



Module 2.2

Exoplanets

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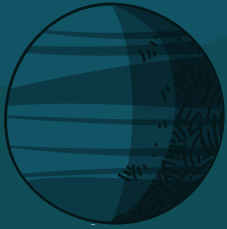
CONTENT OF THIS MODULE

1. What are exoplanets and its types?
2. How do we find exoplanets?
3. Famous exoplanet mission
4. NASA Exoplanet Archive
5. Gaia Data Release

01 Exoplanets and its types

Wend

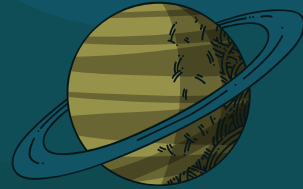
MAJOR TYPES



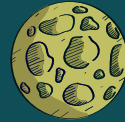
Gas Giants



Super Earth



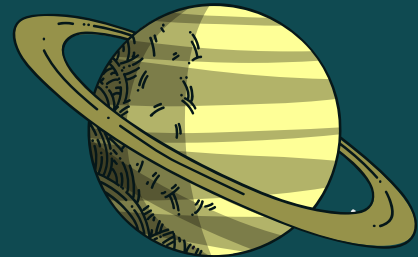
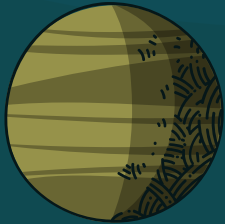
Neptune like



Terrestrial

GAS GIANTS

A gas giant is a large planet mostly composed of helium and/or hydrogen. These planets, like Jupiter and Saturn in our solar system, don't have hard surfaces and instead have swirling gases above a solid core. Gas giant exoplanets can be much larger than Jupiter, and much closer to their stars than anything found in our solar system.



51 Pegasi b

Famous Gas Giant

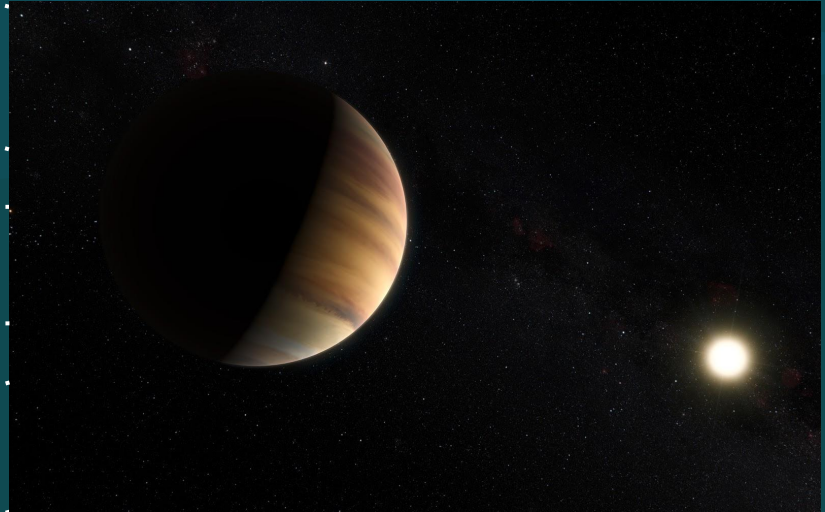
Discovery : 1995

Mass : 0.45 Jupiter Mass

Radius: 1.27 Jupiter Radius

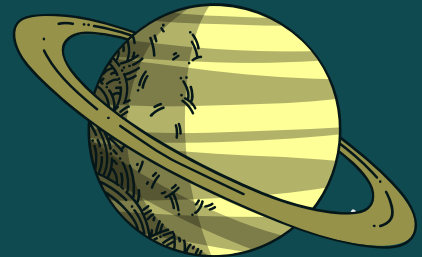
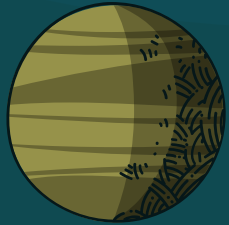
Orbital Period: 4.2 days

Orbits a G-type star



Super Earths

Super-Earths – a class of planets unlike any in our solar system – are more massive than Earth yet lighter than ice giants like Neptune and Uranus, and can be made of gas, rock or a combination of both. They are between twice the size of Earth and up to 10 times its mass.



Bernard's Star b

Famous super-earth

Discovered in 2018

Mass : 3.2 Earth Mass

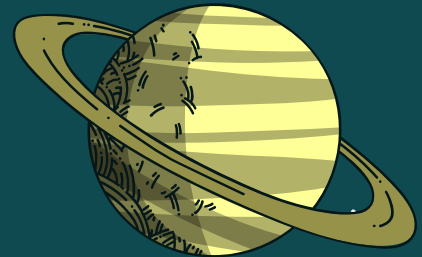
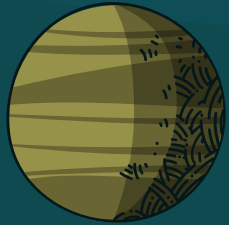
**2nd closest exoplanet (6ly
away)**

Orbital Period~ 233 days



NEPTUNE-LIKE

Neptunian exoplanets are similar in size to Neptune or Uranus in our solar system. Neptunian planets typically have hydrogen and helium-dominated atmospheres with cores of rock and heavier metals.



Kepler 1655 b

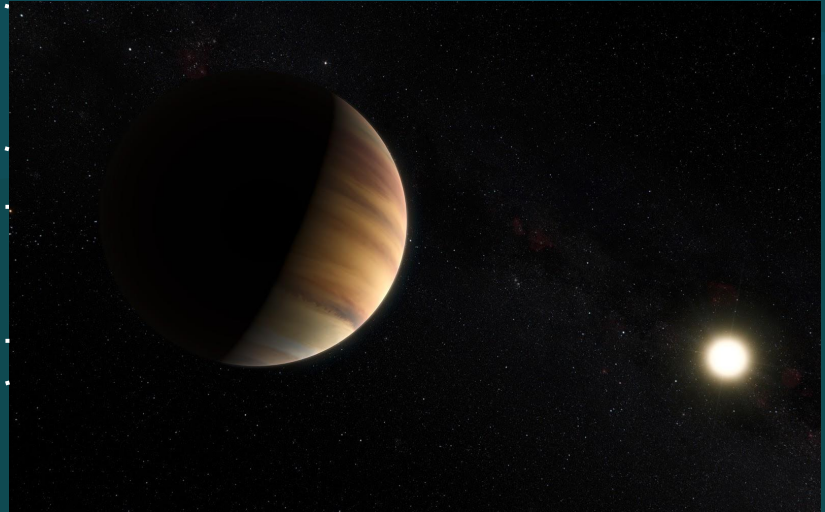
Detection method: Transit

Discovery : 2018

Mass : 5 Sun Mass

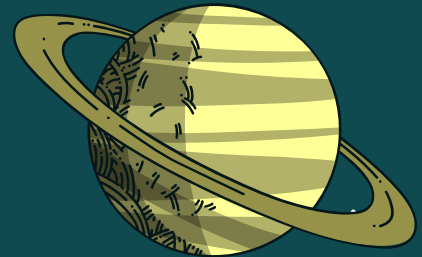
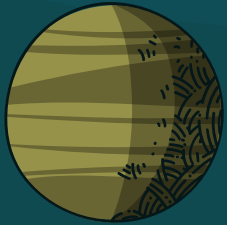
Orbital Radius: 0.103 AU

Orbital Period: 11.9 days



Terrestrial

In our solar system, Earth, Mars, Mercury and Venus are terrestrial, or rocky, planets. For planets outside our solar system, those between half of Earth's size to twice its radius are considered terrestrial and others may be even smaller. Exoplanets twice the size of Earth and larger may be rocky as well, but those are considered super-Earths.



Trappist-1e

Famous terrestrial

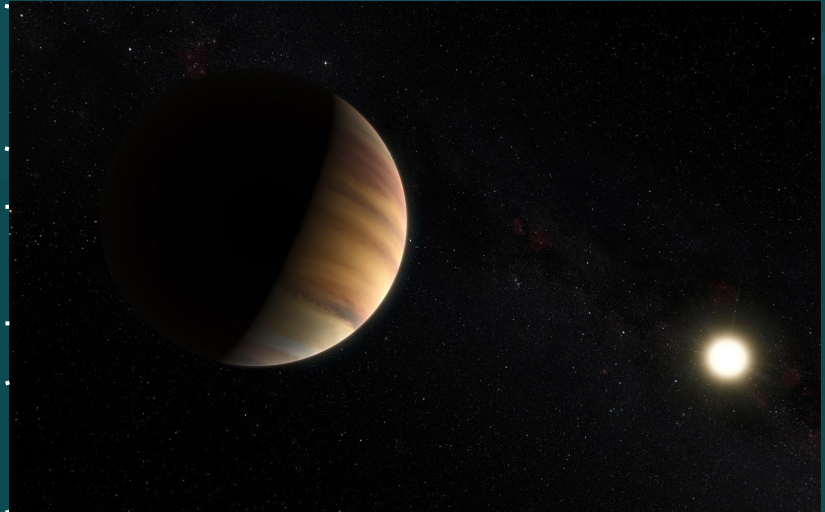
Discovery : 2017

Mass : 0.77 Earth Mass

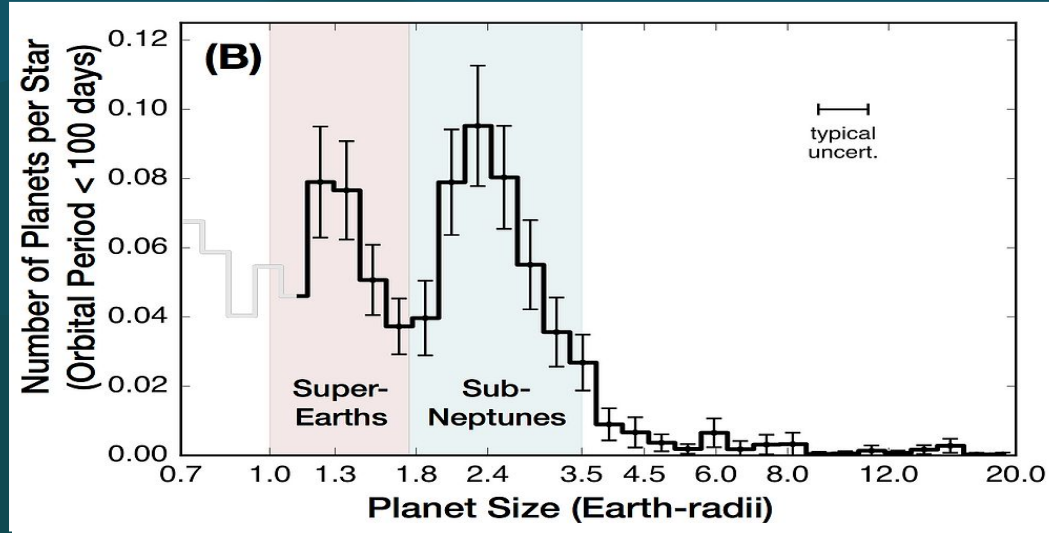
Radius: 0.91 Earth Radius

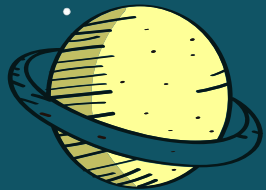
Orbital Period: 6.10 days

Orbital radius:
0.0293 AU



FULTON GAP

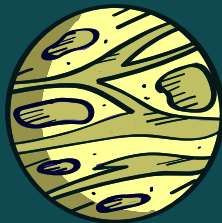
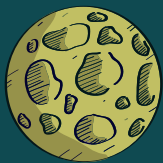
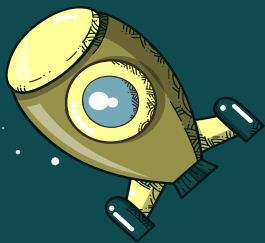
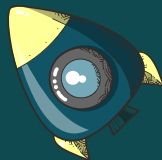




Why Exoplanets?

"Somewhere, something incredible is waiting to be known"

-Carl Sagan



$$N = R_{\star} \times f_p \times n_e \times f_e \times f_i \times f_c \times L$$

The number of
technologically advanced
civilizations in the
Milky Way galaxy

The rate of formation
of stars in the galaxy

The fraction of
those stars with
planetary systems

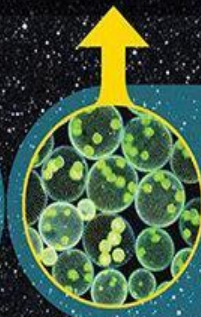
The number of planets,
per solar system,
with an environment
suitable for life

The fraction
of suitable planets
on which life
actually appears

The fraction
of life-bearing planets
on which intelligent life
emerges

The fraction of civilizations
that develop a technology that
releases detectable signs of
their existence into space

The length of time
such civilizations release
detectable signals
into space



$$A = N_{ast} \times$$

\times

$$f_{bt}$$

The number of technological
species that have formed
over the history of
the observable universe

The number of habitable
planets in a given volume
of the universe

The likelihood of a
technological species arising
on one of these planets

HABITABLE ZONE SIZE

X-RAY
IRRADIANCE

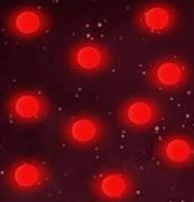
RELATIVE
ABUNDANCE

LONGEVITY

M



400x



100
Billion
Years

K

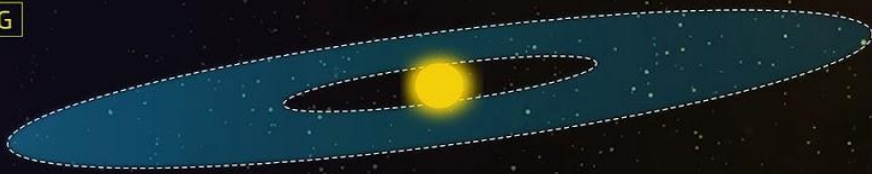


25x



40
Billion
Years

G



1x



10
Billion
Years

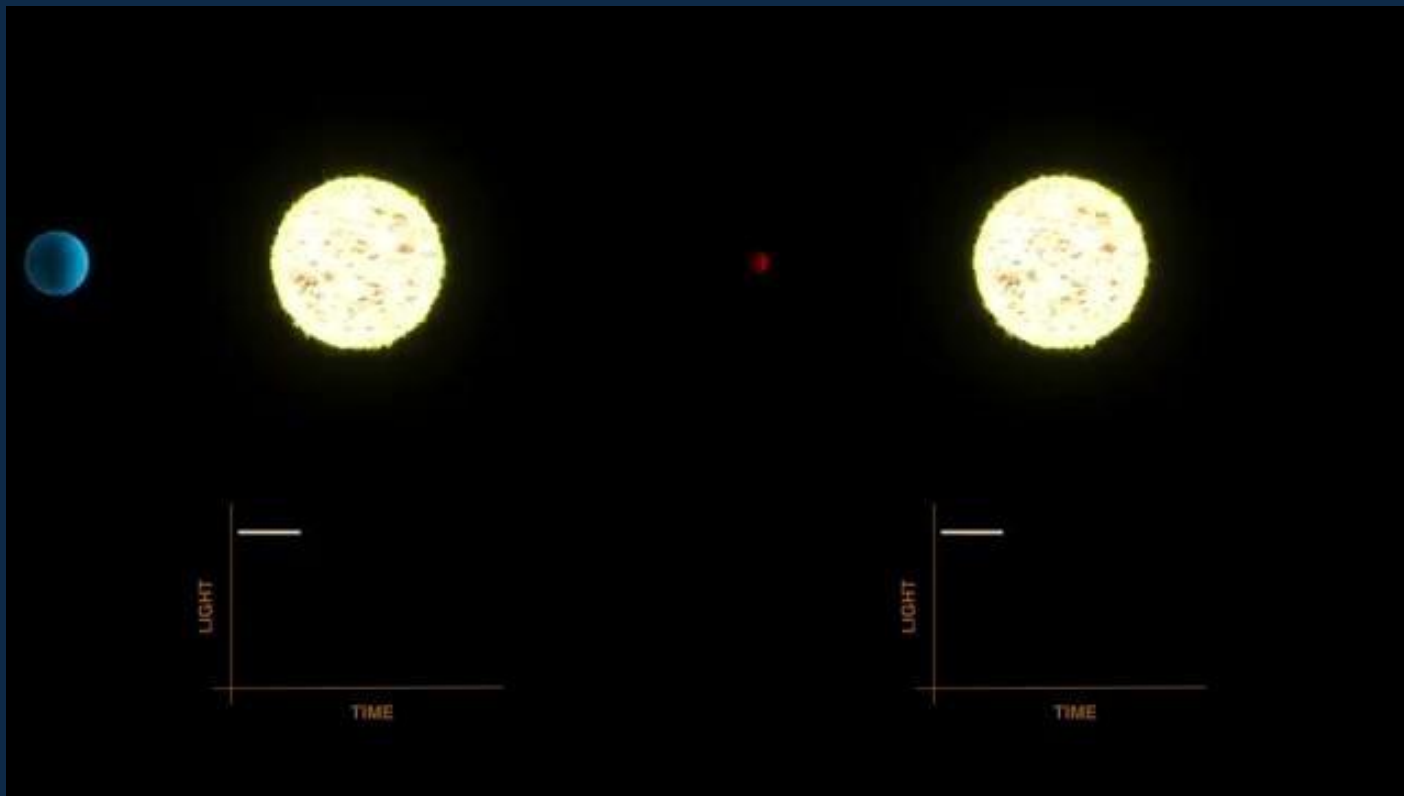
02 EXOPLANET DETECTION

- Transit
- Radial Velocity
- Gravitational Microlensing
- Direct Imaging
- Transit-timing variations
- Eclipse-timing variations
- Pulsar Timing
- Orbital Brightness Modulation
- Pulsation Timing Variation
- Disk Kinematics
- Astrometry

Single-Planet Transit



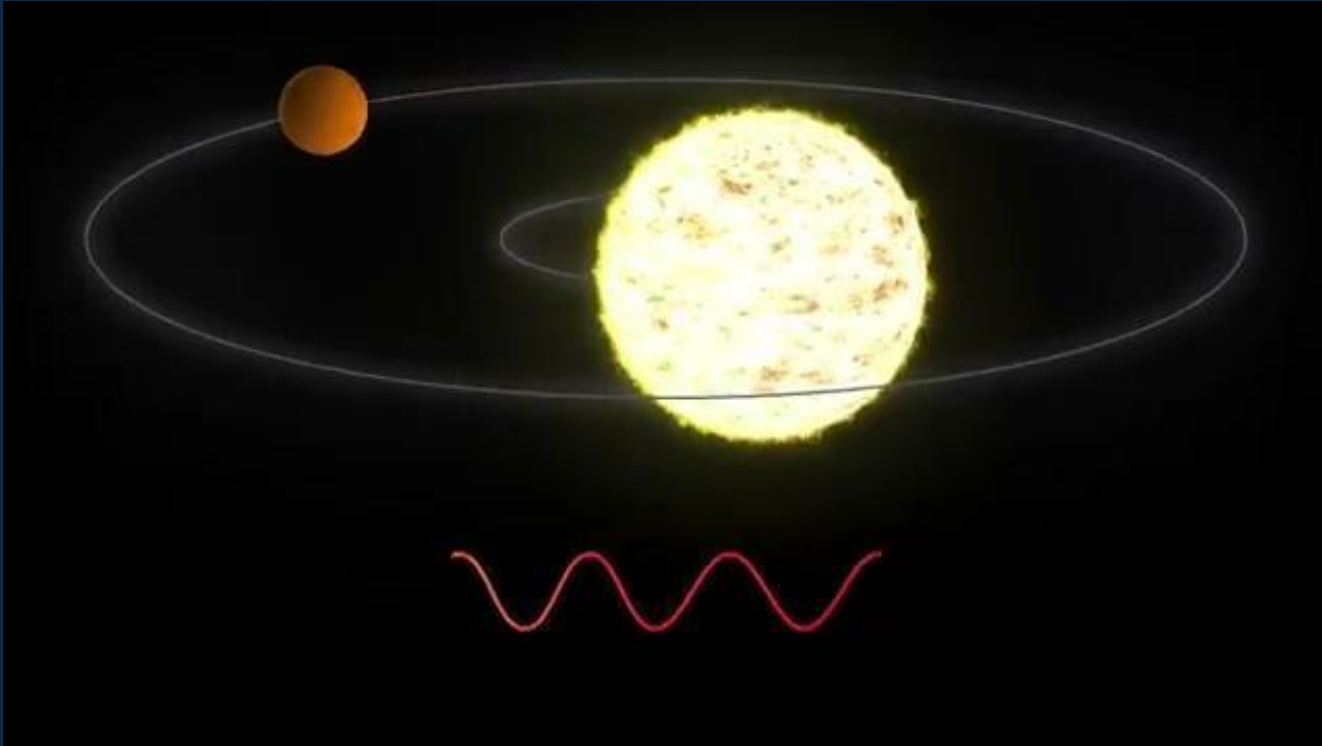
Transit: Different Planet Size



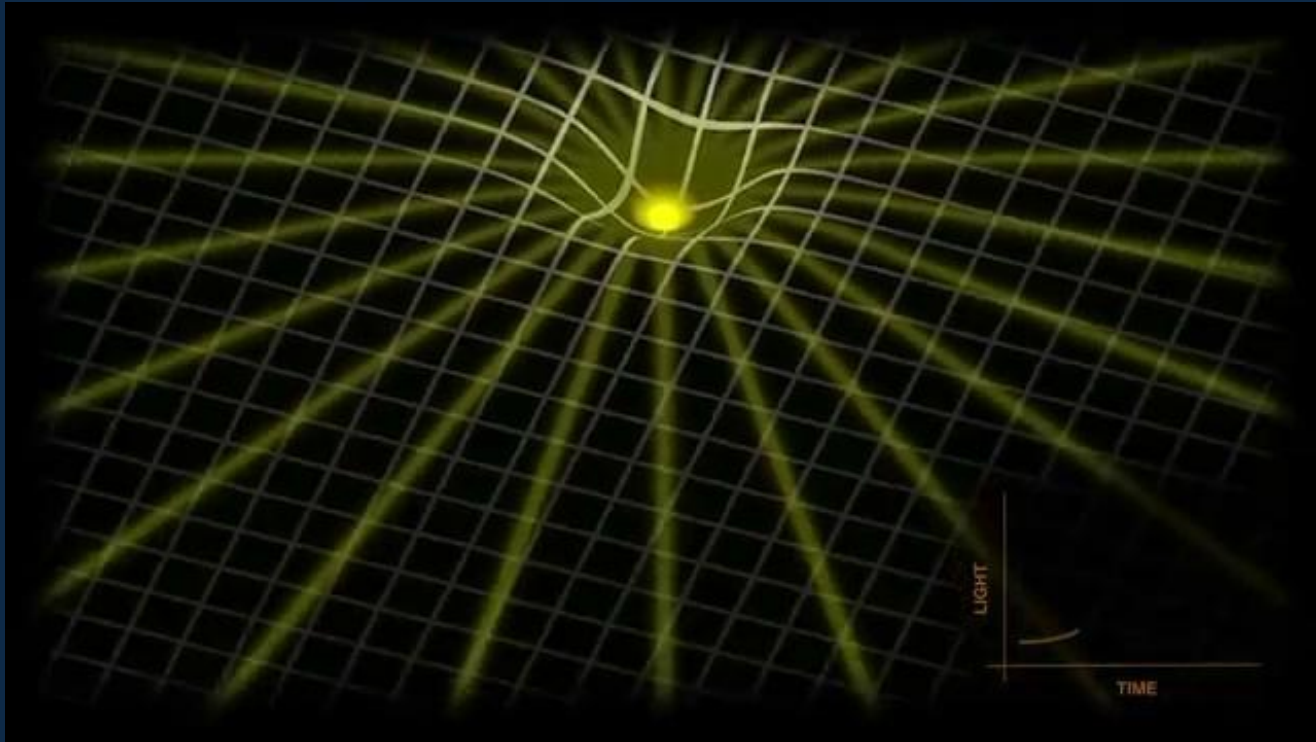
Multiple Planet Transit



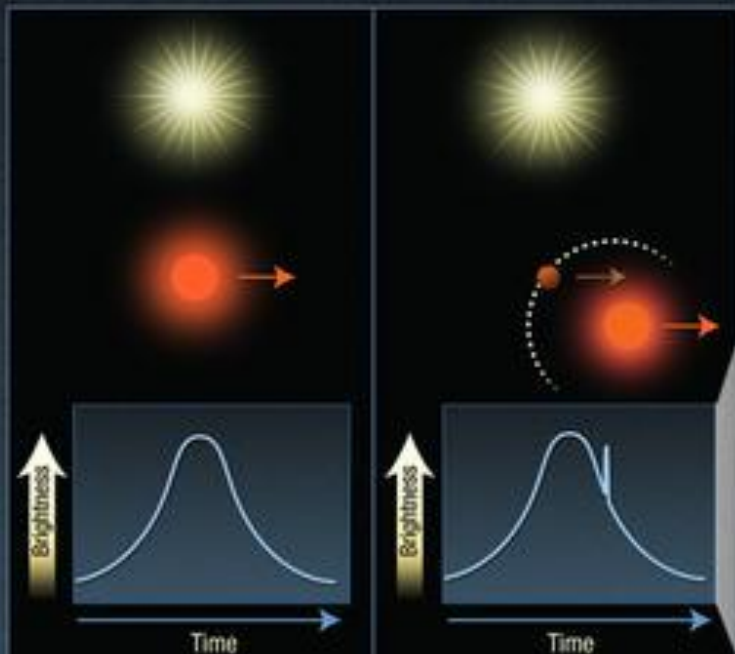
Radial Velocity



Gravitational Microlensing

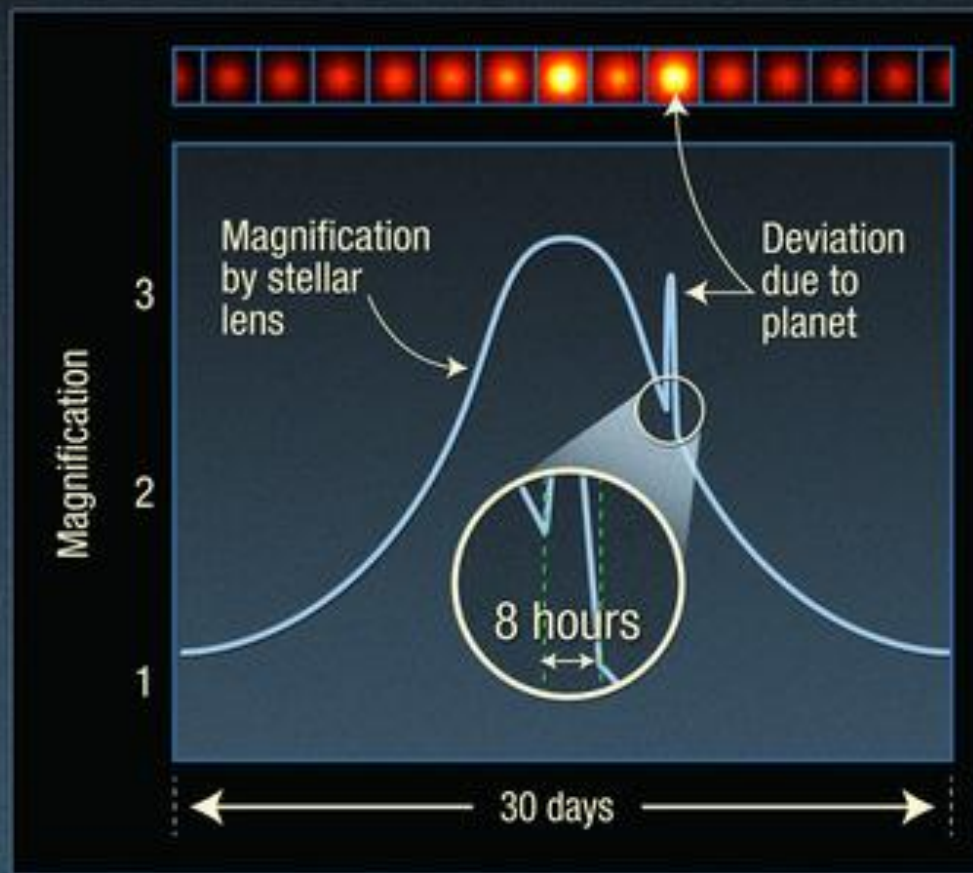


Extrasolar planet detected by gravitational microlensing

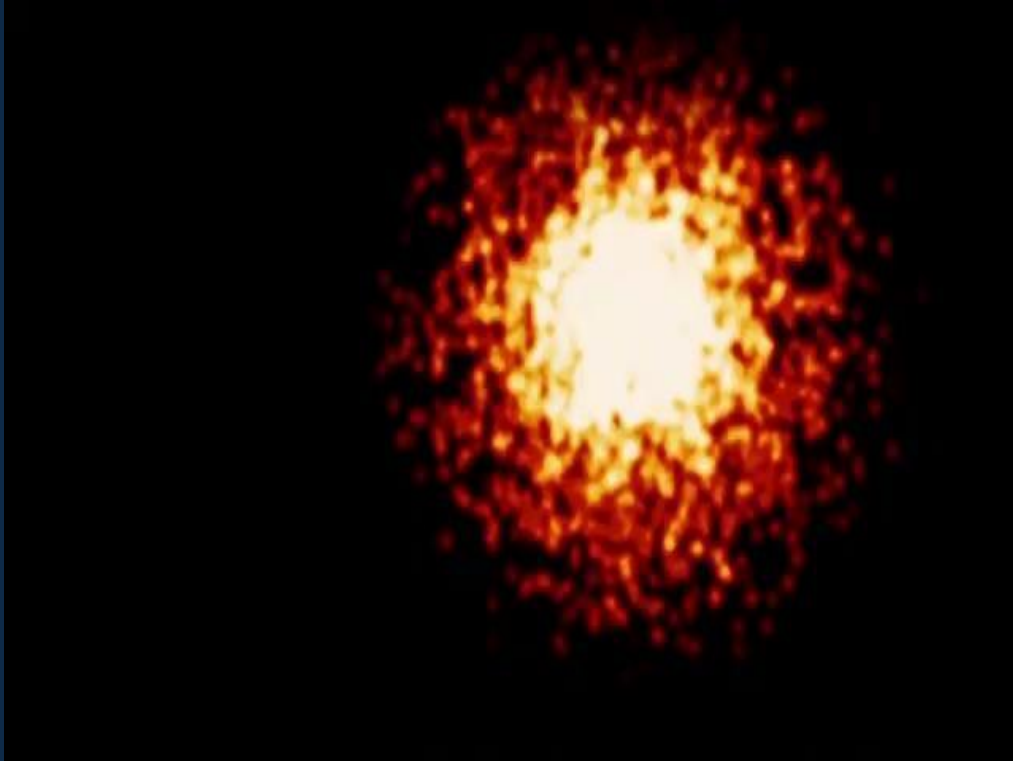


1 When a foreground star (red) passes in front of a background star, it brightens the light of the background star. The gravitational field of the foreground star warps space to create a gravitational lens that magnifies light.

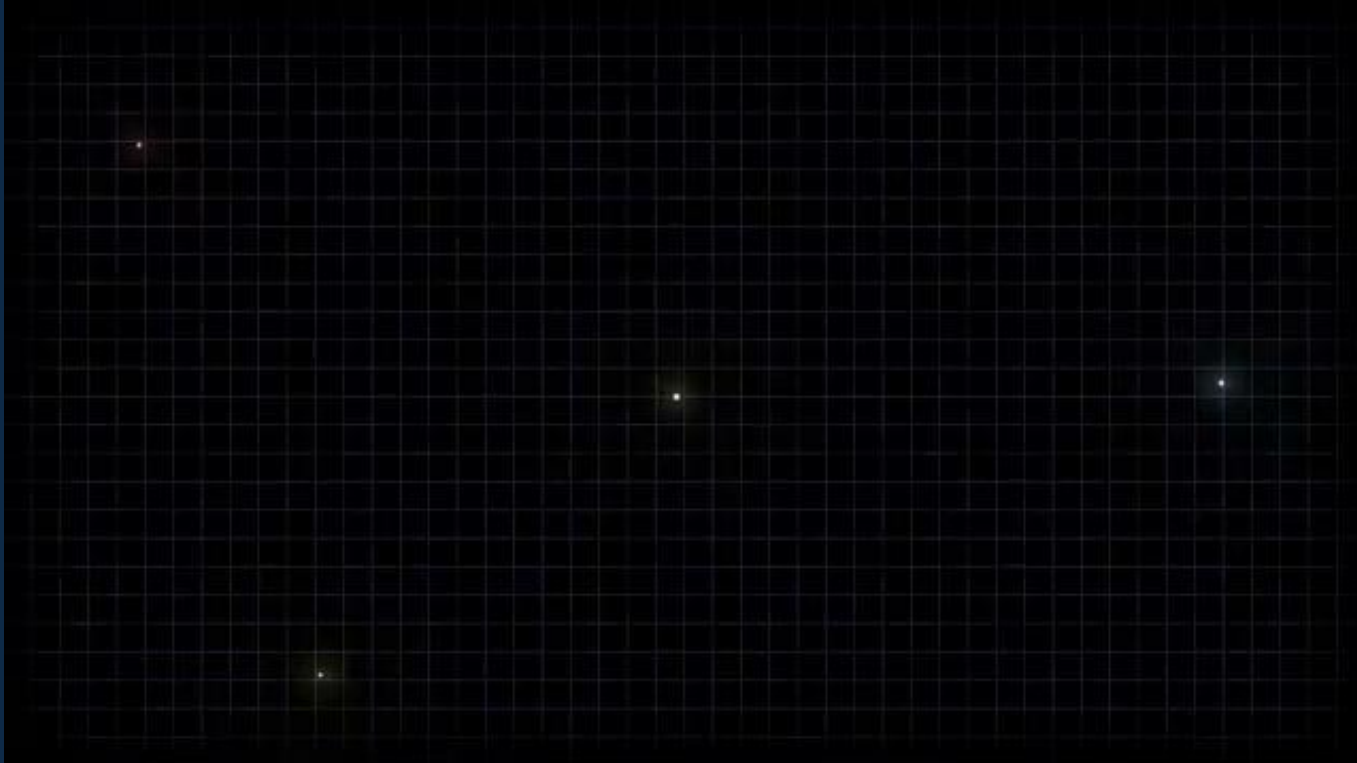
2 If a planet is orbiting the foreground star, it, too, will gravitationally lens the background star for a shorter duration.



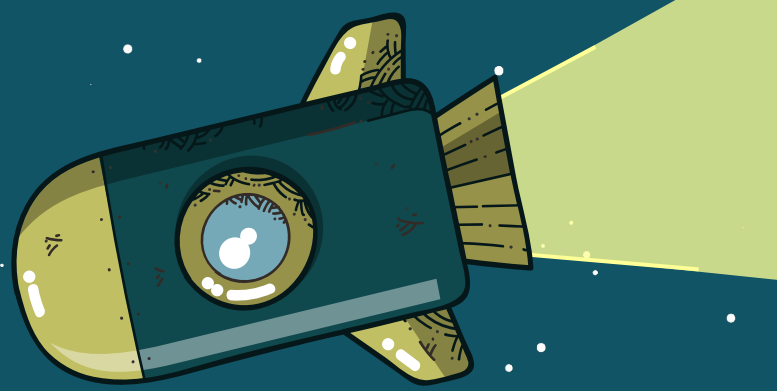
Direct Imaging



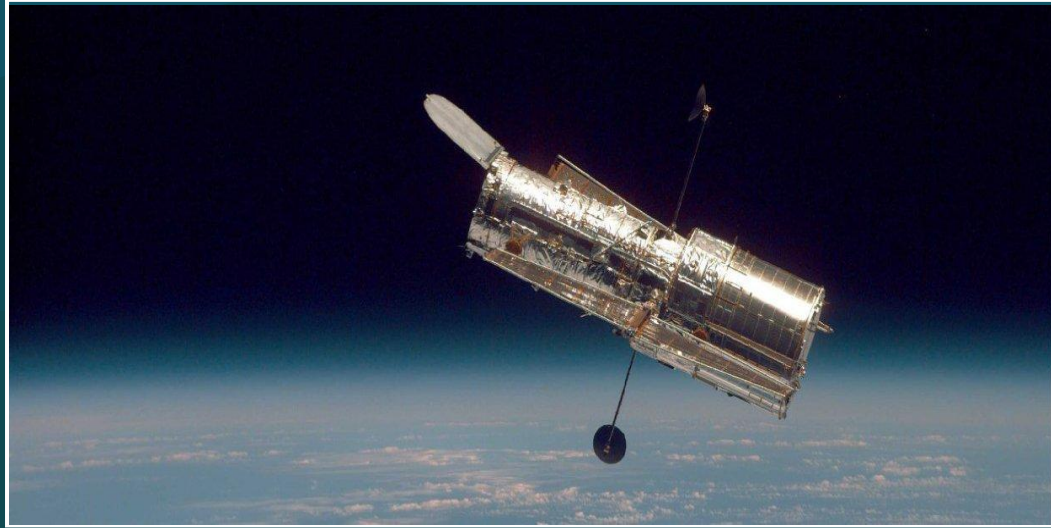
Astrometry



03. FAMOUS EXOPLANET MISSIONS

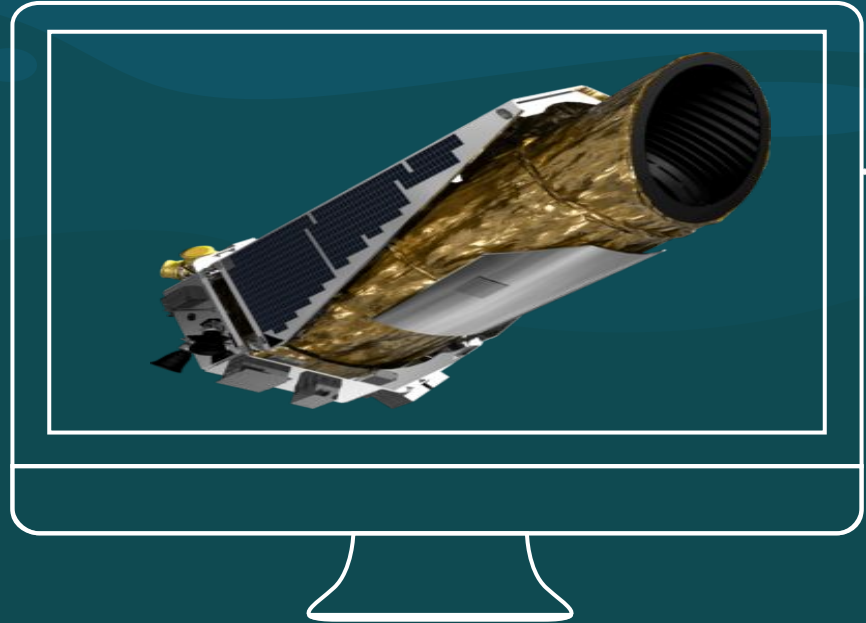


HUBBLE SPACE TELESCOPE



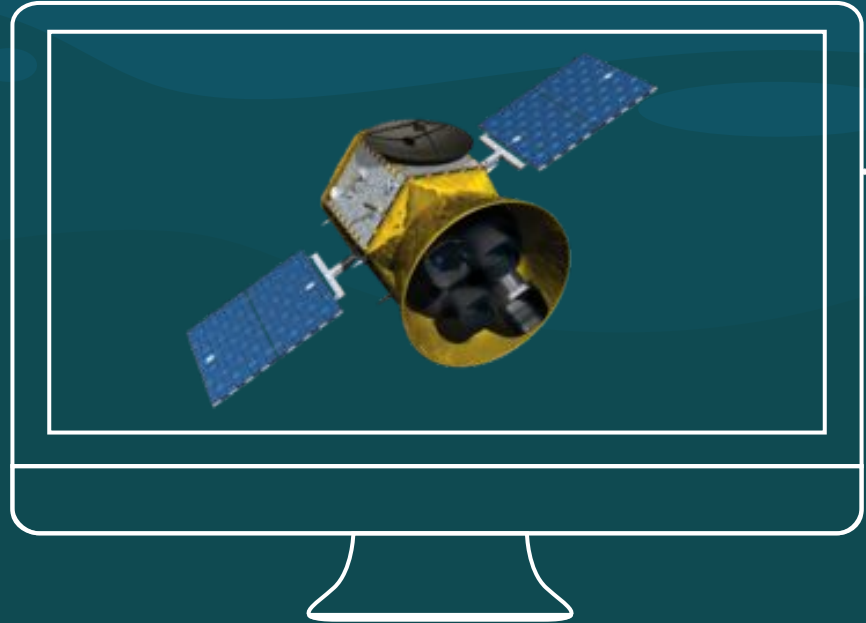
KEPLER MISSION

- Kepler Space Telescope was launched in March 7, 2009 by NASA.
- It surveyed Milky-Way to detect earth-size exoplanets near Goldilocks zone of the star.
- Kepler observed 530,506 stars and discovered 2662 planets
- It got retired on October 30 2018
- The extension of mission was announced on June 2016 beyond the expected exhaustion of fuel



TRANSITING EXOPLANET SURVEY SATELLITE

- Launched in April 18, 2019 by NASA's Falcon 9 rocket
- It covered an area 400 times more than that covered during Kepler mission.
- With TESS it is possible to study the mass, size, density and orbit of large cohort of small planets, including rocky planets in the habitable zone of host stars.
- It was planned for 2 years but is continuing for more than 3 years now.



JAMES WEBB SPACE TELESCOPE

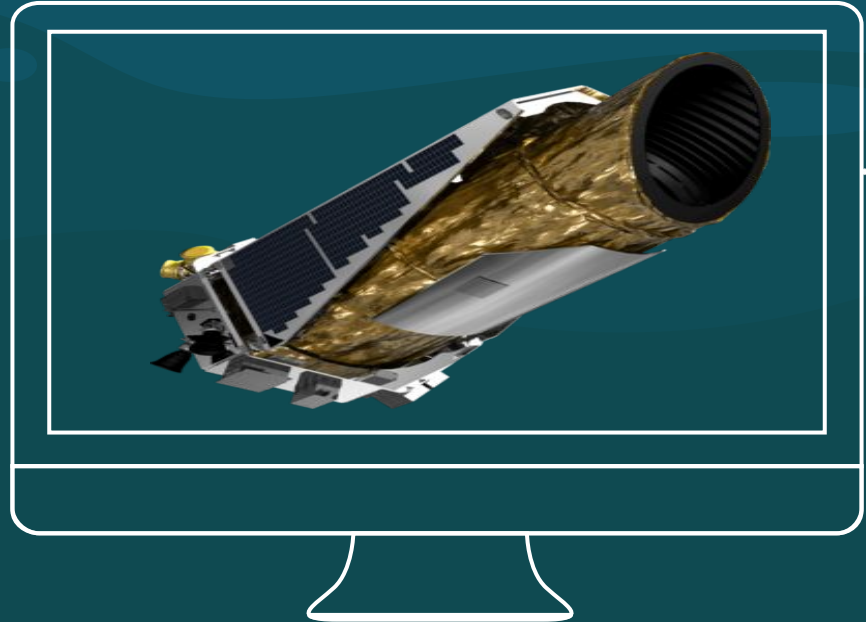
- It is expected to be launch in November 2021 by Ariane 5 rocket.
- Major mission:

To search for light from first stars and galaxies that formed the Universe after Big Bang

To study the formation and evolution of galaxy

To understand the formation of star and planetary system

To study planetary system and origin of life



SPECIAL THANKS

