

General Relativity Fall 2019

Homework 4 solutions

Exercise 1: Index manipulation

(i) If the tensor $T_{\alpha\beta}$ is symmetric, show that $T^\alpha{}_\beta = T_\beta{}^\alpha$.

$$\begin{aligned} T^\alpha{}_\beta &= g^{\alpha\lambda} T_{\lambda\beta} \quad [\text{by definition}] \\ &= g^{\alpha\lambda} T_{\beta\lambda} \quad [T_{\alpha\beta} \text{ is symmetric}] \\ &= T_\beta{}^\alpha \quad [\text{by definition}]. \end{aligned} \tag{1}$$

(ii) Given a rank (0,2) tensor $T_{\alpha\beta}$, what is the rank of the tensor $T_{\alpha\beta} T_\gamma{}^\sigma T^{\beta\gamma}$?

Rank (1,1): there is one free index up and one free index down (all others are contracted).

How about $T_{\alpha\beta} T_\gamma{}^\alpha T^{\beta\gamma}$?

Rank (0,0), i.e. this is a scalar, because all indices are contracted.

Suppose that in some basis, the components of a tensor $T_{\alpha\beta}$ are given by

$$\begin{pmatrix} T_{00} & T_{01} & \cdots \\ T_{10} & T_{11} & \\ \vdots & & \ddots \end{pmatrix} = \begin{pmatrix} -1 & 0 & 2 & 3 \\ 1 & 0 & 2 & 1 \\ -1 & 3 & 0 & -1 \\ -1 & 0 & -2 & 3 \end{pmatrix}.$$

(iii) Explicitly write the components of the tensors $T_{(\alpha\beta)}$ and $T_{[\alpha\beta]}$ in this basis (using the same matrix convention).

$$\begin{aligned} \begin{pmatrix} T_{(00)} & T_{(01)} & \cdots \\ & T_{(11)} & \\ \vdots & & \ddots \end{pmatrix} &= \begin{pmatrix} -1 & 1/2 & 1/2 & 1 \\ & 0 & 5/2 & 1/2 \\ & & 0 & -3/2 \\ & \text{(symmetric)} & & 3 \end{pmatrix}. \\ \begin{pmatrix} T_{[00]} & T_{[01]} & \cdots \\ & T_{[11]} & \\ \vdots & & \ddots \end{pmatrix} &= \begin{pmatrix} 0 & -1/2 & 3/2 & 2 \\ & 0 & -1/2 & 1/2 \\ & & 0 & 1/2 \\ & \text{(antisymmetric)} & & 0 \end{pmatrix}. \end{aligned}$$

(iv) Compute $T^\alpha{}_\alpha$ and $T_{\alpha(\beta} T_\gamma{}^\alpha T^{\beta\gamma]}$.

Since we are not given the metric components in the basis, **we cannot compute** $T^\alpha{}_\alpha = g^{\alpha\beta} T_{\alpha\beta}$.

There is no need to explicitly write any calculation for $T_{\alpha(\beta} T_\gamma{}^\alpha T^{\beta\gamma]}$: the contraction of a symmetric pair of indices with an antisymmetric vanishes:

$$X_{(\beta\gamma)} Y^{[\beta\gamma]} = X_{(\gamma\beta)} Y^{[\beta\gamma]} \quad [\text{by symmetry of the lower two indices}] \tag{2}$$

$$= -X_{(\gamma\beta)} Y^{[\gamma\beta]} \quad [\text{by antisymmetry of the upper two indices}] \tag{3}$$

$$= -X_{(\beta\gamma)} Y^{[\beta\gamma]} \quad [\text{renaming dummy indices}]. \tag{4}$$

Hence $T_{\alpha(\beta} T_\gamma{}^\alpha T^{\beta\gamma]} = 0$.

Exercise 2: Covariant derivatives

(i) Consider two covariant derivatives ${}_A\nabla$ and ${}_B\nabla$, with associated connection coefficients ${}_A\Gamma_{\nu\sigma}^\mu$ and ${}_B\Gamma_{\nu\sigma}^\mu$. Prove that the difference ${}_A\Gamma_{\nu\sigma}^\mu - {}_B\Gamma_{\nu\sigma}^\mu$ transforms as a tensor, even though neither connection separately is a tensor.

This simply follows from the cancellation of the non-tensor part (the second derivatives of coordinates) when taking the difference.

(ii) In class we explicitly derived the expression for the covariant derivative of a dual vector field, starting from the covariant derivative of a vector field. With the same procedure, derive the explicit expression for the components of the covariant derivative of a rank-(0, 2) tensor field, $\nabla_\mu T_{\nu\sigma}$, in a coordinate basis. Do the same thing for a rank-(1, 1) tensor field, i.e. derive the explicit expression for $\nabla_\mu T^\nu{}_\sigma$ in a coordinate basis.

Let us pick two arbitrary vector fields X^α and Y^β . By definition of a rank (0, 2) tensor field, the quantity $T(X, Y) = T_{\mu\nu}X^\mu Y^\nu$ is a scalar field. Therefore, we have

$$\nabla_\lambda(T_{\mu\nu}X^\mu Y^\nu) = \frac{\partial}{\partial x^\lambda}(T_{\mu\nu}X^\mu Y^\nu) = \frac{\partial T_{\mu\nu}}{\partial x^\lambda}X^\mu Y^\nu + T_{\mu\nu}\left(\frac{\partial X^\mu}{\partial x^\lambda}Y^\nu + X^\mu\frac{\partial Y^\nu}{\partial x^\lambda}\right), \quad (5)$$

where in the second equality we just used Leibniz' rule for partial derivatives. On the other hand, using Leibniz rule for the covariant derivative, we have

$$\begin{aligned} \nabla_\lambda(T_{\mu\nu}X^\mu Y^\nu) &= (\nabla_\lambda T_{\mu\nu})X^\mu Y^\nu + T_{\mu\nu}((\nabla_\lambda X^\mu)Y^\nu + X^\mu(\nabla_\lambda Y^\nu)) \\ &= (\nabla_\lambda T_{\mu\nu})X^\mu Y^\nu + T_{\mu\nu}\left(\frac{\partial X^\mu}{\partial x^\lambda}Y^\nu + \Gamma_{\lambda\sigma}^\mu X^\sigma Y^\nu + X^\mu\frac{\partial Y^\nu}{\partial x^\lambda} + X^\mu\Gamma_{\lambda\sigma}^\nu Y^\sigma\right). \end{aligned} \quad (6)$$

Equating the two, we see that partial derivatives of X and Y drop out, and we get

$$(\nabla_\lambda T_{\mu\nu})X^\mu Y^\nu + T_{\mu\nu}(\Gamma_{\lambda\sigma}^\mu X^\sigma Y^\nu + X^\mu\Gamma_{\lambda\sigma}^\nu Y^\sigma) - \frac{\partial T_{\mu\nu}}{\partial x^\lambda}X^\mu Y^\nu = 0. \quad (7)$$

Renaming dummy indices, we may factorize $X^\mu Y^\nu$. Since the equality holds for any vectors X, Y , we conclude that the quantity multiplying it vanishes, i.e.

$$\nabla_\lambda T_{\mu\nu} = \frac{\partial T_{\mu\nu}}{\partial x^\lambda} - \Gamma_{\lambda\mu}^\sigma T_{\sigma\nu} - \Gamma_{\lambda\nu}^\sigma T_{\mu\sigma}. \quad (8)$$

The proof for a rank-(1, 1) tensor is essentially identical.

(iii) A metric-compatible covariant derivative satisfies $\nabla_\alpha g_{\beta\gamma} = 0$. Assuming the covariant derivative is moreover torsion-free, prove that the connection coefficients must be the Christoffel symbols.

Applying the above expression to the metric tensor gives

$$0 = \nabla_\lambda g_{\mu\nu} = \frac{\partial g_{\mu\nu}}{\partial x^\lambda} - \Gamma_{\lambda\mu}^\sigma g_{\sigma\nu} - \Gamma_{\lambda\nu}^\sigma g_{\mu\sigma} \Rightarrow \frac{\partial g_{\mu\nu}}{\partial x^\lambda} = \Gamma_{\lambda\mu}^\sigma g_{\sigma\nu} + \Gamma_{\lambda\nu}^\sigma g_{\mu\sigma}, \quad (9)$$

where we used the symmetry of the metric tensor in the last expression. Guided by the expression for the Christoffel connection, let us compute the sum

$$\begin{aligned} \frac{\partial g_{\mu\lambda}}{\partial x^\nu} + \frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial x^\lambda} &= \Gamma_{\nu\mu}^\sigma g_{\sigma\lambda} + \Gamma_{\nu\lambda}^\sigma g_{\sigma\mu} + \Gamma_{\mu\nu}^\sigma g_{\sigma\lambda} + \Gamma_{\mu\lambda}^\sigma g_{\sigma\nu} - \Gamma_{\lambda\mu}^\sigma g_{\sigma\nu} - \Gamma_{\lambda\nu}^\sigma g_{\mu\sigma} \\ &= (\Gamma_{\nu\mu}^\sigma + \Gamma_{\mu\nu}^\sigma) g_{\sigma\lambda} + (\Gamma_{\nu\lambda}^\sigma - \Gamma_{\lambda\nu}^\sigma) g_{\sigma\mu} + (\Gamma_{\mu\lambda}^\sigma - \Gamma_{\lambda\mu}^\sigma) g_{\sigma\nu} \\ &= 2\Gamma_{(\mu\nu)}^\sigma g_{\sigma\lambda} + 2\Gamma_{[\nu\lambda]}^\sigma g_{\sigma\mu} + 2\Gamma_{[\mu\lambda]}^\sigma g_{\sigma\nu} = 2\Gamma_{(\mu\nu)}^\sigma g_{\sigma\lambda} = 2\Gamma_{\mu\nu}^\sigma g_{\sigma\lambda}, \end{aligned} \quad (10)$$

since we assumed the torsion vanishes. Multiplying by $\frac{1}{2}g^{\delta\lambda}$, we then find

$$\Gamma_{\mu\nu}^\delta = \frac{1}{2}g^{\delta\lambda}\left(\frac{\partial g_{\lambda\mu}}{\partial x^\nu} + \frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial x^\lambda}\right), \quad (11)$$

which is indeed the Christoffel symbol.

Exercise 3: Non-relativistic, perturbed geodesics

(i) Suppose a particle moves on a geodesic at low velocity, i.e. such that, in the coordinates $\{x^\mu\}$, $v^i = dx^i/dt \ll 1$, where as usual $t \equiv x^0$. Compute the coordinate acceleration, i.e. d^2x^i/dt^2 , at second order in the velocity, i.e. neglecting terms of order v^3 and higher. Specifically, find the coefficients A^i, B_j^i and C_{jk}^i such that

$$\frac{d^2x^i}{dt^2} = A^i + B_j^i v^j + C_{jk}^i v^j v^k + \mathcal{O}(v^3). \quad (12)$$

For question (i), keep the Christoffel symbols when they appear, to keep compact expressions (i.e. do not write their explicit expression in terms of the metric). Note that it is d^2x^i/dt^2 that we want to compute here, not $d^2x^i/d\tau^2$.

Chain rule!

$$\frac{d^2x^i}{dt^2} = (dt/d\tau)^{-1} \frac{d}{d\tau} \left((dt/d\tau)^{-1} \frac{dx^i}{d\tau} \right) = (dt/d\tau)^{-2} \frac{d^2x^i}{d\tau^2} - (dt/d\tau)^{-3} \frac{dx^i}{d\tau} \frac{d^2t}{d\tau^2}. \quad (13)$$

Now let's recall that $dx^i/d\tau = (dt/d\tau)dx^i/dt = (dt/d\tau)v^i$, so we find

$$\frac{d^2x^i}{dt^2} = (dt/d\tau)^{-2} \left[\frac{d^2x^i}{d\tau^2} - v^i \frac{d^2t}{d\tau^2} \right]. \quad (14)$$

We can now plug in the geodesic equation:

$$\begin{aligned} \frac{d^2x^i}{dt^2} &= -(dt/d\tau)^{-2} \left[(\Gamma_{00}^i - v^i \Gamma_{00}^0)(dt/d\tau)^2 + 2(\Gamma_{0j}^i - v^i \Gamma_{0j}^0) \frac{dt}{d\tau} \frac{dx^j}{d\tau} + (\Gamma_{jk}^i - v^i \Gamma_{jk}^0) \frac{dx^j}{d\tau} \frac{dx^k}{d\tau} \right] \\ &= -[(\Gamma_{00}^i - v^i \Gamma_{00}^0) + 2(\Gamma_{0j}^i - v^i \Gamma_{0j}^0)v^j + (\Gamma_{jk}^i - v^i \Gamma_{jk}^0)v^j v^k]. \end{aligned} \quad (15)$$

So far this is an *exact expression*. We can put it in the desired form by defining

$$\boxed{A^i \equiv -\Gamma_{00}^i, \quad B_j^i \equiv \delta_j^i \Gamma_{00}^0 - 2\Gamma_{0j}^i, \quad C_{jk}^i = 2\delta_{(j}^i \Gamma_{k)0}^0 - \Gamma_{jk}^i}. \quad (16)$$

Note that I purposefully wrote the last term in a clearly symmetric way.

Now suppose that moreover, the metric components are nearly Minkowski in this coordinate system, i.e. $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$, where $|h_{\mu\nu}| \ll 1$. You can expect that the inverse metric components will also be close to Minkowski: $g^{\mu\nu} = \eta^{\mu\nu} + k^{\mu\nu}$, with $|k^{\mu\nu}| \ll 1$.

(ii) Compute $k^{\mu\nu}$ explicitly in terms of $h_{\mu\nu}$, at first order (i.e. neglecting terms quadratic in $h_{\mu\nu}$).

We enforce that $g^{\mu\nu}g_{\nu\sigma} = \delta_\sigma^\mu$ and expand to linear order in perturbations:

$$\delta_\sigma^\mu = (\eta^{\mu\nu} + k^{\mu\nu})(\eta_{\nu\sigma} + h_{\nu\sigma}) = \eta^{\mu\nu}\eta_{\nu\sigma} + \eta^{\mu\nu}h_{\nu\sigma} + k^{\mu\nu}\eta_{\nu\sigma} + \mathcal{O}(h^2). \quad (17)$$

The first term on the right-hand-side is δ_σ^μ , thus we find

$$k^{\mu\nu}\eta_{\nu\sigma} = -\eta^{\mu\nu}h_{\nu\sigma}. \quad (18)$$

Multiplying on the right by the inverse-Minkowski metric, we find

$$\boxed{k^{\mu\lambda} = -\eta^{\mu\nu}h_{\nu\sigma}\eta^{\sigma\lambda}}. \quad (19)$$

Important: $h_{\mu\nu}$ is not a tensor. Only the full metric $g_{\mu\nu}$ is a general tensor, and $\eta_{\mu\nu}$ is just a symbol, not a tensor (it only behaves as a tensor for Lorentz transformations). Thus, $h^{\mu\nu}$ is meaningless: we only defined $h_{\mu\nu} \equiv g_{\mu\nu} - \eta_{\mu\nu}$, and it is **only meaningful to raise indices of tensors**.

(iii) Compute the Christoffel symbols at second order in $h_{\mu\nu}$, i.e. neglecting terms cubic in $h_{\mu\nu}$.

The partial derivatives in the Christoffel symbols only apply to the perturbation, so we just need to include terms linear in $h_{\mu\nu}$ in the inverse metric:

$$\Gamma_{\mu\nu}^\lambda = \frac{1}{2}(\eta^{\lambda\sigma} - \eta^{\lambda\rho}h_{\rho\delta}\eta^{\delta\sigma})(h_{\mu\sigma,\nu} + h_{\nu\sigma,\mu} - h_{\mu\nu,\sigma}). \quad (20)$$

(iv) Write explicit expressions of A^i, B_j^i and C_{jk}^i in terms of $h_{\mu\nu}$, at first order in $h_{\mu\nu}$ only (i.e. neglecting terms quadratic in $h_{\mu\nu}$ and higher). You will thus not need the full expression for the Christoffel symbols derived in (iii).

At first order in $h_{\mu\nu}$, we just need the first term in the first parenthesis. We can then substitute $\eta^{i\sigma} = \delta^{i\sigma}$ and $\eta^{0\lambda} = -\delta^{0\lambda}$, and find

$$A^i = \frac{1}{2}h_{00,i} - h_{0i,0}, \quad (21)$$

$$B_j^i = -\frac{1}{2}h_{00,0} \delta_j^i - h_{ij,0} + 2h_{0[j,i]}, \quad (22)$$

$$C_{jk}^i = -h_{00,(j} \delta_{k)}^i + \frac{1}{2}h_{jk,i} - h_{i(j,k)}. \quad (23)$$

Note the use of symmetrization/ antisymmetrization brackets to keep compact expressions.