

DESI 3PCF Tutorial

Zack Slepian
University of Florida

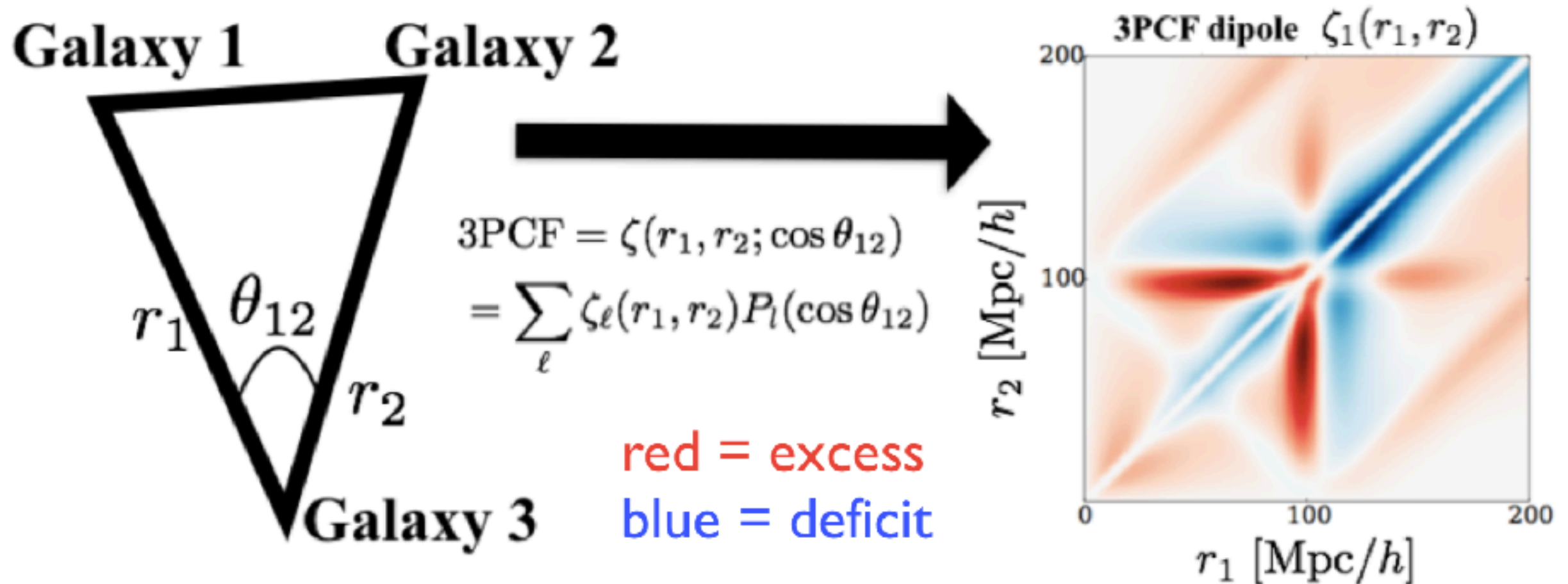
THE 3-POINT CORRELATION FUNCTION

Leading-order statistic sensitive to
information beyond the Gaussian
Random Field

Galaxy clustering has a 3PCF

Can we use it to sharpen the BAO
standard ruler?

A NEW BASIS: MULTIPOLES



Color shows # of triangles with given side lengths; angle dependence is projected onto Legendre polynomial P_{ℓ}
Proposed by Szapudi 2004 but not used
Developed by myself and collaborators

The 3PCF is computationally
expensive

if you have 10 friends, choosing two
to take to lunch is $10 \times 9 = 90$

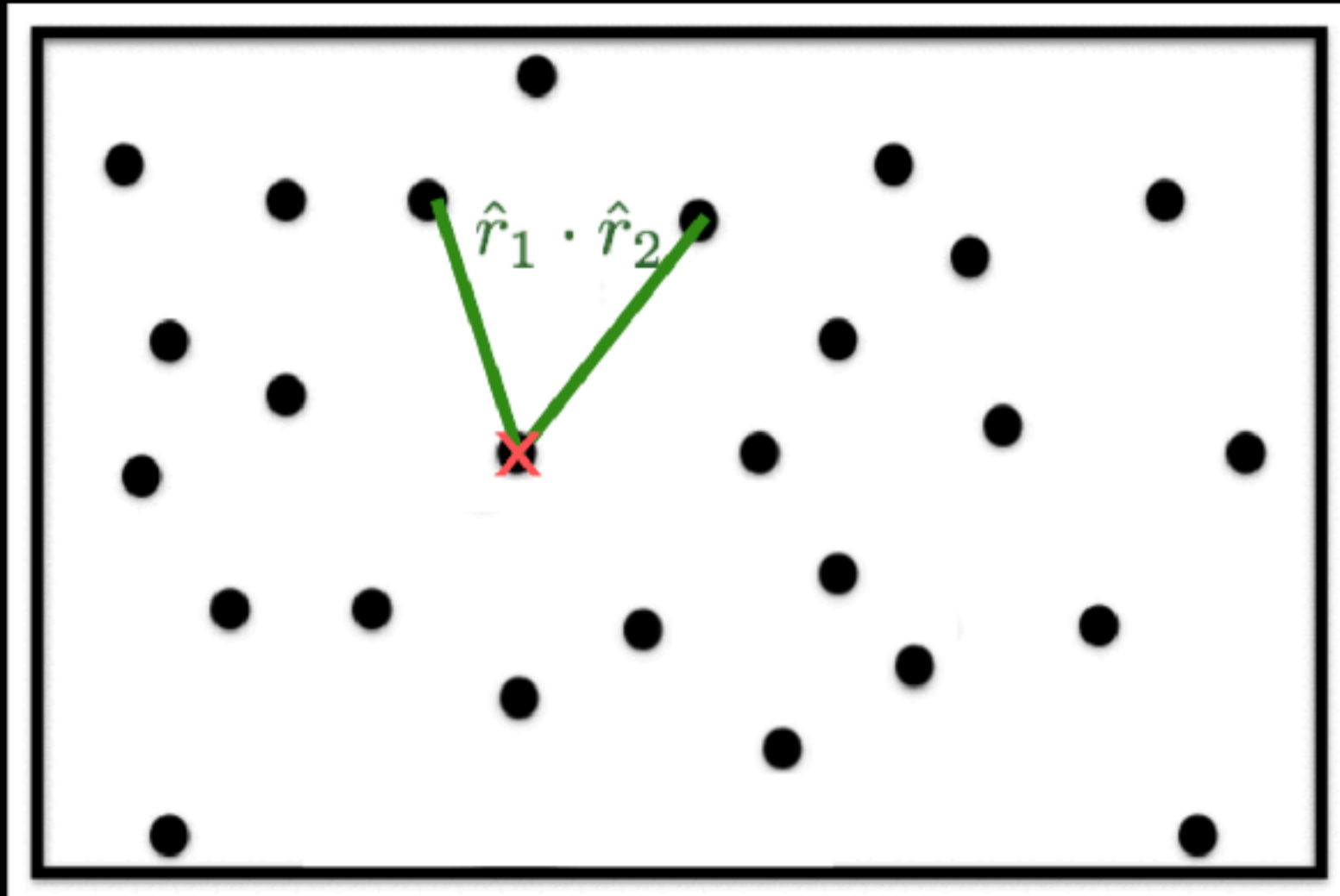
Choosing 3 is
 $10 \times 9 \times 8 = 720$
—you are double-booked!

For large N , 3PCF scales as N^3 vs.
2PCF's N^2

THE N^2 3PCF ALGORITHM

Compute about a particular galaxy given
by \mathbf{s} first, then average over all galaxies

$$\bar{\zeta}_l(r_1, r_2; \vec{s}) = \frac{2l+1}{(4\pi)^2} \int d\Omega_1 d\Omega_2 \delta(\vec{s}) \bar{\delta}(r_1; \hat{r}_1; \vec{s}) \bar{\delta}(r_2; \hat{r}_2; \vec{s}) P_l(\hat{r}_1 \cdot \hat{r}_2)$$



looks like
order N^2
about
each
galaxy

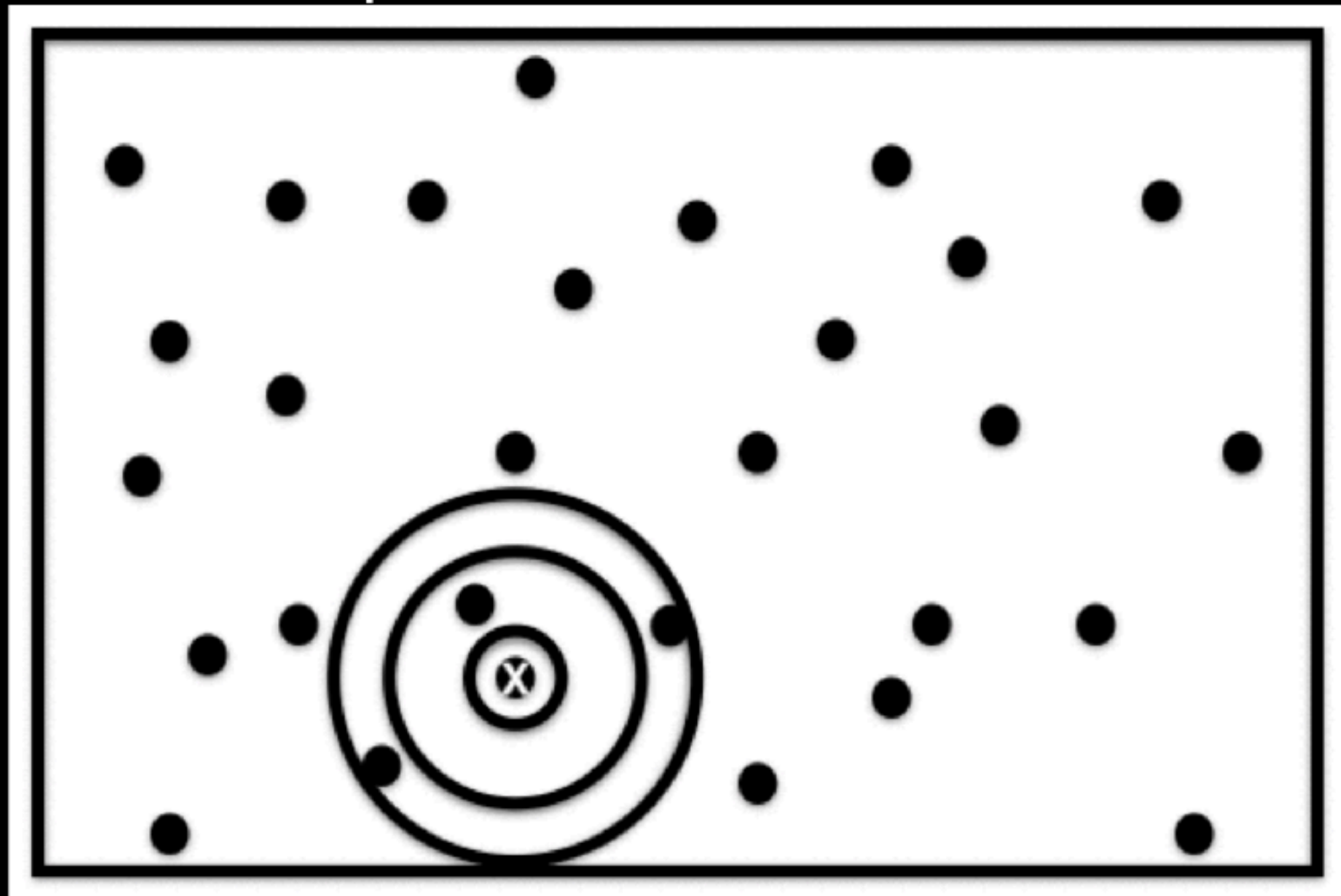
so overall
order
 N^3

SPHERICAL HARMONIC ADDITION THEOREM ALLOWS FACTORIZATION

$$P_l(\hat{r}_1 \cdot \hat{r}_2) = \frac{4\pi}{2l+1} \sum_{m=-l}^l Y_{lm}(\hat{r}_1) Y_{lm}^*(\hat{r}_2)$$

Assemble 3PCF multipole moments from spherical harmonic coefficients, which depend only on direction to one galaxy each

Around each galaxy, compute a_{lm}
in spherical shells/radial bins



$$a_{lm}(r; \vec{s}) = \sum_{\text{gals } j \text{ in bin}} Y_{lm}^*(\hat{r}_j)$$

NOW ORDER N ABOUT
EACH GALAXY, OVERALL
 N^2

SPEED OF THE ALGORITHM

500X FASTER THAN A TRIPLET
COUNT FOR 700,000 GALAXY TEST
CASE

ONLY 6X SLOWER THAN
COMPUTING A 2-POINT FUNCTION

Contact/References/Links

zslepian@ufl.edu

References

Algorithm paper:

<https://arxiv.org/abs/1506.02040>

FT-based version:

<https://arxiv.org/abs/1506.04746>

And further description of implementation of the FT version:

<https://arxiv.org/abs/1711.09907>

A few application papers

To BOSS data:

<https://arxiv.org/abs/1512.02231>

<https://arxiv.org/abs/1607.06097>

To MHD turbulence:

<https://arxiv.org/abs/1711.09907>