

# Spatial and Time Cluster Analysis of Fermi-LAT High Energy Gamma Ray Photons

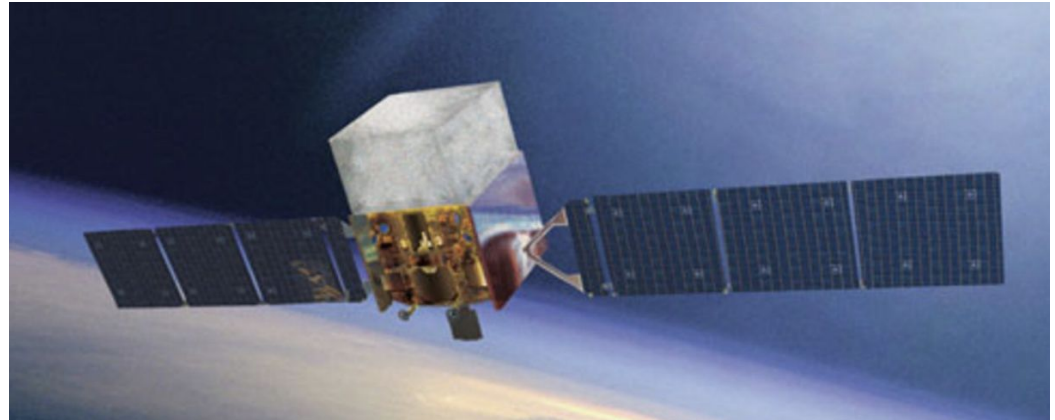
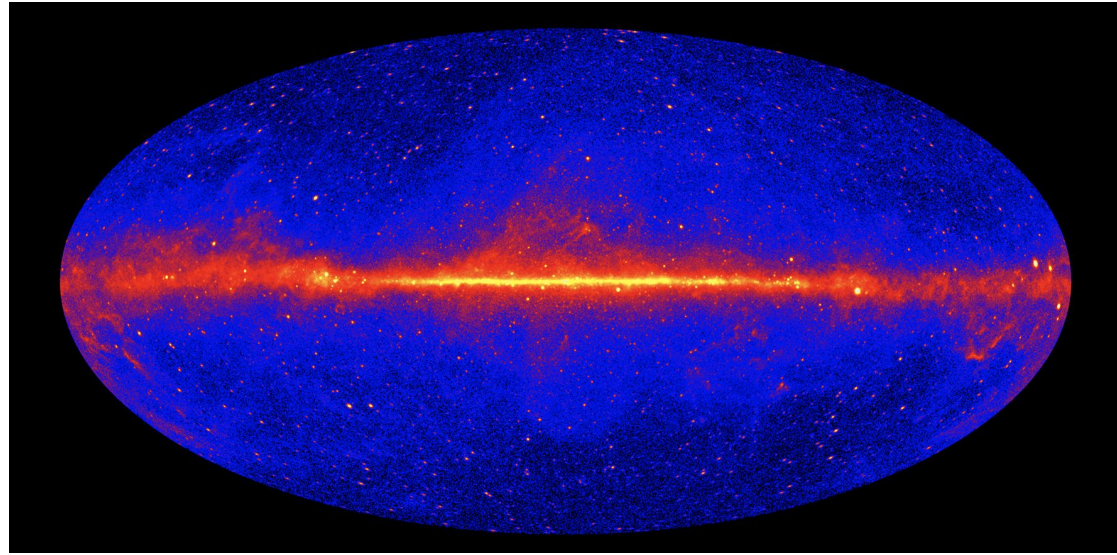
Summer Project Presentation

Honey Htun

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# Fermi - LAT

- **Energy Range:** 20 MeV to  $> 300$  GeV
- **Provide event-resolved observational data**
- **Events described by tuples:** sky coordinates, arrival time, energy
- **Capabilities:** Maps  $\gamma$ -ray sources, studies cosmic bursts, pulsars, blazars, dark matter
- **Data Products:** 3FHL, 4FGL

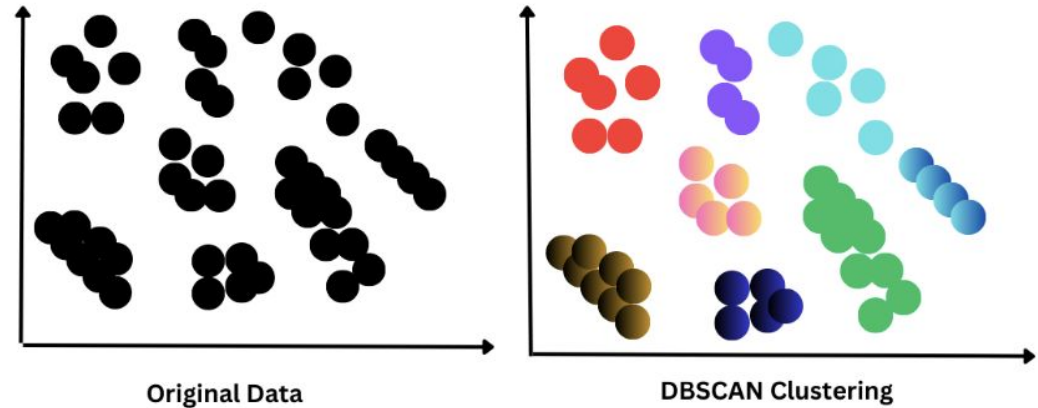


# Motive and Objectives

- Large Number of data points and modest resolution with all sky survey
- Identify Spatial and Temporal Clusters from Fermi photon counts images
- Using AI or ML automation to detect spatial clusters
- Follow up CTA Cherenkov small patch observations

# DBSCAN (Unsupervised Density Based Algorithm to identify Spatial Clusters)

- Clusters spatial data affected by background noise
- Discriminates signal (cluster) from background (noise) based on the local density of the event



Reference image :

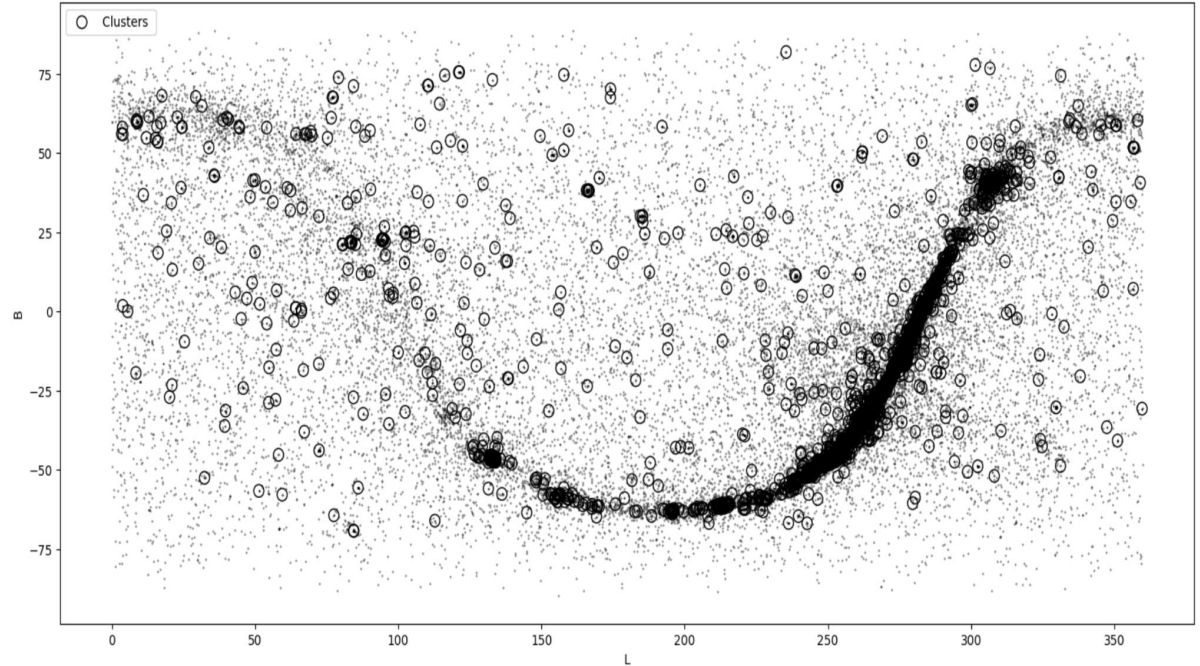
<https://medium.com/@jayaramganesh238/dbscan-clustering-dea27873ed30>

# Initial Scan of (photons with energy $> 50$ GeV)

DBSCAN Input

Number of photons

Search radius : Directly  
correlated with PSF



# Modification from DBSCAN output

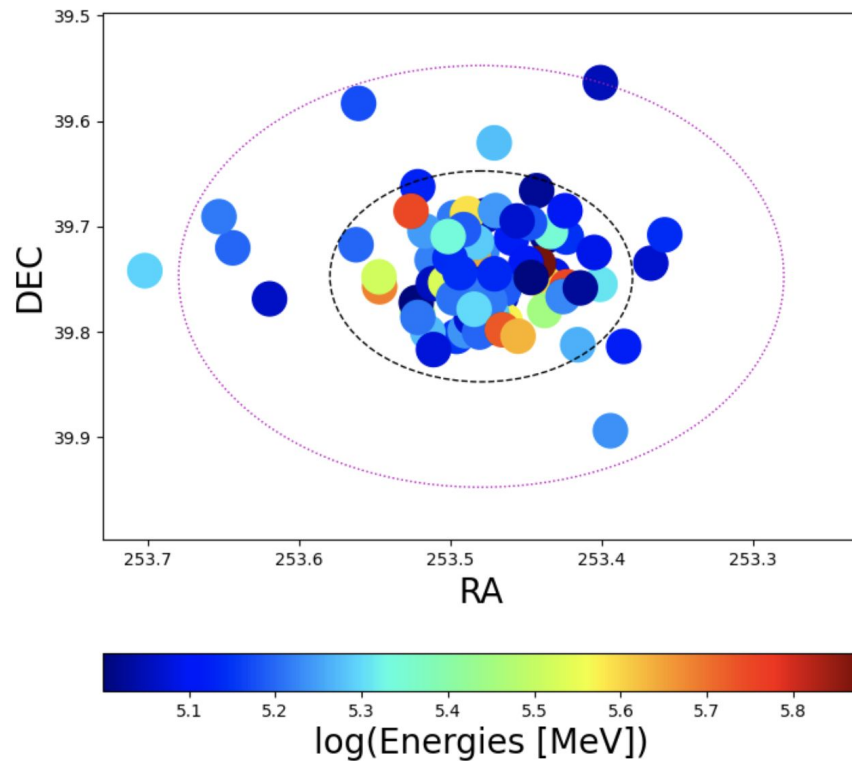
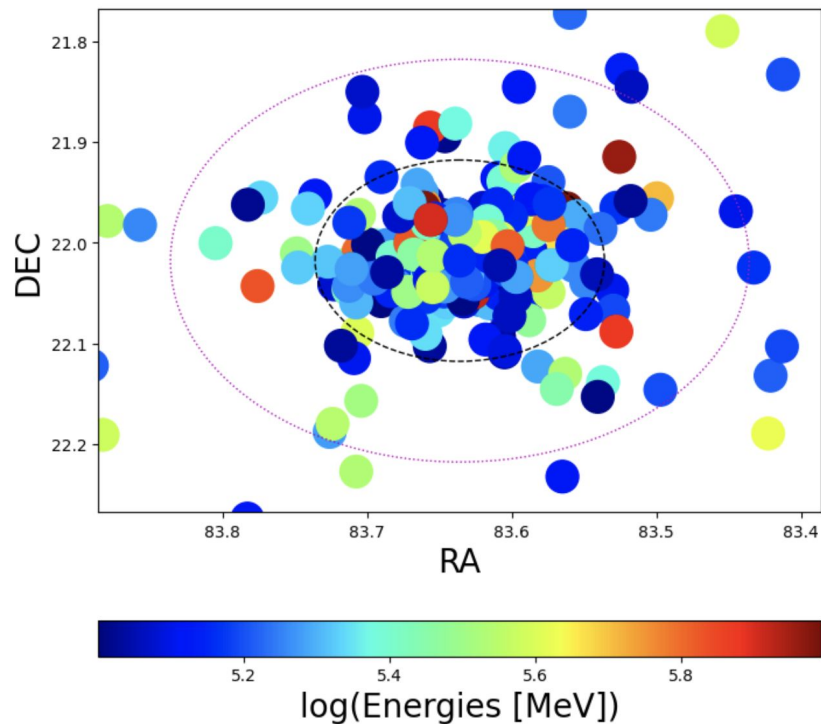
- Data Storage in a Tuple
- Coordinates of the cluster center
- Members of the cluster (location, arrival times, energy, total numbers)

```
def cluster_from_X(X,r, n , energy, time): #  
    db = DBSCAN(eps=r, min_samples=n).fit(X)
```

```
return cluster_centers , clusters, energies, arrival_time
```

```
def cluster_finder_updated(data, g_r, eg_r, g_n, eg_n):
```

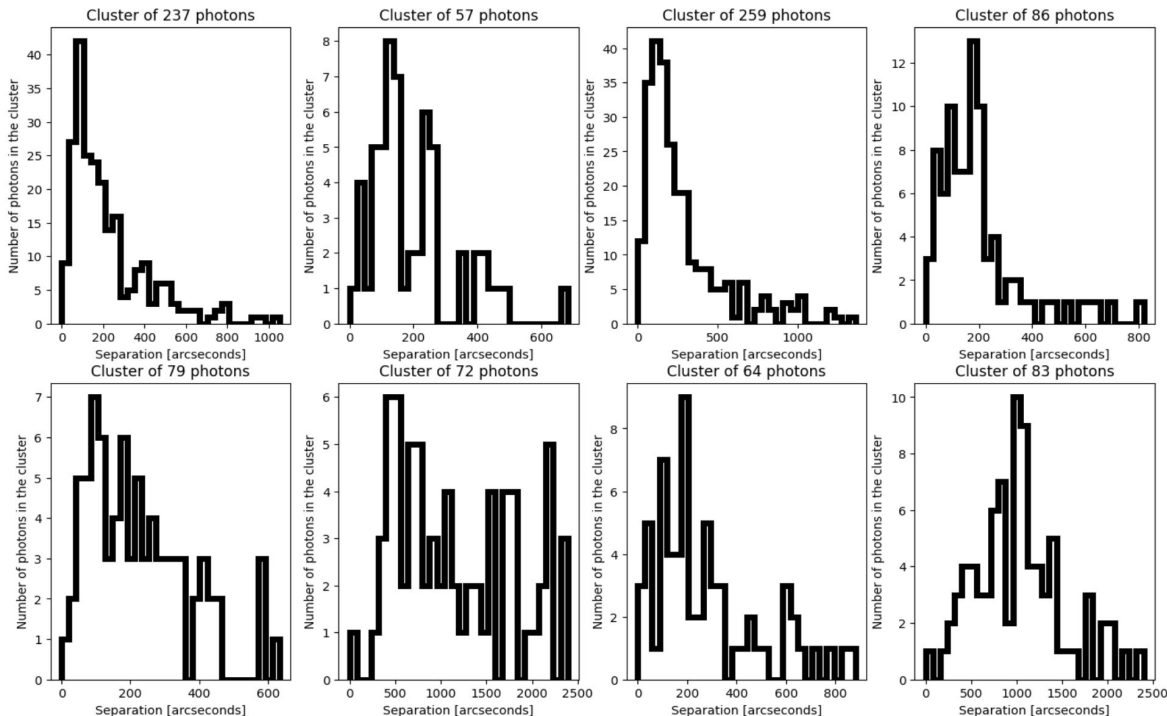
# Clusters Examples (0.1 and 0.2 degree circles)



# Radial Distributions - Point sources as well as Extended

Next step:

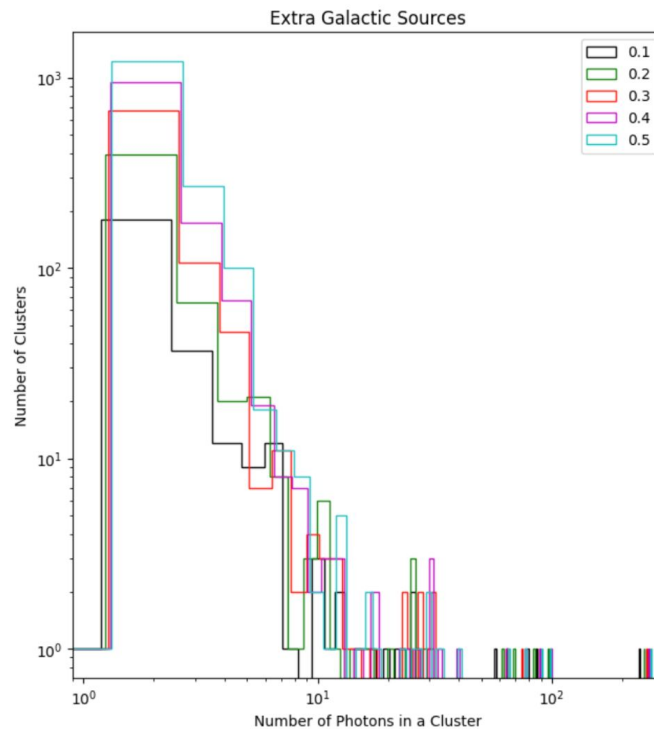
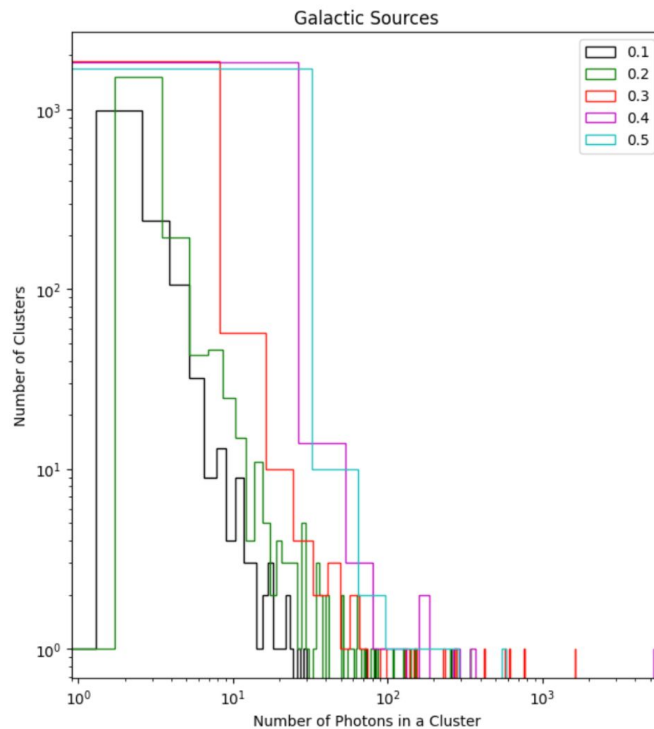
To find good  
input  
parameters





# Varying the Cluster Radius

Varying the search Radius[degrees]



# Time Clustering

For 78 months, probability of 2 photons occurring in 3°

$$P(k) = \lambda^k \cdot \frac{e^{-\lambda}}{k!}$$

$$k = 2 \lambda^2 \cdot \frac{e^{-\lambda}}{2!} = 0.25$$

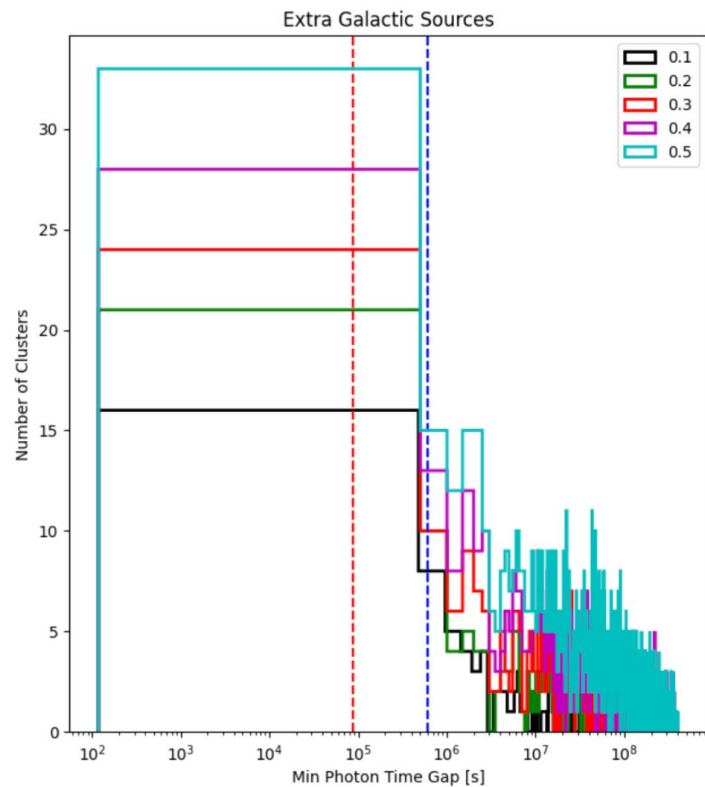
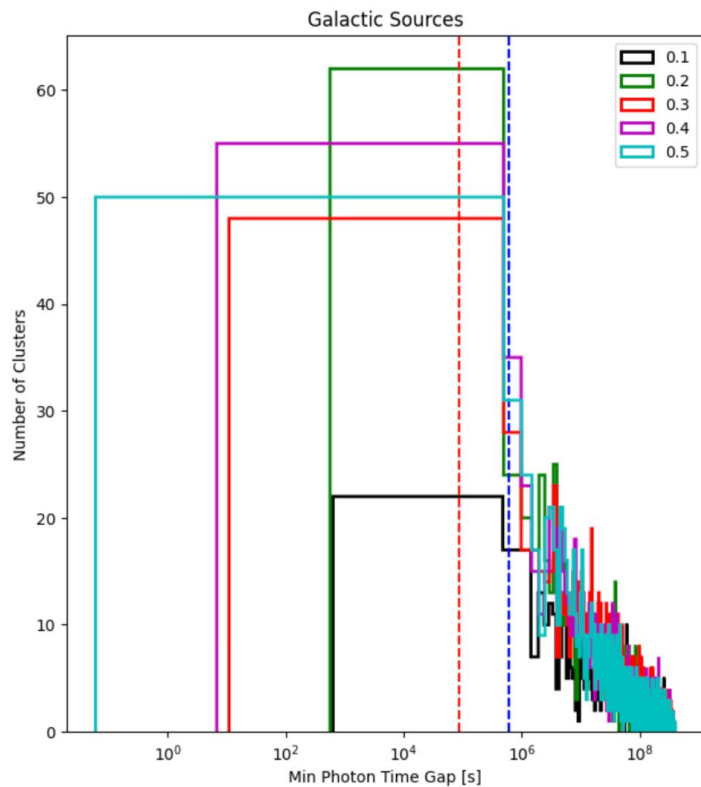
For one day :  $k = 0.00001$

$$k = 3 \lambda^3 \cdot \frac{e^{-\lambda}}{3!} = 0.21$$

For one week :  $k = 0.007$

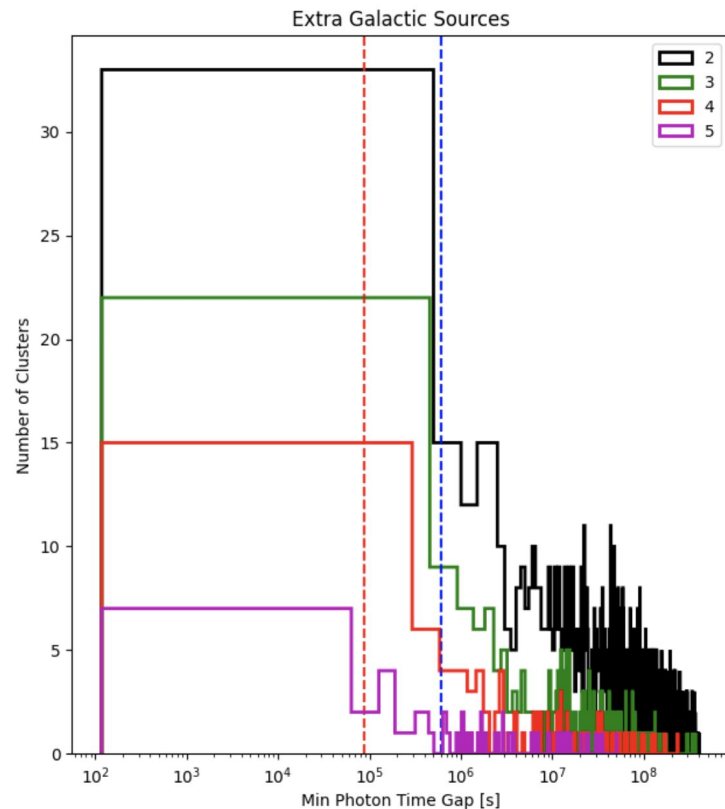
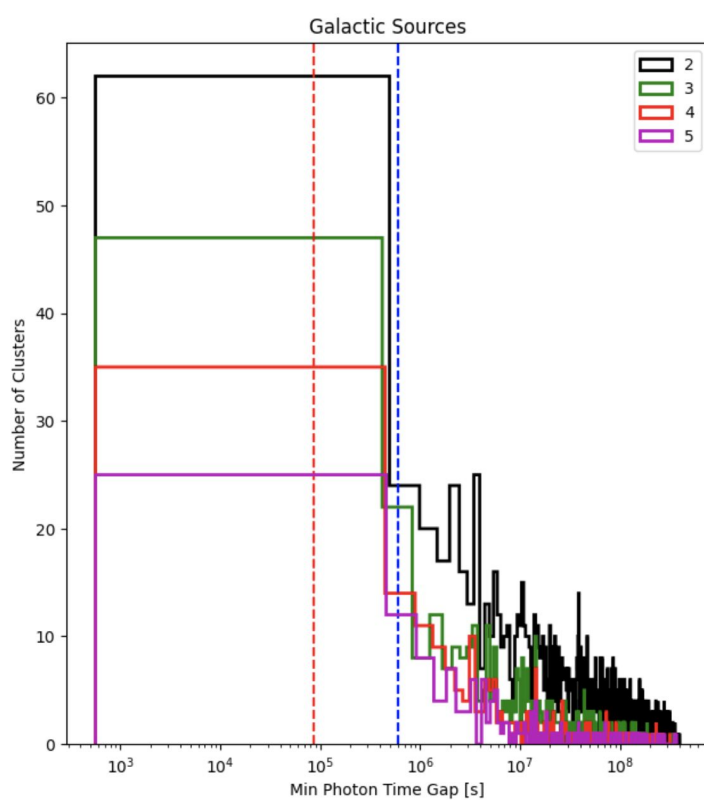
(Reference :  
Principe G. Master  
Thesis, 2015)

# Varying the Cluster Radius



# Varying the Photon Numbers with fixed Cluster Radius

$0.2^\circ$  for  $|b| < 10^\circ$  and  $0.5^\circ$  for  $|b| > 10^\circ$



# Final Parameters

## Galactic

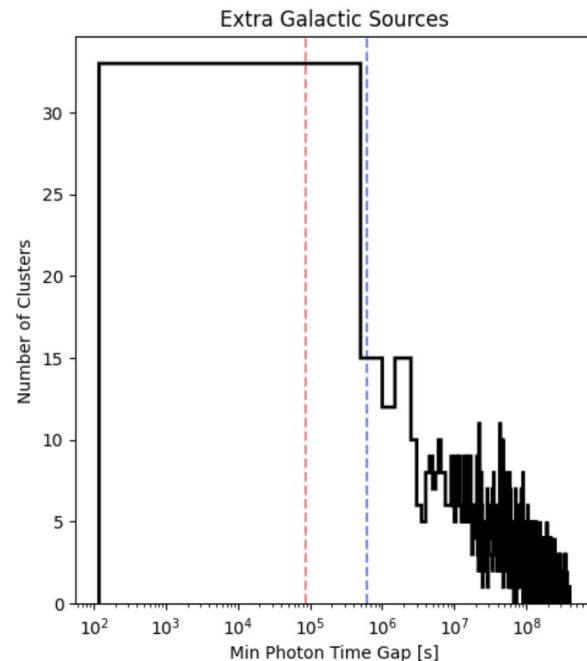
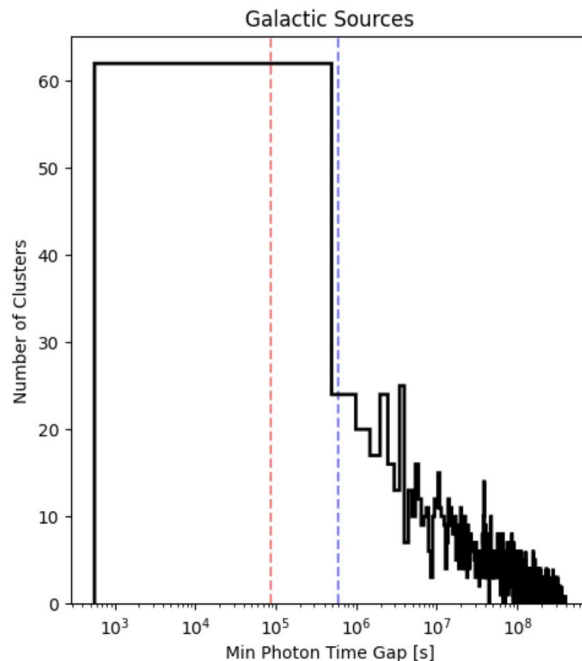
Cluster Radius : 0.2

Min Photon Number : 2

## Extra - Galactic

Cluster Radius : 0.5

Min Photon Number : 2



# Source Catalogs to check counterpart presence

## 3FHL (Third Catalog of Hard Fermi-LAT Sources)

- Covers energy range from 10 GeV to 2 TeV
- Includes 1,556 sources
- Utilizes 7 years of Fermi-LAT data
- Focuses on sources with high-energy  $\gamma$ -ray emissions
- Key for studying extreme  $\gamma$ -ray phenomena

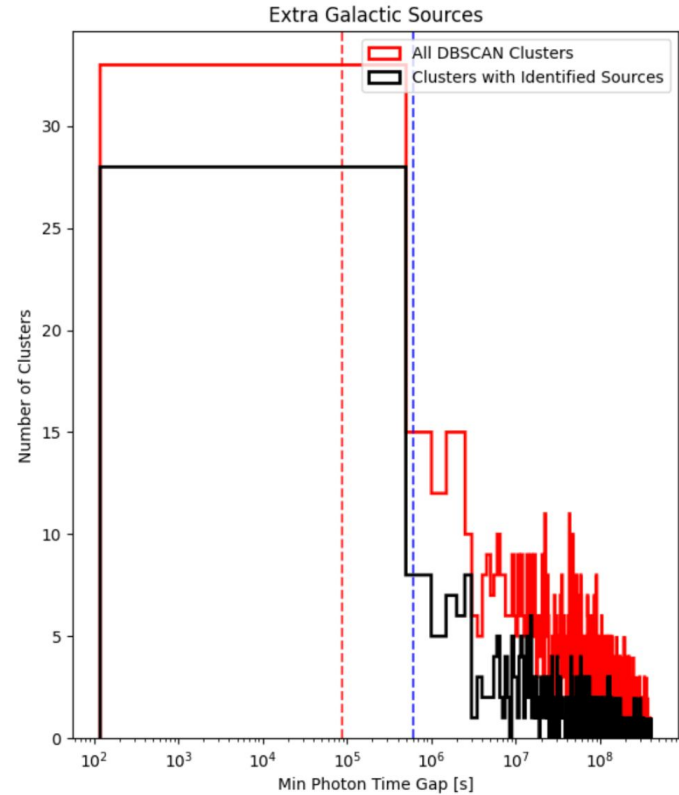
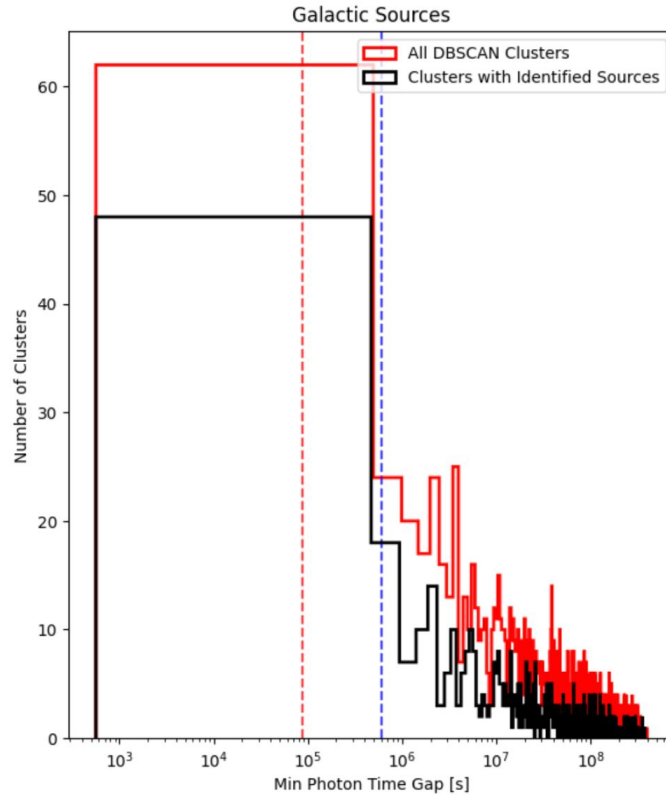
## 4FGL (Fourth Fermi-LAT Catalog)

- Covers energy range from 50 MeV to 1 TeV
- Includes over 5,000 sources
- Utilizes 8 years of Fermi-LAT data
- Comprehensive catalog with improved source localization
- Important for multi-wavelength and multi-messenger astronomy

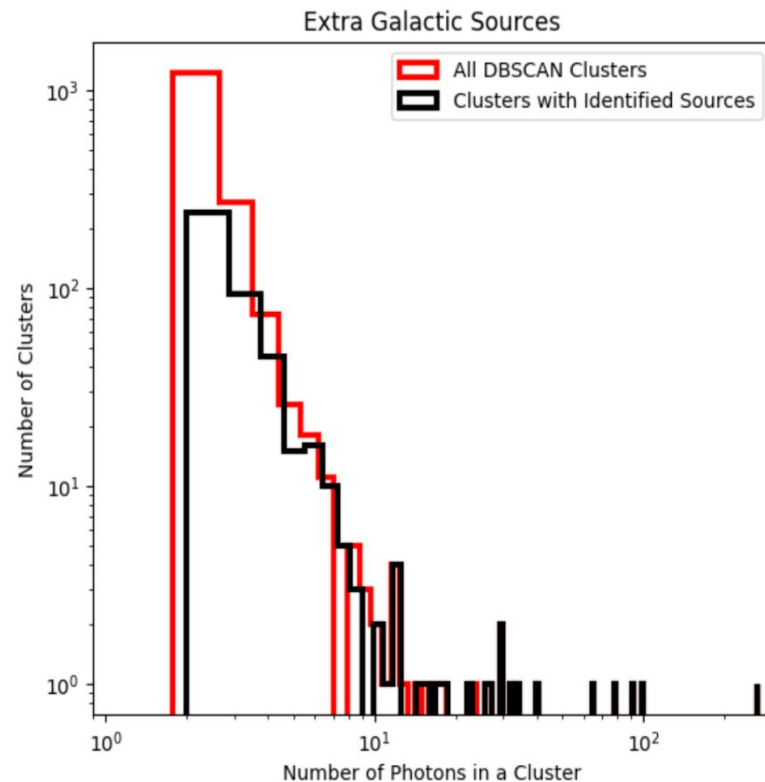
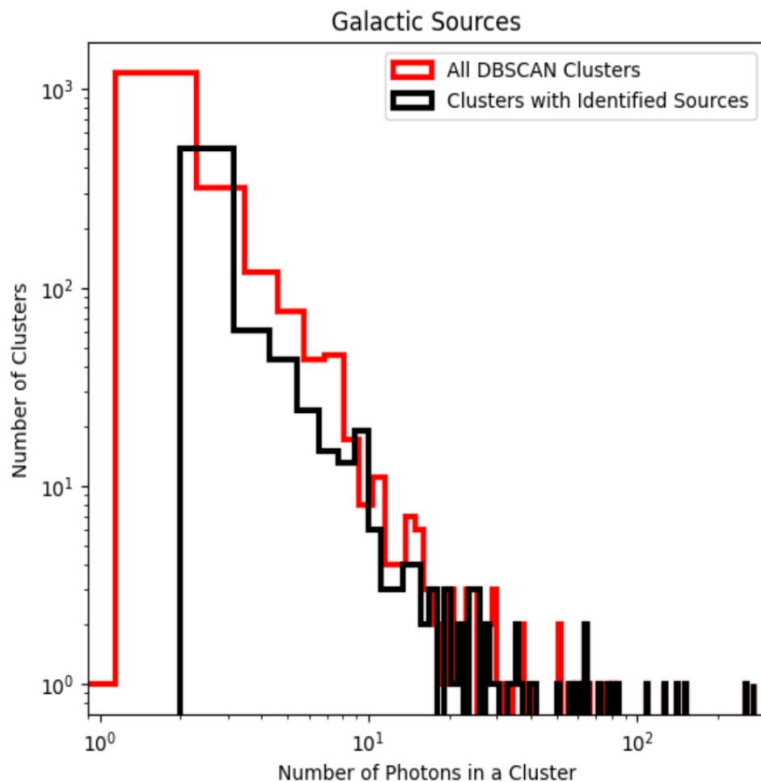
## TeV Catalogs

- Collects data from ground-based observatories (e.g., H.E.S.S., MAGIC, VERITAS)
- Covers energy range above 1 TeV
- Includes sources detected at very high energies
- Energetic astrophysical processes
- Complements Fermi-LAT catalogs for a full spectrum analysis

# Counterparts in 3FHL, 4FGL and TeV catalogs

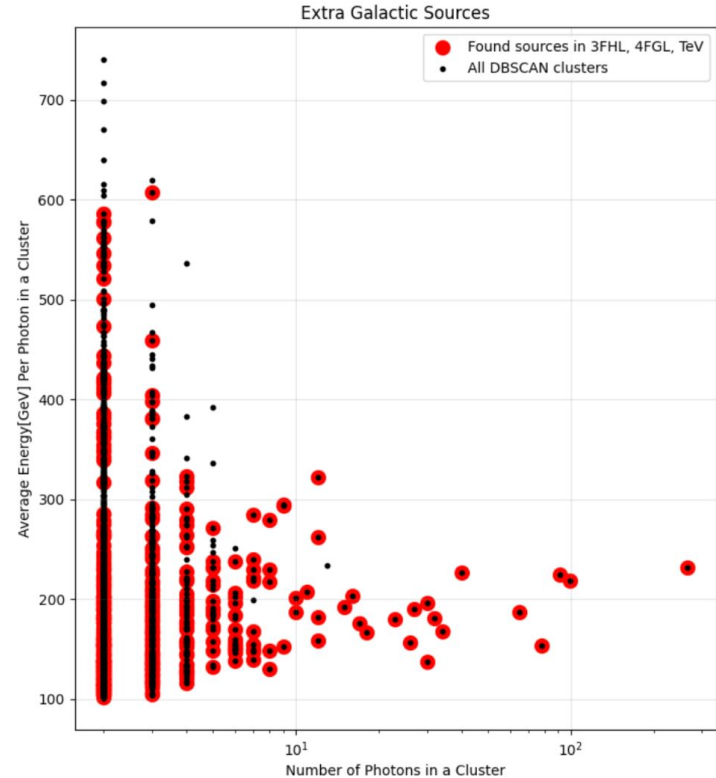
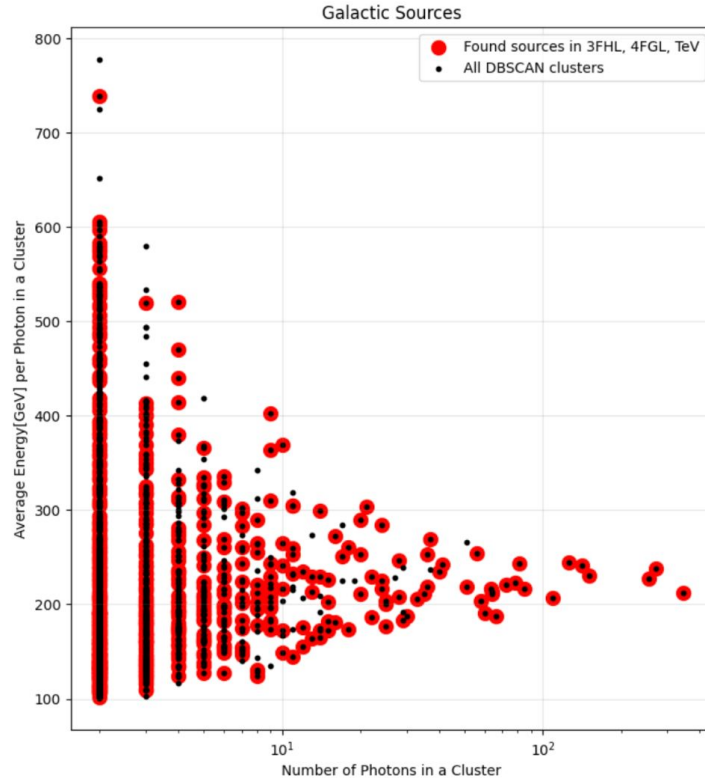


# Photon Distributions

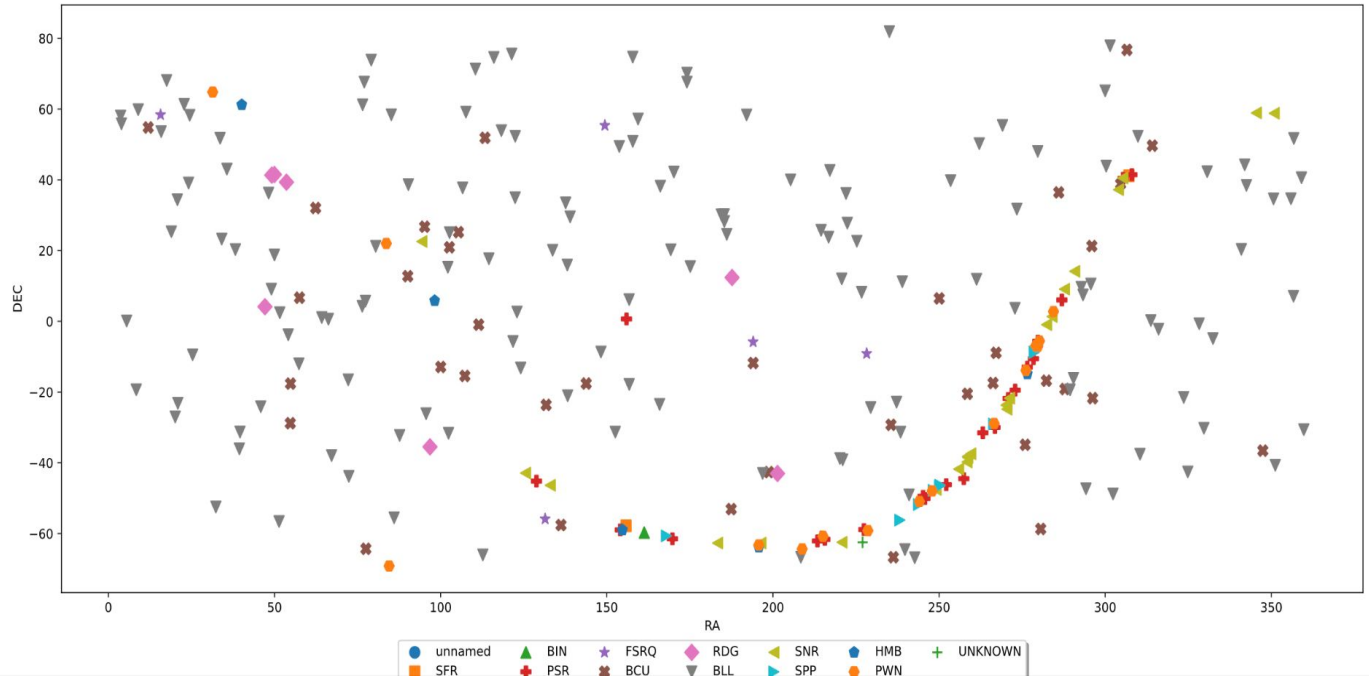




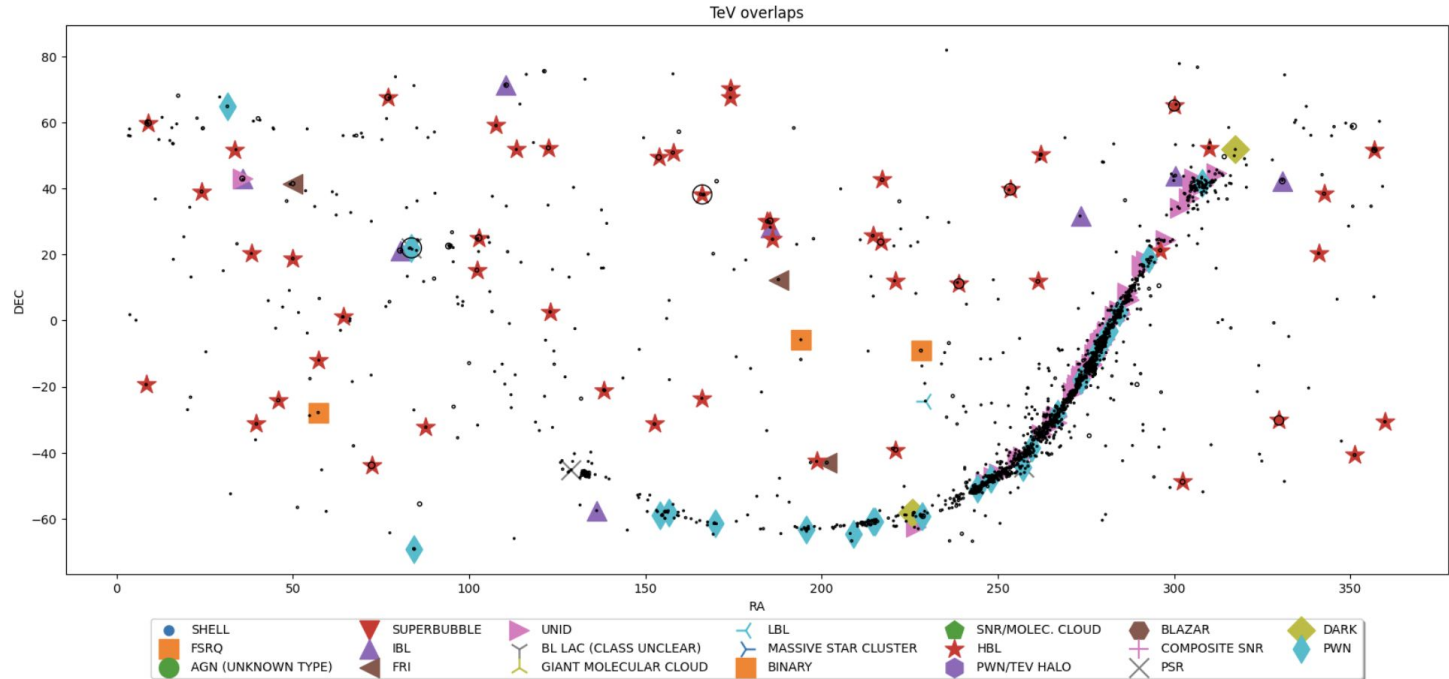
# Photon Numbers vs Average Energy inside the Clusters



# Type of overlapping 3FHL sources/clusters

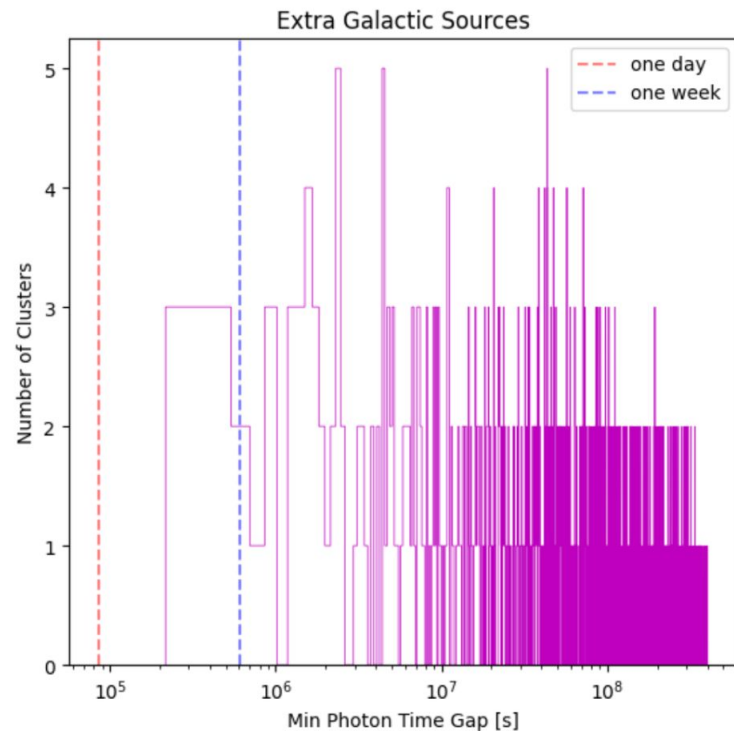
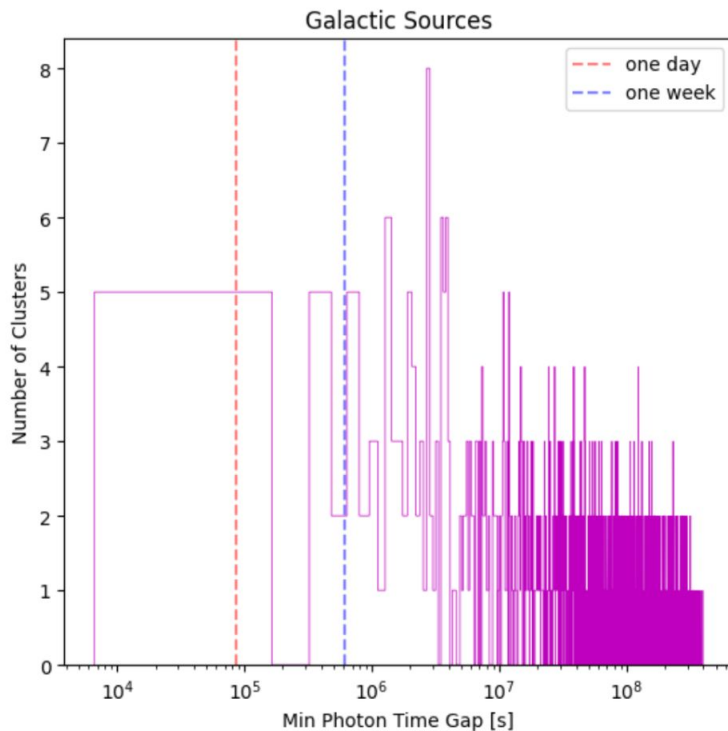


# Type of Overlapping TeV sources

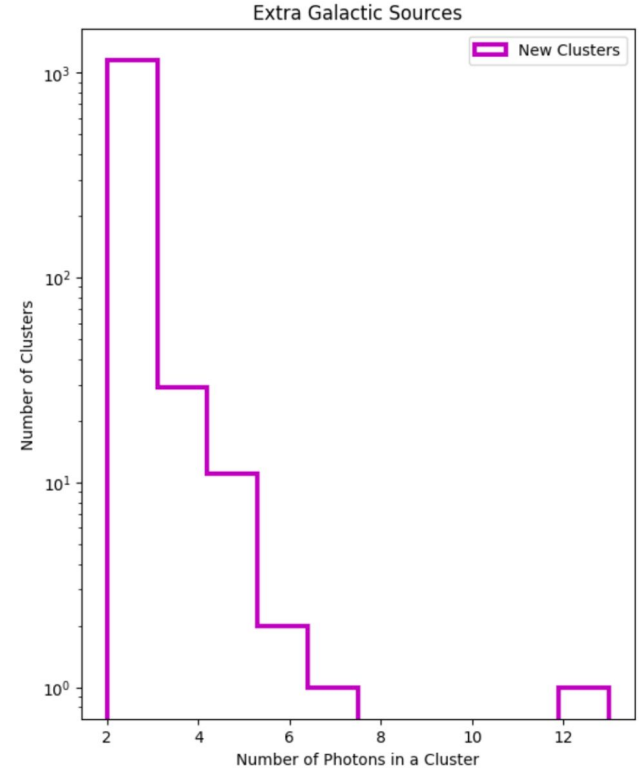
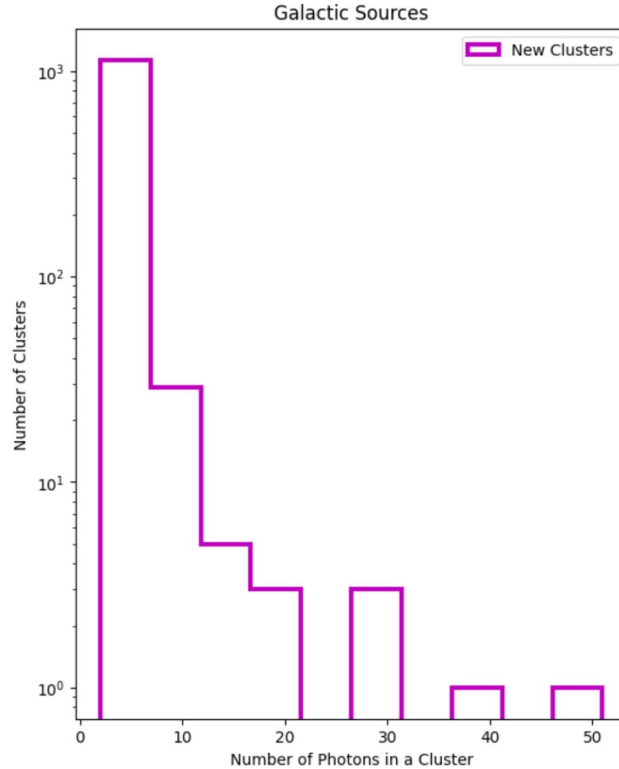


# Time Clustering of New Spatial Clusters

Clusters with no counterparts



# Photon Number Distribution in New Clusters

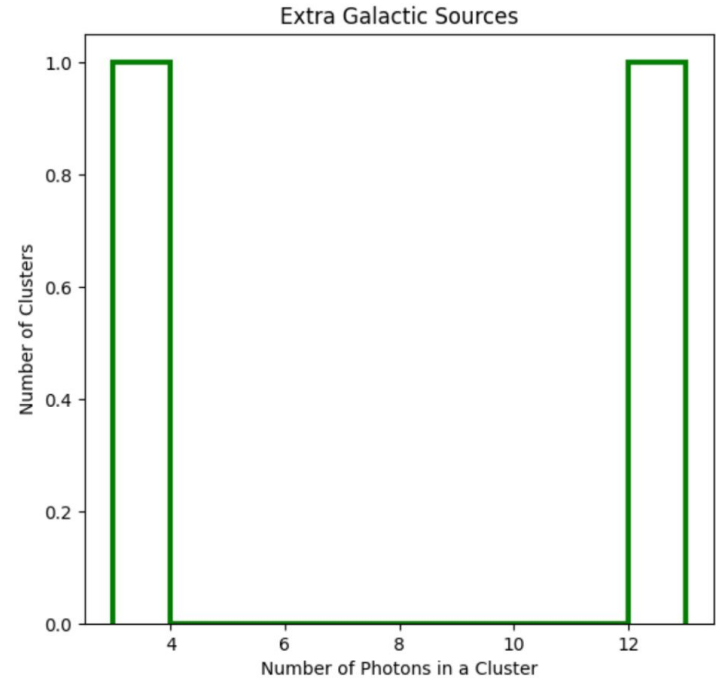
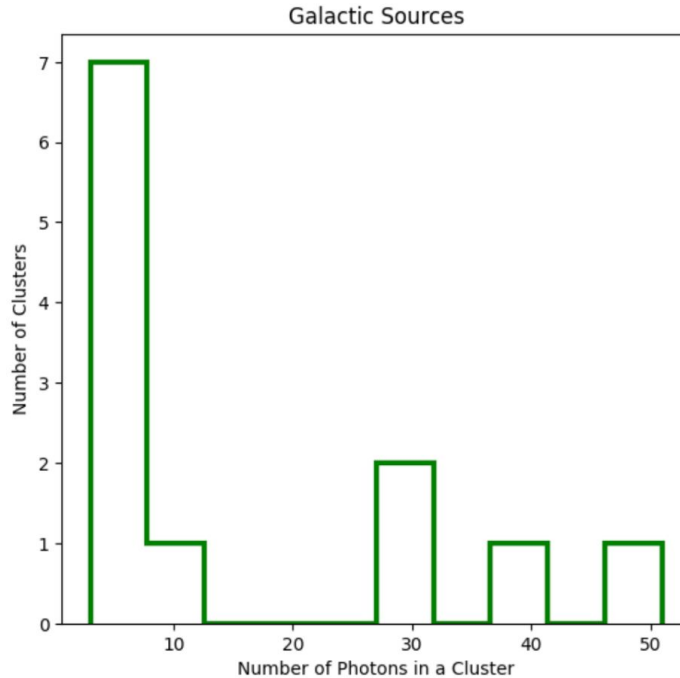


# New Clusters

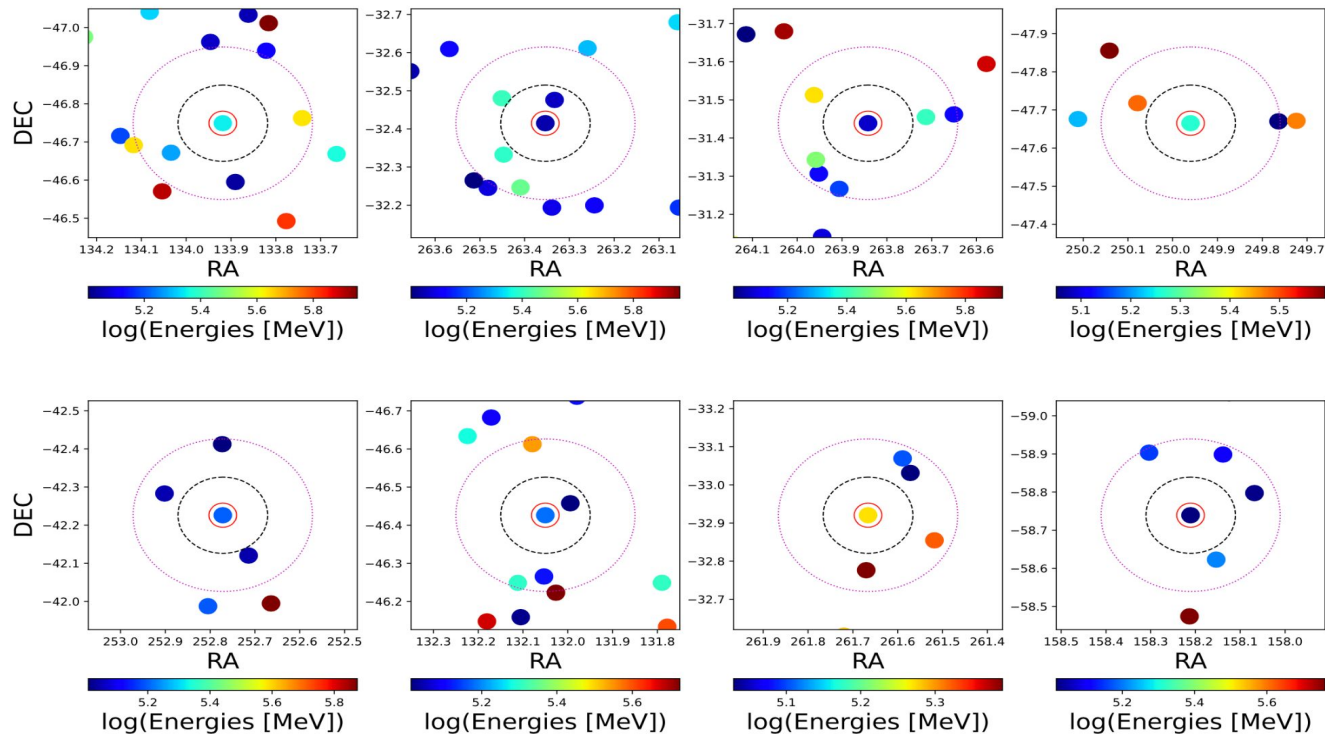
Total :	Galactic	Extra Galactic
Time Gap < 1 week :	12	7
N(Photons) > 2	338	221
N(Photons) > 2 & Time Gap < 1 week :	12	2

# Spatial and Temporal Clustering Final Results

New Time and Space Clusters

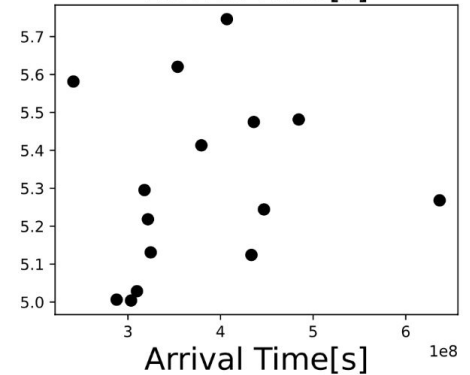
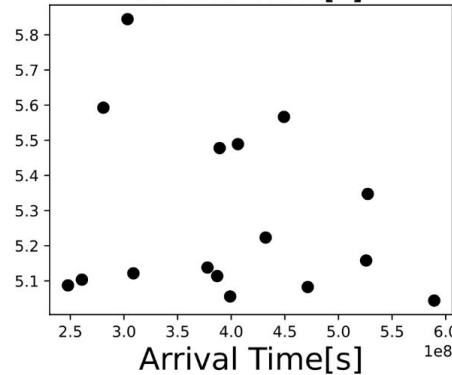
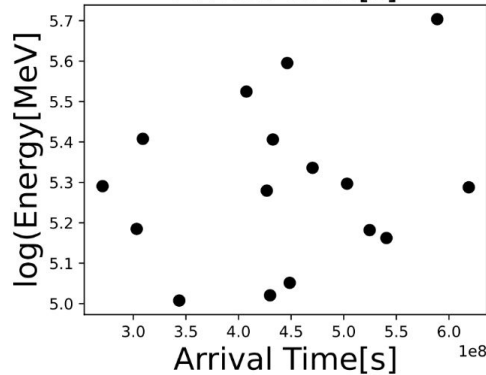
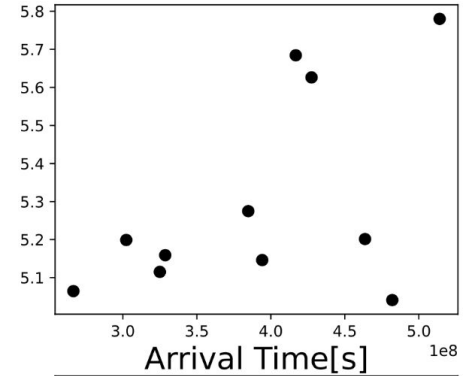
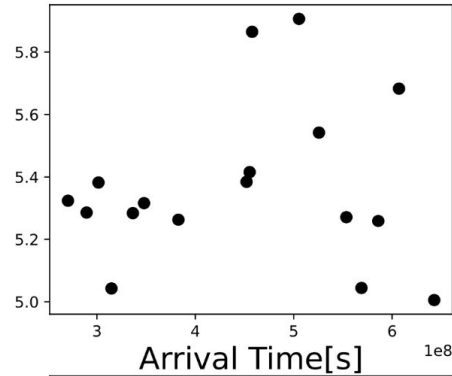
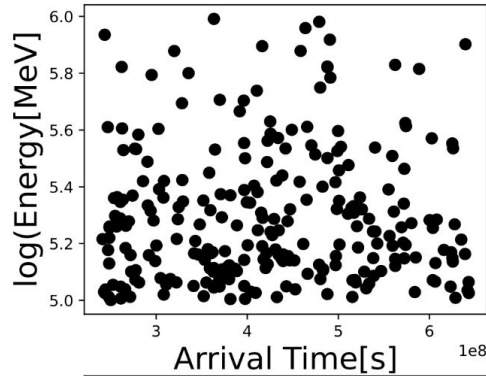


# New Photon Clusters

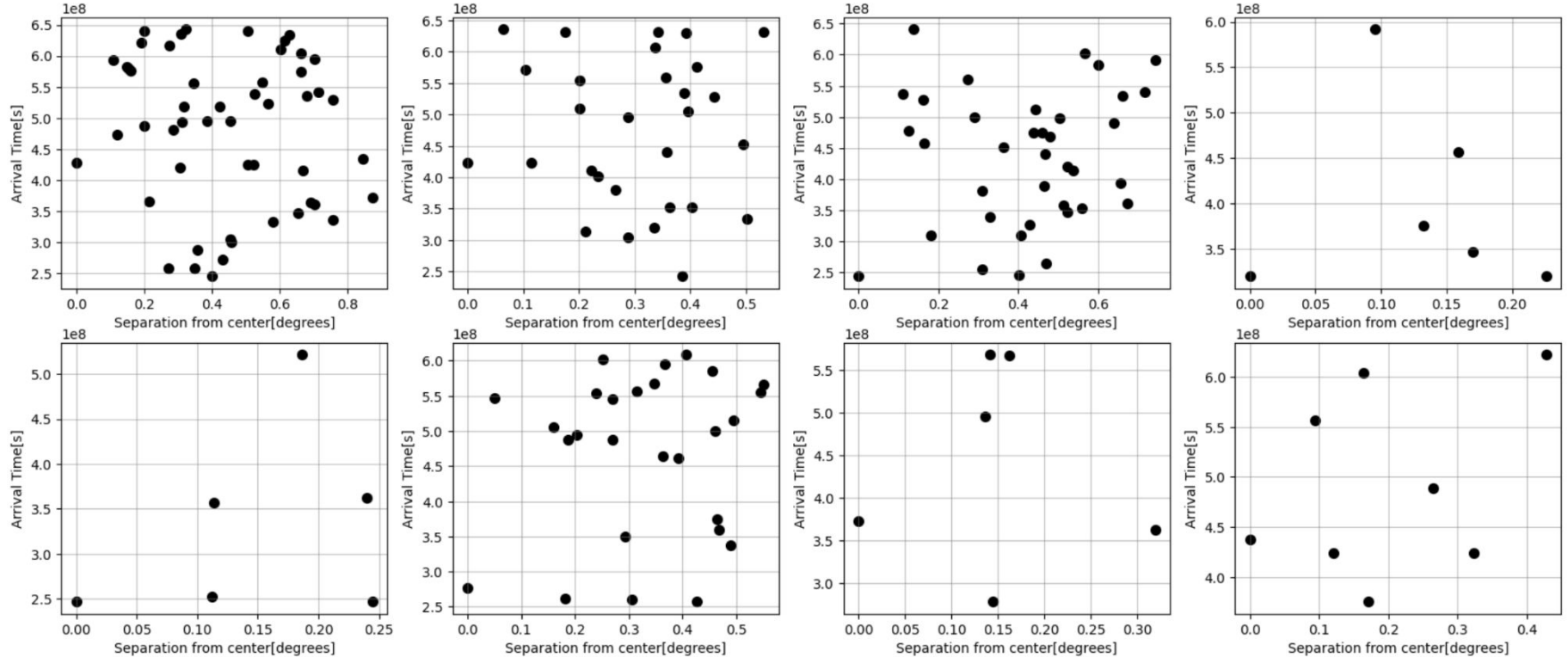




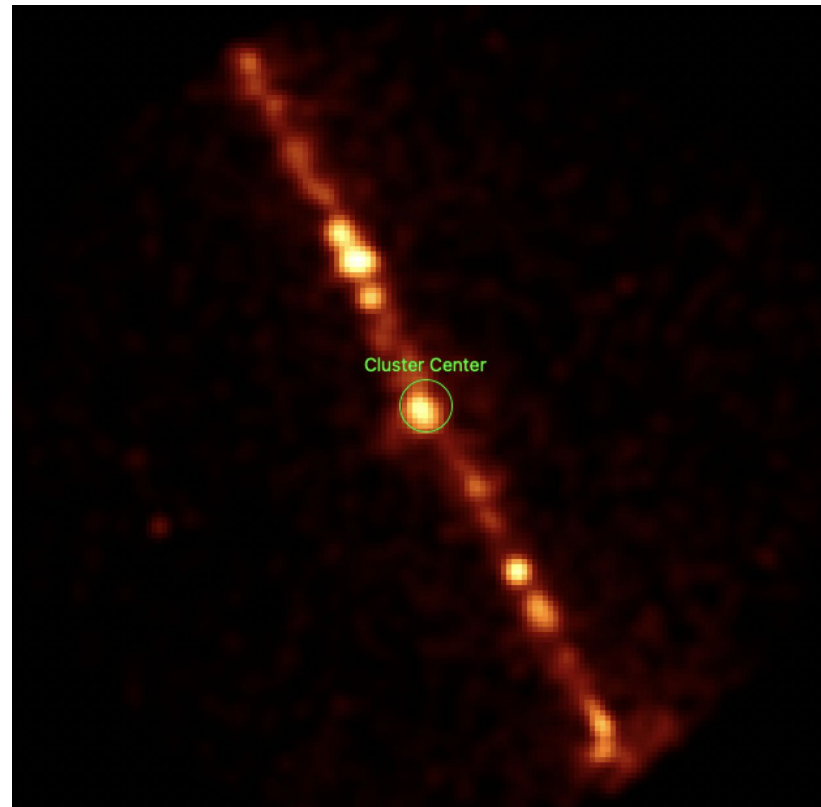
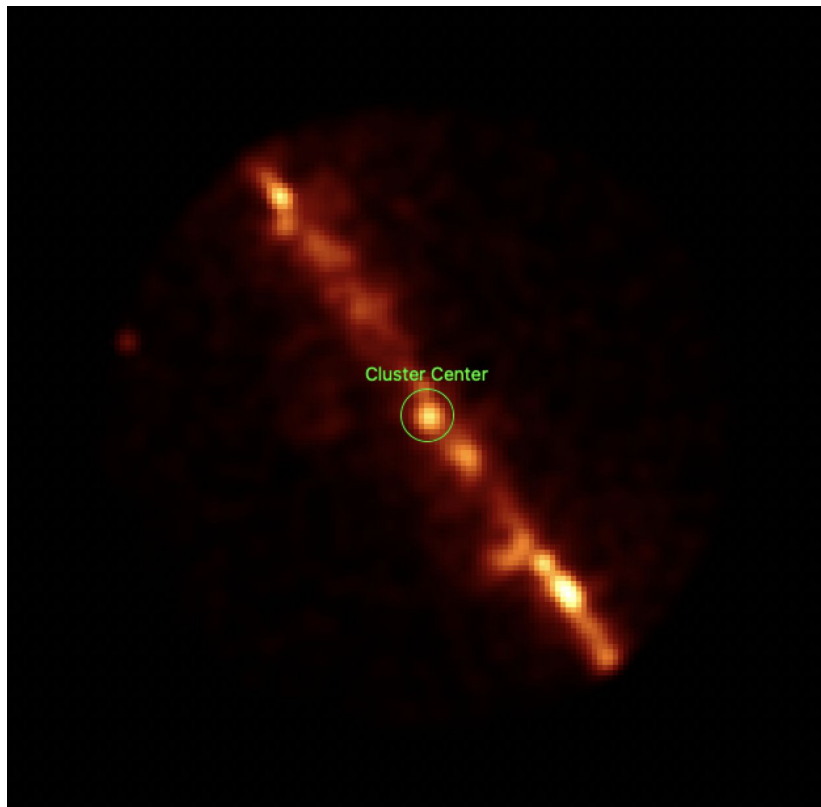
# Light Curves of some of the New clusters



# Arrival Times vs Separation from Center



## Cmaps of some the regions



# Future work (next semester)

Matching Energy Cuts

More Fermipy Analysis

More Statistical Analysis work

Publishing a catalog