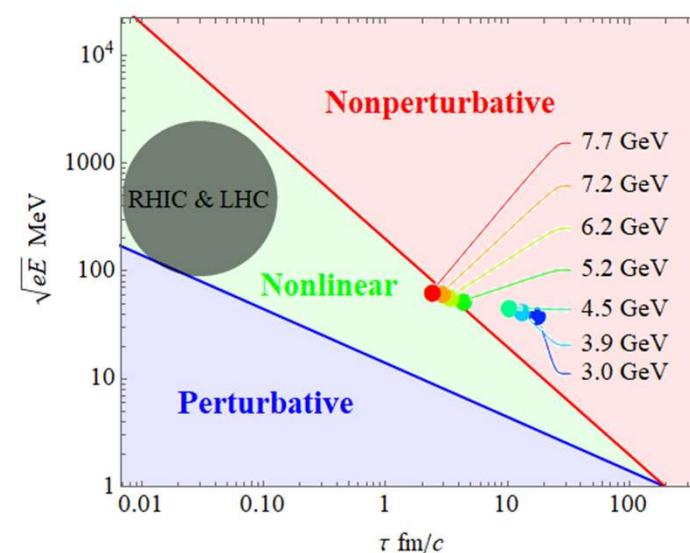
[HT, Jinno, Kitazawa, Nara, 2409.07685]



- (2) Such an intermediate-energy regime is also EXTREMELY interesting for studying physics of strong EM field

[HT, Nishimura, Ohnishi, 2402.17136]

[HT, 2501.18171]



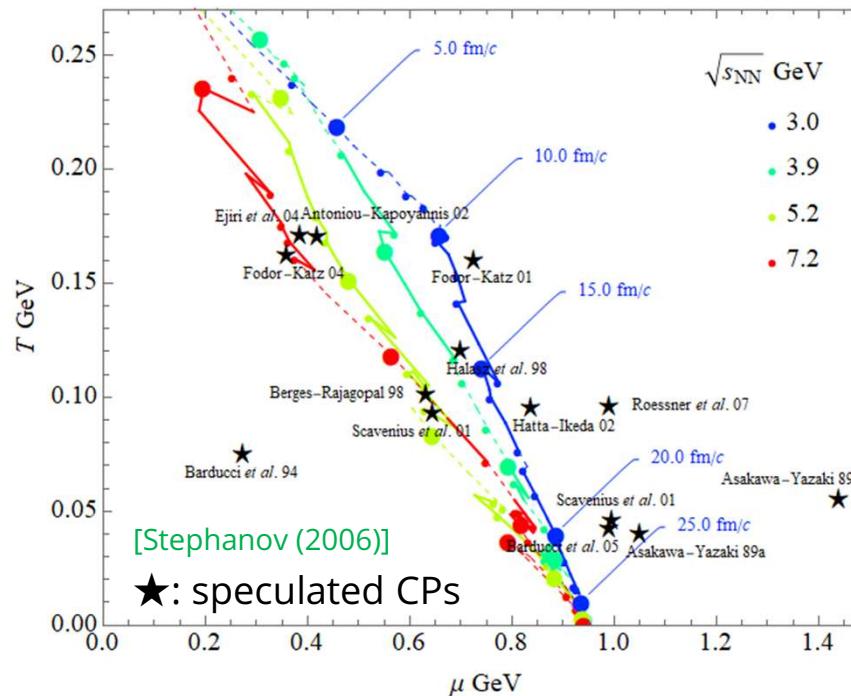
# Backup

# T and $\mu$

## ✓ What are the corresponding temperature T and chemical potential $\mu$ ?

Naïve estimate from the classical ideal gas EoS, e.g.,  $P = nT$ :

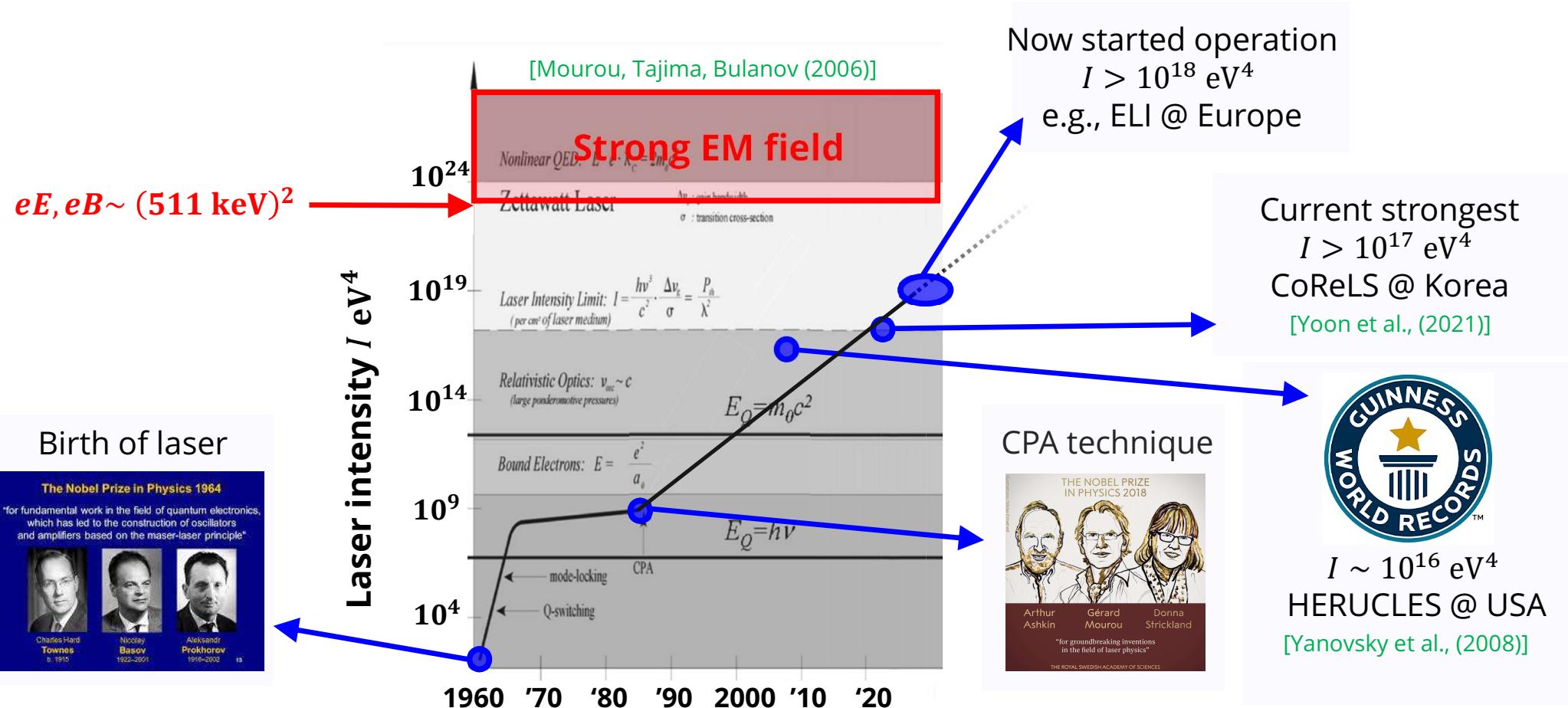
( $\Leftarrow$  no visible changes from those obtained by the numerical fit of the spectrum with Fermi dist)



⇒ Good news for critical-point search:

the trajectories pass the speculated critical points in the best energy range

# Development of intense laser



# How $E \cdot B \neq 0$ emerges

