

# Low energy effective description of dark $Sp(4)$ theory with matter in non fundamental representation

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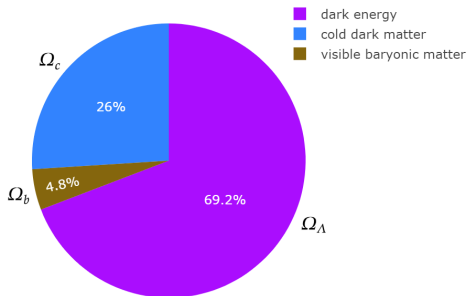
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**NAWI Graz**  
Natural Sciences

# Let's start with the usual story

There is a non-negligible non-visible matter component in the universe



No experimentally verified description on the fundamental level so far

# Cusp vs. Core problem

Data from the DDO 154 dwarf galaxy

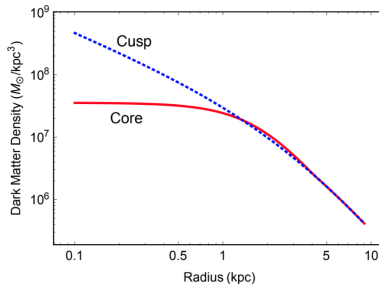
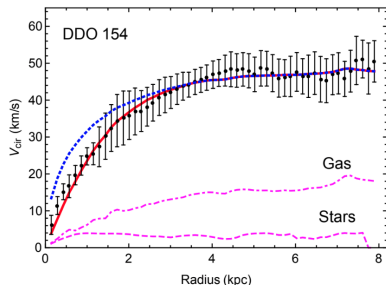


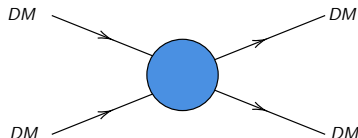
Figure: Taken from the talk on Dark QCD of [Murayama (2022)]

# Self interacting dark matter

Introduction of self interactions within the dark sector may solve these problems as shown by N-body simulations. [\[arXiv:astro-ph/9909386v2\]](#)

- Required self interaction cross section:

$$\frac{\sigma}{m} = 0.1 - 1.0 \frac{\text{cm}^2}{g}$$

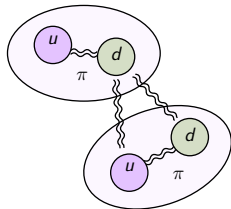


Constraints:

- Bullet cluster constraint:  $\frac{\sigma}{m} \lesssim 0.7 \frac{\text{cm}^2}{g}$  [\[arXiv:astro-ph/0704.0261\]](#)

# Dark matter as composite states

Cold dark matter may consist of composite states of an additional confining, non-abelian gauge sector in a chirally broken phase.



- Natural implementation of dark matter self interactions.
- Potential velocity dependence of self interaction may relate problems at different scales.
- Accidental flavor symmetry may stabilize dark matter.

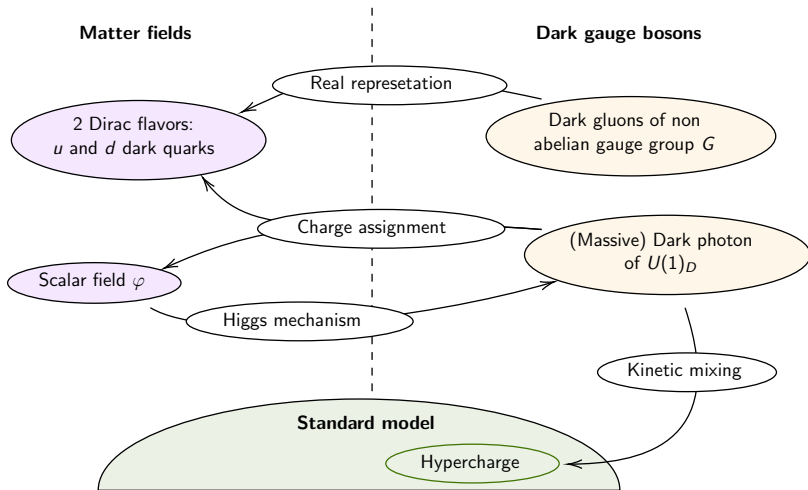
## Technical aspect:

Low energy effective description of the relevant parts of the spectrum.

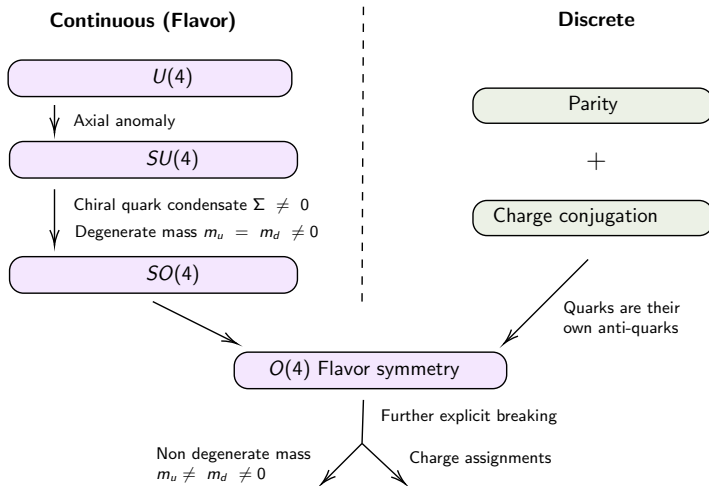
# Dark matter model in the UV

arXiv:hep-ph/1512.07917

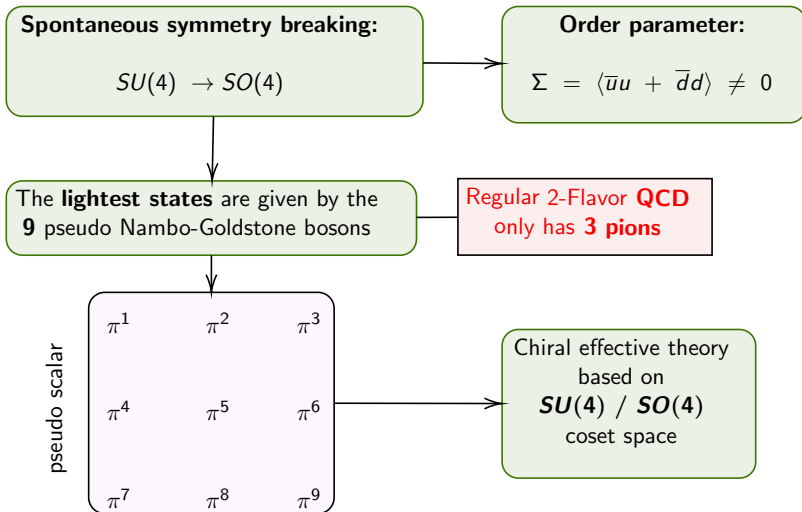
arXiv:hep-ph/2005.01515



# Symmetries and breaking patterns of fermion sector



# Lightest states - pion dark matter





# Explicit breaking via non degenerate mass

Flavor symmetry reduces to:

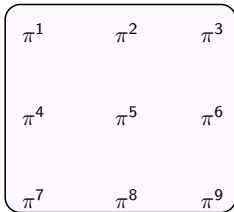
$$O(2) \times O(2)$$

$\pi^3$  is not protected  
via symmetry and may  
decay to SM via  
dark photon

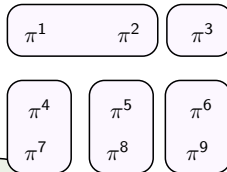
**degenerate mass**

**non degenerate mass**

pseudo scalar



$$m_u \neq m_d$$

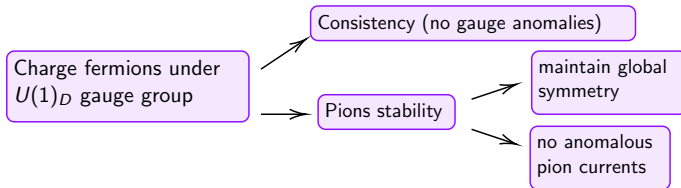


Particles within  
flavor doublet  
relate via  
C-conjugation

**NO DARK MATTER**

# $U(1)_D$ charge assignments in mass degenerate scenario

## Guidance principles:



## Charge assignments of quarks and pions :

There is only one physically distinct charge assignment in the UV that provides stable dark pions

$u$	$d$
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 $e_D = +1/2$

$\pi^1$	$\pi^2$	$\pi^3$
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 $e_D = 0$

Flavor symmetry  
 $SU(2)$   
remains

$\pi^4$	$\pi^5$	$\pi^6$
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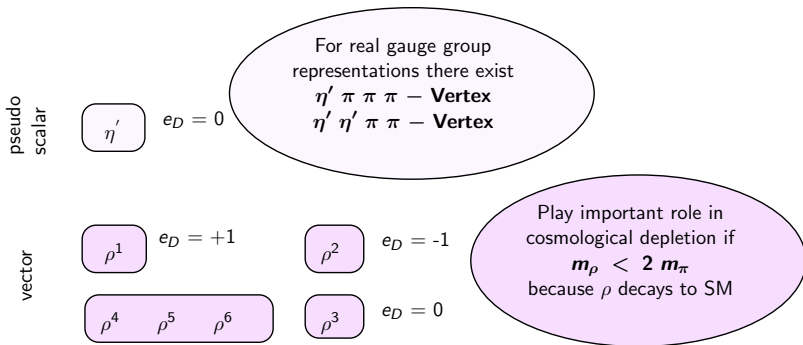
 $e_D = +1$

$\pi^7$	$\pi^8$	$\pi^9$
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 $e_D = -1$

## Further particles to be considered

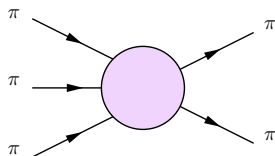
Dependent on the mass-spectrum further particles will take part in the cosmological depletion of dark matter.



Dark matter depletion via self-cannibalizing  $3 \rightarrow 2$  process.

$$\frac{N_C}{f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \text{Tr}\{\pi \partial_\mu \pi \partial_\nu \pi \partial_\rho \pi \partial_\sigma \pi\}$$

Results from topological  
Wess-Zumino-Witten terms.



Standard construction of these terms not possible since

$$\pi_4(SU(4)/SO(4)) \neq 0$$

# What was done so far

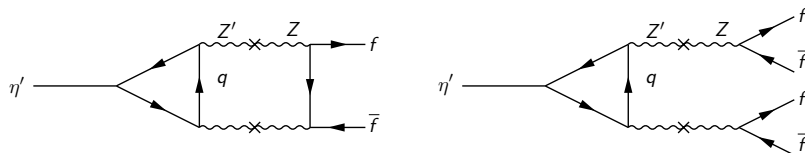
- Investigation of UV-theory
  - Identification of relevant structures
  - Investigation of symmetries and flavor structure
  - Construction of gauge invariant interpolating operators.
- Modelling the effective description
  - First order chiral effective Lagrangian for pions under  $U(1)_D$ .
  - Inclusion of the  $\eta'$ -meson in the effective description.

# Further steps in this work

- Construction of topological terms.
- Investigation of the freeze out process and the competing processes:
  - Pion annihilation decay via  $Z'$
  - Pion depletion via  $\eta$ -decay
  - Pion depletion via  $\rho$ 's
  - $3 \rightarrow 2$  self-cannibalizing of pions
- Investigation of self interaction strength.

Bonus content

Dependent on the mass  $m_{\eta'}$  and lifetime  $\tau_{\eta'}$  of the  $\eta'$  the dark matter scenario might be sufficiently altered or spoiled.



$\eta'$ -decay may lower the relic abundance of dark matter significantly and may render the model invalid.

Limiting cases:

- $\eta'$  decays almost instantly  
 $\Rightarrow$  **NO DARK MATTER**
- $\tau_{\eta'} \approx$  age of universe.  
 $\Rightarrow$  Pion abundance not affected.  
BBN or CMB constraints on  $m_{\eta'}$ .



# Interpolating operators for pions

	$\overline{\Psi}_C T_n^\Psi \Psi + (\overline{\Psi}_C T_n^\Psi \Psi)^*$	$J^D$		$\overline{\Psi}_C T_N^\pi \Psi + (\overline{\Psi}_C T_N^\pi \Psi)^*$	$I_3$	$B$
$\pi_1$	$\frac{1}{\sqrt{2}} (\bar{u}\gamma^5 d + \bar{d}\gamma^5 u)$	$1^-$	$\pi^A$	$\bar{u}\gamma^5 d$	1	0
$\pi_2$	$\frac{1}{\sqrt{2}} (\bar{u}\gamma^5 d - \bar{d}\gamma^5 u)$	$1^-$	$\pi^B$	$\bar{d}\gamma^5 u$	-1	0
$\pi_3$	$\frac{1}{\sqrt{2}} (\bar{u}\gamma^5 u - \bar{d}\gamma^5 d)$	$1^-$	$\pi^C$	$\frac{1}{\sqrt{2}} (\bar{u}\gamma^5 u - \bar{d}\gamma^5 d)$	0	0
$\pi_4$	$\frac{1}{2} (\bar{u}_C\gamma^5 u + \bar{u}\gamma^5 u_C)$	$1^-$	$\pi^D$	$\frac{1}{\sqrt{2}} \bar{u}_C\gamma^5 u$	-1	-1
$\pi_5$	$\frac{1}{2} (\bar{d}_C\gamma^5 d + \bar{d}\gamma^5 d_C)$	$1^-$	$\pi^E$	$\frac{1}{\sqrt{2}} \bar{d}\gamma^5 d_C$	1	-1
$\pi_6$	$\frac{1}{\sqrt{2}} (\bar{u}\gamma^5 d_C + \bar{u}_C\gamma^5 d)$	$1^-$	$\pi^F$	$\bar{u}\gamma^5 d_C$	0	-1
$\pi_7$	$\frac{1}{2} (\bar{u}_C\gamma^5 u - \bar{u}\gamma^5 u_C)$	$1^-$	$\pi^G$	$\frac{1}{\sqrt{2}} \bar{u}\gamma^5 u_C$	1	1
$\pi_8$	$\frac{1}{2} (\bar{d}\gamma^5 d_C - \bar{d}_C\gamma^5 d)$	$1^-$	$\pi^H$	$\frac{1}{\sqrt{2}} \bar{d}_C\gamma^5 d$	-1	1
$\pi_9$	$\frac{1}{\sqrt{2}} (\bar{u}\gamma^5 d_C - \bar{u}_C\gamma^5 d)$	$1^-$	$\pi^I$	$\bar{u}_C\gamma^5 d$	0	1