

Quantum Algorithms for Quantum Field Theories

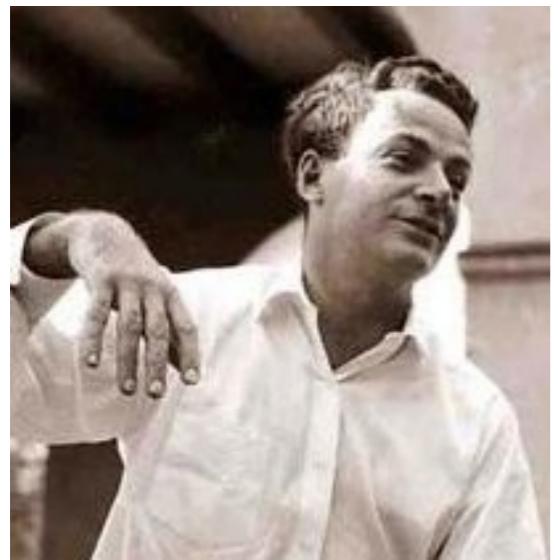
Stephen Jordan

Joint work with
Keith Lee
John Preskill

Science, 336:1130 (2012)

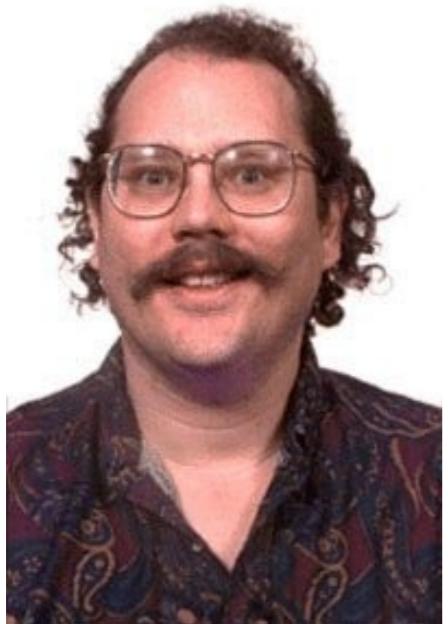


Jan 24, 2012



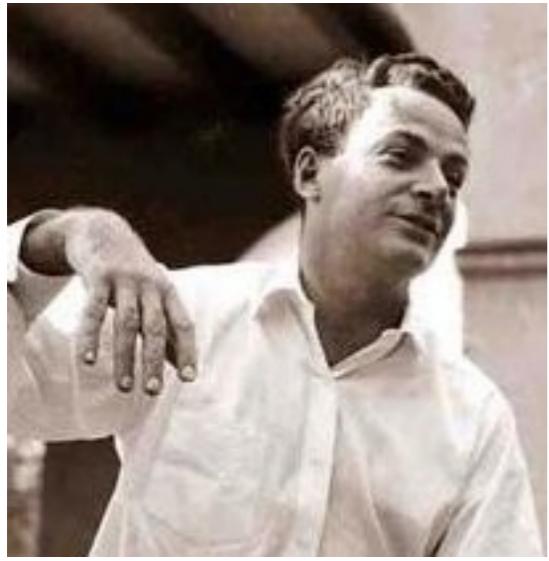
The full description of quantum mechanics for a large system with R particles has too many variables. It cannot be simulated with a normal computer with a number of elements proportional to R .

-Richard Feynman, 1982



An n -bit integer can be factored on a quantum computer in $O(n^2)$ time.

-Peter Shor, 1994



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Are there any systems that remain hard to simulate even with quantum computers?

Quantum Simulation

Condensed-matter lattice models:

[Lloyd, 1996]

[Abrams, Lloyd, 1997]

[Berry, Childs, 2012]

Many-particle Schrödinger and Dirac Equations:

[Meyer, 1996]

[Zalka, 1998]

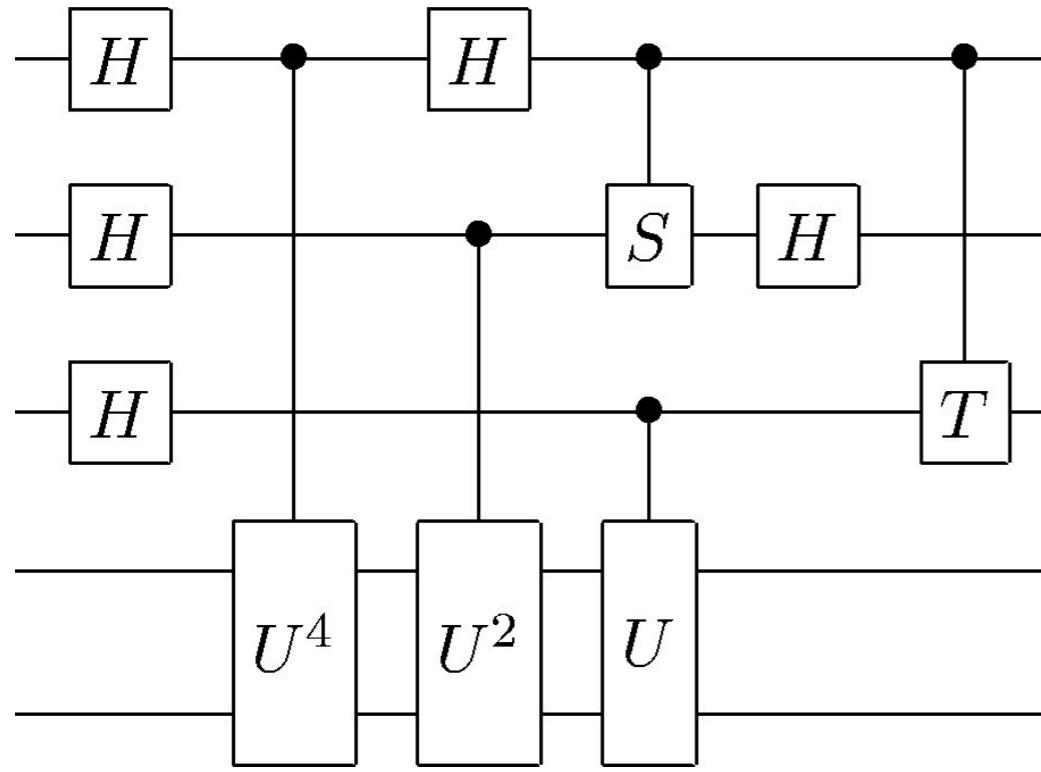
[Taylor, Boghosian, 1998]

[Kassal, S.J., Love, Mohseni, Aspuru-Guzik, 2008]

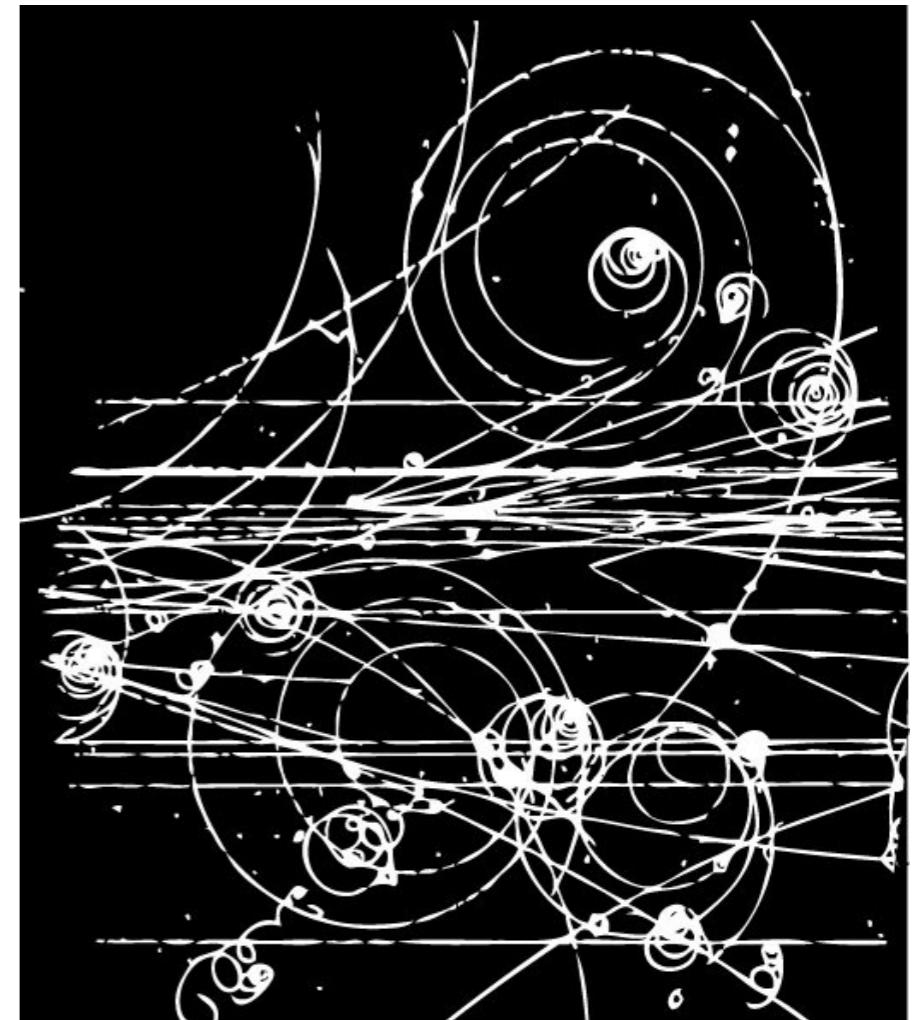
Quantum Field Theory

- Much is known about using quantum computers to simulate quantum systems.
- Why might QFT be different?
 - Field has infinitely many degrees of freedom
 - Relativistic
 - Particle number not conserved
 - Formalism looks different

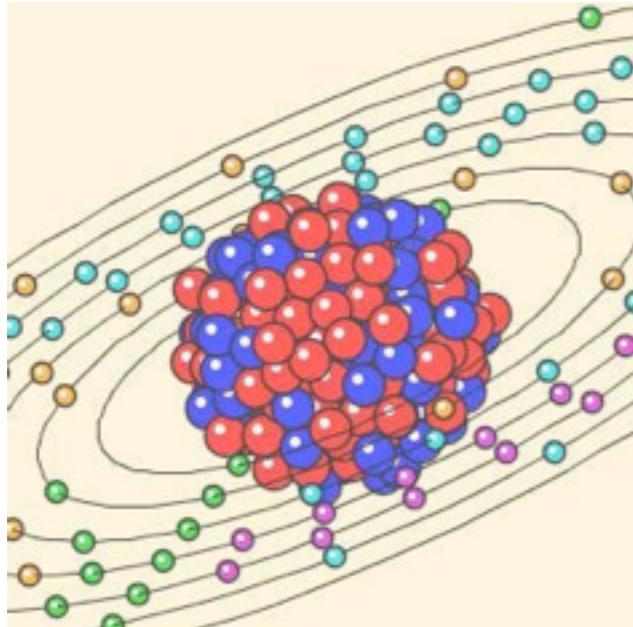
What is the computational power of our universe?



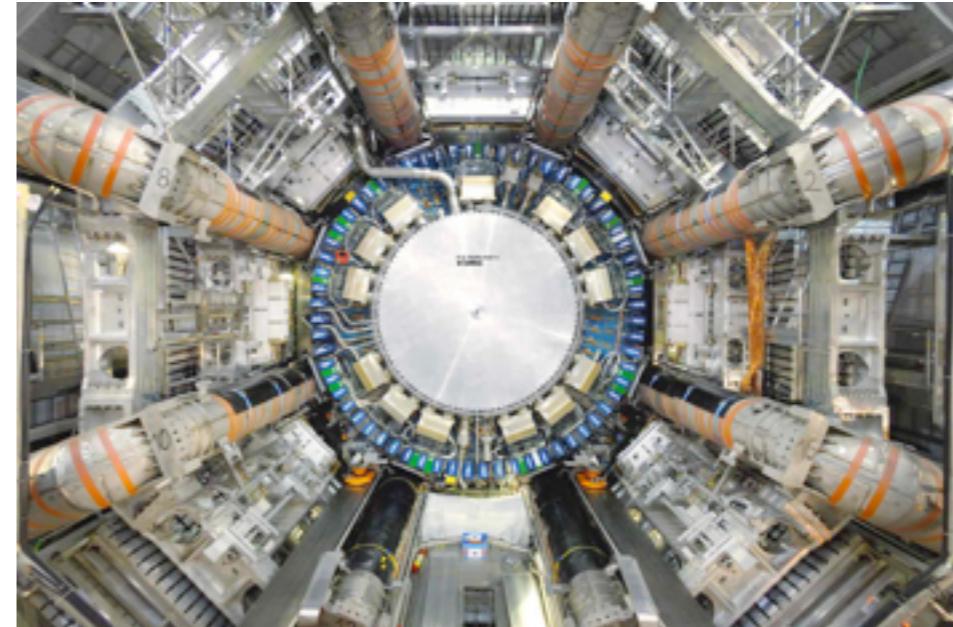
simulate



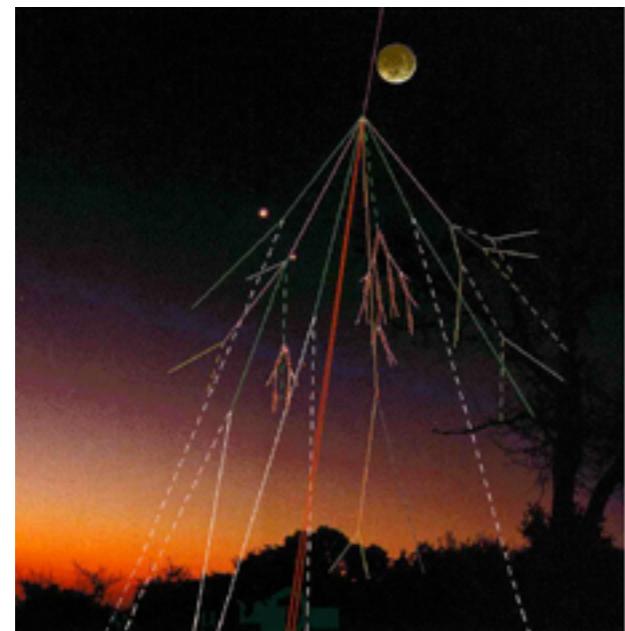
When do we need QFT?



Nuclear Physics



Accelerator Experiments

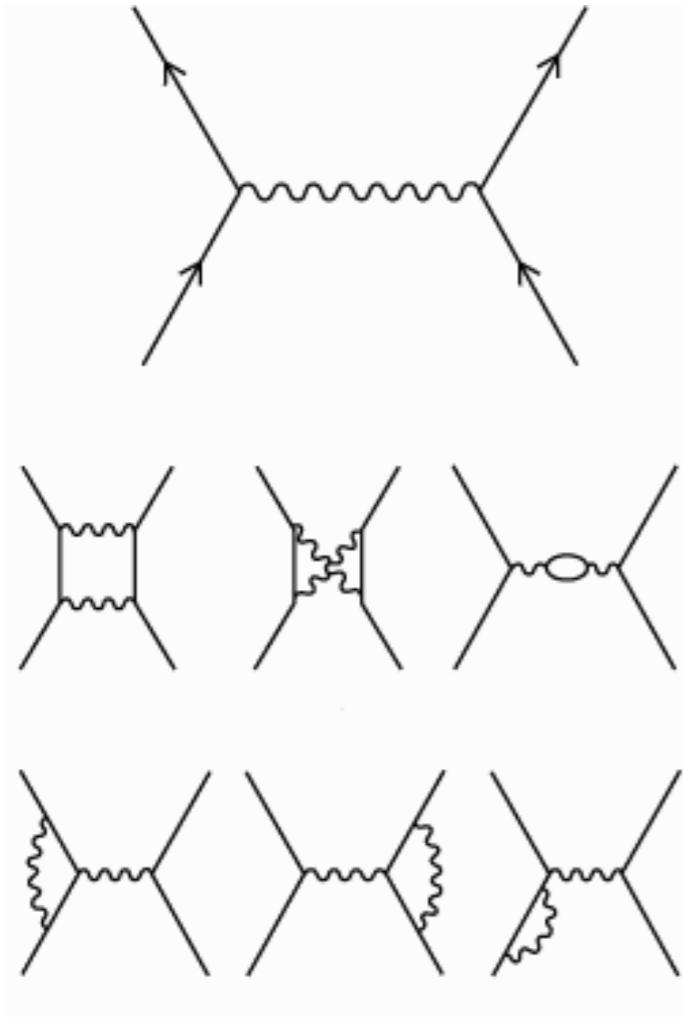


Cosmic Rays

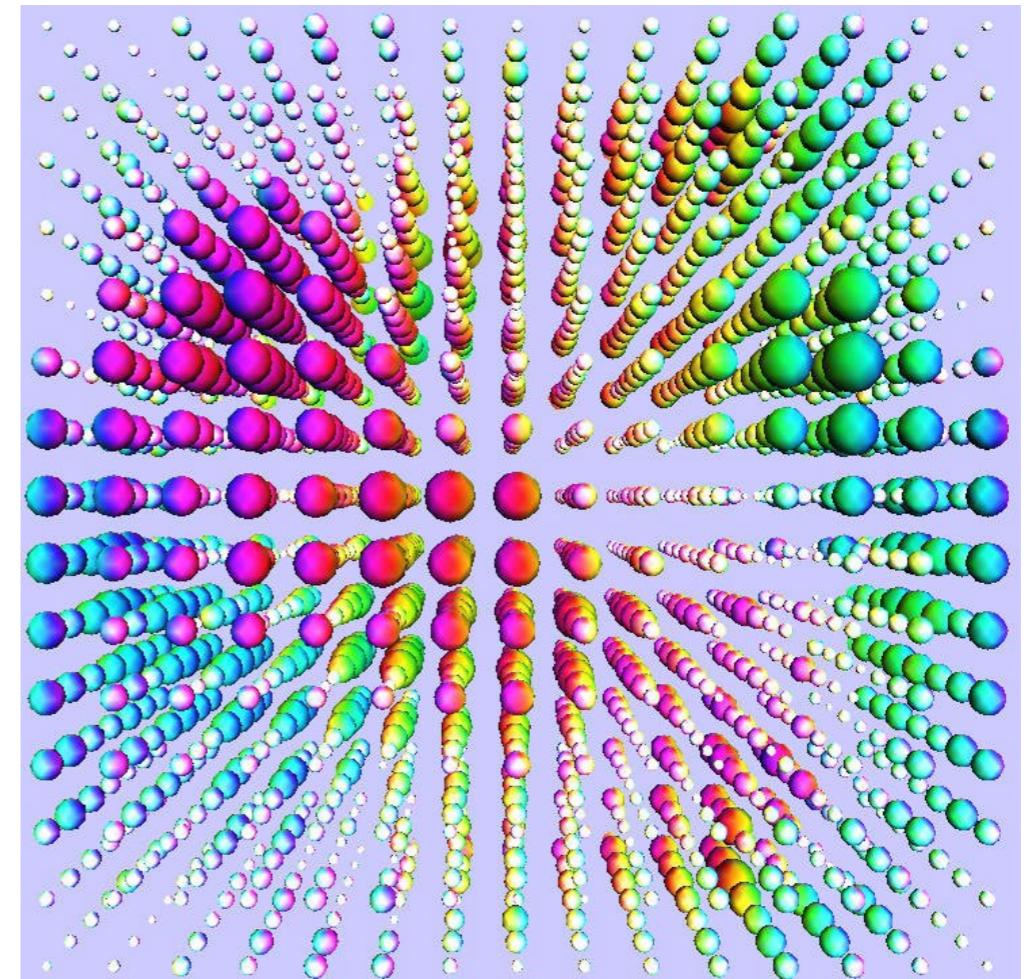
→ Whenever quantum mechanical and relativistic effects are both significant.

Classical Algorithms

Feynman diagrams



Lattice methods



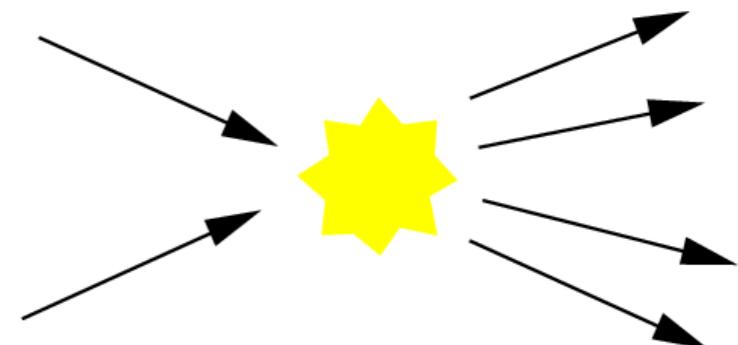
Break down at strong
coupling or high precision

Cannot calculate
scattering amplitudes

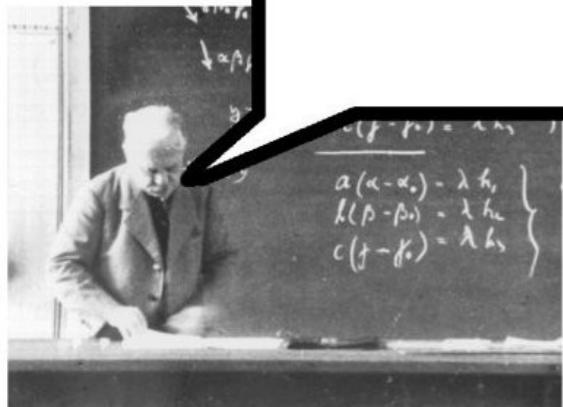
A QFT Computational Problem

Input: a list of momenta of incoming particles

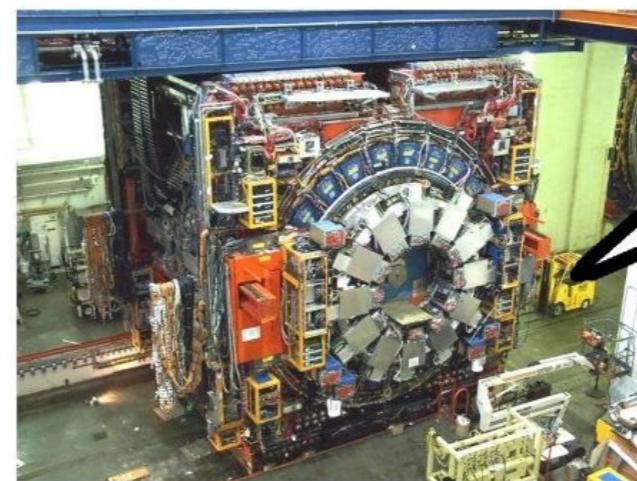
Output: a list of momenta of outgoing particles



S-matrix



Particle accelerator



I will present a polynomial-time quantum algorithms to compute scattering probabilities in the ϕ^4 and Gross-Neveu models with nonzero mass

These are simple models that illustrates some of the main difficulties in simulating a QFT:

- Discretizing spacetime
- Preparing initial states

ϕ^4 -theory

Lagrangian density

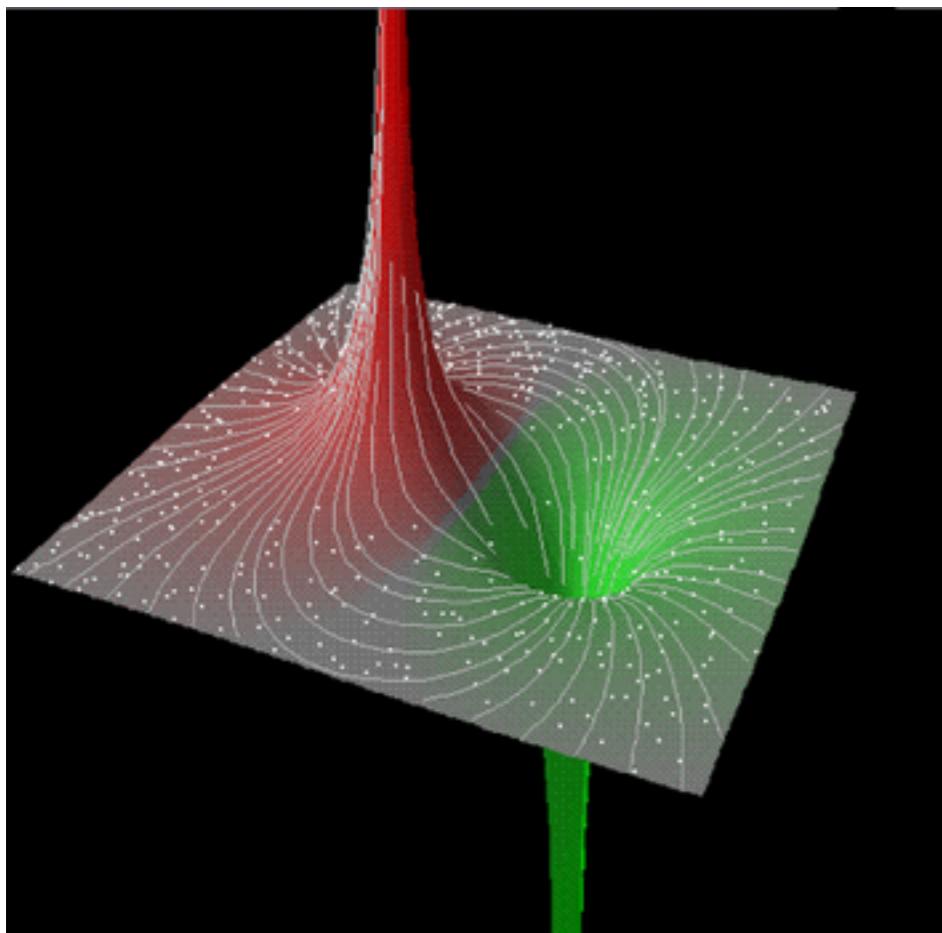
$$\mathcal{L} = \frac{1}{2} \partial^\mu \phi \partial_\mu \phi - m^2 \phi^2 - \frac{\lambda}{4!} \phi^4$$

For quantum simulation we prefer
Hamiltonian formulation (equivalent)

$$H = \int d^d x \left[\pi^2 + (\nabla \phi)^2 + m^2 \phi^2 + \lambda \phi^4 \right]$$

$$[\phi(x), \pi(y)] = i\delta^{(d)}(x - y)$$

Quantum Fields



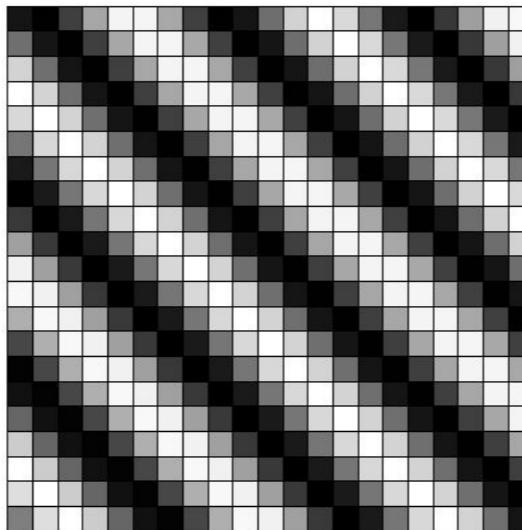
A classical field is described by its value at every point in space.

$$E(r) = \frac{1}{4\pi\epsilon_0} \frac{q}{r^2}$$

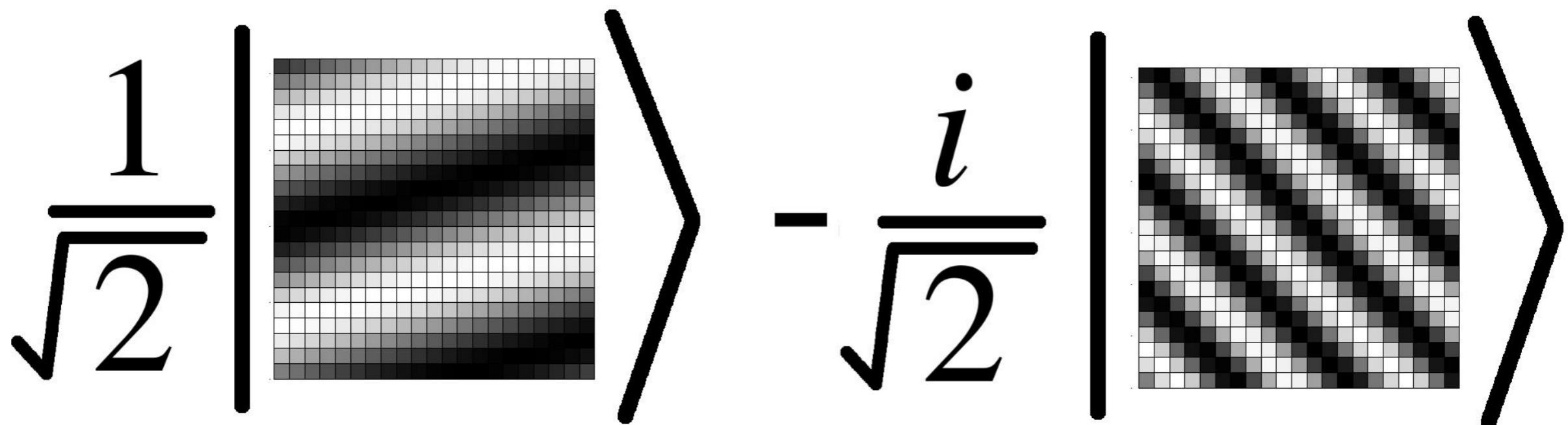
A quantum field is a superposition of classical field configurations.

$$|\Psi\rangle = \int \mathcal{D}[E] \Psi[E] |E\rangle$$

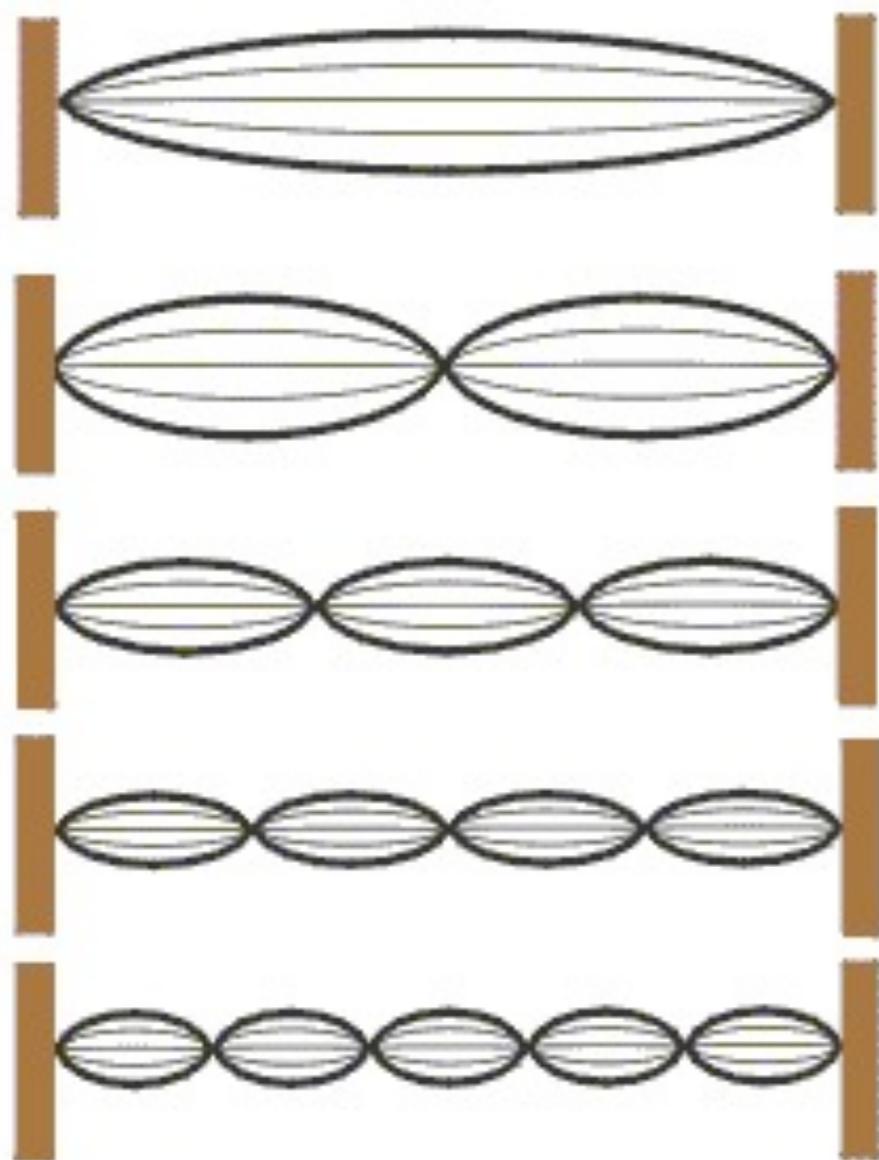
A configuration of the field is a list of field values,
one for each lattice site.



A quantum field can be in a superposition of
different field configurations.



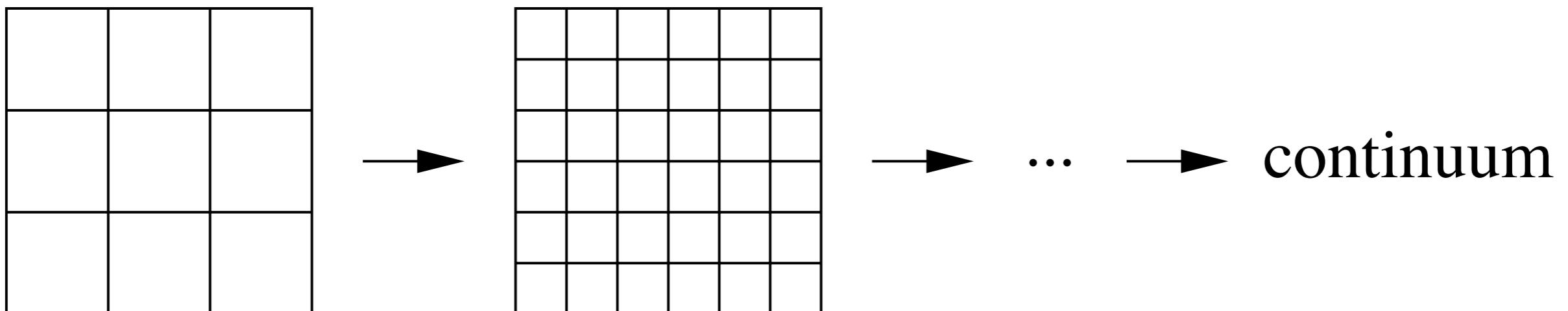
Particles Emerge from Fields

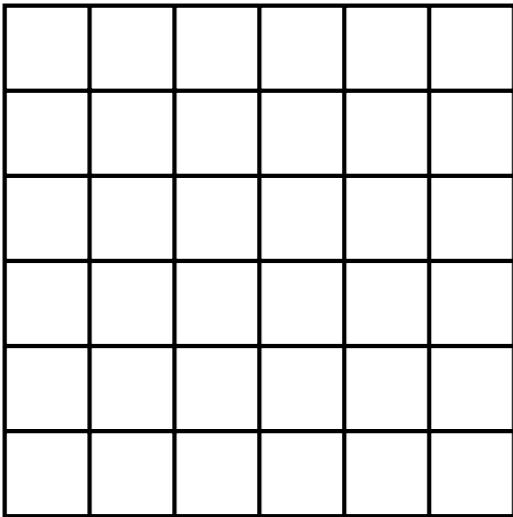


Particles of different energy
are different resonant
excitations of the field.

Lattice cutoff

Continuum QFT = limit of a sequence of theories on successively finer lattices



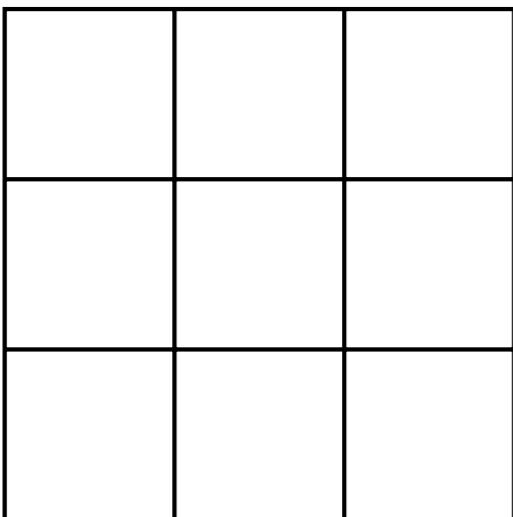


Mass: m

Interaction strength: λ



Coarse grain

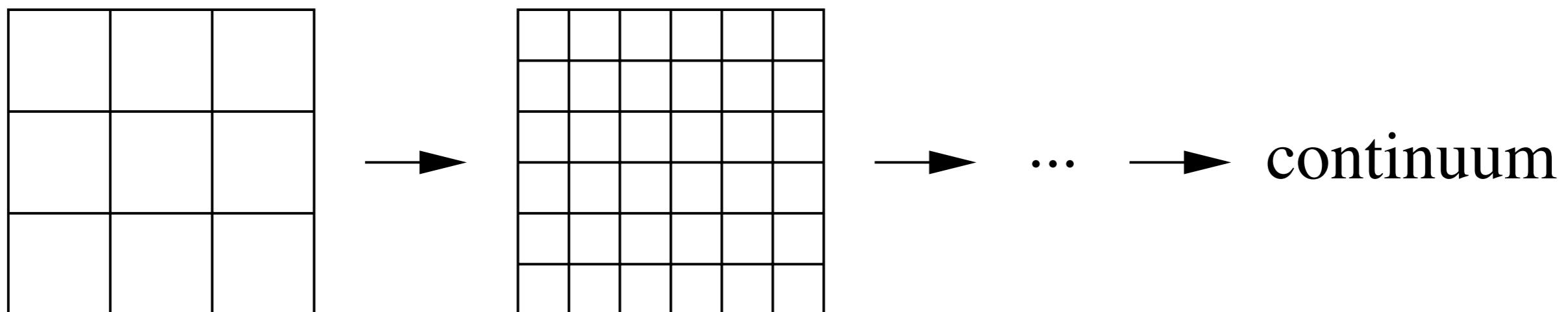


Mass: m'

Interaction strength: λ'

Lattice cutoff

Continuum QFT = limit of a sequence of theories on successively finer lattices



m and λ are functions of lattice spacing!

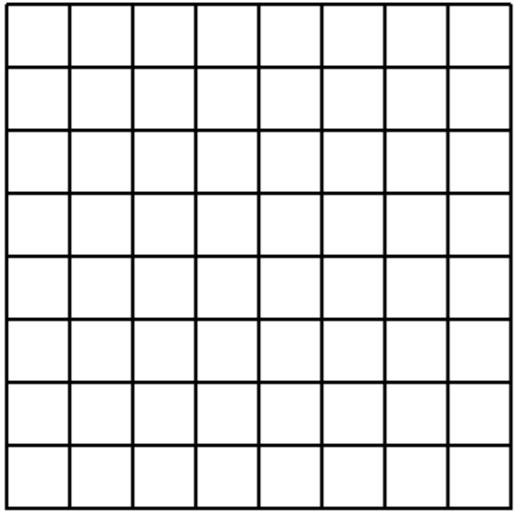
Discretization Errors

- Renormalization of m and λ make discretization tricky to analyze
- In ϕ^4 -theory, in $d=1,2,3$, discretization errors scale as a^2

$$\text{---} \circ \text{---} = \frac{(-i\lambda_0)^2}{6} \iint \frac{d^D k}{(2\pi)^D} \frac{d^D q}{(2\pi)^D} \frac{i}{(k^0)^2 - \sum_i \frac{4}{a^2} \sin^2 \left(\frac{ak^i}{2} \right) - m^2} \frac{i}{(q^0)^2 - \sum_i \frac{4}{a^2} \sin^2 \left(\frac{aq^i}{2} \right) - m^2} \\
 \times \frac{i}{(p^0 + k^0 + q^0)^2 - \sum_i \frac{4}{a^2} \sin^2 \left(\frac{a(p^i + k^i + q^i)}{2} \right) - m^2} \quad (207)$$

$$= \frac{i\lambda_0^2}{3} \int_0^1 \int_0^1 \int_0^1 dx dy dz \delta(x + y + z - 1) \iint \frac{d^D k}{(2\pi)^D} \frac{d^D q}{(2\pi)^D} \frac{1}{D^3}, \quad (208)$$

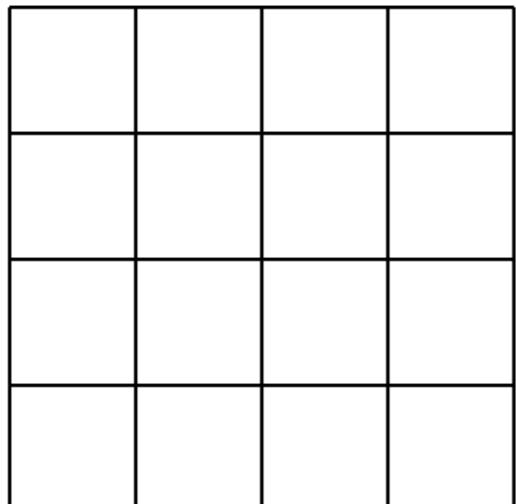
...it's complicated



$$H = \frac{1}{2} \sum_{x \in \Omega} a^d [\pi^2 + (\nabla \phi)^2 + m^2 \phi^2 + \lambda \phi^4]$$



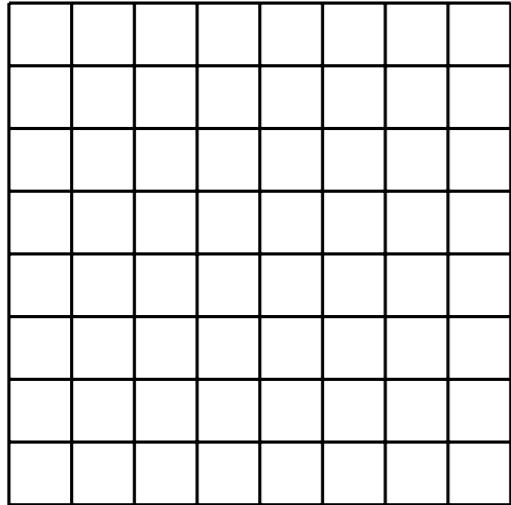
Coarse grain



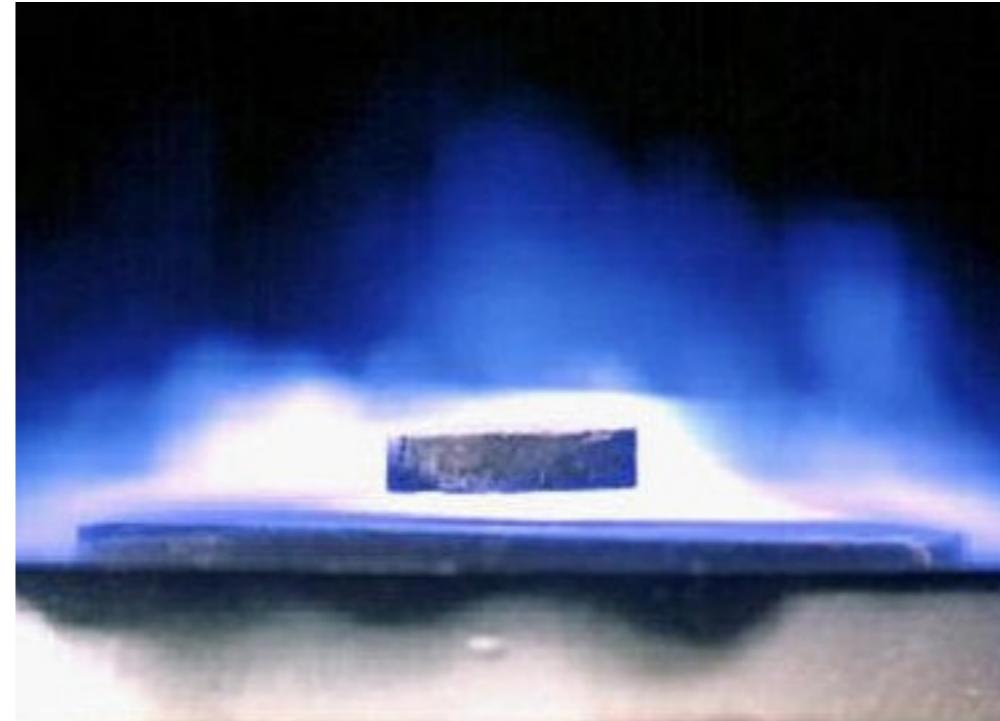
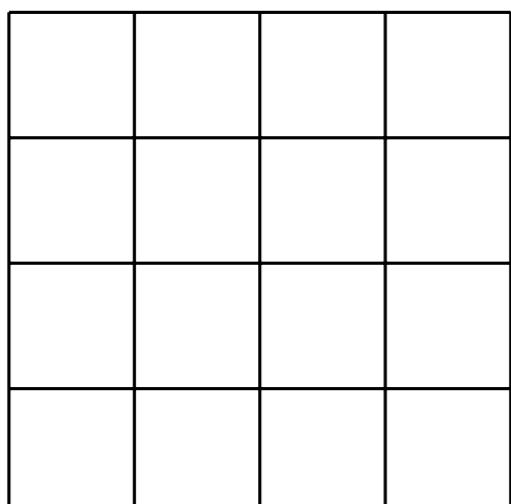
$$H_{\text{eff}} = \frac{1}{2} \sum_{x \in \Omega'} (2a)^d [\pi^2 + (\nabla' \phi)^2 + m_{\text{eff}}^2 \phi^2 + \lambda_{\text{eff}} \phi^4 + g \phi^6 + \dots]$$

Simulation converges as a^2

Condensed Matter



↓
RG



There is a fundamental lattice spacing.

But:

We may save qubits by simulating a coarse-grained theory.

After imposing a spatial lattice we have a many-body quantum system with a local Hamiltonian

Simulating the time evolution in polynomial time is a **solved problem**

Standard methods scale as N^2 . We can do N .

- Discretizing spacetime
- Preparing initial states

With particle interactions turned off, the model is exactly solvable.

1. Prepare non-interacting vacuum (Gaussian)
2. Prepare wavepackets of the non-interacting theory
3. Adiabatically turn on interactions
4. Scatter
5. Adiabatically return to non-interacting theory to make measurements

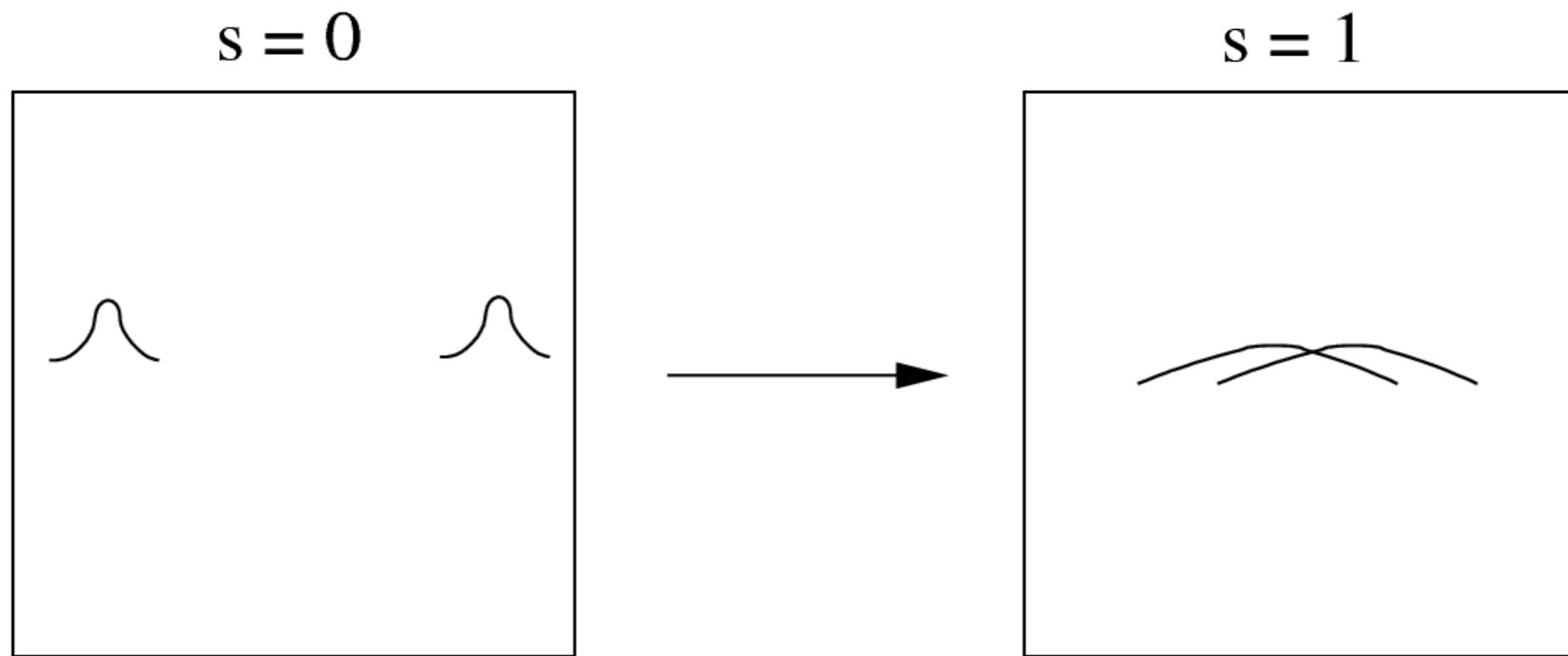
Adiabatically Turn on Interaction

$$H(s) = \sum_{x \in \Omega} \left[\pi^2 + (\nabla \phi)^2 + m^2 \phi^2 + s \lambda \phi^4 \right]$$

Use standard techniques to simulate $H(s)$ with s slowly varying from 0 to 1

This **almost** works....

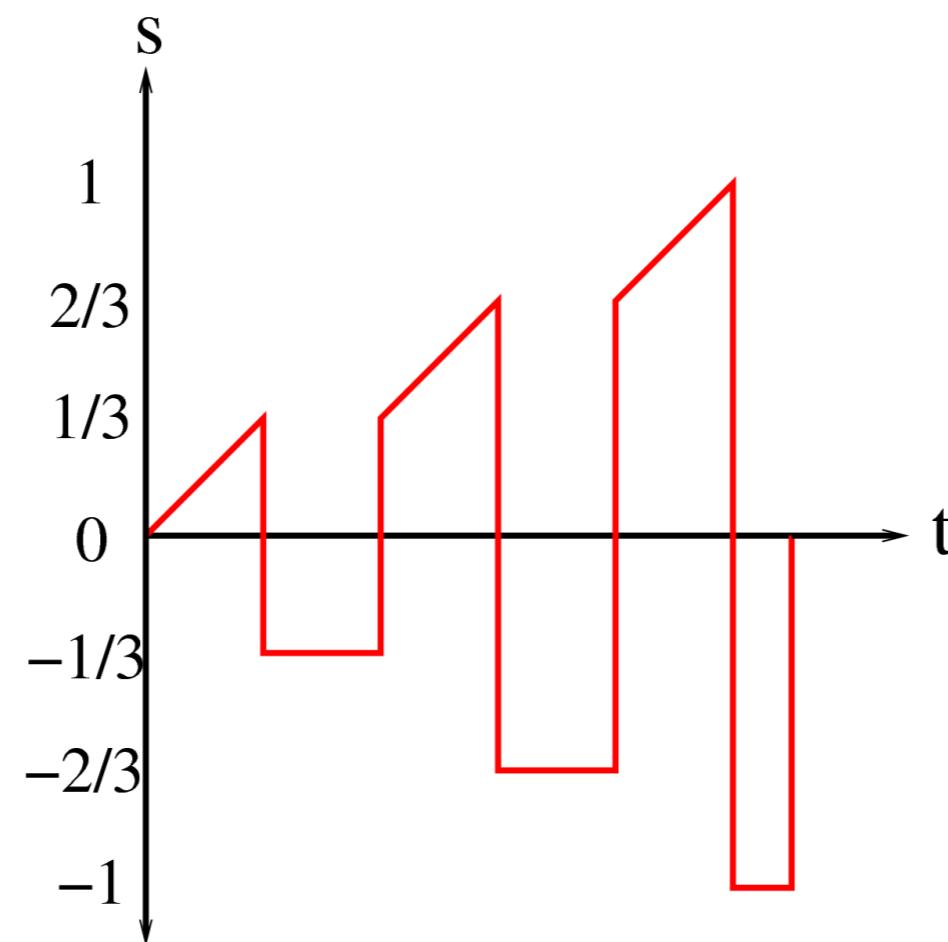
During the slow time evolution



The wavepackets propagate and broaden.

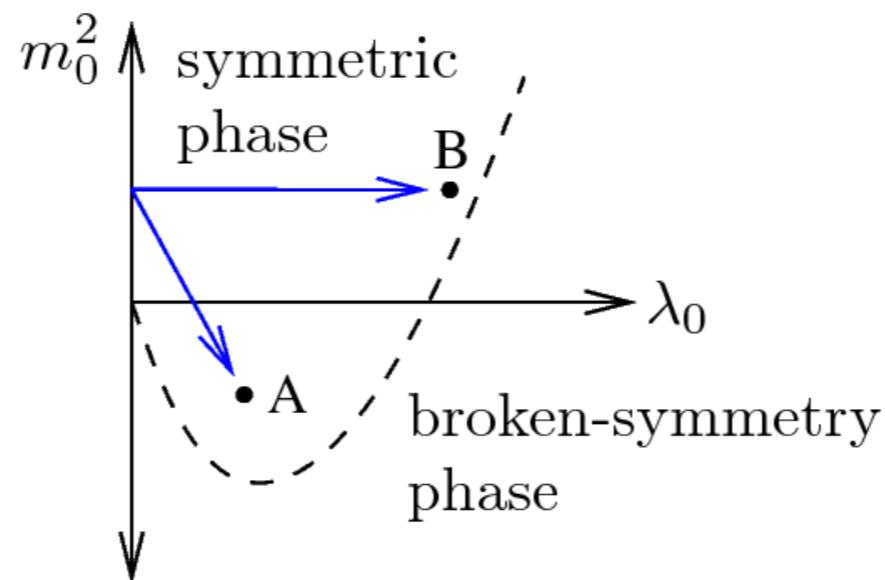
Solution: intersperse backward time evolutions with time-independent Hamiltonians

This winds back the dynamical phase on each eigenstate, without undoing the adiabatic change of eigenbasis



Strong Coupling

ϕ^4 -theory in 1+1 and 2+1 dimensions has a quantum phase transition in which the $\phi \rightarrow -\phi$ symmetry is spontaneously broken



Near the phase transition perturbation theory fails and the gap vanishes.

$$m_{\text{phys}} \sim (\lambda_c - \lambda_0)^\nu$$

$$\nu = \begin{cases} 1 & d = 1 \\ 0.63\dots & d = 2 \end{cases}$$

Complexity

Weak Coupling:

$d = 1$	$(1/\epsilon)^{1.5}$
$d = 2$	$(1/\epsilon)^{2.376}$
$d = 3$	$(1/\epsilon)^{5.5}$

Strong Coupling:

	$\lambda_c - \lambda_0$	p	n_{out}
$d = 1$	$\left(\frac{1}{\lambda_c - \lambda_0}\right)^9$	p^4	n_{out}^5
$d = 2$	$\left(\frac{1}{\lambda_c - \lambda_0}\right)^{6.3}$	p^6	$n_{\text{out}}^{7.128}$

Fermions:

- Fermion doubling problem
- Free vacuum different from Bosonic case

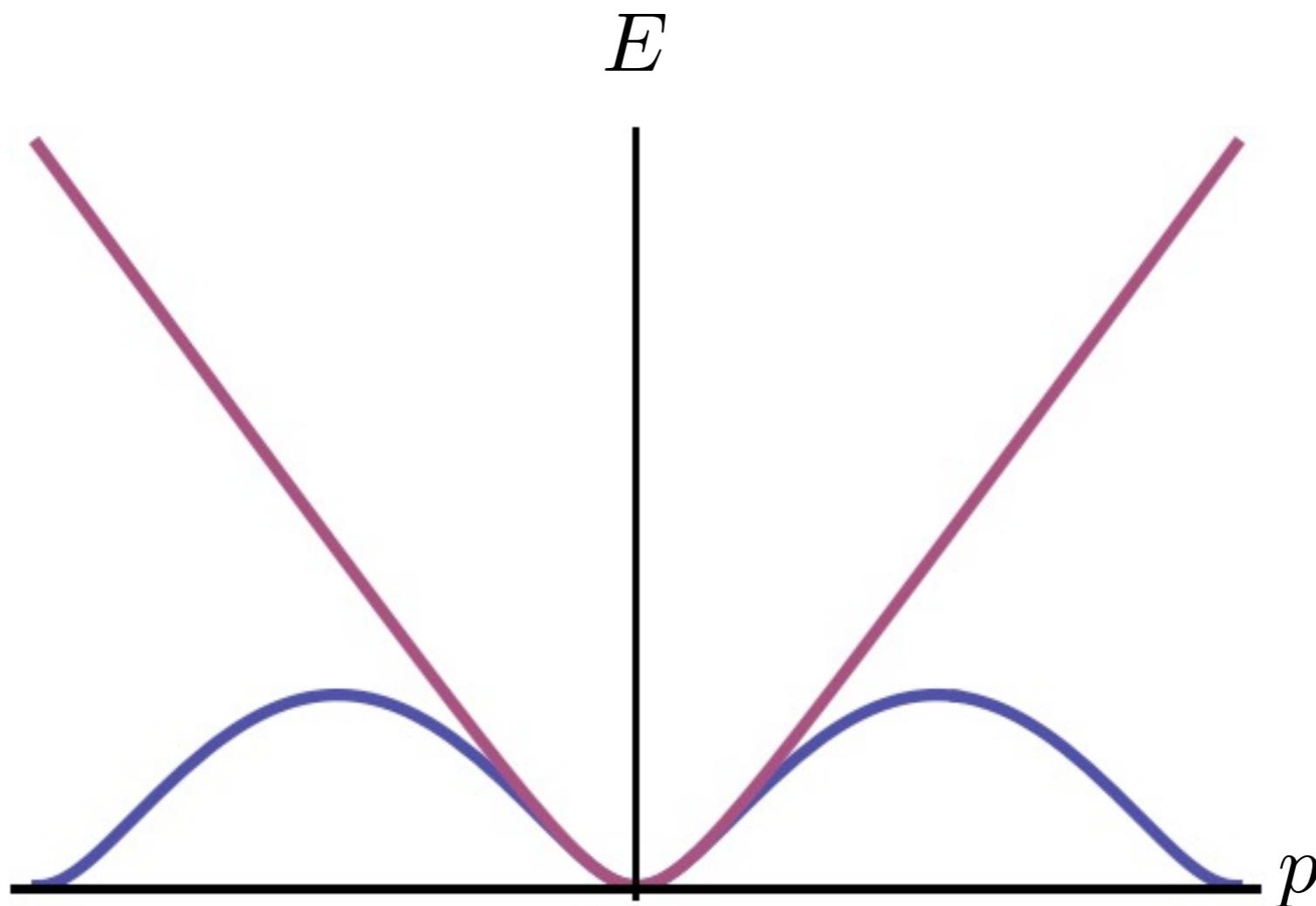
Gross-Neveu:

$$H = \int dx \left[\sum_{j=1}^N \bar{\psi}_j \left(m_0 - i\gamma^1 \frac{d}{dx} \right) \psi_j + \frac{g^2}{2} \left(\sum_{j=1}^N \bar{\psi}_j \psi_j \right)^2 \right]$$

Fermion Doubling Problem

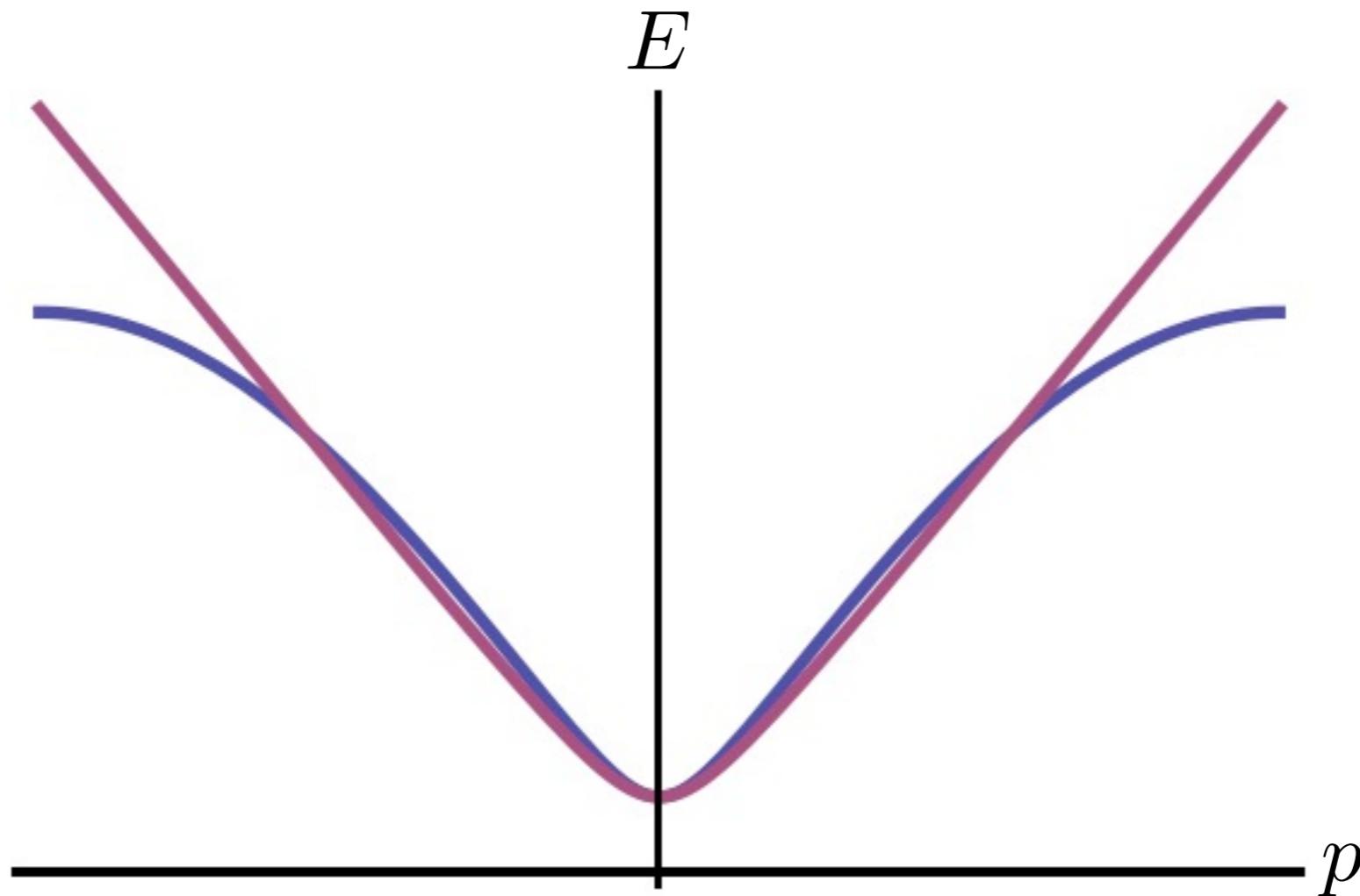
$$\frac{d\psi}{dx} \rightarrow \frac{\psi(x+a) - \psi(x-a)}{2a}$$

$$\sqrt{p^2 + m^2} \rightarrow \sqrt{\sin^2 p + m^2}$$



Wilson Term

$$H \rightarrow H - \frac{r}{2a} \sum_x \bar{\psi} (\psi(x+a) - 2\psi(x) + \psi(x-a))$$



Preparing Fermionic Vacuum

$$H = \sum_x \bar{\psi}(x) m \psi(x)$$

↓ Adiabatic

$$H = \sum_x \bar{\psi}(x) \left(m + \frac{d}{dx} \right) \psi(x)$$

↓ Adiabatic

$$H = \sum_x \bar{\psi}(x) \left(m + \frac{d}{dx} \right) \psi(x) + \frac{g^2}{2} (\bar{\psi}(x) \psi(x))^2$$

Eventual goal:

Simulate the standard model in polynomial time with quantum circuits.

Solved problems:

ϕ^4 -theory [[Science, 336:1130 \(2012\)](#)]

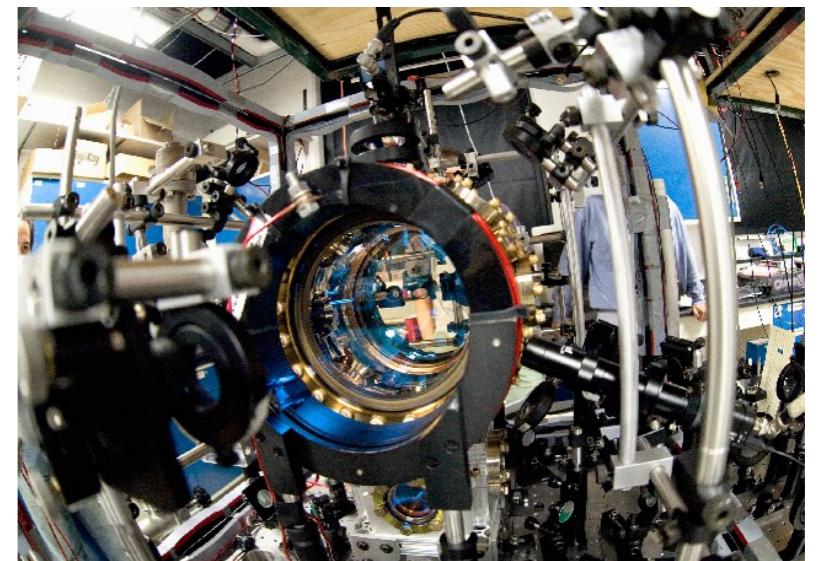
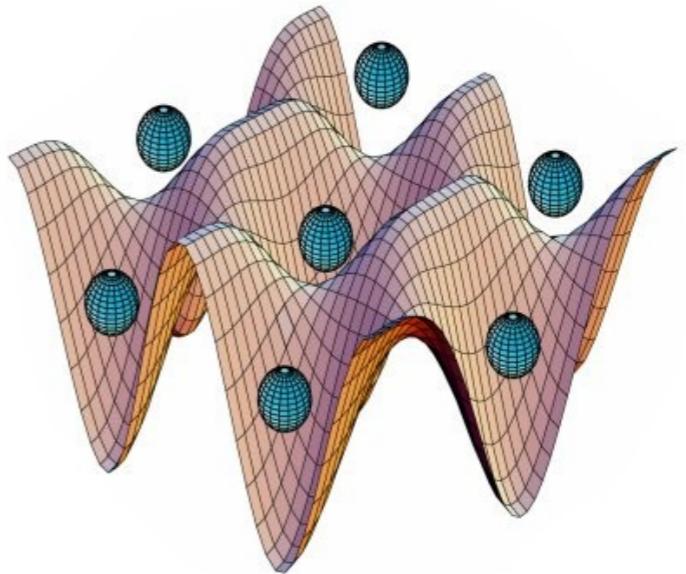
Gross-Neveu [[S.J., Lee, Preskill, *in preparation*](#)]

Open problems:

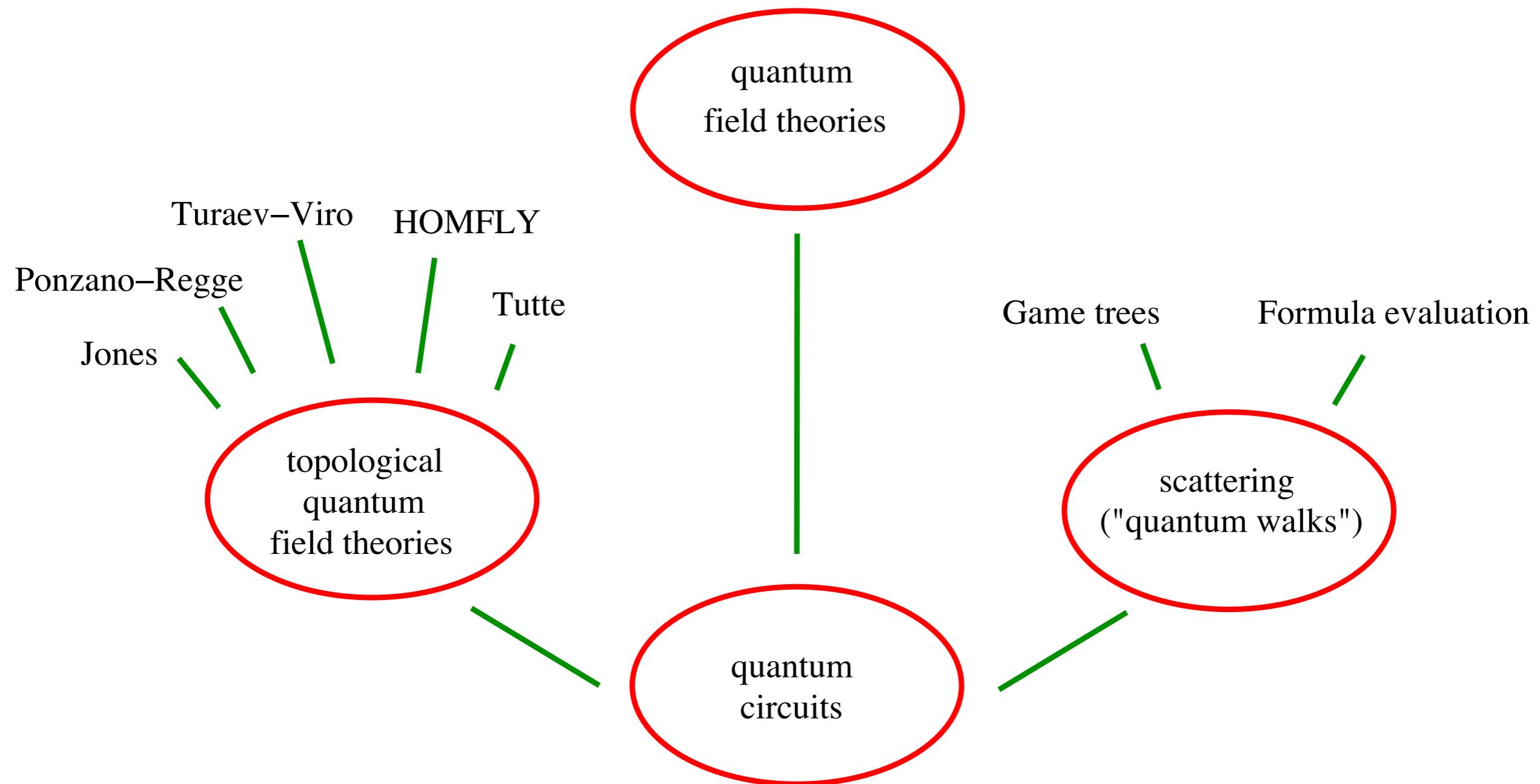
gauge symmetries, massless particles,
spontaneous symmetry breaking,
bound states, confinement, chiral fermions

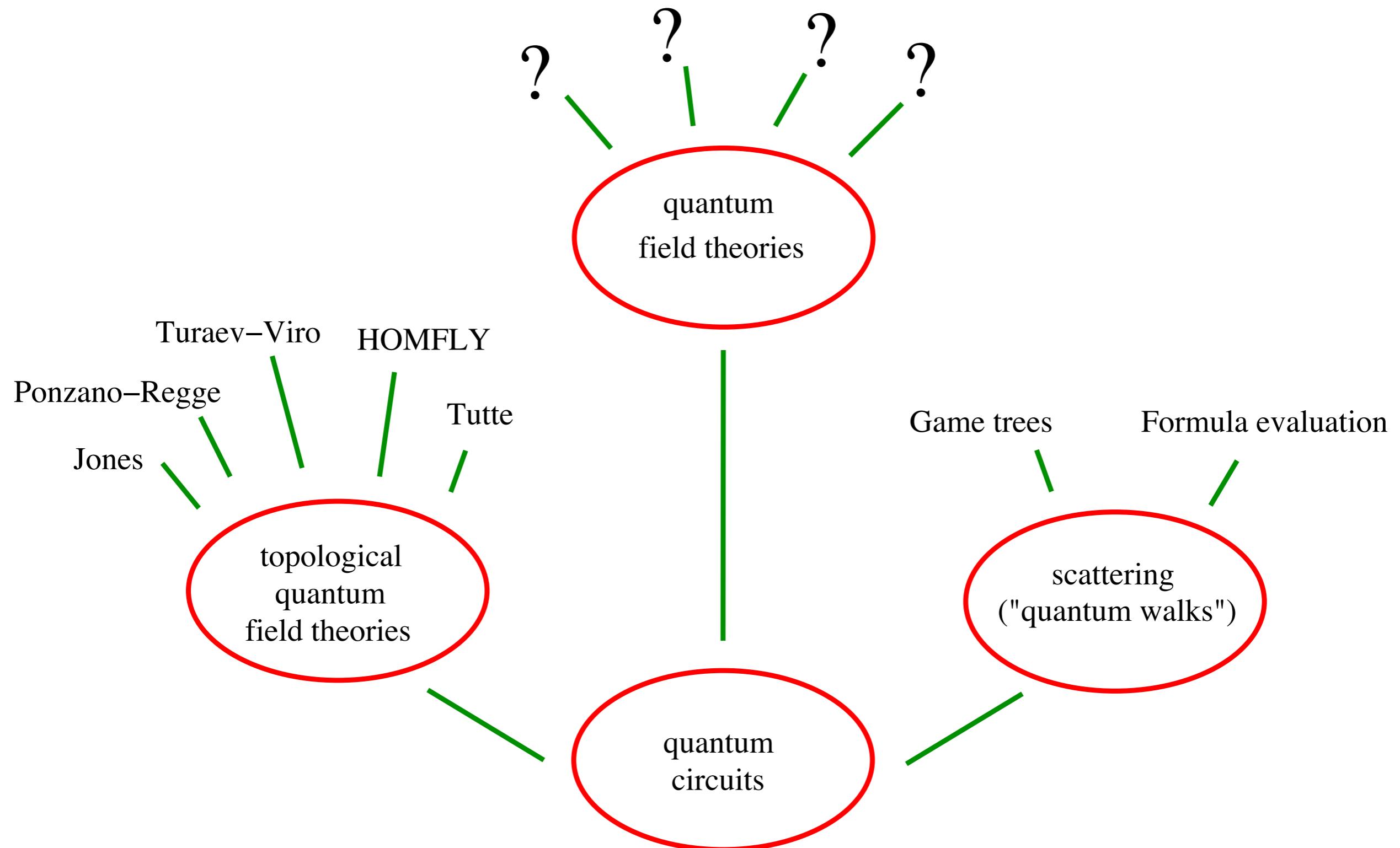
Analog Simulation

- No gates: just implement a Hamiltonian and let it time-evolve
- Current experiments do this!



Broader Context







What I'm trying to do is get you people who think about computer simulation to see if you can't invent a different point of view than the physicists have.

-Richard Feynman, 1981



In thinking and trying out ideas about “what is a field theory” I found it very helpful to demand that a correctly formulated field theory should be soluble by computer... It was clear, in the ‘60s, that no such computing power was available in practice.

-Kenneth Wilson, 1982

Conclusion

Quantum computers can simulate scattering in ϕ^4 -theory and the Gross-Neveu model.

There are many exciting prospects for quantum computation and quantum field theory to contribute to each other's progress.

I thank my collaborators:



Thank you for your attention.