

PHY401-Nuclear and Particle Physics

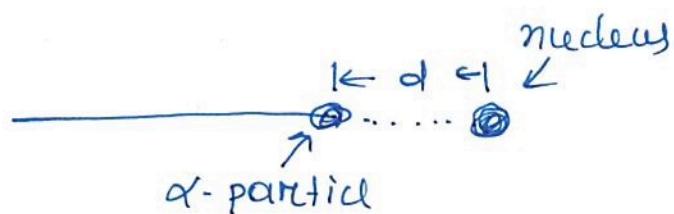
Academic Year:	2024-25	Instructor:	Dr. Satyajit Jena,
Class	MS21	Lecture	Nuclear Properties

Nuclear properties

Based on previous discussion, we will start discussing the properties of nucleus. In this we must first establish the size of nucleus. As a historical event I discussed about the importance of Rutherford scattering experiment & the result out of its findings. Let's find out some common facts about Rutherford scattering.

One of the ways to determine the size of object is to look into a scattering of small object into the object under observation. In that sense if Rutherford was using α -particle as probe (due to the availability) he must choose a nucleus bigger than α -particle precisely for same reason, he chose the Gold or higher element metal foil for his experiment.

Starting with head-on collision of α -particle with the nucleus.



Say at a distance d from the nucleus α -particle starts reversing its direction due to the coulomb repulsion. Why?

Because α -particle is at charge $\geq 2e$ nucleus is five charged $\Rightarrow Z_n$.

Therefore α -particle being at charges will suffer a coulomb repulsion as it comes closer to the approaches to nucleus with +ve charges. Even eventually it will stop at particle distance say d from nucleus & start coming back. In other word we can say at d the kinetic Energy of α -particle, T , becomes equal to coulomb energy.

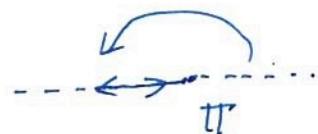
Then

$$T = \frac{ZeZ_e}{4\pi\epsilon_0 d} = \frac{z^2 e^2}{4\pi\epsilon_0 d}$$

or

$$d = \frac{z^2 e^2}{4\pi\epsilon_0 T}$$

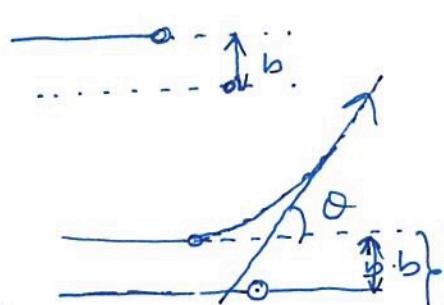
What is the scattering angle θ ?



$\boxed{\theta = \pi}$ for such situation.

What would be the situation if the α -particle is not in the line of sight?

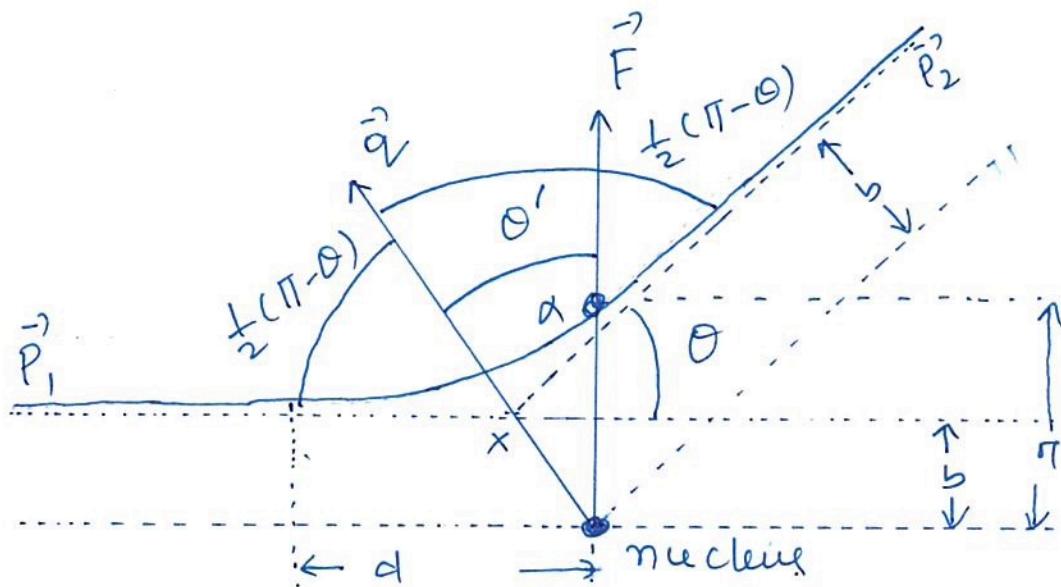
If the line of incidence of the α -particle is at a distance b from the nucleus, then the scattering angle will be smaller.



$b \rightarrow$ This is called impact parameter.

By using Newton's Second law of motion, coulomb's law for the force between α -particle and nucleus, and conservation of angular momentum, we can calculate the relation between $\theta, d \& b$. i.e. the relation between impact parameter, scattering angle & distance of columbic deflection.

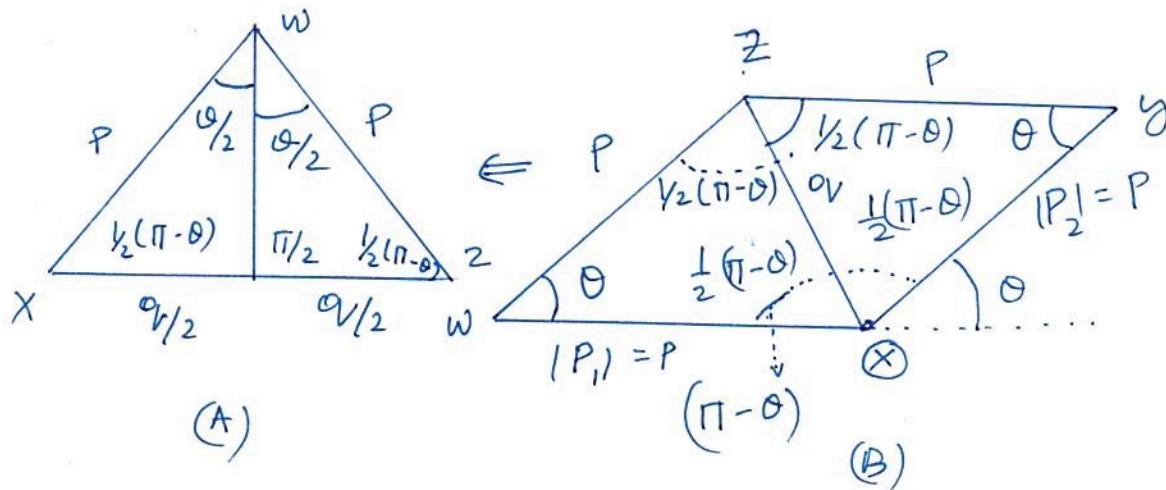
lets draw the picture it carefully.



Here, note that if $\theta = \pi$, b becomes zero $\Rightarrow b=0$
That is if b increases the α -particle 'glances' the nucleus so that the scattering angle decreases.

Assuming elastic scattering, we can approximate that the magnitude of initial momentum \vec{P}_1 & final momentum \vec{P}_2 of α -particle will be of same value, P
 \vec{v} , therefore is the resultant vector of \vec{P}_1 & \vec{P}_2

To there with \vec{P}_1 ; \vec{P}_2 , the momentum change vector of will form an isoscele triangle with angle θ between P_1 & P_2 (initial & final) vectors result.



Making more geometrical analysis we can make fig (B) from the scattering pictures, this then to ~~not~~ have more clear pictures we can make \oplus figure (A) which is a isosceles triangle. Now using pure geometry we can write.

$$\sin\left(\frac{\theta}{2}\right) = \frac{\left(\frac{q}{2}\right)}{P}.$$

Just before clarification, the \vec{q} is the recoil momentum vector, & the direction of this vector is along the line joining the nucleus to closest approach of the α - particle.

Since the target nucleus is heavier than α -particle we can neglect its recoil. we also neglect all relativist effects.

By Newton's second law, the rate of change of momentum in the direction of q is the component of the force acting on the α -particle due to the electric charge of the nucleus. Thus by coulomb's law the magnitude of the force is

$$F = \frac{zZe^2}{4\pi\epsilon_0 r^2}$$

Putting Kinetic energy T derived earlier.

The position of α -particle is in terms of 2-D polar coordinates (r, θ') with the nucleus at origin & $\theta' = 0$ chosen to be the point of closest approach

$$T = \frac{zZe^2}{4\pi\epsilon_0} \frac{1}{d} \Rightarrow \left[\frac{zZe^2}{4\pi\epsilon_0} = Td \right]$$

$$\Rightarrow \boxed{F = \frac{Td}{r^2}}$$

The component of this force in the direction of q , vector will be

$$F_q(t) = \frac{Td}{r^2} \cos(\theta'(t))$$

Thus the change of momentum can be given by

$$q = \int \frac{dq}{dt} = \int F_q(t) dt$$

$$\Rightarrow q = \int \frac{zZe^2}{4\pi\epsilon_0} \frac{1}{r^2} \cos \theta'(t) dt$$

we can replace integration of time over the θ' using

$$dt = \frac{d\theta'}{\dot{\theta}'}$$

$\dot{\theta}'$ can be obtained from the conservation of angular momentum.

$$L = m_d r^2 \dot{\theta}'$$

m_d = mass of α -particle.

2 the initial angular momentum is given by

$$L = b p$$

$$\Rightarrow \boxed{\dot{\theta}' = \frac{bp}{m_d r^2}}$$

Now we can rewrite the q as

$$q = \int \frac{z Z e^2}{4\pi G_0} \frac{1}{r^2} \cos \theta' dt$$

$$= \int \frac{z Z e^2}{4\pi G_0} \frac{1}{r^2} \cos \theta' \frac{d\theta'}{\dot{\theta}'}$$

$$= \int \frac{z Z e^2}{4\pi G_0} \frac{1}{r^2} \frac{\cos \theta' d\theta'}{\left(\frac{bp}{m_d r^2} \right)}$$

$$= \int \frac{T d}{r^2} \frac{r^2 m_d}{bp} \cos \theta' d\theta'$$

$$= \int \frac{p^2 d}{2 m_d r^2} \frac{r^2 m_d}{bp} \cos \theta' d\theta'$$

$$T = \frac{p^2}{2 m_d}$$

$$\boxed{q = \int \frac{p d}{2b} \cos \theta' d\theta'}$$

Kinetic energy
in terms of
momentum of
 α -particle.

Let's perform the integration of R.H.S

④ Before integration let's decide what is the limit of θ' , This can move from $-\frac{1}{2}(\pi - \theta)$ to $+\frac{1}{2}(\pi - \theta)$

$$q = \int_{-\frac{1}{2}(\pi - \theta)}^{+\frac{1}{2}(\pi - \theta)} \frac{pd}{2b} \cos \theta' d\theta'$$

$$q = \frac{pd}{2b} [2 \sin(\frac{1}{2}(\pi - \theta))]$$

Earlier we calculated $\frac{q}{p} = 2 \sin(\frac{\theta}{2})$

$$\Rightarrow \frac{q}{p} = \frac{d}{2b} [2 \sin\{\frac{1}{2}(\pi - \theta)\}]$$

$$\Rightarrow 2 \sin(\frac{\theta}{2}) = \frac{d}{b} \sin(\frac{1}{2}(\pi - \theta))$$

$$\Rightarrow \frac{\sin \frac{\theta}{2}}{\cos \frac{\theta}{2}} = \frac{d}{2b} \quad \begin{aligned} &\cos(\pi/2 - \theta/2) \\ &= \cos \frac{\theta}{2} \end{aligned}$$

$$\Rightarrow \boxed{\tan(\frac{\theta}{2}) = \frac{d}{2b}}$$

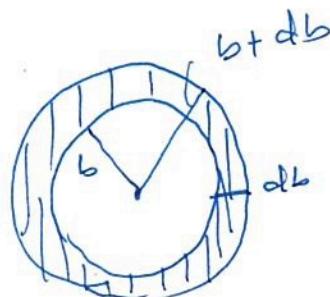
Above calculation was for the single α -particle but in reality, the ~~same~~ experiment was conducted with α -source. This α -source emits large number of α -particles. Thus we must have a way to calculate θ & the scattering angle of all these. But since all these α -particle will have different lines of incidence, their scattering angle will also be very different.

Thus getting the σ_s calculated for each of them individually is not possible. Then we can calculate the total number of α -particle passing in a particular area per unit time. This is precisely we call FLUX of α -particle.

~~FLUX~~

Let's say $dN(b)$ is the number of α -particle within the impact parameter between b & $b+db$. Thus the $dN(b)$ will be the flux multiplied by the area between two concentric circle.

Flux ϕ is number of particle per unit area. per second.



$$\text{Area} = 2\pi b \times db$$

$$dN(b) = \phi 2\pi b db.$$

Earlier we calculated

$$\tan \frac{\theta}{2} = \frac{d}{2b} \Rightarrow$$

$$\Rightarrow b = \frac{d}{2} \frac{1}{\tan(\frac{\theta}{2})} = \frac{d}{2} \cot(\theta/2)$$

$$\begin{aligned} \Rightarrow \frac{db}{d\theta} &= \frac{d}{2} \frac{d}{d\theta} \cot(\theta/2) = \frac{d}{2} \frac{d}{dx} \cot x \frac{dx}{d\theta} \quad (x = \theta/2) \\ &= \frac{d}{2} \cdot -\operatorname{cosec}^2 x \frac{dx}{d\theta} \end{aligned}$$

$$\Rightarrow \frac{db}{d\theta} = -\frac{d}{2} \csc^2 \frac{\theta}{2} \frac{d(\frac{\theta}{2})}{d\theta}$$

$$= -\frac{d}{2} \csc^2 \frac{\theta}{2} \frac{1}{2} \frac{d\theta}{d\theta}$$

$$\Rightarrow \frac{db}{d\theta} = -\frac{d}{4} \frac{1}{\sin^2(\frac{\theta}{2})}$$

$$\Rightarrow \boxed{db = -\frac{d}{4 \sin^2(\frac{\theta}{2})} d\theta}$$

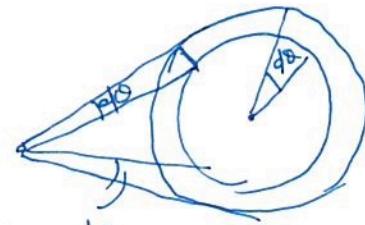
No $dN(b) = \phi 2\pi b db$ can be written in terms of θ & $d\theta$ as.

$$dN(\theta) = \phi 2\pi b \left(-\frac{d}{4 \sin^2 \frac{\theta}{2}}\right) d\theta$$

$dN(\theta)$

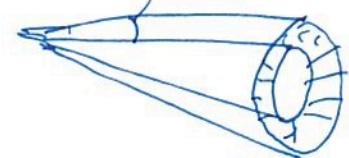
This the number of α -particle between the angle area θ & $\theta + d\theta$. What will call \Rightarrow "solid angle".

Let's use $\tan \frac{\theta}{2} = \frac{d}{2b}$



$$\Rightarrow dN(\theta) = -\phi 2\pi \frac{d}{4 \sin^2 \frac{\theta}{2}} \frac{d}{2 \tan \frac{\theta}{2}} d\theta d\theta$$

$$= -\phi \frac{\pi d^2}{4} \frac{\cos(\frac{\theta}{2})}{\sin^3(\frac{\theta}{2})} d\theta$$



$dN(\theta)$ is a number, so this -ve sign indicates that number should decrease as you increase θ & similarly increase the b θ will decrease.

The differential cross-section is defined by $\frac{d\sigma}{d\Omega}$, with respect to the scattering angle, which is the number of scatterings between θ and $\theta + d\theta$ per unit flux per unit range of angle.

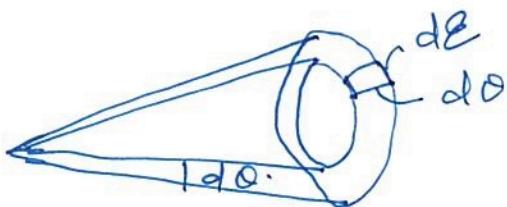
$$\frac{d\sigma}{d\Omega} = \frac{dN(\theta)}{\Phi d\Omega}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = \pi \frac{d^2}{4} \frac{\cos(\frac{\theta}{2})}{\sin^3(\frac{\theta}{2})}$$

The differential cross-section is given by the number within $\theta + d\theta$. What exactly it is? this angle area is nothing but the differential solid angle $d\Omega$

we know that

$$\Omega = \frac{a}{R^2} \left[\begin{array}{l} \text{area} \\ \text{of sphere} \end{array} \right]$$

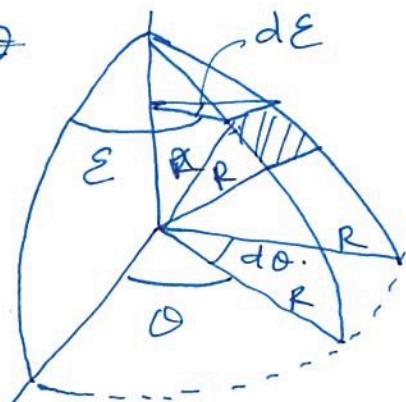


using polar co-ordinate system

we can calculate that

$$d\Omega = (R \sin \theta d\phi)(R d\theta)$$

$$d\Omega = \frac{(R \sin \theta d\phi)(R d\theta)}{R^2}$$



$$\Rightarrow d\Omega = \sin \theta d\phi d\theta$$

$$\Rightarrow d\Omega = \sin\left(\frac{\theta}{2} + \frac{\phi}{2}\right) d\theta d\phi = 2 \sin\left(\frac{\theta}{2}\right) \cos\left(\frac{\theta}{2}\right) d\theta d\phi$$

ϵ is azimuthal angle.

Therefore ~~the~~ The differential cross section with respect to differential solid angle can be written as

$$\textcircled{*} \quad N = \phi \sigma$$

$$\boxed{\frac{dN}{d\Omega} = \phi \frac{d\sigma}{d\Omega}}$$

Thus we need to write $\frac{d\sigma}{d\Omega}$ in terms of $\frac{d\sigma}{d\Omega}$
which is more meaningful.

Thus

$$\frac{d\sigma}{d\Omega} = \pi \frac{d^2}{4} \frac{\cos(\theta/2)}{\sin^3(\theta/2)}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = \frac{d\sigma}{d\Omega} \frac{d\Omega}{d\theta}$$

$$\frac{d\Omega}{d\theta} = 2 \sin(\theta/2) \cos \frac{\theta}{2} d\theta$$

$$= 2 \sin \frac{\theta}{2} \cos \frac{\theta}{2} 2\pi$$

$$\therefore = 4\pi \sin \left(\frac{\theta}{2} \right) \cos \left(\frac{\theta}{2} \right)$$

$$\Rightarrow \frac{d\sigma}{d\Omega} \frac{d\Omega}{d\theta} = \frac{d\sigma}{d\theta}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} 4\pi \sin \left(\frac{\theta}{2} \right) \cos \left(\frac{\theta}{2} \right) = \pi \frac{d^2}{4} \frac{\cos(\theta/2)}{\sin^3(\theta/2)}$$

$$\Rightarrow \boxed{\frac{d\sigma}{d\Omega} = \frac{d^2}{16} \frac{1}{\sin^4(\theta/2)}}$$

$\int d\theta$ is for
total range of
 $\theta + d\theta$.
i.e. all azimuthal
angle.

The differential cross section have dimension of an area and this is usually quoted as barn

$$1 \text{ barn} = 10^{-28} \text{ m}^2$$

$$\Rightarrow 1 \text{ millibarn} = 1 \text{ mb} = \cancel{10^{-28}} 10^{-31} \text{ m}^2$$

$$\Rightarrow 10^{-15} \text{ m} = 1 \text{ fm} \Rightarrow$$

$$\boxed{1 \text{ fm}^2 = 10 \text{ mb}}$$

The difficulties of Rutherford scattering was multiple small angle scattering. Thus to reduce the multiple scattering one needs to make very thin foil. But no matter how thin you make it will certainly ~~have~~ go through multiple scattering.

Say A is mass number then the total number of nuclei per unit area can be given as

$$f \delta \perp \text{Amp}$$

f is density
 δ is thickness.
 A is atomic mass.

Fraction of α particle scattered from small interval of solid angle will be

$$\frac{\delta n}{n} = f \delta \perp \text{Amp} \frac{d\sigma}{d\Omega} d\Omega$$

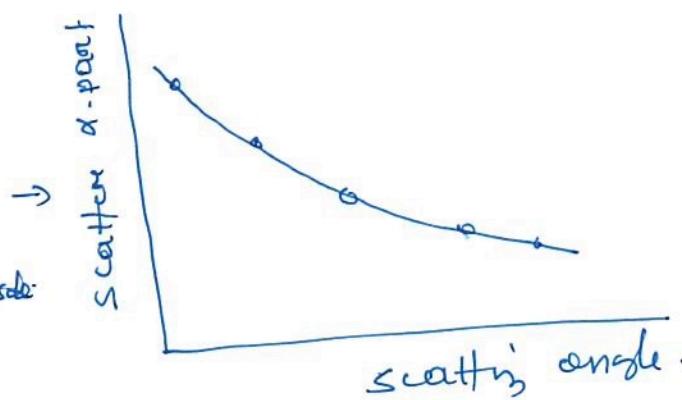
Solid angle is defined such that an area element dA at a distance r from the scattering center subtends a solid angle.

$$d\Omega = \frac{dA}{r^2}$$

Putting a detector with $d\Omega = \frac{dA}{r^2}$, the α -particle can be detected

Experimental result

This theoretical result compares very well with the data taken by Geiger & Marsden Marsden



The coulomb force can be written in terms of constant i :

$$\bullet \boxed{\frac{e^2}{4\pi\epsilon_0} = \alpha \hbar c}$$

$$\hbar c = 197.3 \text{ MeV fm}$$

$$\Rightarrow \frac{e^2}{4\pi\epsilon_0} = \frac{197.3}{137} \text{ MeV fm}$$

$$d = \frac{197.3}{137} \frac{ze}{T} \text{ fm}$$

T is in ~~MeV~~ MeV
which is experimental value.

$$d \approx 10^{-14} \text{ m i.e. } 10 \text{ fm.}$$

To be able to reach the distance of closest approach in order of fm, one needs to increase the energy of α -particle to very large kinetic energy. Thus there are deviations from the Rutherford scattering formulae for very high energy α -particle. A quantum mechanical treatment is required to encounter it.

Finding the nuclear size:

standard way of determining the shape(size) of an object is to look into scattered radiation, the similar way we see any object via scattered light; Doing so, we know the Rutherford's scattering formula where the scattering of α -particles from nuclei is modeled from the Coulomb's force & treated as an orbit. By treating the scattering process statistically for the cross-section of interaction with nucleus (a point like charge Ze), all the quantities of scattering can be found out:

$$\frac{d\sigma}{d\Omega} = \frac{d^2}{16 \sin^4(\frac{\theta}{2})} = \left(\frac{Ze^2}{4\pi \epsilon_0 k_E} \right)^2 \frac{1}{\sin^4 \frac{\theta}{2}}$$

$$N(\theta) = \frac{N_i n L Z^2 k^2 e^4}{4\pi^2 k_E^2 \sin^4(\theta/2)}$$

where

$\theta \rightarrow$ scattering angle

$N_i \rightarrow$ Number of α -particles

$n \rightarrow$ number of target atoms per unit volume

$L \rightarrow$ thickness of the target

$Z \rightarrow$ atomic number of target atom.

$k \rightarrow$ Coulomb constant

$e \rightarrow$ charge of electron

$k_E \rightarrow$ kinetic energy of α -particle

~~It is really difficult to produce α -particle with stuff~~

To see the object and its details, the wavelength of the reradiation must be smaller than the dimensions of the object. otherwise the effects of diffraction will partially or completely obscure the image.

From well known formula

$$\lambda = \frac{h c}{E}$$

λ = wavelength
 h = plank's const.
 E = is energy.

one can see that de-Broglie wave-length is inversely proportional to the its energy.

one can use this fact to resolve the structure of the particle under test if the de-Broglie wave length of a test particle is comparable smaller than the size of particle under test. On contrary if the energy of the test particle is too low, due to its large de-Broglie wave-length it will just bend around the object under study and no its internal structure will be resolved.

Previous class we derived that, as per Rutherford's α -scattering experiment, the size of nucleus or the distance of closest approach $d \approx 10 \text{ fm}$.

so for the nuclei with 10 fm , we require $\lambda \leq 10 \text{ fm}$. correspondingly to $P \geq 100 \text{ MeV/c}$, thus, it is difficult to produce α -particle with sufficient energy to probe particle's size of $\leq 10 \text{ fm}$,

Thus one can use high energy electrons. It is comparatively easier to produce electron beam with 100 MeV to 1 GeV using some of the accelerators, and it can be analyzed with a precise spectrometer to select only those electrons that are elastically scattered from the selected nuclear target.

Considering electron as projectile, it would certainly move faster, thus relativistic treatment is required for the Rutherford's finding, which was first calculated by Mott

$$\left(\frac{d\sigma}{d\Omega}\right)_{\text{Mott}} = \left(\frac{d\sigma}{d\Omega}\right)_{\text{Rutherford}} \left[1 - \frac{v^2}{c^2} \sin^2\left(\frac{\theta}{2}\right)\right]$$

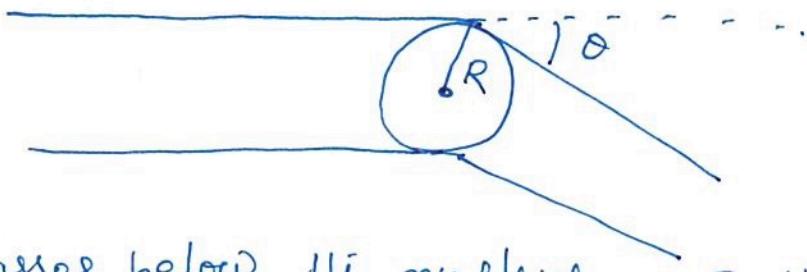
$$\Rightarrow \left(\frac{d\sigma}{d\Omega}\right)_{\text{mott}} = \frac{d^2}{(16 \sin^4 \frac{\theta}{2})} \left(1 - \beta^2 \sin^2\left(\frac{\theta}{2}\right)\right)$$

The cross-section for scattering from a point-like target is given by the Rutherford scattering formula. If the target has a finite spatial extent, the cross-section can be divided into two factors (1) cross-section times (2) squared of a term called "Electric Form Factor", a correction factor; This takes care of the spatial extent & shape of the target. Thus the probability amplitude for a point like scattered gets modified by the form factor.

$$\left.\frac{d\sigma}{d\Omega}\right|_{\text{expt}} = \left.\frac{d\sigma}{d\Omega}\right|_{\text{Mott}} \left[\frac{F(p^2)}{F(q^2)}\right]^2$$

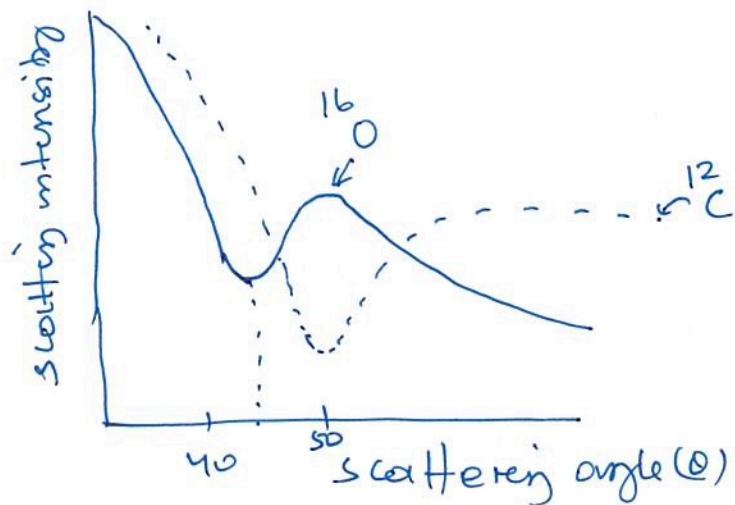
q^2 is the momentum transferred by the electrons in the scattering. Its magnitude is related to the scattering angle. This mathematically is Form Factor is defined as the ratio by which the scattering cross-section is reduced when the charge Ze is spread out over a finite volume.

To understand the structure of the electric form factor we need to recall that the electron has a de-Broglie wavelength $\lambda = h/p$ and when this wavelength is of the order of the nuclear radius we get the diffraction pattern. Simple example that the nucleus were a solid sphere of radius R with an infinite potential inside the sphere & zero outside, so that the electron can't penetrate the sphere. The wave that passes over nucleus travels a distance $\approx 2R \sin\theta$ further than the -



wave that passes below the nucleus. If this difference is equal to $\pi/2, 3\pi/2, \dots$ then we get destructive interference. At these angles the differential cross-section vanishes.

Taking these into consideration, early experiments at SLAC Stanford linear Accelerator center published results in Phy. Rev. 113, 666, (1959).



420 MeV electrons scattering from ^{16}O & ^{12}C was reported.
(The paper reported other results like 240, 360 MeV electron data was also presented.)

If you consider 400 MeV electrons, the $\lambda = \frac{1240}{420} \text{ fm} \approx 3 \text{ fm}$
putting this into the first minima one can find the radii ≈ 4

$$R_{\text{O}^{16}} = 2.6 \text{ fm}$$

$$R_{\text{C}^{12}} = 2.3 \text{ fm}$$

But experiment could not produce the results of ~~the~~ destructive minima. That is these minima do not fall to zero like diffracting minima seen with light incident on an opaque disk. That refers. nucleus does not have sharp boundary.

The above measurements are only rough measurement considering the fact that potential scattering is a three dimensional problem only approximately related to the diffraction by a two dimensional sphere.

The real case is a little more complicated than above. A proper quantum mechanical treatment (analogous to diffraction in optics) shows that the electric form factor is actually the Fourier transform of the charge distribution.

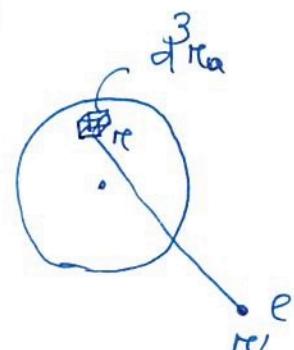
Quantum Mechanical Treatment of elastic scattering of electrons with a finite shape of charge distribution

From previous discussions, it is clear that nucleus is not a point like object, rather a extended object of finite size. Within this size the charges are distributed charges means the protons.

Consider the charges are distributed uniformly with charge density $f(r)$.

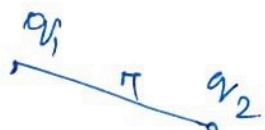
What is the charges of a small volume element d^3r can be calculated as

$$[f(r) d^3r]$$



Say an electron is at r' . & what is the potential energy is

$$V = \frac{q_1 q_2}{4\pi \epsilon_0 r}$$



$$\Rightarrow V = \frac{(f(r) d^3r)(-e)}{4\pi \epsilon_0 |r - r'|}$$

for all volume you integrate it.

Now let us look into the electron scattering with the nucleus.

$$\psi_i = e^{i \vec{k} \cdot \vec{r}} \quad (\text{wavy line})$$



$$\psi_f = e^{i \vec{k}' \cdot \vec{r}}$$

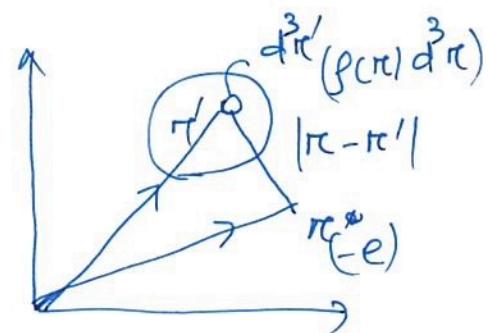
Say the ψ_i be the wavefunction of electron before scattering & it becomes ψ_f , the final wavefunction of electron after scattering.

Since the scattering is elastic scattering the magnitude of \vec{k} should be equal to the magnitude of \vec{k}' . The magnitude of kinetic energy remain same. only the direction changes.

What made the change of electron wavefunction ψ_i to ψ_f ? It is the Coulomb interaction !!. Recall the potential energy that we discussed earlier.

The interaction energy can be written as

$$H' = \int \frac{(\rho(r') dr')(-e)}{4\pi G_0 |r-r'|} \quad \text{all value.}$$



This interaction will change the electron wavefunction ψ_i to ψ_f . What is the probability to find it.

$$P \propto |\langle \psi_f | H' | \psi_i \rangle|^2$$

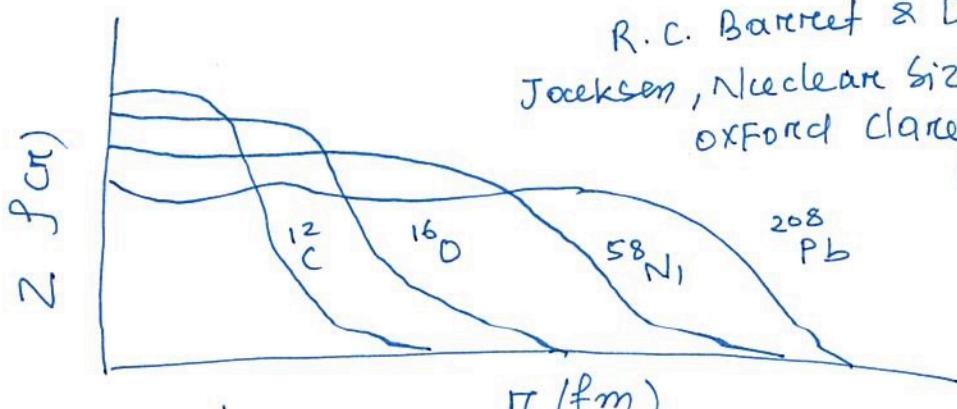
This is nothing but the integration over entire space

$$\begin{aligned}
 P &= \int \psi_f^*(\vec{r}) H' \psi_f(\vec{r}) d^3r \\
 &= \int e^{i(\vec{k}-\vec{k}')\cdot \vec{r}} H' d^3r \\
 &= \int e^{i(\vec{k}-\vec{k}')\cdot \vec{r}} d^3r \left[\int f(r') \frac{d^3r'(-e)}{4\pi r_0^3 (r-r')} \right]
 \end{aligned}$$

This is the probability how an electron scatters from nucleus. For this we should know the charge distribution, But this is what do not know. Thus we look into the experimental data, that it ~~has~~ experimentally measured the probability. Before we discuss about the experimental results, let's little bit visualise it. The electron scattering will only be the coulomb interaction, why because it does not interact strongly. Thus the experimental outcomes will mostly be the coulombic interaction.

At this stage the problem is reversed, the probabilities are available from the experiments. The question is although we know the total charges (Ze) but don't know the charge density $f(r)$.

R. C. Barrett & D. F. Jackson, Nuclear Size & Structure, Oxford Clarendon 1977.



Radial charge distribution of ^{12}C , ^{16}O , ^{58}Ni , ^{208}Pb

The above figure shows a more quantitative determination of the relationship between nuclear radius & total charge, based on electron scattering. What we observe that for different nuclei, for carbon, light nucleus the graph is starting higher & for heavier nuclei it is starting at lower charge density.

What does this means? For lighter nucleus more charge density at center, heavier nucleus less charge density (this is purely coming from the experimental data). And the radius upto which density is almost constant, is ~~is~~ smaller for lighter & larger for heavier nuclei.

By the time these experiments were done, neutrons were known. Since electrons ~~do~~ interact ~~only~~ with proton these curves are reflection of proton distribution or we call them charge density. But in same volume \Rightarrow there are presence of neutrons.

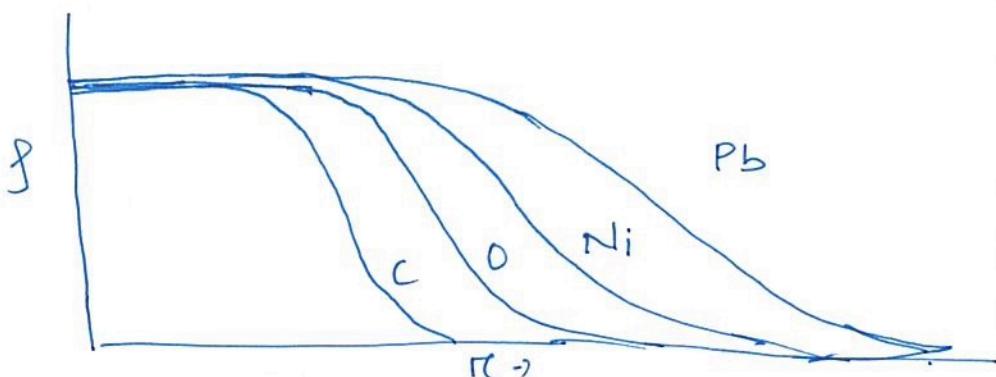
Thus, this distribution also depends how neutron ~~are~~ charge distributed. Let make an analogical explanation.

What is the mass ~~of~~ of nucleus. It is the sum of total proton & neutron. Or we can say the total number of nucleon. & we call it mass number A

Now we can derive a mass density f , which is

$$f = f_p \left(\frac{r}{2}\right)$$

And if we redo the curve with same data but the mass density. The curve looks.



This describes how mass is distribution across the nucleus. Can we find a function to fit this data?

The charge distribution should take a form.

$$f(r) = \frac{f_0}{1 + e^{(r-R)/\delta}}$$

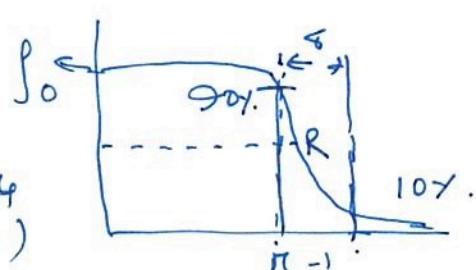
where f_0 is the matter density at center

$R \rightarrow$ nuclear radius (~~density falls to $\frac{1}{2}$ of f_0~~)

$\delta \rightarrow$ surface depth which

measures the range in r over which the charge density / or matter density changes from the order of its value at the center to the much smaller than this value.

This distribution is known as Wood-Saxon formula or Model.

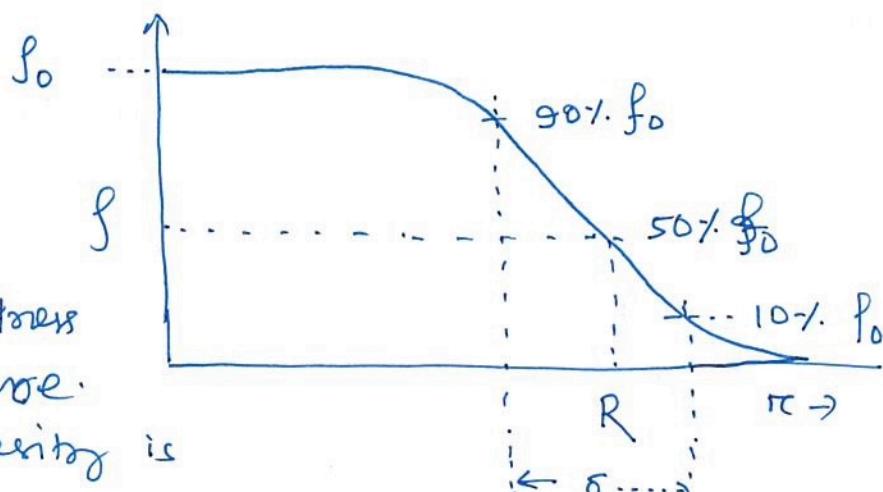


Let us analyze little bit to find various quantities.

R = average radius

R is ~~50%~~ of ~~50%~~

$\delta \rightarrow$ measure of stiffness
of density change.
i.e. how fast the density is
changing from 90% to 10% of its central value f_0



Experimentally, $f_0 = 0.17 \text{ fm}^3$

at 90% of f_0 :

$$0.9f_0 = \frac{f_0}{1 + e^{(r-R)/\delta}}$$

$$\Rightarrow \frac{1}{0.9} = 1 + e^{(r-R)/\delta}$$

$$\Rightarrow 0.1 = e^{(r-R)/\delta}$$

$$\Rightarrow \ln(0.1) = \ln[e^{(r-R)/\delta}]$$

$$\Rightarrow \frac{r-R}{\delta} = -2.303$$

$$\ln(0.1) = -2.303$$

$$\Rightarrow \boxed{r = R - 2.303\delta}$$

at 10% f_0 :

$$0.1f_0 = \frac{f_0}{1 + e^{(r-R)/\delta}}$$

$$\Rightarrow \frac{1}{0.1} = 1 + e^{(r-R)/\delta}$$

$$\ln$$

$$\ln 9 = 2.197$$

$$\Rightarrow 9 = e^{(r-R)/\delta} \Rightarrow \ln 9 = \ln[e^{(r-R)/\delta}] \Rightarrow r = R + 2.197\delta$$

After fitting various data, one can find out the value of δ :

i.e. for almost all nuclear range.

$$\boxed{\delta = 0.4 \rightarrow 0.5 \text{ fm}}.$$

Solving Woods Saxon $F(r)$, one can find out & incorporating experimental data.

$$r_c = r_{c0} A^{1/3}$$

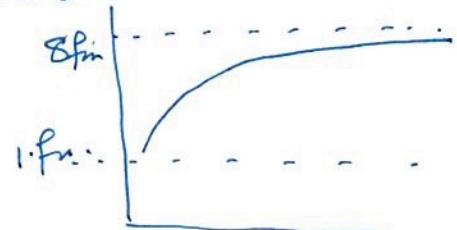
where

$$\boxed{r_{c0} = 1.18 \text{ fm}}$$

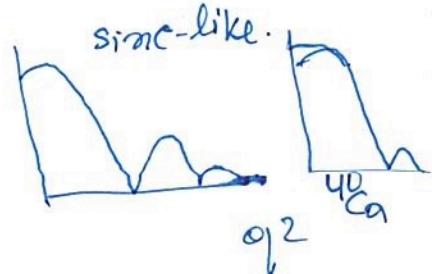
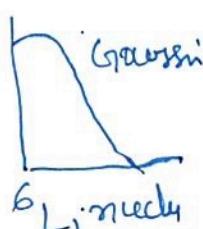
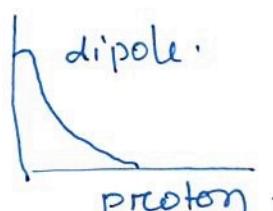
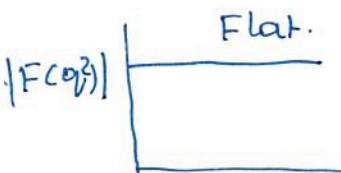
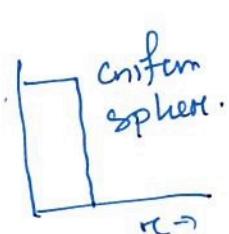
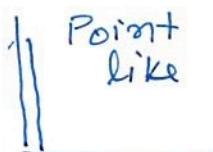
$$\boxed{s_0 = 0.17 \text{ nucleon/fm}^3}$$

If you try to draw the radius of nucleus vs the mass number you will get a result similar to

This gives interesting result that there is not much change in radius for higher mass numbers.



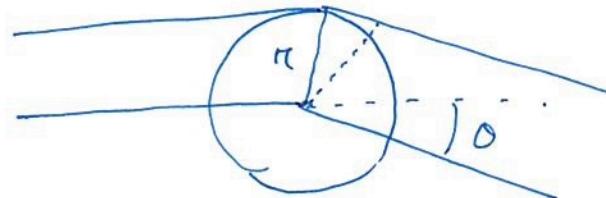
Experimentally what you will find that the skin thickness $\delta = 0.4 - 0.5 \text{ fm}$ is same for all nuclei connecting to $F(r) \propto |F(\alpha_f^2)|^2$



what we learn from the above exercise??

Qualitatively, the charge distribution is almost similar all across. And experiment precisely verified these calculations. And the diffraction pattern is shown in electron scattering of nucleus. The reason for this is that (reson for interference pattern), the part of wave front that passes through the nucleus at a distance r from the center and is scattered through an angle θ travels a further distance than the part of the wave that passes through the center, by an amount proportional to r and therefore suffers a phase change (relative to the part of the wave passing through the center)

This phase change also depends on the scattering angle θ and is equal to $\frac{qr}{\hbar}$



This means that different part of wave front suffer a different phase change (just as in optical diffraction). These different amplitudes are summed to get the total amplitude at some scattering angle θ and this gives rise to the diffraction pattern. The contribution to amplitude part from the part of wave front passed at a distance r from the center of the nucleus is proportional to charge density $f_p(r)$ at r or matter density $f(r)$ at r . The total scattering amplitude is therefore the sum of the amplitude from all these different parts, which is what the integral in some equation above

~~* Other ways of calculating the Radii of Nucleus.~~

One can use purely quantum mechanical approach to calculate the radius of nucleus using the similar analogy that of quantum mechanical treatment of Bohr's atomic structure. i.e. ~~using the~~ this you can find in details in some of the books.

Before going into such calculation let me briefly discuss about the Quantum theory of Hydrogen atom (we will not discuss this in class - and this is a supplementary material)

Hydrogen system is a system

of one proton, one electron & the electrostatic (Coulomb) potential that holds them together.

The potential energy in this case is

$$V = -\frac{e^2}{4\pi G_D \pi} \quad \left[\text{attractive potential between} \right]$$

Notice that it is a function π , not (x, y, z)

we can simply write $\pi^2 = x^2 + y^2 + z^2$

$$r = \sqrt{x^2 + y^2 + z^2}$$

By going (θ, ϕ) , one can go away with it, but things becomes more complicated. The appropriate approach is to let the ~~spherical~~ symmetry of the potential guide in solving the spherically symmetric potential to make use of spherical polar coordinates.

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Time independent Schrödinger equation

$$H|\vec{E}_i\rangle = E_i |\vec{E}_i\rangle$$

E_i are eigen values & $|\vec{E}_i\rangle$ are energy Eigenstates

By changing the notation $E_i \rightarrow E$ & $|\vec{E}_i\rangle \rightarrow \Psi$, the 3-dimensional Schrödinger equation can be written as

$$\left(-\frac{\hbar^2}{2m} \nabla^2 + V \right) \Psi = E \Psi$$

∇^2 is the Laplacian operator.

$$\frac{\partial^2 \Psi}{\partial x^2} + \frac{\partial^2 \Psi}{\partial y^2} + \frac{\partial^2 \Psi}{\partial z^2} + \frac{2m}{\hbar^2} (E - V) \Psi = 0$$

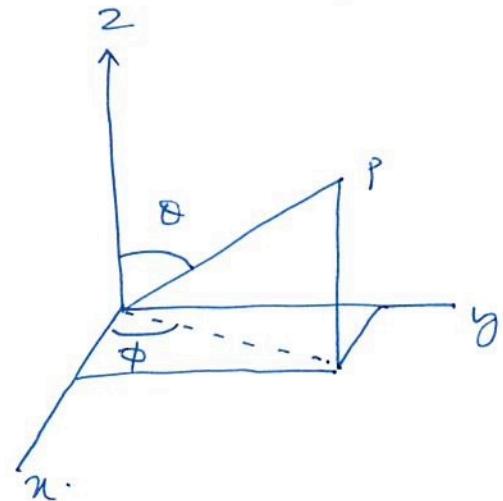
we can do a spherical polar co-ordinate transformation

$$\theta = \cos^{-1} \left(\frac{z}{\sqrt{x^2 + y^2 + z^2}} \right)$$

$$\phi = \tan^{-1} \left(\frac{y}{x} \right)$$

Hence the co-ordinate is (r, θ, ϕ)

and are related to (x, y, z) &
can be expressed as



$$x = r \sin \theta \cos \phi$$

$$y = r \sin \theta \sin \phi$$

$$z = r \cos \theta$$

so the Schrödinger's equation in spherical polar coordinates

$$\text{as } \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \Psi}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial \Psi}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \Psi}{\partial \phi^2} + \frac{2m}{\hbar^2} (E - V) \Psi = 0$$

plugging V & multiplying both side $\pi^2 \sin^2 \theta$

$$\sin^2 \theta \frac{\partial}{\partial r} \left(r^2 \frac{\partial \Psi}{\partial r} \right) + \sin \theta \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial \Psi}{\partial \theta} \right) + \frac{\partial^2 \Psi}{\partial \phi^2} + \frac{2m \pi^2 \sin^2 \theta}{\hbar^2} \left(\frac{e^2}{4\pi \epsilon_0 r} + E \right) \Psi = 0$$

This equation does not look nice.

This equation gives the wave function ψ for the electron in the hydrogen atom. That means if we solve for ψ in principle we know everything about the hydrogen atom. When we solve one dimensional ψ , we get one quantum number & for above one is three dimensional equation; we can get detail information about hydrogen atom.

The solution to this equation for electron in the hydrogen atom by separation of variables. This is usually done to ease the complexity of linear algebraic equations, differential equations, coupled differential equations etc.

When we have an equation like the one above, we like to see if we can "separate" the variable; i.e. "split" the equation into different parts, with only one variable in each part. Equation becomes simplified if we write.

$$\psi(\pi, \theta, \phi) = R(\pi) \Theta(\theta) \Phi(\phi) = R\Theta\Phi$$

The schrodinger equation will become:

$$\frac{\partial \psi}{\partial \pi} = \Theta\Phi \frac{\partial R}{\partial \pi} = \Theta\Phi \frac{dR}{d\pi}$$

$$\frac{\partial \psi}{\partial \theta} = R\Phi \frac{\partial \Theta}{\partial \theta} = R\Phi \frac{d\Theta}{d\theta}$$

$$\frac{\partial^2 \psi}{\partial \phi^2} = R\Theta \frac{\partial^2 \Phi}{\partial \phi^2} = R\Theta \frac{d^2 \Phi}{d\phi^2}$$

where the partial derivatives become full derivatives because R , Θ , and Φ depend on π , θ and ϕ only.

To separate the variables, plug $\psi = R\Theta\Phi$ into the Schrodinger's equation and divide by ~~$R\Theta\Phi$~~ $R\Theta\Phi$

The Schrödinger's equation will become

$$\frac{\sin^2\theta}{2} \frac{d}{dr} \left(r^2 \frac{dR}{dr} \right) + \frac{\sin\theta}{\theta} \frac{d}{d\theta} \left(\sin\theta \frac{d\Theta}{d\theta} \right) + \frac{1}{\Phi} \frac{d^2\Phi}{d\phi^2} + \frac{2m\epsilon^2 \sin^2\theta}{h^2} \left(\frac{e^2}{4\pi\epsilon_0 r} + E \right) = 0$$

Note that we have separated out the ϕ variable, in term

$$\frac{1}{\Phi} \frac{d^2\Phi}{d\phi^2}$$

is a function of ϕ only. Let's put it over on the right hand side of the equation. This gives us.

$$\frac{\sin^2\theta}{R} \frac{d}{dr} \left(r^2 \frac{dR}{dr} \right) + \frac{\sin\theta}{\Theta} \frac{d}{d\theta} \left(\sin\theta \frac{d\Theta}{d\theta} \right) + \frac{2m\epsilon^2 \sin^2\theta}{h^2} \left(\frac{e^2}{4\pi\epsilon_0 r} + E \right) = -\frac{1}{\Phi} \frac{d^2\Phi}{d\phi^2}$$

This equation looks like.

$$f(r, \theta) = g(\phi)$$

f is a function of r and θ only, and g is a function of ϕ only. If an r or θ shows up on the LHS, it can't be satisfied, because r and θ never show up on the RHS. Similarly if ϕ shows up on the RHS, the equation can't be satisfied because ϕ never shows up on the LHS. The only way for the equation to be satisfied is for

$$f(r, \theta) = \text{a constant, independent of } r, \theta, \text{ and } \phi = g(\phi)$$

we have taken a complex equation & using r , θ , and ϕ separation method separated it into two equations, one in r and θ , and the other in ϕ only.

It is a general practice that some differential equations can be solved only if certain conditions are satisfied. These conditions have led us to quantum numbers. In this context, we can use a similar concept for $-\frac{1}{\Phi} \frac{d^2\Phi}{d\phi^2}$

$$\Rightarrow \boxed{-\frac{1}{\Phi} \frac{d^2\Phi}{d\phi^2} = m_l^2}$$

Now The previous equation: we set the LHS equal to m_e^2 , divide by $\sin^2\theta$ and rearrange, we get,

$$\pi \frac{1}{R} \frac{d}{d\pi} \left(R^2 \frac{dR}{d\pi} \right) + \frac{2m_e^2}{\hbar^2} \left(\frac{e^2}{4\pi\epsilon_0 R} + E \right) = \frac{m_e^2}{\sin^2\theta} - \frac{1}{\Theta \sin\theta} \frac{d}{d\theta} \left(\sin\theta \frac{d\theta}{d\phi} \right)$$

Hence we again separated variables. The LHS is a function of π only, and the RHS is a function of θ only. Again, the only way to satisfy this equation is for LHS = a constn = RHS

Conditions on the constant will arise from the solution of the differential equations. this constant = $l(l+1)$

Now the equations can be written:

$$\frac{d^2\Phi}{d\phi^2} + m_l^2 \Phi = 0 \quad - (A)$$

$$\frac{1}{\sin\theta} \frac{d}{d\theta} \left(\sin\theta \frac{d\theta}{d\phi} \right) + \left[l(l+1) - \frac{m_l^2}{\sin^2\theta} \right] \theta = 0 \quad - (B)$$

$$\frac{1}{R} \frac{d}{d\pi} \left(R^2 \frac{dR}{d\pi} \right) + \left[\frac{2m}{\hbar^2} \left(\frac{e^2}{4\pi\epsilon_0 R} + E \right) - \frac{l(l+1)}{\pi^2} \right] R = 0 \quad - (C)$$

At this stage we manage to make it ~~more~~ lot more simpler because the "big" partial differential equation into three simpler ones, each of which is a function of a single variable and each of which is much "easier" to handle. At this stage the Schrodinger's equation for the hydrogen atom into three separate equations, one for each variable, R, θ, ϕ .

But we need to now solve three equation instead of one equation. But this is lot more easier than were starting equation. let's quickly see the solutions-

we write down

Quantum numbers are constants which identify each solution to Schrödinger's equation.

First quantum number: comes by solving the differential eqn for Φ

$$\frac{d^2\Phi}{d\theta^2} + m_l^2 \Phi = 0$$

The general solution to above equation will be

$$\Phi(\theta) = A e^{im_l \theta}$$

Now we know that θ and $\theta + 2\pi$ represent a single point in space, we must have

$$A e^{im_l \theta} = A e^{im_l (\theta + 2\pi)}$$

This happens only for $m_l = 0, \pm 1, \pm 2, \pm 3, \dots$

m_l is called the magnetic quantum number.

Solving the differential equation Θ will give ~~two~~ 2nd quantum number

$$\frac{1}{\sin\theta} \frac{d}{d\theta} \left(\sin\theta \frac{d\Theta}{d\theta} \right) + \left[l(l+1) - \frac{m_l^2}{\sin^2\theta} \right] \Theta = 0$$

it solves the term $l(l+1) - \frac{m_l^2}{\sin^2\theta}$

it turns out that from differential equations that the equation for Θ can be solved only if l is an integer greater than equal to the absolute value of m_l .

l is called orbital quantum number & the requirement on l can be restated as $m_l = 0, \pm 1, \pm 2, \dots, \pm l$.

The radial differential equation is

$$\frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{dR}{dr} \right) + \left\{ \left[\frac{2m}{\hbar^2} \left(\frac{e^2}{4\pi\epsilon_0 r} + E \right) \right] - \frac{l(l+1)}{r^2} \right\} R = 0$$

It can be solved only for energies E which satisfy the same condition as we found on the energies for

for the Bohr atom:

$$E_n = -\frac{me^4}{32\pi^2\epsilon_0\hbar^2n^2} = \frac{E_1}{n^2} \quad n = 1, 2, 3 \dots$$

n is known as the principal quantum number.

By summarizing all above we can write the solution of Schrödinger equation for the hydrogen atom must be of the form.

$$\Psi = R_{nl} \Theta_{lm_l} \Phi_{m_l}$$

where n, l, m_l being the quantum numbers discussed above. If we write it concisely.

First coordinate . second coordinate	$\Phi(\phi) = e^{im\phi} \quad m \Rightarrow \text{integer}$ $\frac{1}{sm\theta} \frac{\partial}{\partial \theta} sm\theta \frac{\partial \Phi}{\partial \theta} + (\lambda - \frac{m^2}{sm\theta}) \Phi = 0$ $\lambda = l(l+1) \text{ with } l = 0, 1, 2 \dots$ $m = -l, -l+1 \dots l-1, l$	Angular part .
---	---	----------------

Third coordinate .

$\frac{1}{R} \left[\frac{d}{dr} r^2 \frac{dR}{dr} + \frac{2mr^2}{\hbar^2} (E - V(r)) R \right] = l(l+1)$ $E_n = -\frac{Z^2}{n^2} R_\infty = -\frac{Z^2}{n^2} \left(\frac{e^2}{4\pi\epsilon_0} \right) \frac{me}{2\hbar^2}$ with $n > 0$	radial part .
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Angular wave function

$$L^2 = \frac{\hbar^2}{i} \left[\frac{1}{sm\theta} \frac{\partial}{\partial \theta} sm\theta \frac{\partial}{\partial \theta} + \frac{1}{sm\theta} \frac{\partial^2}{\partial \phi^2} \right].$$

$$L_z = \frac{\hbar}{i} \frac{\partial}{\partial \phi} \quad \vec{L} = (L_x, L_y, L_z)$$

Angular momentum.

There is a class of functions that are simultaneous eigenfunctions.

$$L^2 Y_{lm}(\theta, \phi) = l(l+1) \hbar^2 Y_{lm}(\theta, \phi) \quad \left. \begin{array}{l} l=0, 1, 2 \dots \\ m=-l, -l+1 \dots l \end{array} \right\}$$

$$L_z Y_{lm}(\theta, \phi) = m \hbar Y_{lm}(\theta, \phi) \quad \left. \begin{array}{l} m=-l, -l+1 \dots l \end{array} \right\}$$

Spherical harmonics $Y_{lm}(\theta, \phi)$

$$Y_{00} = \sqrt{\frac{1}{4\pi}} \quad Y_{11} = \sqrt{\frac{3}{8\pi}} \sin \theta e^{i\phi}$$

$$Y_{10} = \sqrt{\frac{3}{4\pi}} \cos \theta \quad Y_{1,-1} = \sqrt{\frac{3}{8\pi}} \sin \theta e^{-i\phi}$$

The radial part.

$$\frac{1}{R} \left[\frac{d}{dr} r^2 \frac{dR}{dr} + \frac{2mr^2}{\hbar^2} (E - V(r)) R \right] = \lambda \quad \lambda = l(l+1)$$

For ~~one~~ solution to ground state i.e. $l=0, m=0$.

$$-\frac{\hbar^2}{2m} \left(R'' + \frac{2}{r} R' \right) - \frac{Ze^2}{4\pi\epsilon_0 r} R = E R$$

$r \rightarrow \infty \Rightarrow$ no density.

$$R(r) = A e^{-r/a}$$

$$R' = -\frac{A}{a} e^{-r/a} = -\frac{R}{a}$$

$$R'' = \frac{A}{a^2} e^{-r/a} = \frac{R}{a^2}$$

$$\Rightarrow -\frac{\hbar^2}{2m} \left(\frac{1}{a^2} - \frac{2}{ar} \right) - \frac{Ze^2}{4\pi\epsilon_0 r} = E$$

Solution for the energy.

$$E = -\frac{\hbar^2}{2ma} = -Z^2 \left(\frac{e^2}{4\pi\epsilon_0} \right)^2 \frac{me}{2\hbar^2}$$

Ground state in the Bohr's model $n=1$.

Prefactor for $1/r$

$$\frac{\hbar^2}{ma} - \frac{Ze^2}{4\pi\epsilon_0} = 0$$

Solution for the length scale parameter

$$a_0 = \frac{4\pi\epsilon_0 \hbar^2}{Ze^2 m} \text{ Bohr radius.}$$

Thus the ground state is $n=1, l=0, m=0$, i.e. there is only one quantum state with the wave function for $n=1$ and its Ψ_{100} ; However for $n=2$ we have

$$l=0, m=0$$

$$l=1 \quad m=-1, 0, 1.$$

Thus the allowed state for $n=2$, the states are

$\Psi_{200}, \Psi_{21-1}, \Psi_{210}, \Psi_{211}$: Thus the wave functions are:

$$\Psi_{100} = \frac{1}{\sqrt{\pi}} \frac{1}{a_0^{3/2}} e^{-\pi/a_0} \quad n=1, l=0, m_l=0$$

$$\Psi_{200} = \frac{1}{4\sqrt{2}\pi} \frac{1}{a_0^{3/2}} \left(2 - \frac{\pi}{a_0}\right) e^{-\pi/2a_0} \quad n=2, l=0, m_l=0$$

$$\Psi_{21-1} = \frac{1}{8\sqrt{\pi}} \frac{1}{a_0^{3/2}} \frac{\pi}{a_0} e^{-\pi/2a_0} \sin\theta e^{-i\phi} \quad n=2, l=1, m_l=-1$$

$$\Psi_{210} = \frac{1}{4\sqrt{2}\pi} \frac{1}{a_0^{3/2}} \frac{\pi}{a_0} e^{-\pi/2a_0} \cos\theta \quad n=2, l=1, m_l=0$$

$$\Psi_{211} = \frac{1}{8\sqrt{\pi}} \frac{1}{a_0^{3/2}} \frac{\pi}{a_0} e^{-\pi/2a_0} \sin\theta e^{i\phi} \quad n=2, l=1, m_l=1.$$

usually Ψ_{100} the corresponding orbital name are

$$\Psi_{100} \rightarrow \Psi_{1s}$$

$$\Psi_{200} \rightarrow \Psi_{2s}$$

$$\Psi_{210} \rightarrow \Psi_{2p_z}$$

$$\Psi_{21-1} \rightarrow \Psi_{2p_y}$$

$$\Psi_{211} \rightarrow \Psi_{2p_x}$$

1

Understanding the nuclear volume using hydrogen or hydrogen like nucleus (one electron in orbit)

The potential energy for such a system will be

$$V(r) = -\frac{Ze^2}{4\pi\epsilon_0 r} \quad \left. \begin{array}{l} \text{Note here} \\ \text{this is a} \\ \text{hydrogen like} \\ \text{so } z \neq 1 \end{array} \right\}$$


We need to solve the Schrödinger eq wave function

$$H\Psi = E\Psi$$

$$H = -\frac{\hbar^2}{2m} \nabla^2 - \frac{Ze^2}{4\pi\epsilon_0 r}$$

The ground state of this hydrogen like nuclei.

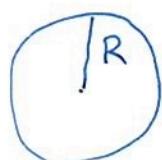
$$\Psi_{100} = \sqrt{\frac{1}{\pi} \left(\frac{z}{a_0} \right)^3} e^{-Zr/a_0} \quad a_0 \leftarrow Bohr's \text{ radius}$$

This wave function doesn't go to zero if you bring it at $r=0$. That means there is a probability of finding the electron at the side of the nucleus. Thus if the ~~stree~~ electron is penetrating the nucleus, the structure of nucleus will affect it. Thus this electron is going after that energy states also.

For point charge the potential

$$V = -\frac{Ze^2}{4\pi\epsilon_0 r} \quad \text{if } r > R$$

For the extended charge distribution in sphere the potential:



$$r > R$$

$$V(r) = \frac{Ze}{4\pi\epsilon_0 r}$$

$$r = R \Rightarrow V(r) = \frac{Ze}{4\pi\epsilon_0 R}$$

What happens to the potential inside surface?

i.e. $r < R$.

$$\Rightarrow V(r) - V(R) = \int_R^r \vec{E} \cdot d\vec{r}$$

Potential at r & potential at R , the potential difference is $\int \vec{E} \cdot d\vec{r}$.

$$\vec{E}(r) = \frac{Ze}{4\pi\epsilon_0 r^2} \left(\frac{R^2 + r^2}{R^3} \right)$$

$$\Rightarrow V(r) - V(R) = \int_R^r \frac{Ze}{4\pi\epsilon_0} \frac{r^2}{R^3} dr$$

$$= -\frac{Ze}{4\pi\epsilon_0 R^3} \left[\frac{r^2 - R^2}{2} \right]$$

$$r \cdot dr = r \cdot r^2 dr = dR$$

$$\Rightarrow V(r) = V(R) - \frac{Ze}{4\pi\epsilon_0 R^3} \left[\frac{r^2 - R^2}{2} \right]$$

$$= \frac{Ze}{4\pi\epsilon_0 R} - \frac{Ze}{4\pi\epsilon_0 R^3} \left(\frac{r^2 - R^2}{2} \right)$$

$$= \frac{Ze}{4\pi\epsilon_0 R} \left[1 - \frac{r^2 - R^2}{2R^2} \right] = \frac{Ze}{4\pi\epsilon_0 R} \left[1 - \frac{1}{2} - \frac{r^2}{2R^2} \right]$$

$$\Rightarrow \boxed{V(r) = \frac{Ze}{4\pi\epsilon_0 R} \left[\frac{3}{2} - \frac{r^2}{2R^2} \right]}$$

Extended charge sphere.

If we find the potential difference from uniform potential charge distributed sphere to a point charge will be.

$$\Delta V = \left[\frac{Ze}{8\pi\epsilon_0 R} \left(3 - \frac{r^2}{R^2} \right) \right] - \left(\frac{Ze}{4\pi\epsilon_0 r} \right)$$

the change in energy

$$\Delta U = + \frac{Ze^2}{8\pi\epsilon_0 R} \Delta V (-e)$$

$$= - \frac{Ze^2}{8\pi\epsilon_0 R} \left(3 - \frac{r^2}{R^2} \right) - \left(\frac{-Ze^2}{4\pi\epsilon_0 r} \right)$$

If the potential energy change by an amount of ΔU to bring point nucleus to a extended nucleus, the hamiltonian will also change $H' = H_0 + H_1$

H_0 = part. like

H_1 = change ΔU

Thus the Schrodinger equation

$$\text{is } (H_0 + H_1) \Psi = (E + \Delta E) \Psi$$

For schrodinger technique provides a methods to find the solution for such a system through the perturbation treatment provided ΔE is very small. This done by knowing the solution for $H_0 \Psi = E \Psi$ & applying that to whole equation. Thus.

$$\langle \Delta E \rangle = \langle \Psi_0 | H' | \Psi_0 \rangle \quad \Psi_0 = \text{solution for } H_0$$

$$= \int \Psi_0^* H' \Psi_0 \frac{d^3r}{4\pi} d\zeta$$

$$\langle \Delta E \rangle = \int \Psi_0^* \left[\frac{ze^2}{4\pi\epsilon_0 r} - \frac{ze^2}{8\pi\epsilon_0 R} \left(3 - \frac{r^2}{R^2} \right) \right] \Psi_0 \frac{d^3r}{4\pi} d\zeta$$

Let's calculate the wave function for the ground state.

$$\Psi_{100} = \Psi_{1s} = \sqrt{\frac{2^3}{\pi a_0^3}} e^{-2\pi/a_0}$$

$$\Delta E = \int \sqrt{\frac{2^3}{\pi a_0^3}} e^{-2\pi/a_0} \left[\frac{ze^2}{4\pi\epsilon_0 r} - \frac{ze^2}{8\pi\epsilon_0 R} \left(3 - \frac{r^2}{R^2} \right) \right] \sqrt{\frac{2^3}{\pi a_0^3}} e^{-2\pi/a_0} \frac{d^3r}{4\pi} d\zeta$$

$$d\zeta = r^2 \sin\theta \, dr \, d\theta \, d\phi$$

$$\Rightarrow \Delta E = \int \frac{2^3}{4\pi a_0^3} e^{-2\pi/a_0} \left[\frac{ze^2}{4\pi\epsilon_0 r} - \frac{ze^2}{8\pi\epsilon_0 R} \left(3 - \frac{r^2}{R^2} \right) \right] r^2 \sin\theta \, dr \, d\theta \, d\phi$$

since the potential difference is calculated from $r = 0$ to R on the outside of the sphere H_1 will become zero. Thus the integral of H will be from 0 to R .

$$\Delta E = \frac{z^3}{\pi a_0^3} \cdot \frac{ze^2}{4\pi G_0} \int_0^R \left[\frac{1}{r} - \frac{1}{2R} \left(3 - \frac{\pi^2}{R^2} \right) \right] r^2 dr \int_0^{2\pi} d\theta \int_0^\pi d\phi \cancel{\frac{2\pi}{r}} \cancel{\frac{2\pi}{r}} \rightarrow 2\pi \cdot 2\pi$$

$$\Delta E = \frac{z^4 e^2}{\pi a_0^3 \epsilon_0} \int_0^R \left(\pi - \frac{3\pi^2}{2R} + \frac{\pi^4}{2R^3} \right) e^{-2\pi/a_0} dr.$$

we could use the approximation for $e^{-2\pi/a_0}$ & its limit to solve or simplify this integrator or one can do proper integration without putting $e^{-2\pi/a_0}$ as the limiting case to make ϕ it disappear

$$e^{-2\pi/a_0} \rightarrow r \sim 10^{-15} \text{ m}$$

$$\Rightarrow e^{-2 \times 10^{-5}} \quad a_0 \sim 10^{-10} \text{ m.}$$

$$\Rightarrow \Delta E = \frac{z^4 e^2}{\pi G_0 a_0^3} \int_0^R \left(\pi - \frac{3\pi^2}{2R} + \frac{\pi^4}{2R^3} \right) dr$$

$$= \frac{z^4 e^2}{\pi G_0 a_0^3} \left[\int_0^R \pi dr - \frac{1}{2} \left[\frac{3\pi^2}{2} R + \frac{\pi^4}{2R} \right] \right]$$

$$= \frac{z^4 e^2}{\pi G_0 a_0^3} \left[\frac{R^2}{2} - \frac{3}{2} \frac{R^3}{3} + \frac{R^5}{2R^3 \cdot 5} \right]$$

$$= \frac{z^4 e^2 R^2}{10\pi G_0 a_0^3} \quad \begin{cases} \text{Mathematica solution to} \\ \int_0^R \left(\frac{1}{r} - \frac{1}{2R} \left(3 - \frac{\pi^2}{R^2} \right) \right) r^2 dr \\ = \frac{R^2}{10} \end{cases}$$

If you do not want use the vanishing of $e^{-2\pi/a_0}$ then I give the simple equation & its mathematical solution

$$\begin{aligned}
 & \int_0^R \left[\frac{\pi^4}{2R^3} - \frac{3\pi^2}{2R} + \pi \right] e^{-2\pi/a_0} d\pi \\
 &= \int_0^R \left[\frac{\pi^4 e^{-2\pi/a_0}}{2R^3} - \frac{3\pi^2 e^{-2\pi/a_0}}{2R} + \pi e^{-2\pi/a_0} \right] d\pi \\
 &= \left[-\frac{3a_0^5}{8R^3} e^{-2\pi/a_0} - \frac{3a_0^4 \pi}{4R^3} e^{-2\pi/a_0} - \frac{3a_0^3 \pi^2}{4R^3} e^{-2\pi/a_0} + \frac{3a_0^3}{8R} e^{-2\pi/a_0} \right. \\
 &\quad - \frac{a^2 \pi^3}{2R^3} e^{-2\pi/a_0} + \frac{3a^2 \pi}{4R} e^{-2\pi/a_0} - \frac{1}{4} a^2 e^{-2\pi/a_0} - \frac{a\pi^4}{4R^3} e^{-2\pi/a_0} \\
 &\quad \left. + \frac{3a\pi^2}{4R} e^{-2\pi/a_0} - \frac{1}{2} a\pi e^{-2\pi/a_0} \right]_0^R \\
 &= \left[-\frac{1}{8R^3} a e^{-2\pi/a} \left[3a^4 + 6a^3 \pi + a^2 (6\pi^2 - 3R^2) + 2a (2\pi^3 - 3\pi R^2 + R^3) \right. \right. \\
 &\quad \left. \left. + 2\pi (\pi - R)^2 (\pi + 2R) \right] \right]_0^R \\
 &= \frac{1}{8R^3} \left[a^2 (3a^3 - 3a_0 R^2 - 3a_0 e^{-2R/a_0} (a_0 + R)^2 + 2R^3) \right]
 \end{aligned}$$

In any case at the end of the day with some approximate

$$\boxed{\Delta E = \frac{Z^4 e^2 R^2}{10 \pi \epsilon_0 a_0^3}}$$

At this stage we will take analogy from the atomic structure like taking the difference of energy of various energy states like E_{1s} & E_{2sp}

$2p$ electron goes to zero at $r =$

$$r = 0$$



So what is the picture for the nuclear system (extended charge sphere)? this we can do with an analogous to the energy gap & when electron comes from $2s2p$ to $1s$ it emits some rays (x-ray) with fix energy.

$2s2p \downarrow$
 $1s \downarrow$

→ This called K_{α} -x-ray

What we found by now is

$$E_{1s} = E_{1s}^{\text{Point}} + \Delta E$$

$$E_{1s} = E_{1s}^{\text{Point}} + \frac{z^4 e^2 R^2}{10 \pi \epsilon_0 q_0^3}$$

If nucleus were a point particle, the expression for K_{α} -x-ray

$$E_{K_{\alpha}} = E_{2p} - E_{1s}$$

$$= E_{2p} - E_{1s}^{\text{Point}} - \frac{z^4 e^2 R^2}{10 \pi \epsilon_0 q_0^3}$$

$2p$ state exists
at $r=0$ i.e.
we can say
 $E_{2p} = E_{2p}^{\text{Point}}$

let's find an elements with same protons but several neutrons (isotopes). Then find Energy of K_{α} x-ray.

$$E_{K_{\alpha}}(A') = E_{2p} - E_{1s}^{\text{Point}} - \frac{z^4 e^2 R'^2}{10 \pi \epsilon_0 q_0^3} \quad A' \text{ is}$$

A' is isotope for A

$$E_{K_{\alpha}}(A) = E_{2p} - E_{1s}^{\text{Point}} - \frac{z^4 e^2 R^2}{10 \pi \epsilon_0 q_0^3}$$

E_{2p} does not change for A' to A as energy is not changed for E_{2p} .

E_{1s}^{Point} also does not change for A' to A as

it is a point charge like state.

only change is in last term $\frac{z^4 e^2 R'^2}{10 \pi \epsilon_0 q_0^3} \rightarrow \frac{z^4 e^2 R^2}{10 \pi \epsilon_0 q_0^3}$

Now subtract these two energies. $E_{K\alpha}^{(A')} - E_{K\alpha}(A)$

$$(\Delta E_{K\alpha}) = [E_{K\alpha}(A) - E_{K\alpha}(A')]$$

This is a measurable quantity as there are x-rays & you put proper detectors to count it & find the energy.

$$(\Delta E_{K\alpha}) = \frac{Z^4 e^2}{10 \pi G_0 a_0^3} (R^2 - R'^2)$$

we do not need the point structure of nucleus. At this stage we do not know R' & R , but we know Z & A . Now using $R = R_0 A^{1/3}$ from the simple volumetric calculation

$$\Delta E_{K\alpha} = \frac{Z^4 e^2}{10 \pi G_0 a_0^3} \left[R_0^2 (A'^{1/3})^2 - R_0^2 (A^{1/3})^2 \right]$$

$$\Delta E_{K\alpha} = \frac{Z^4 e^2}{10 \pi G_0 a_0^3} R_0^2 \left[A'^{2/3} - A^{2/3} \right]$$

If we measure several x-rays $\Delta E_{K\alpha}$ & try to draw a line, it will be a straight line with $\left(\frac{Z^4 e^2}{10 \pi G_0 a_0^3} R_0^2 \right)$ as the slope of equation

Now fitting the curve one can find out what should be R_0 .

The straight line goes as function of $A^{2/3}$ of various isotopes of some Z

This can also be done nicely with muonic x-ray which is beyond the scope of our course.

