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# How to Generate A Gravitational Waveform Template Bank?

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# Matched Filtering

## □ Matched filtering:

- A widely used method to detect Gravitational Waves(GWs)
- Match between waveform templates and the signal
- More correlation, more likely to have such waveform in signal
- Need to use different templates to find the closest one
  - Optimal SNR:  $\rho_{\text{opt}}^2 = (h | h)$
  - **Need a pre-computed template bank**

## • Inner product:

$$(h | g) \equiv 4\Re \int_0^{+\infty} \frac{\tilde{g}^*(f)\tilde{h}(f)}{S_n(f)} df$$

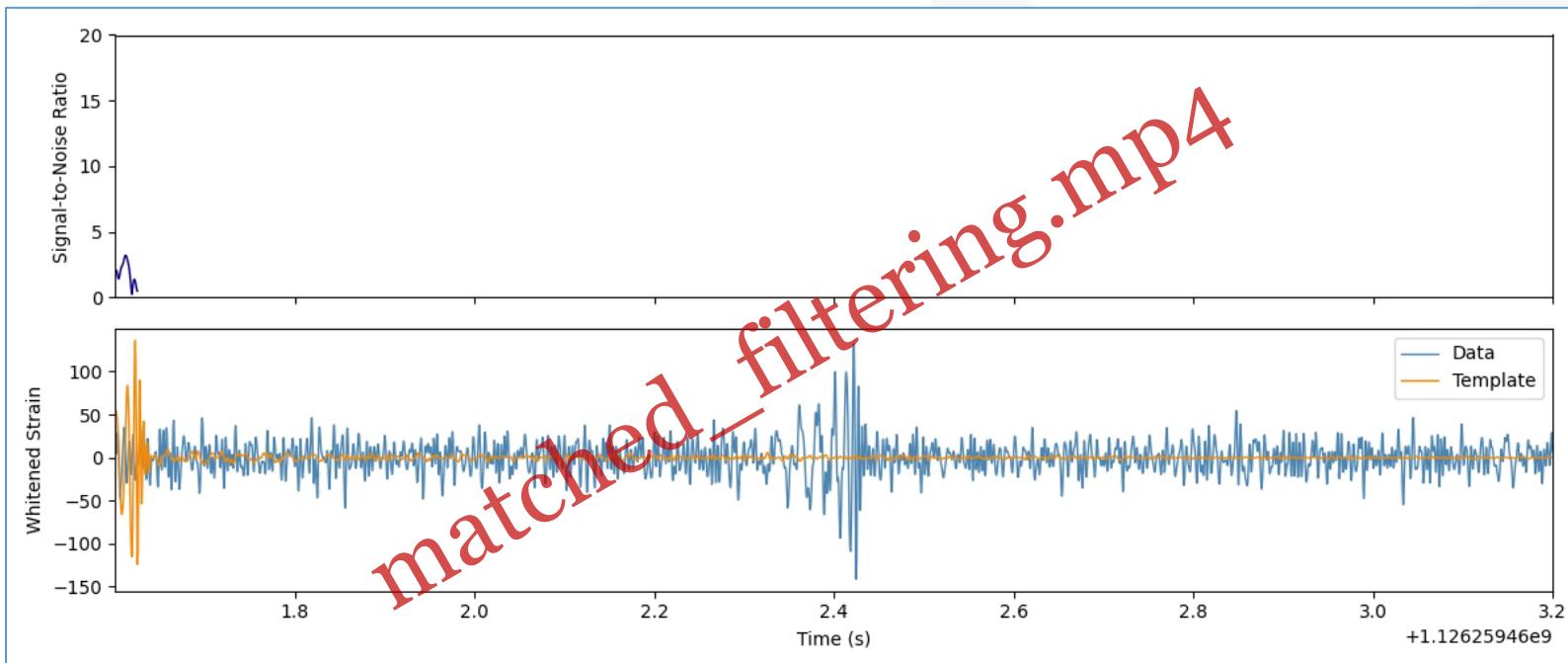


Illustration of matched filtering method (GW150914 as an example)



# Template Bank

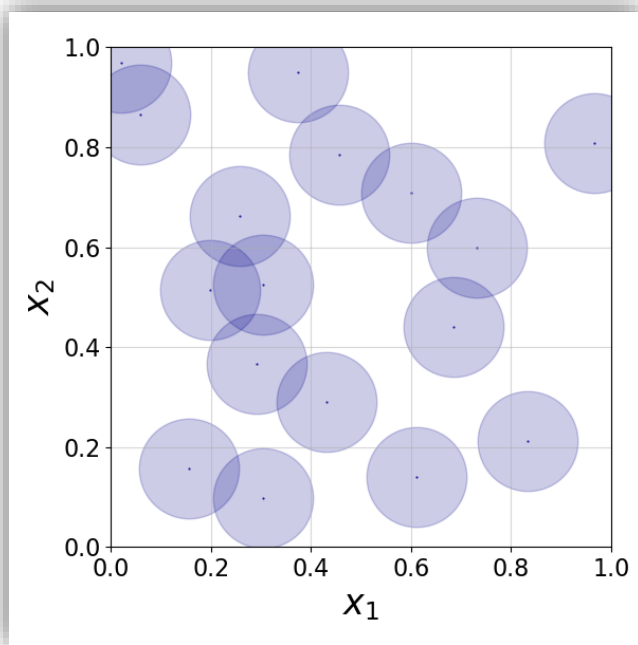
## □ Validity

- Any GW signal in its parameter space:  $FF \geq M$
- $M = 0.97$ : minimal match

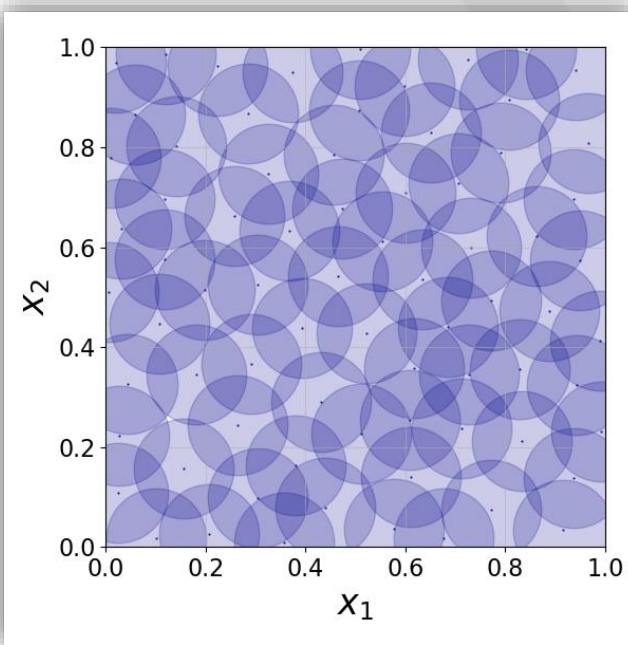
## □ Redundancy

- More computational cost

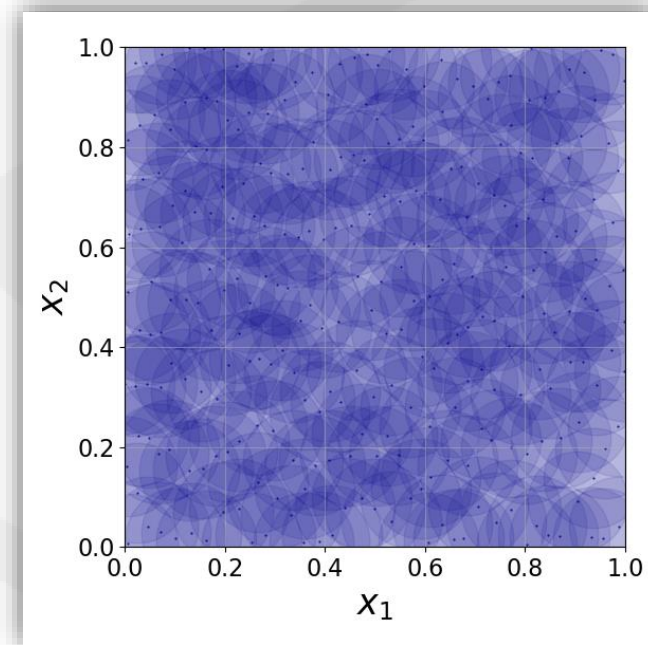
## □ Take Euclidean 2D space as an example:



Too few templates



“Just right”



Too many templates

## • Inner product:

$$(h | g) \equiv 4\Re \int_0^{+\infty} \frac{\tilde{g}^*(f)\tilde{h}(f)}{S_n(f)} df$$

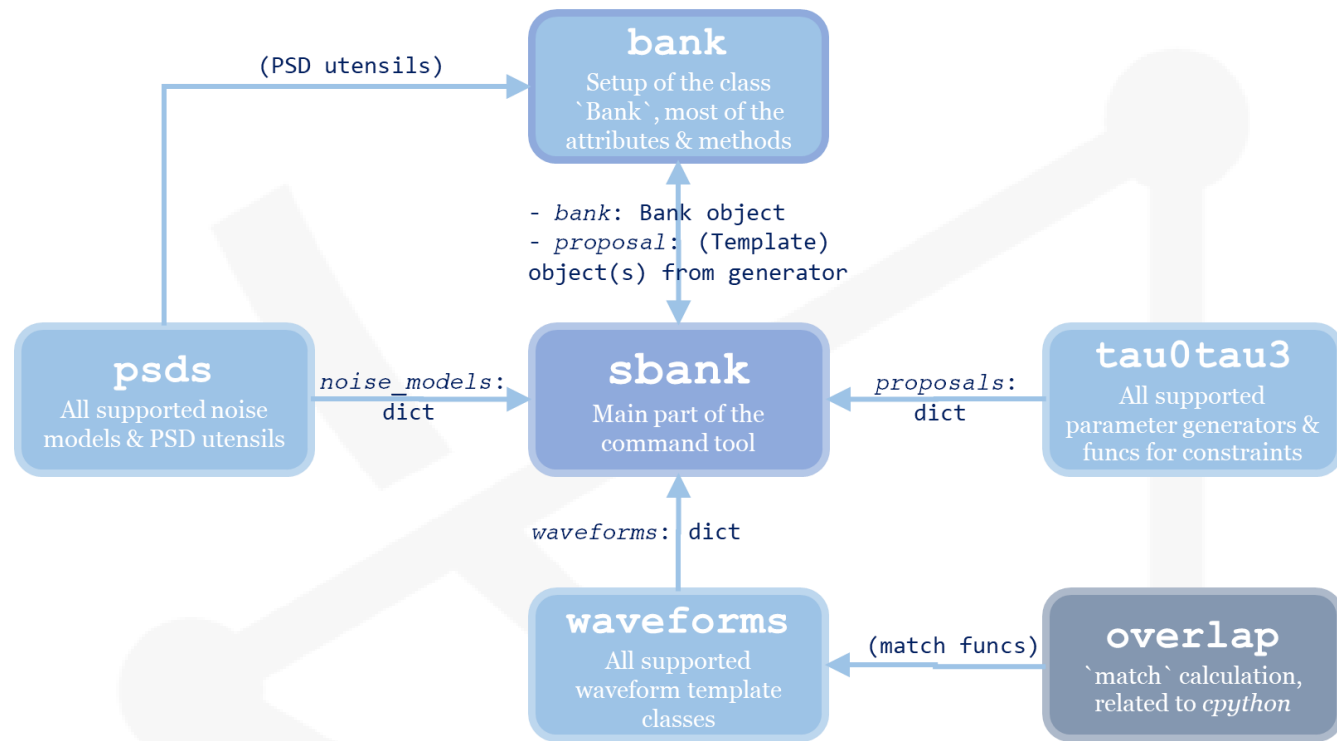
## • Fitting Factor (FF):

$$FF(\lambda^\mu) \equiv \max_{\lambda^\mu} \frac{(h(\lambda^\mu) | h(\lambda^{\mu'}))}{\sqrt{(h(\lambda^\mu) | h(\lambda^\mu))(h(\lambda^{\mu'}) | h(\lambda^{\mu'}))}}$$



# Template Bank Generation

- ❑ A similar question with covering N-dimensional curved space (*signal manifold*)
- ❑ Hard to know optimal placement algorithms (exact metrics), especially higher dimension and curved space
- ❑ sbank package
  - Stochastic template bank placement
- ❑ A(n over-)simplified version of sbank
  - Only kept the key algorithms
  - Only consider Euclidean(-like) 2D space to make it simple and be illustrated easily



Structure of sbank package



# Key Algorithm: Covering

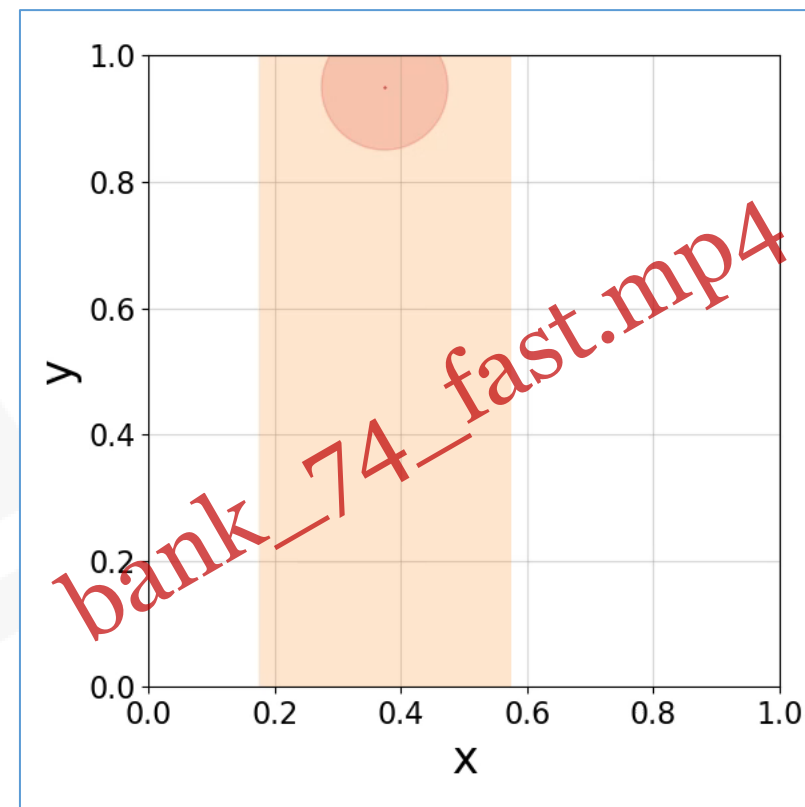
□ How should we regard one template as near the other one?

- Euclidean(-like) 2D space:
  - Calculate the proper distance between two points
  - Covering regions are drawn as circles(ellipses)
  - (Here max\_distance=0.1)

➤ Real GW template bank:

➤ Overlap:

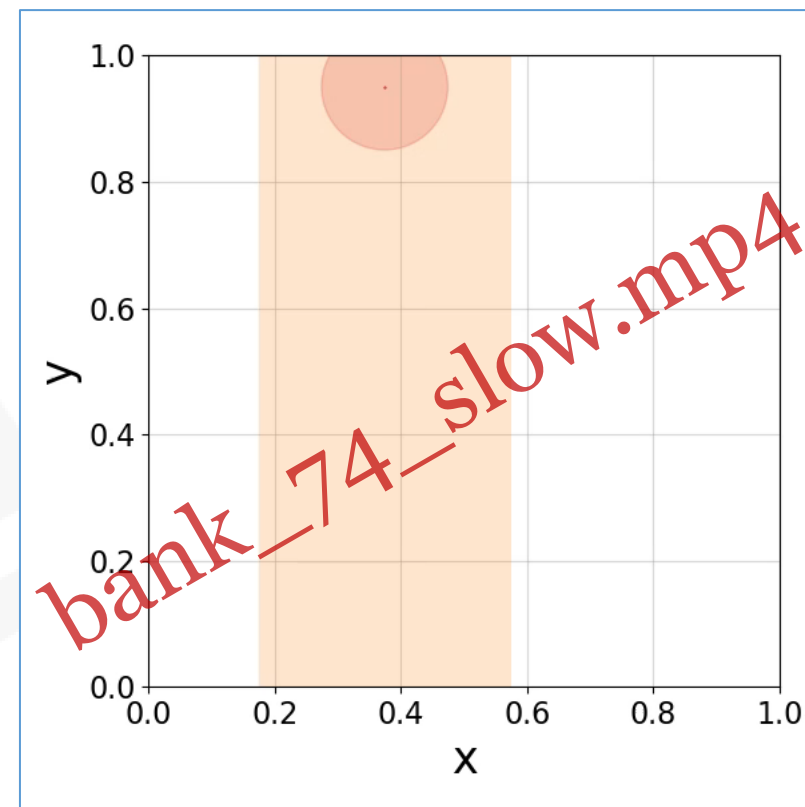
$$O(h(\lambda^\mu), h(\lambda^{\mu'})) \equiv \frac{(h(\lambda^\mu) | h(\lambda^{\mu'}))}{\sqrt{(h(\lambda^\mu) | h(\lambda^\mu))(h(\lambda^{\mu'}) | h(\lambda^{\mu'}))}}$$





# Key Algorithm: Covering

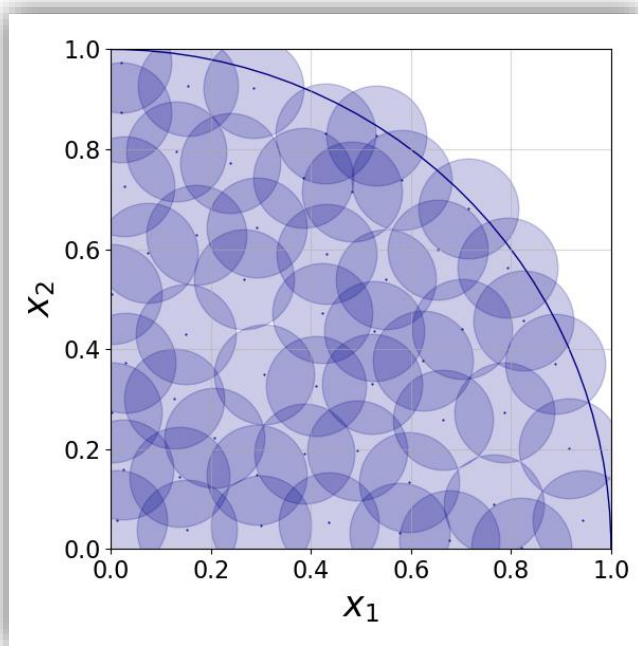
- ❑ How can we filter proposals when generating a bank?
  - We are not expected to calculate overlaps between a proposal and ALL templates in a bank
  - Neighborhood parameter and neighborhood size
    - Templates are sorted by this parameter
- Euclidean(-like) 2D space:
  - (e. g.) Neighborhood parameter: the value of  $x$ 
    - (Here `nhood_size=0.2`)
- Real GW template bank:
  - Neighborhood parameter:
    - $\tau_0$ : A function of chirp mass and lower frequency
    - Since Chirp mass is a dominant parameter in GW



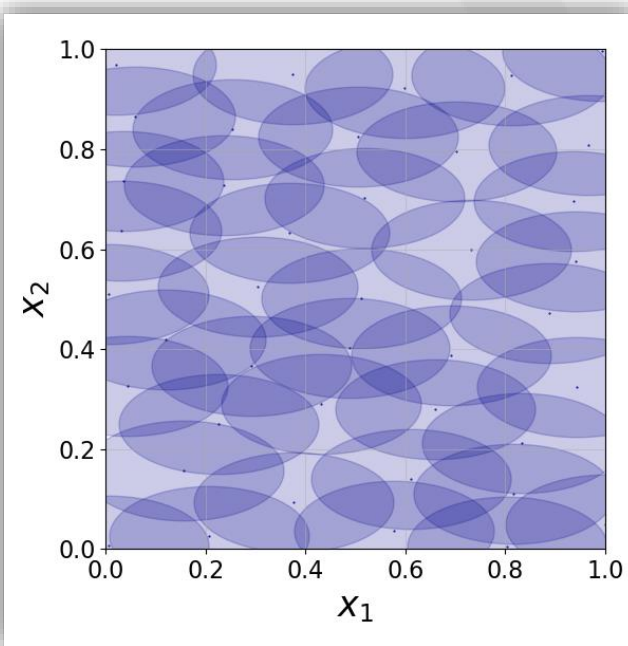


## Further Discussions

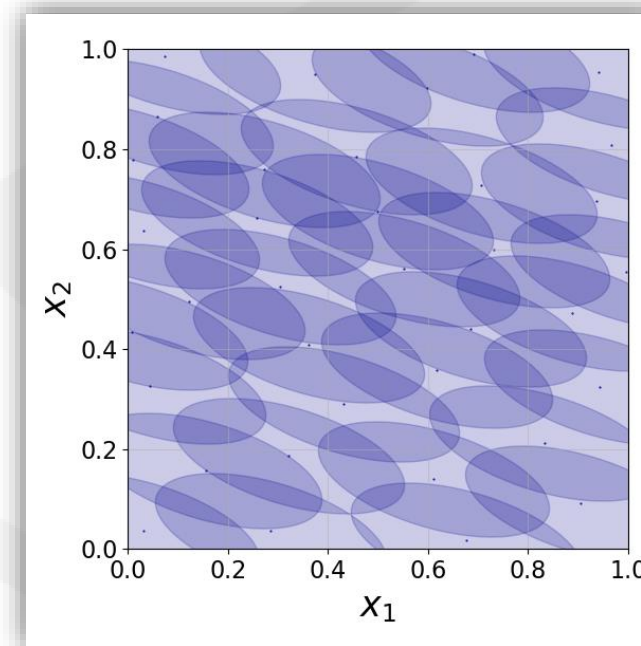
- ❑ Does the distribution of trial points(proposals) effect template distribution?
  - Almost none. Better distribution can help converge faster
- Real GW template bank:
  - Need to consider the physical distribution when stochastically propose a parameter



Cartesian metric but a polar generator

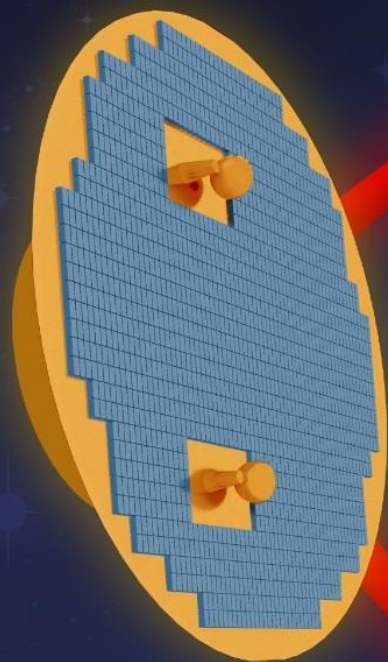


Metric:  $[[1/4, 0], [0, 1]]$



Metric:  $[[1/4, 1/4], [1/4, 1]]$





地球  
Earth

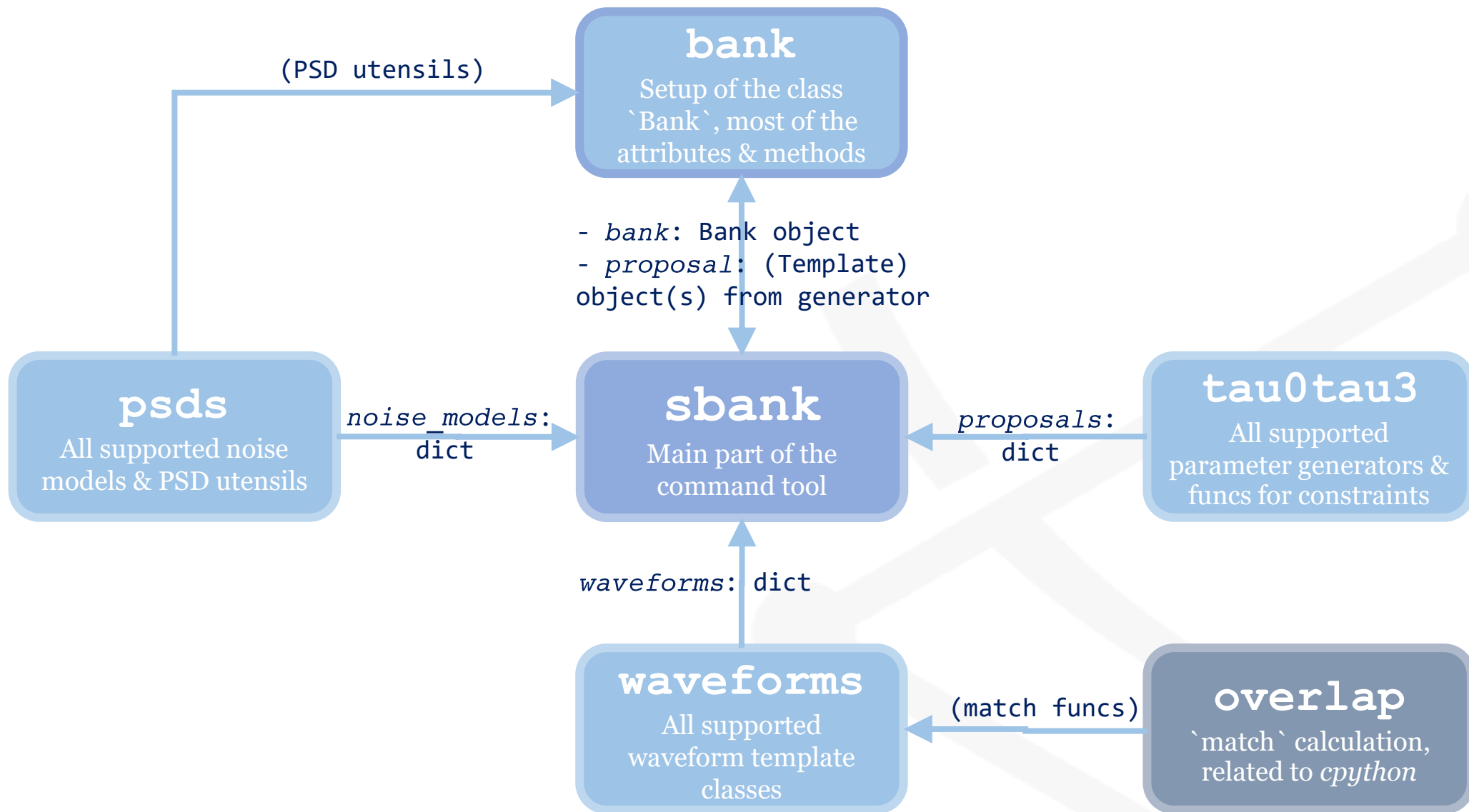


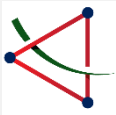
Thanks





# `sbank` structure





# `sbank\_simplified` structure

