- 1. It's a classical midel of the thormodynamic behaviour of gardes (typically i'deal gas)
- 2. The model assumes large number of identical microscopic particles Catoms, molecules) which are in rapid constant and random motion
 - 3. The size of the particles are assumed to be much smaller than their interatomic distances.
 - 4. The particles underso random elastic collisions between themselves and with the enclosing walls.
 - 5. The basic vorsion of the model describes ideal gas and consider no other interaction between the particles.
 - 6. The number of particles are so large that a statistical treatment of the problem is justified.

 This is called the thermodynamic limit.
 - I. collisions between the particles are strictly binary.

 There is no three-body on higher-order interactions.
 - 8. The Collisions no forces on each other rie, no potential energy term.
 - 9. The Kinetic theory of gas explains the mainscopic properties of gases such as waterness pressure, viscosity, thermal conductivity, mass diffusivity, etc.

$$\Gamma(n) = \int_{0}^{\infty} e^{-x} x^{n-1} dx$$

Some standard integrals $\Gamma(n) = \int_{0}^{\infty} e^{-x} x^{n-1} dx \qquad \begin{cases} \text{In is a principle of } \\ \text{only number}. \end{cases}$

This is Called gamma function.

T(h) = (n-0! For n is a possitive intesor

In general:
$$\Gamma(n) = (n-1) \Gamma(n-1)$$

$$I(x) = \int_{0}^{\infty} \int_{0}^{\infty} e^{-\alpha x^{2}} x^{n} dx$$

$$I(A) = \frac{1}{2} \Gamma\left(\frac{n+1}{2}\right) \alpha^{-(n+0)/2}$$

$$I(0) = \frac{1}{2}\Gamma(\frac{1}{2})a^{-1/2} = \frac{1}{2}\sqrt{\pi}a^{-1/2}$$

$$T(1) = \frac{1}{2}\Gamma(1) \propto^{-1} = \frac{1}{2} \propto^{-1}$$

$$I(1) = \frac{1}{2}\Gamma(1) \propto^{-1} = \frac{1}{2} \propto^{-1}$$

$$I(2) = \frac{1}{2}\Gamma(\frac{3}{2}) \sqrt{3}^{1/2} = \frac{1}{2} \sqrt{2}\Gamma(\frac{1}{2}) \sqrt{3}^{1/2} = \frac{1}{4}\sqrt{11} \sqrt{3}^{1/2}$$

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The Boltzmann factor (e-BEr)

Consider the syptem A is in contact with a heat reservoir A', where AKA'.

The energy of the combined system A LA' is card EdsE). Constant in some range of energy (E') and EdsE).

$$A \rightarrow E_r$$
, $A' \rightarrow E'$
 $A + A' = A^{(i)} \rightarrow CE^{(i)} \rightarrow E^{(i)} + GE$

=> Ext E = E(0) | The system A in the microsotate or has enough Ex which can exchange enough with E.

The probability (b) of occurrence in the ensemble of situation where A is in the state or is given by the number of one micro states accessible to A (c)

$$P_p = o'\Omega'(E^{(0)} - E_r) ; \Omega' = No of all microstates$$

Since En KE (0), we can write

$$\ln \Omega'(E^{(0)} - E_{1}) = \ln \Omega'(E^{(0)}) - \left[\frac{\partial \ln \Omega'}{\partial E'}\right] E_{1}$$

$$\ln \Omega'(E^{(0)} - E_{1}) = \ln \Omega'(E^{(0)}) - \left[\frac{\partial \ln \Omega'}{\partial E'}\right] E_{2}$$

$$= \ln \Omega'(E^{(0)}) - \beta E_{r}$$

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$$P_r = c' \Omega(E^{(0)}) e^{-\beta E_r}$$

Porbability of finding the System A in the microstate or within a energy Ero Within a range of energy (SE). container can be specified by specificiny that the possition of the particle lies in the range of to rotor rie, in the volume element dr = drdy dr



- · And by specifying ets momentum (F) in the range F to F+dF.
- -. Probability of finding the particle in the rayse

 (\$\overline{\pi}\$ to \$\overline{\pi}\$+\d\overline{\pi}\$) and (\$\overline{\pi}\$ to \$\overline{\pi}\$+\d\overline{\pi}\$) is $P($\overline{r},\overline{p}) d$^3\overline{\pi}$ \alpha \end{arg} = \beta P | 2m \ d^3\overline{\pi}$ d$^3\overline{\pi}$$
 - · Since 19 = p/m, we can re-express the above relation in terms of 15 and 7.
 - · Since the above enpression is a probability, if we multiply it by the total no. of particle, it gives the mean number of particles in this position and momentum range.

Let's say N = total no. of particle

NP(\bar{p} , \bar{p}) $d^3\bar{r}$ $d^3\bar{p}$ = $f(\bar{p}, \bar{v})$ $d^3\bar{p}$ $d^3\bar{v}$ with which is the mean number particles with position between \bar{p} and \bar{p} to \bar{d} and velocity between \bar{p} and \bar{p} to \bar{d}

-: $f(\bar{r}, \bar{v}) d^3 \bar{r} d^3 \bar{v} = C e^{\beta m \bar{v}/2} d^3 \bar{r} d^3 \bar{v}$ Therefore, $\int_{\bar{r}} \int_{\bar{v}} f(\bar{r}, \bar{v}) d^3 \bar{r} d^3 \bar{v} = N$

$$C \int_{\bar{p}} \int_{\bar{u}} e^{-\beta m \bar{u}^2/2} d^3\bar{p} d^3\bar{v} = N$$

Since the integrand does not depend on F, we get

$$= 7 \text{ eV} \left(\frac{2\pi K_0 T}{m}\right)^{3/2} = N$$

$$e \vee \int e^{-\beta m \sqrt{2}} d^{3}b = N$$

$$e \vee \left(\frac{2\pi K_{0}T}{m}\right)^{3/2} = N$$

$$e = \sqrt{\left(\frac{m}{2\pi K_{0}T}\right)^{3/2}} = n\left(\frac{m}{2\pi K_{0}T}\right)^{3/2}$$

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$$e = \sqrt{\frac{m}{2\pi K_{0}T}}$$

$$B = K_BT$$
 $K_B = Boltzmann$

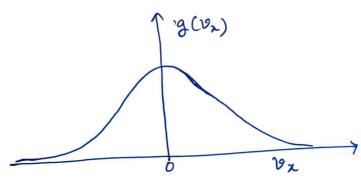
constant

$$f(\overline{\nu})d^{3}\overline{\nu}d^{3}\overline{\nu} = n\left(\frac{m}{2\pi K_{3}T}\right)^{3/2}e^{-\frac{m \nu^{2}}{2K_{3}T}}d^{3}\overline{\nu}d^{3}\overline{\nu}$$

Maxwell velocity distribution. 18 This

If we divide the above expression by dir, $f(\bar{\nu}) d^3\bar{\nu} = n \left(\frac{m}{2\pi K_A T}\right)^{3/2} e^{-\frac{m \hat{\nu}^2}{2K_B T}} d^3\nu$

This is the mean number of molecules Det unit volume with velocity in the range To and To+do.



· Distribution of Velocity component: we can define g(v2)dvx = Sushing

This is the mean number of molecules per unit volume With 12 in the range between 102 and 12+d 102.

$$g(v_{\lambda}) dv_{\chi} = n \left(\frac{m}{2\pi k_{b}T}\right)^{3/2} \int_{v_{\delta}} e^{-\frac{m}{2k_{b}T}} \left(v_{\lambda}^{\gamma} + v_{\delta}^{\gamma} + v_{\delta}^{\gamma}\right) dv_{\lambda} dv_{\delta} dv_{\delta}$$

$$= n \left(\frac{m}{2\pi k_{b}T}\right)^{3/2} \int_{v_{\delta}} e^{-\frac{mv_{\delta}}{2k_{b}T}} dv_{\delta} \int_{v_{\delta}} e^{-\frac{mv_{\delta}}{2k_{b}T}} dv_{\delta} e^{-\frac{mv_{\delta}}{2k_{b}T}} dv_{\delta}$$

$$= n \left(\frac{m}{2\pi k_{b}T}\right)^{3/2} \left(\frac{2\pi k_{b}T}{m}\right)^{3/2} \left(\frac{2\pi k_{b}T}{m}\right)^{3/2} \left(\frac{2\pi k_{b}T}{m}\right)^{3/2} e^{-\frac{mv_{\delta}}{2k_{b}T}} dv_{\delta}$$

$$= n \left(\frac{m}{2\pi k_{b}T}\right)^{3/2} e^{-\frac{mv_{\delta}}{2k_{b}T}} dv_{\delta}.$$

$$\int g(v_{\delta}) dv_{\delta} = n \left(\frac{m}{2\pi k_{b}T}\right)^{3/2} e^{-\frac{mv_{\delta}}{2k_{b}T}} dv_{\delta}.$$

$$\int g(v_{\delta}) dv_{\delta} = n$$

$$\int g(v_{\delta}) dv_{\delta} = n$$

Mean velocity distribution $\langle v_n \rangle = \frac{1}{n} \int_{-\infty}^{\infty} (v_x) dv_x = 0$ $\langle v_n^* \rangle = \frac{1}{n} \int_{-\infty}^{\infty} (v_x) dv_x = 0$ $= \frac{2x!}{n} \frac{1}{n} \frac{\Gamma(\pm)}{2} \left(\frac{m}{2k_B T} \right)^{-1}$ $= \frac{2x}{n} \times \left(\frac{m}{2\pi k_B T} \right)^{1/2} \frac{1}{2} \Gamma\left(\frac{3}{2} \right) \left(\frac{m}{2k_B T} \right)^{-3/2}$ $= \left(\frac{m}{2\pi k_B T} \right)^{1/2} \frac{1}{2} \frac{1}{2} \sqrt{\pi} \left(\frac{m}{2k_B T} \right)^{-3/2}$

· This result can be directly obtained using the equipartition theorem (\frac{1}{2} m \nu_2^2) = \frac{1}{2} k_B T

 $= \frac{1}{4} \times \frac{2k_BT}{m} = \frac{k_BT}{m}$