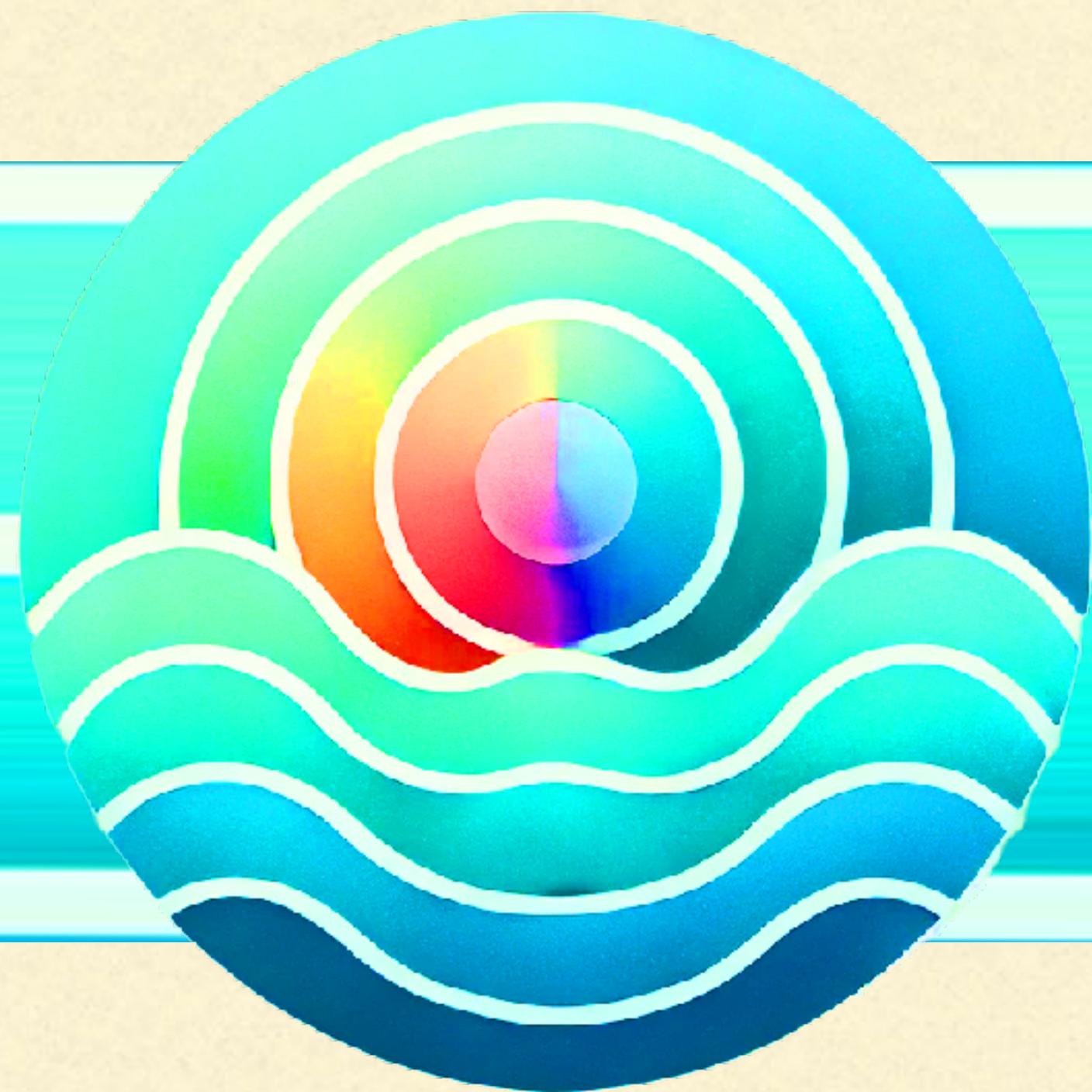


Break

- Thank you for not using your cell phones during the lectures.
- **Brain rot:** A condition of mental foginess, lethargy, reduced attention span, and cognitive decline that results from an overabundance of screen time.



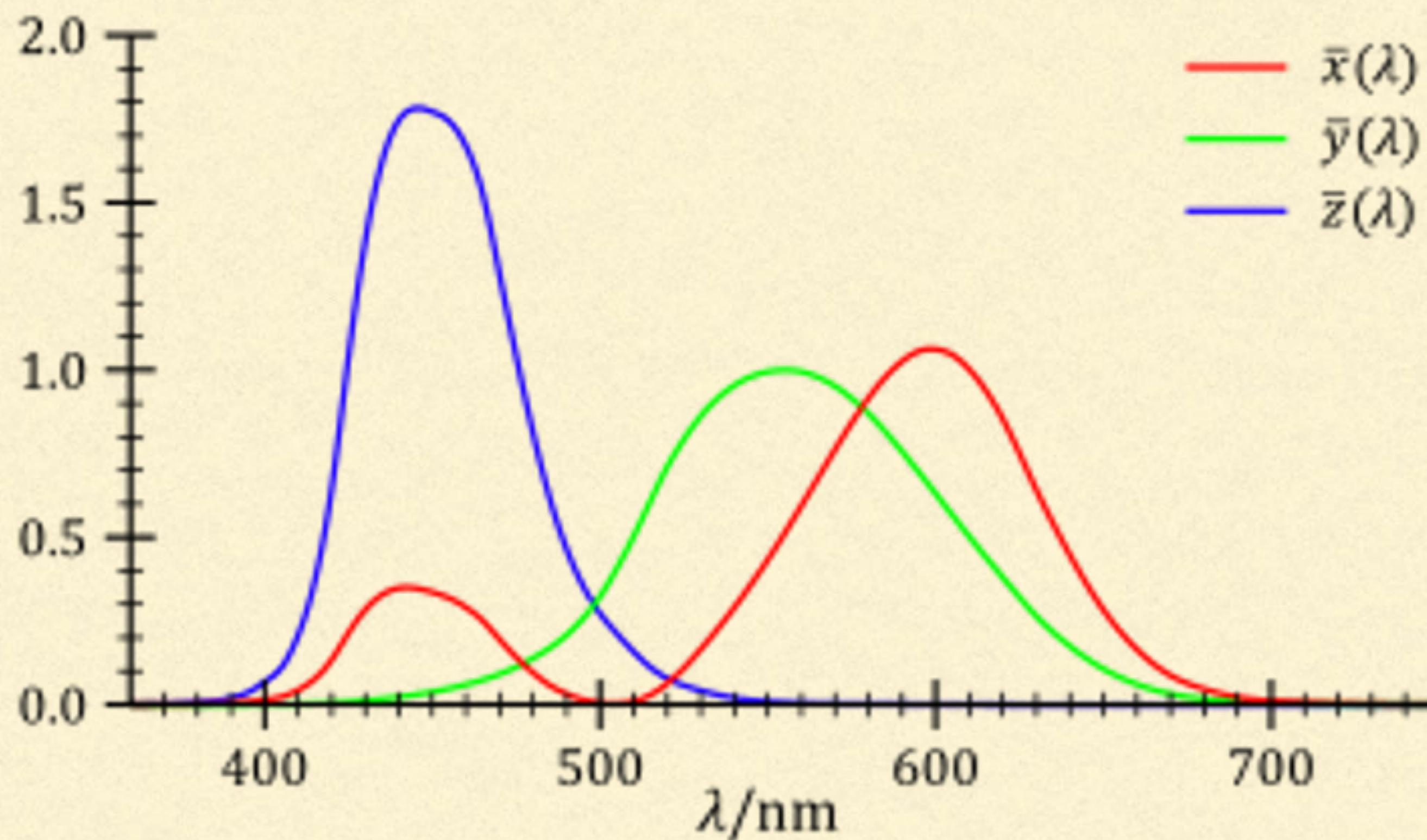


UNDERWATER COLORIMETRY

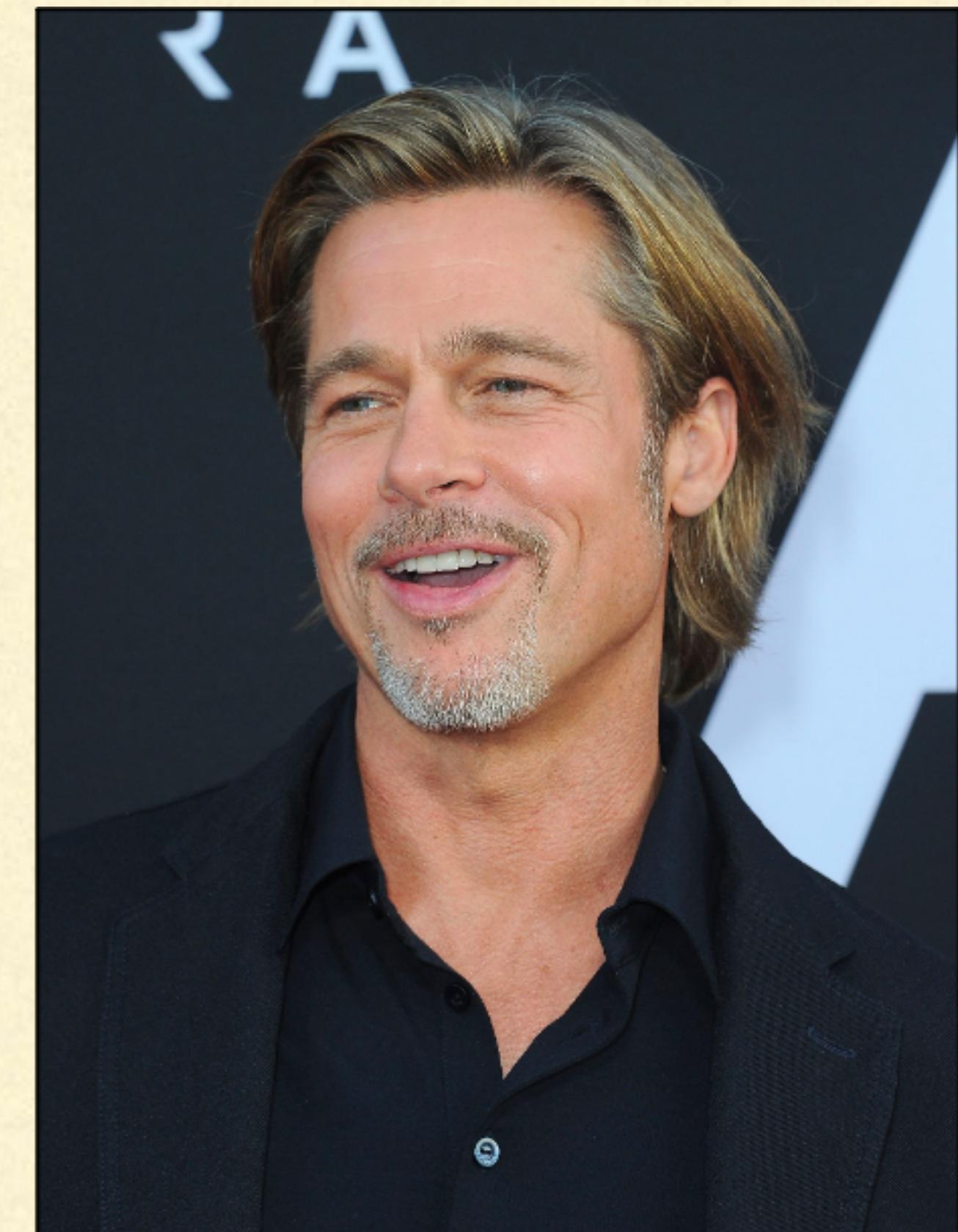
BASIC COLORIMETRY - II

Dr. Derya Akkaynak | dakkaynak@univ.haifa.ac.il

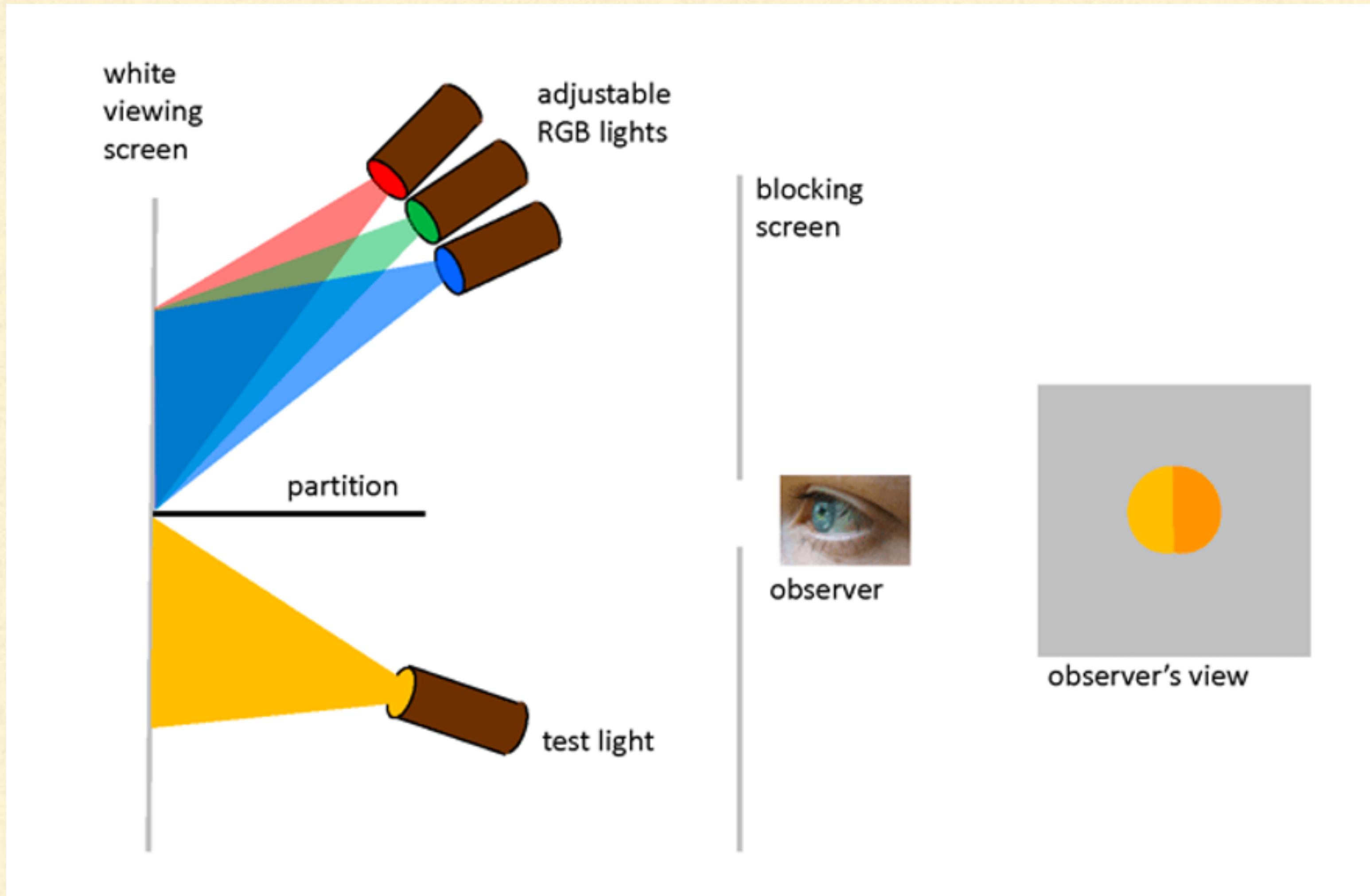
CIE 1931 STANDARD OBSERVER



Spectral sensitivity of a human with “normal” color vision



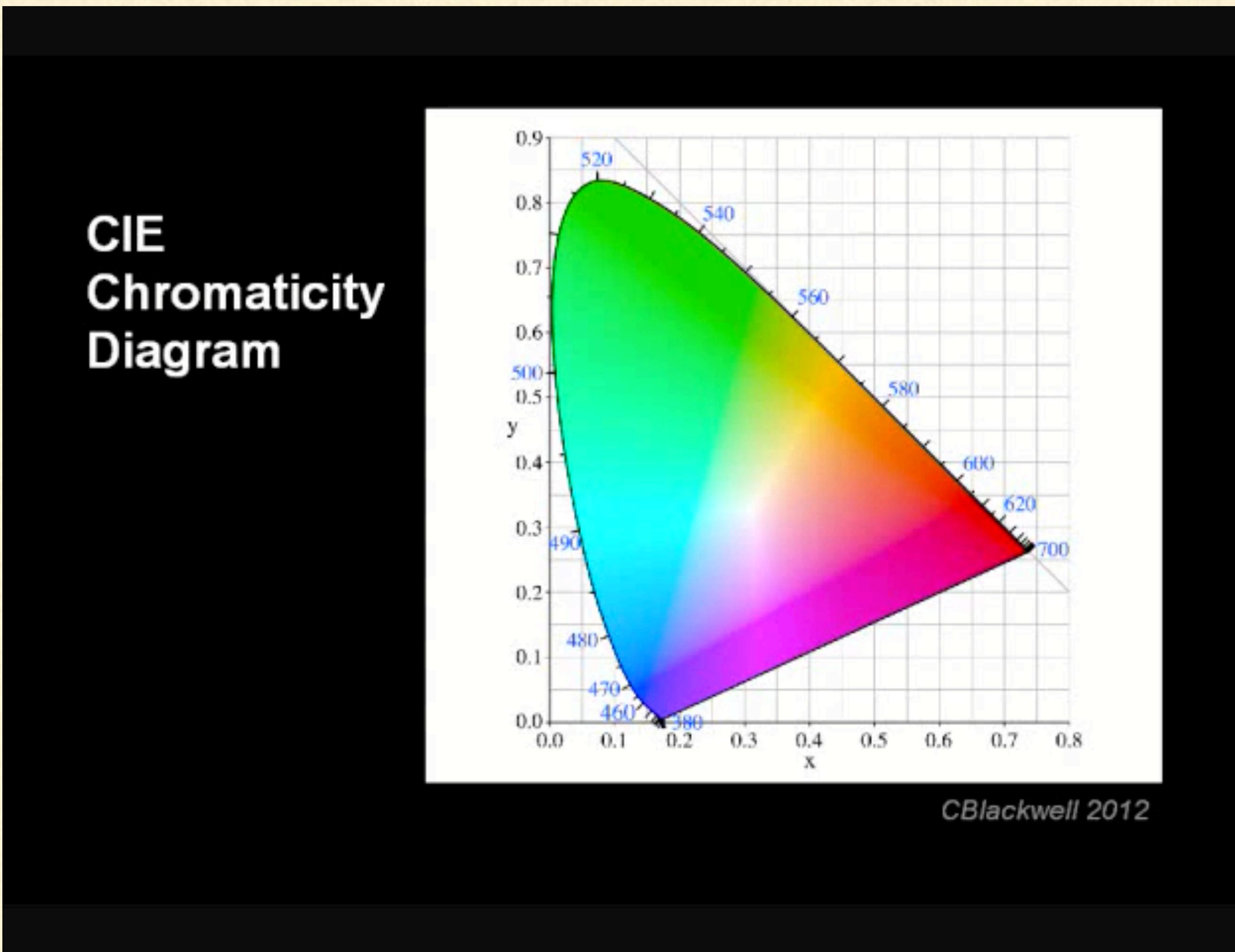
COLOR MATCHING EXPERIMENTS



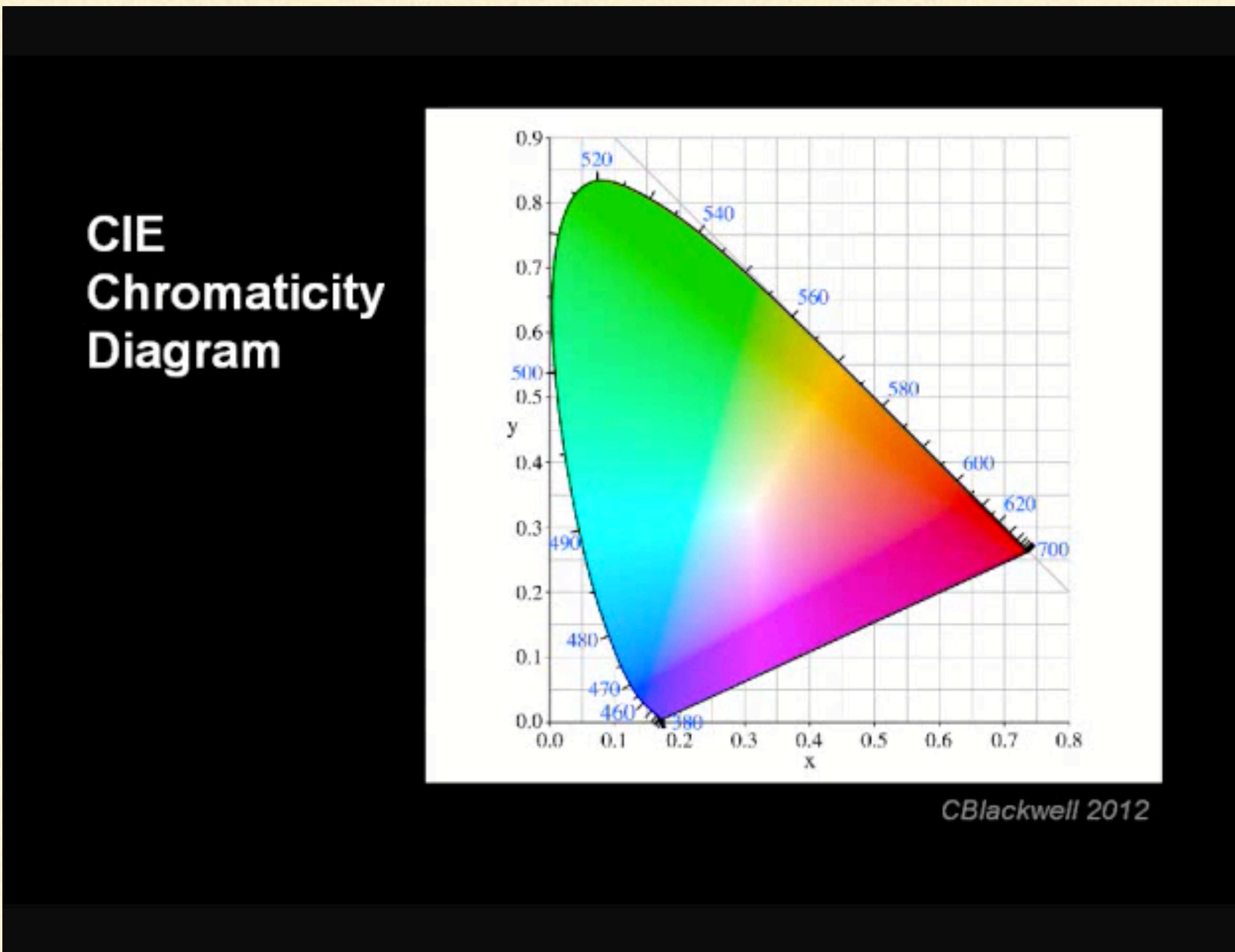
Try for yourself!

<https://rgbcmyk.com.ar/1931/>

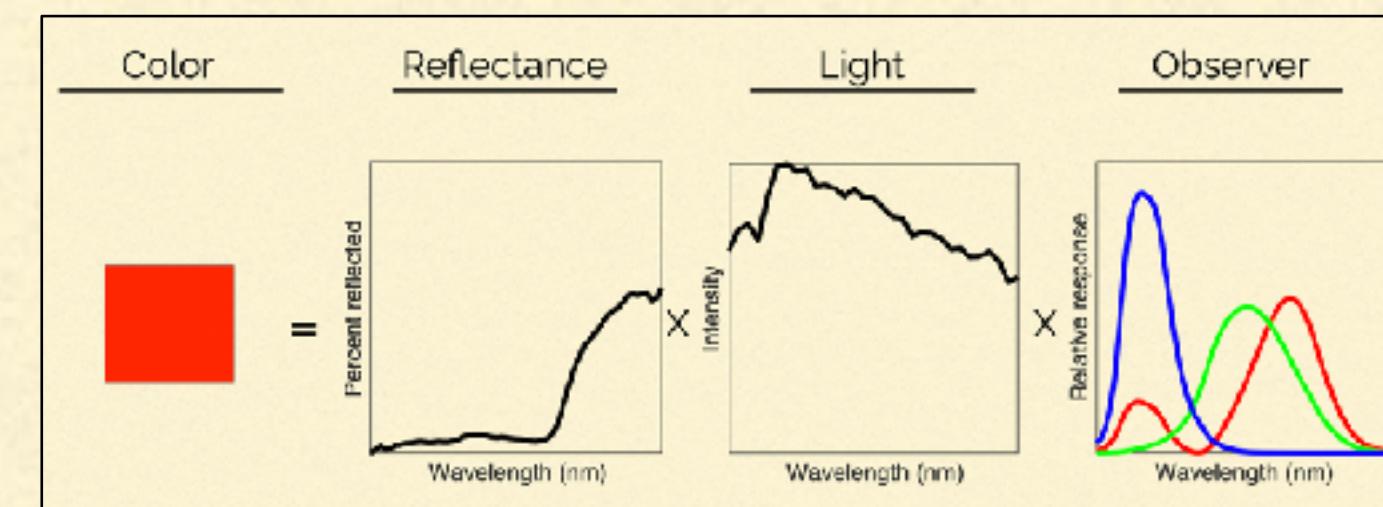
HOW WAS THE CIE 1931 STANDARD OBSERVER DERIVED?



HOW WAS THE CIE 1931 STANDARD OBSERVER DERIVED?

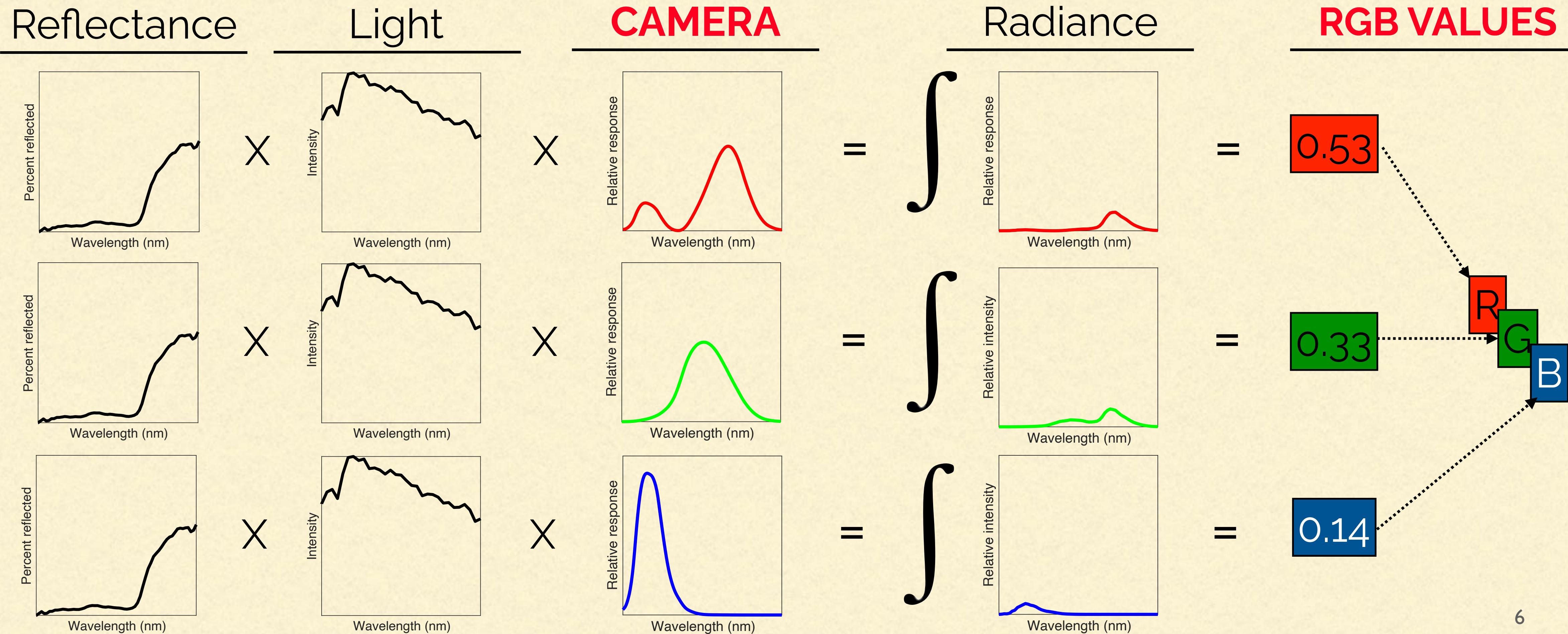


RGB Values

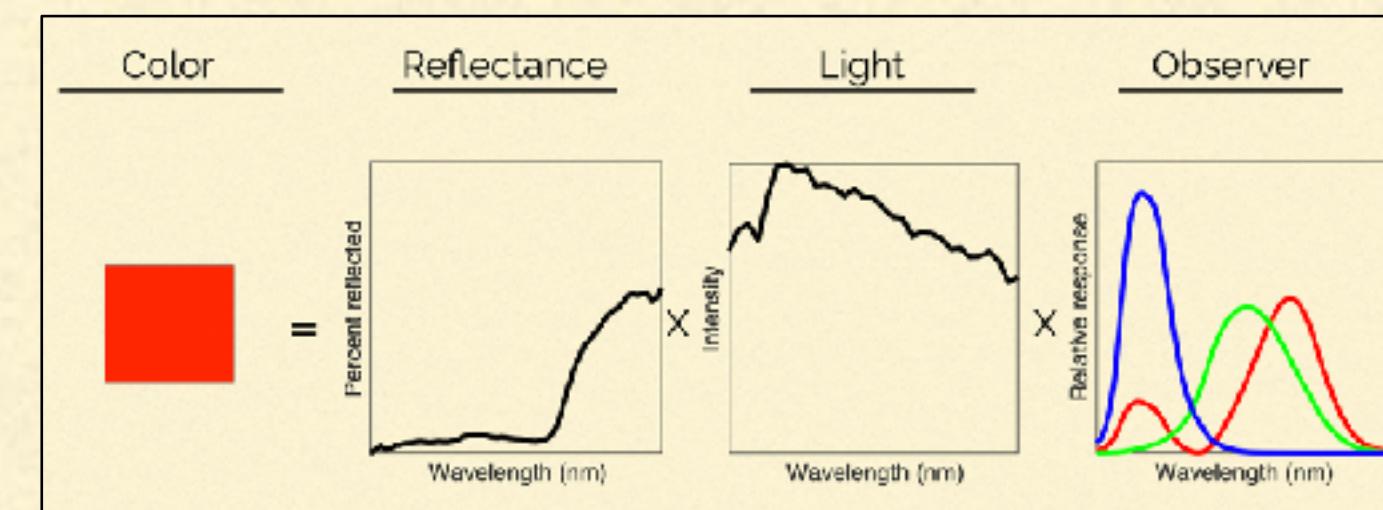


$$Color = \frac{1}{\kappa} \int_{\lambda_1}^{\lambda_2} \rho(\lambda) E(\lambda) S(\lambda) d\lambda$$

κ : exposure-related constant

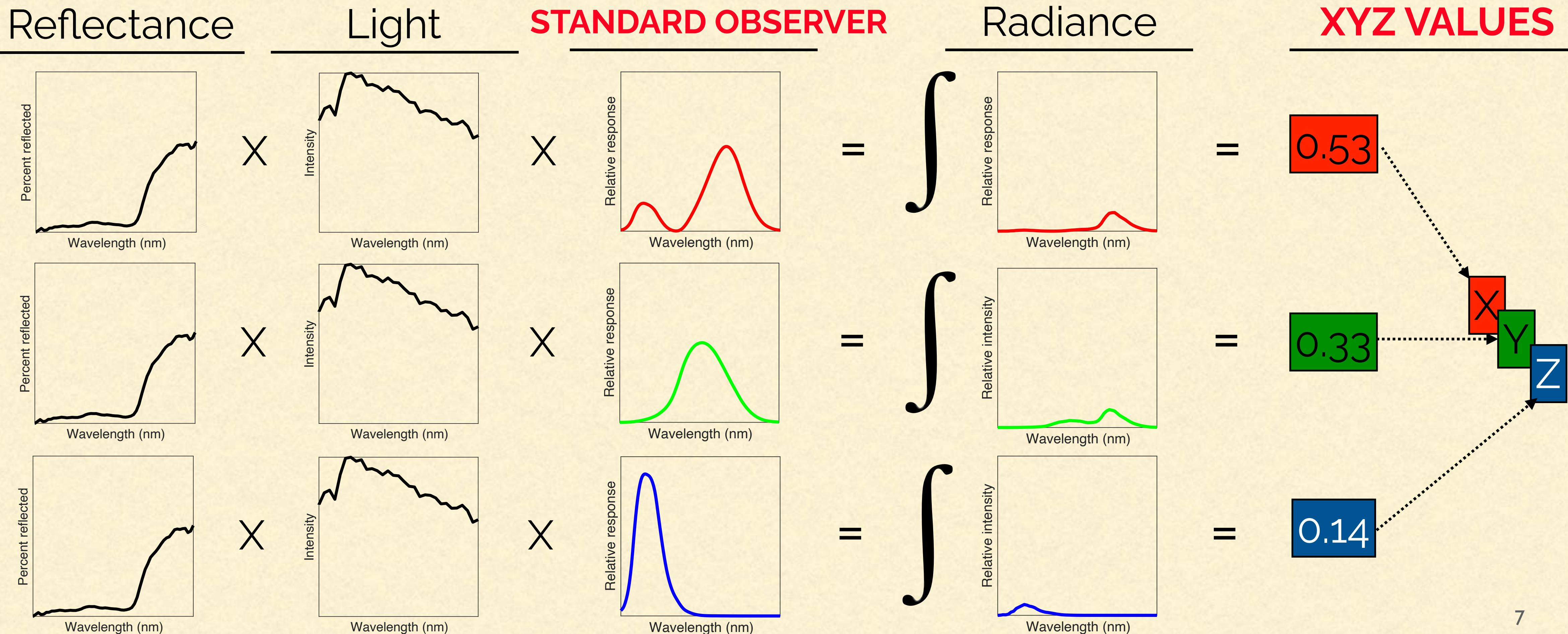


XYZ Values

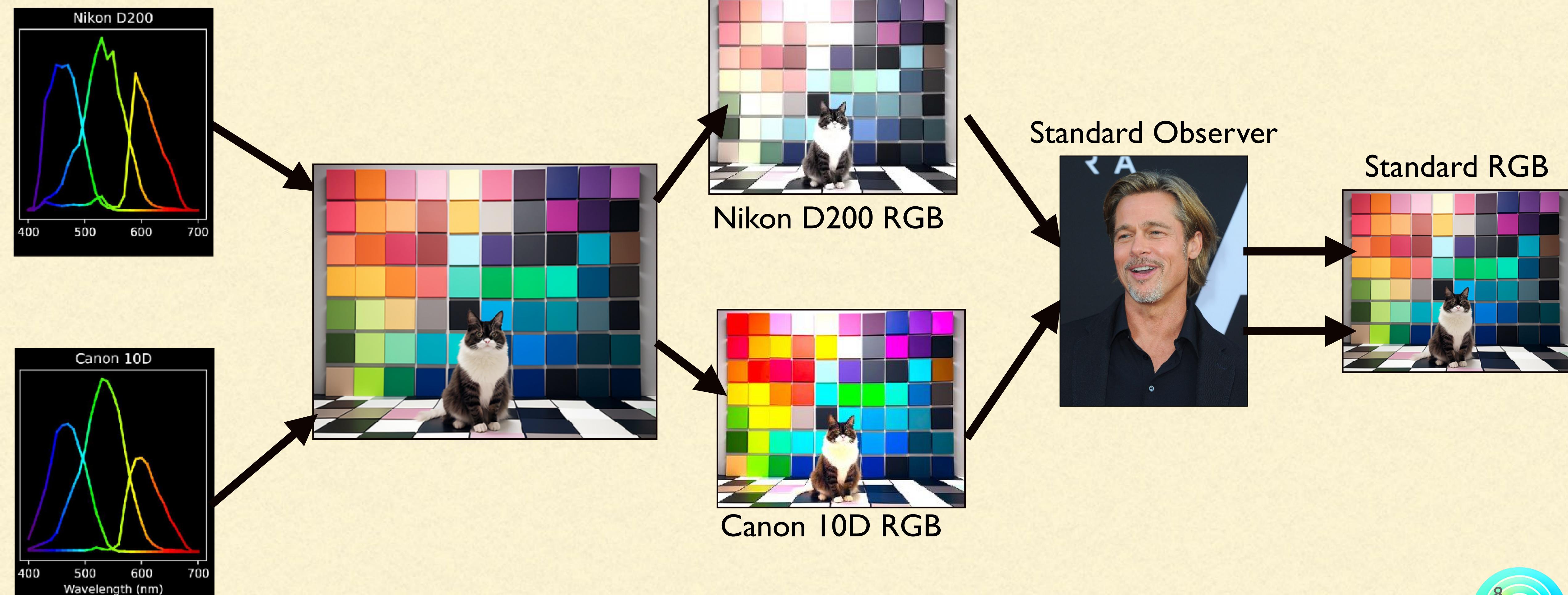


$$Color = \frac{1}{\kappa} \int_{\lambda_1}^{\lambda_2} \rho(\lambda) E(\lambda) S(\lambda) d\lambda$$

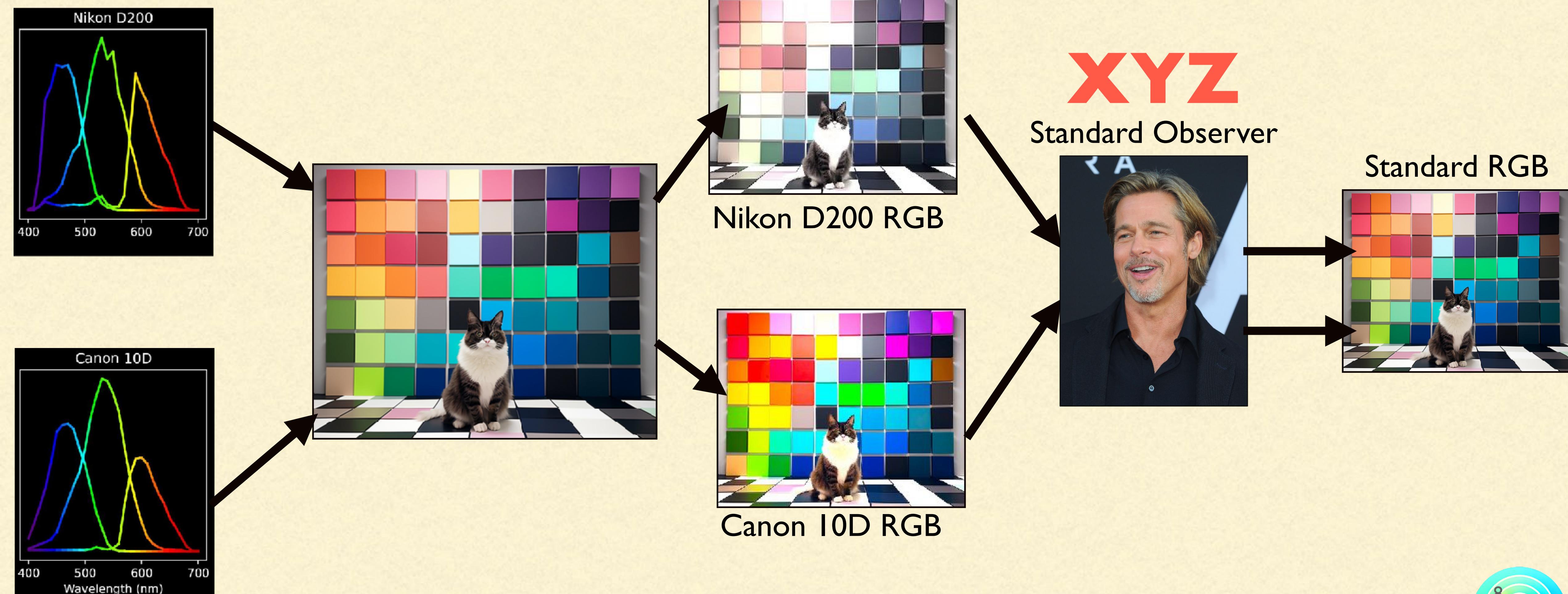
κ : exposure-related constant



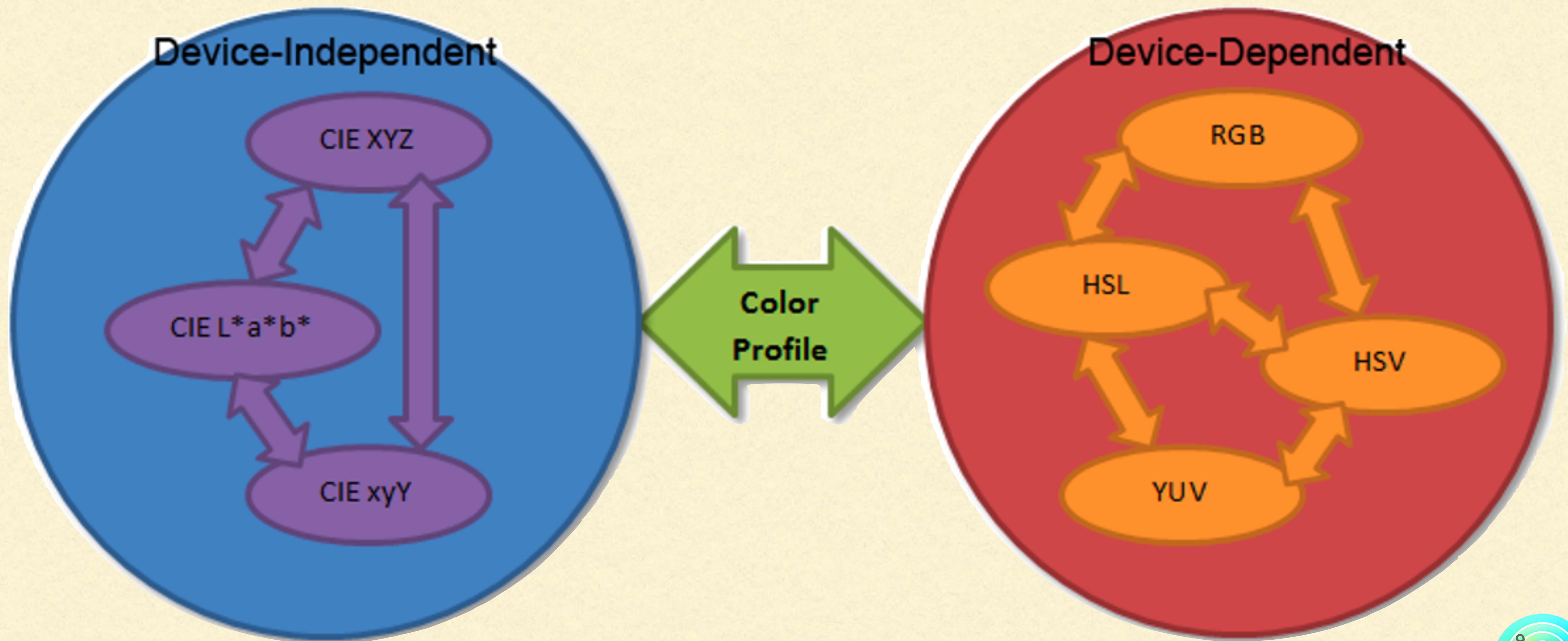
HOW ARE COLORS STANDARDIZED BETWEEN CAMERAS?



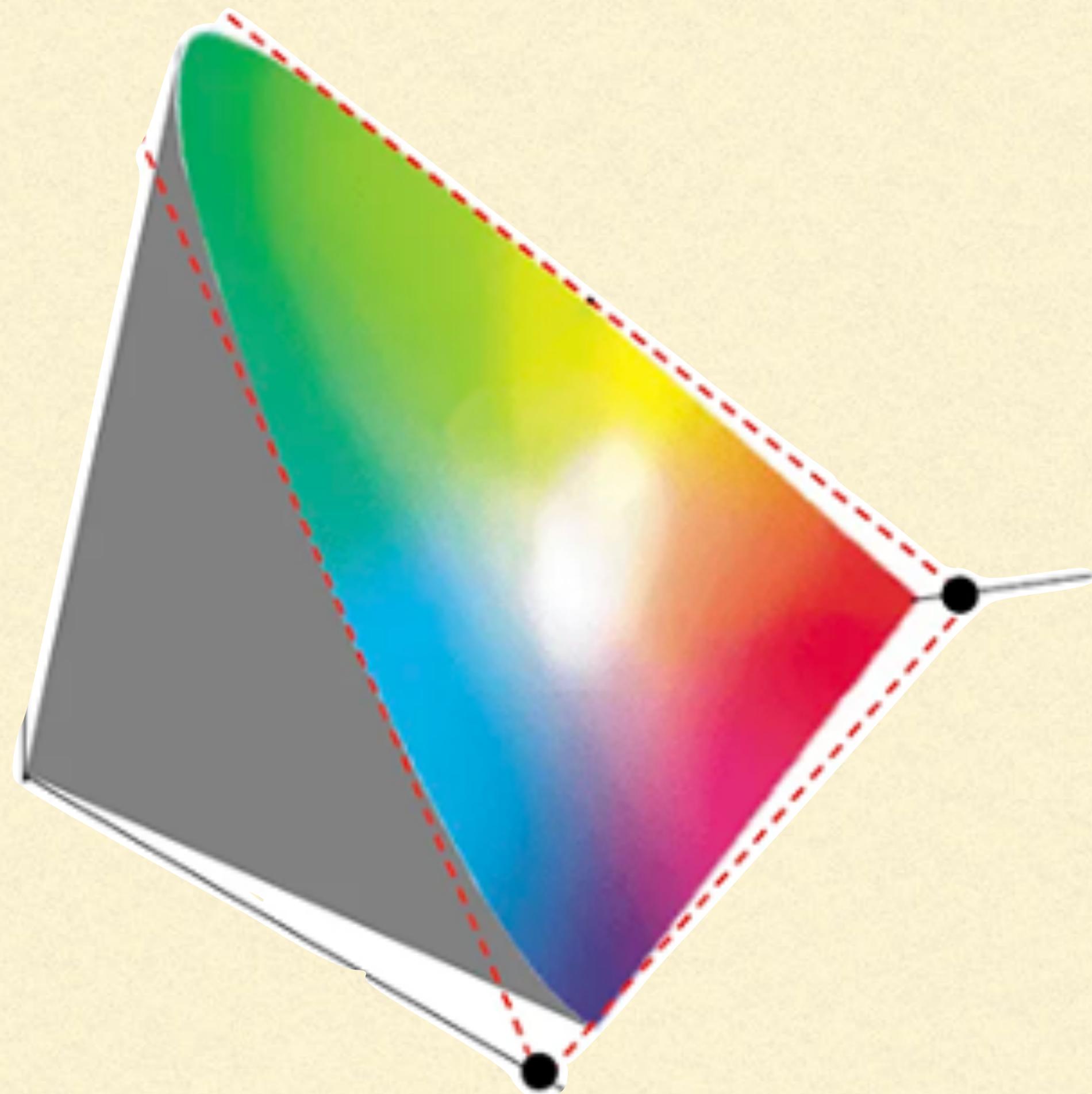
HOW ARE COLORS STANDARDIZED BETWEEN CAMERAS?



COLOR SPACES



1931 CIE XY CHROMATICITY DIAGRAM

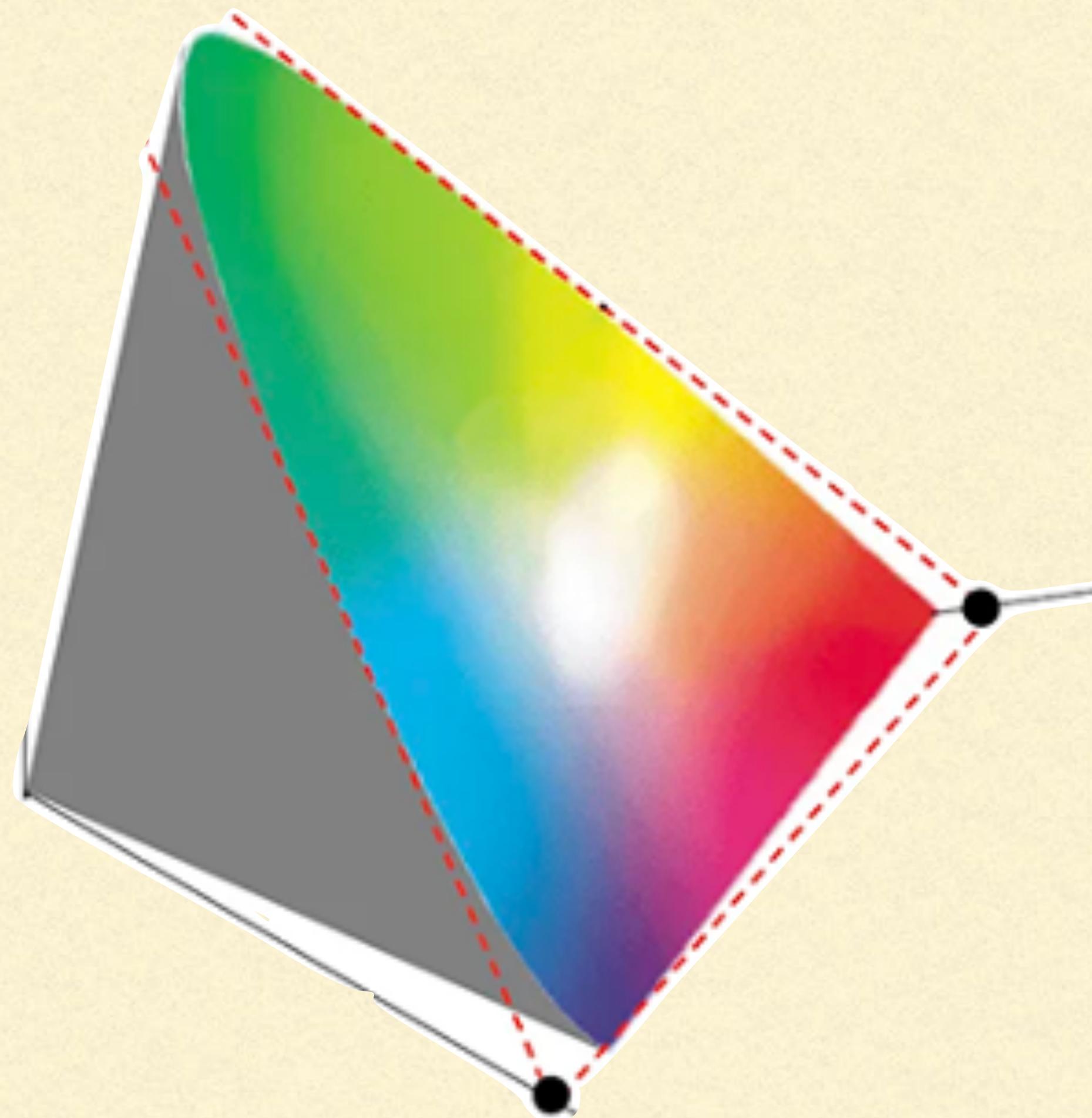


XYZ color solid

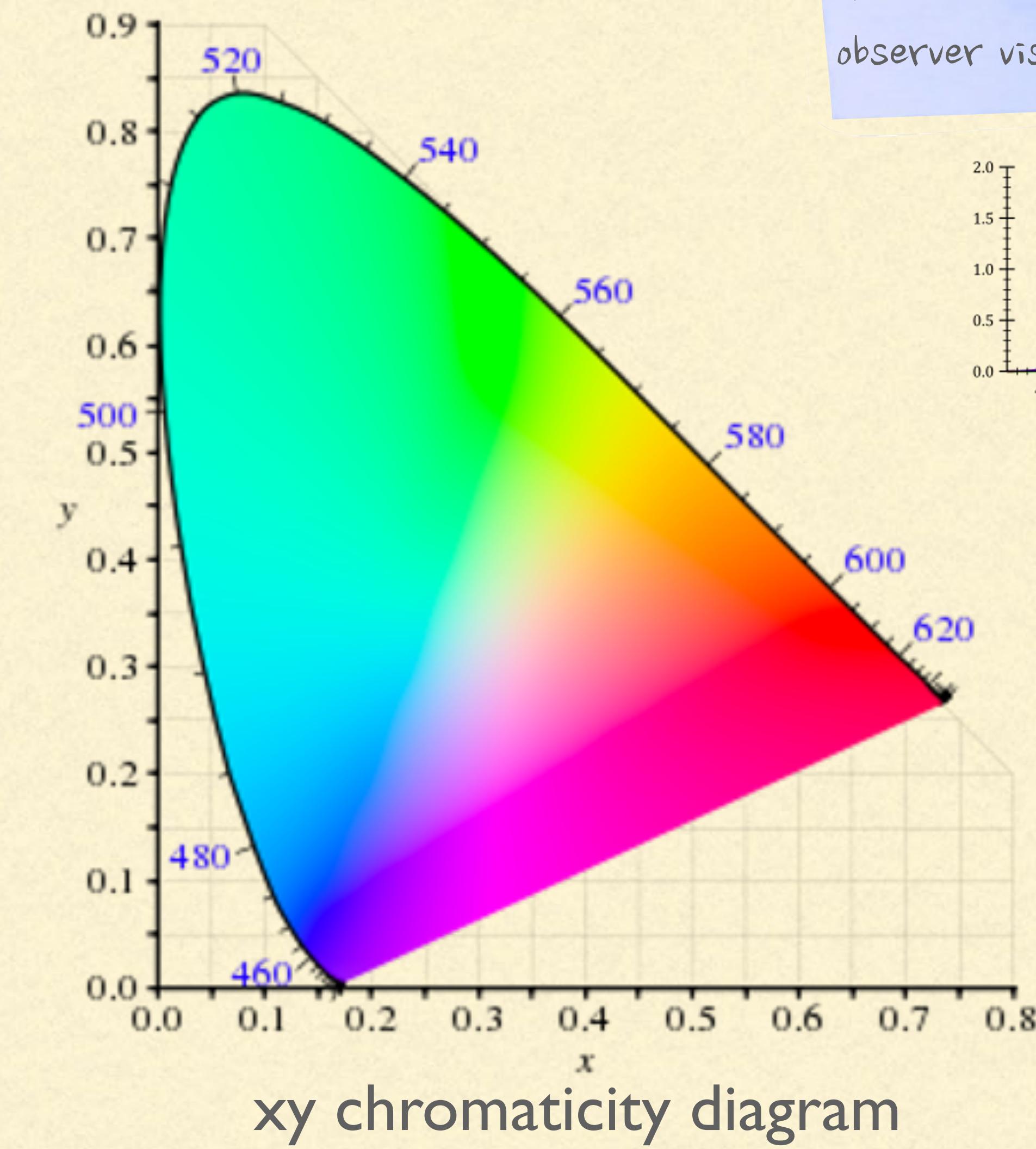
The chromaticity diagram is considered to be a gamut of all colors visible a viewer with the CIE standard observer visual system.



1931 CIE XY CHROMATICITY DIAGRAM

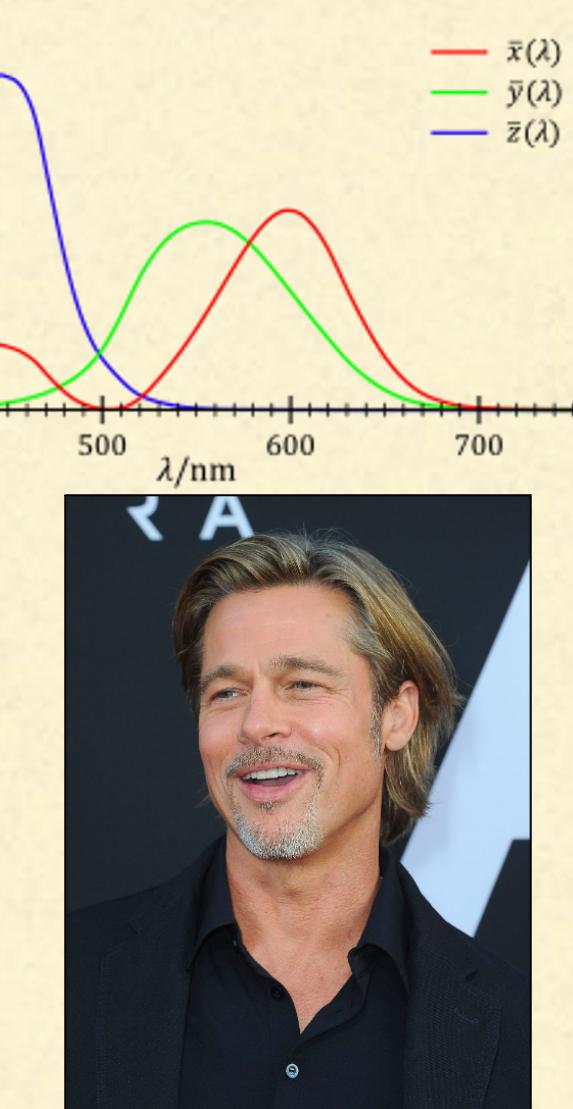


XYZ color solid

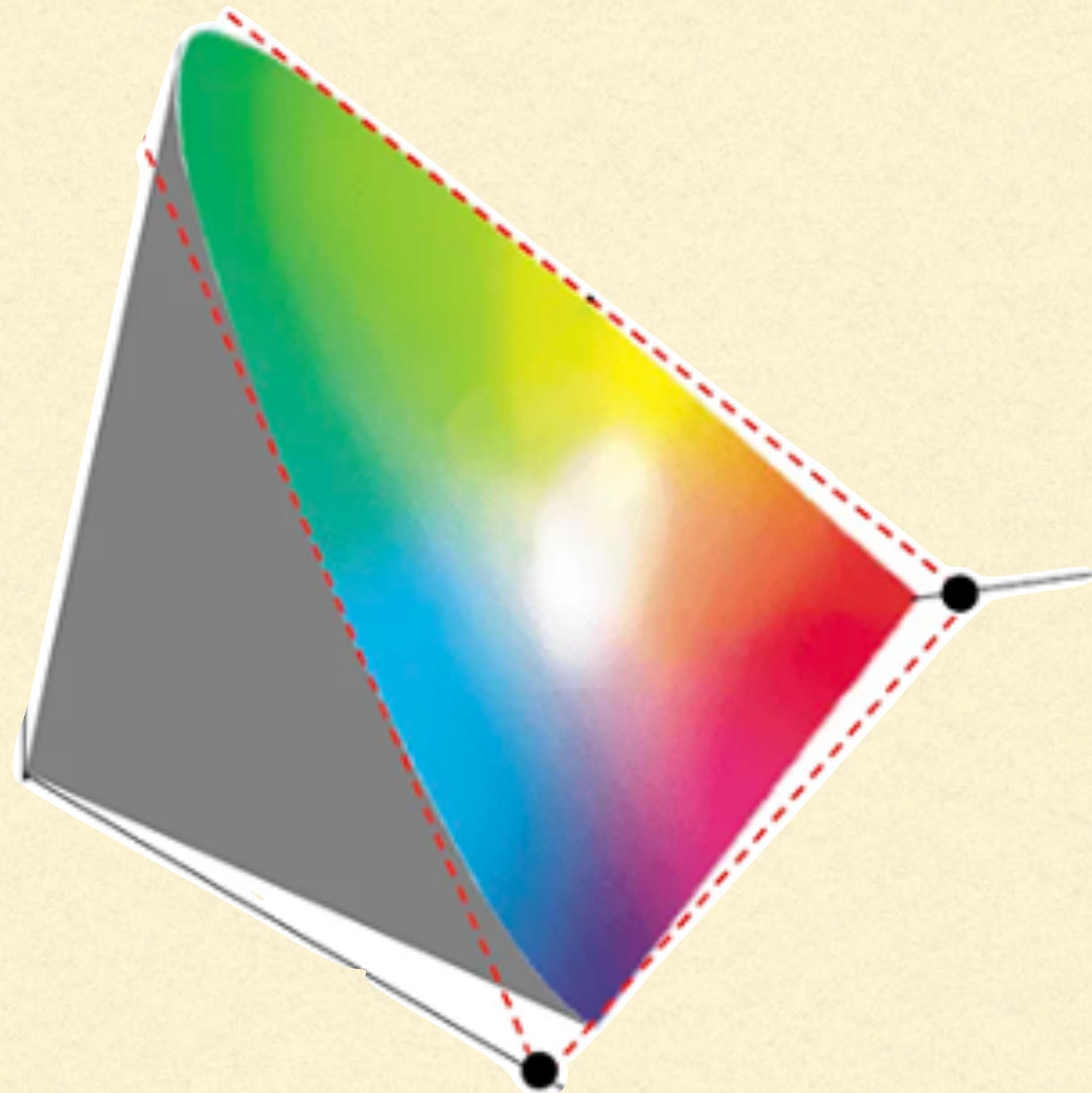


xy chromaticity diagram

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1931 CIE XY CHROMATICITY DIAGRAM

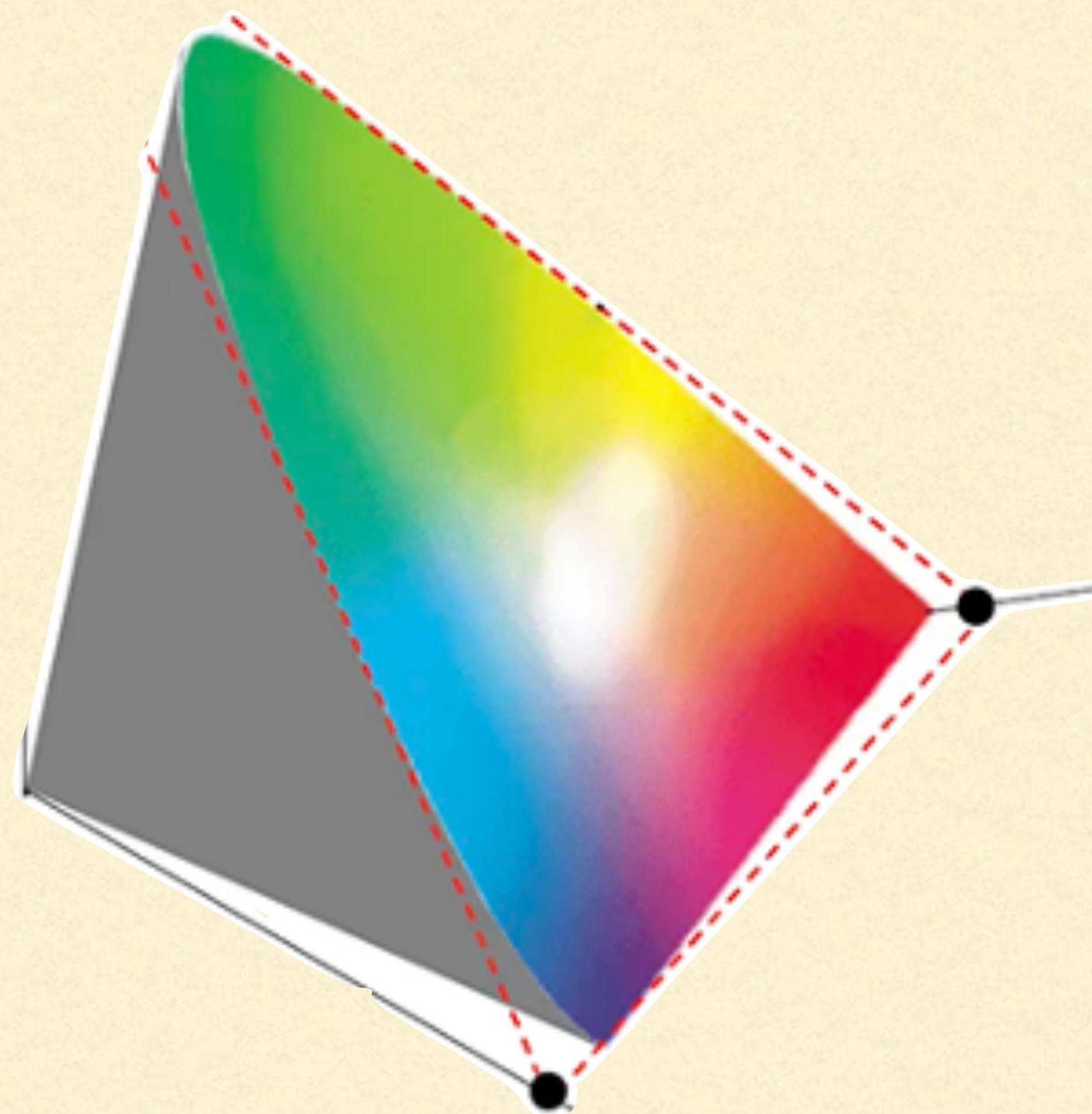


XYZ color solid

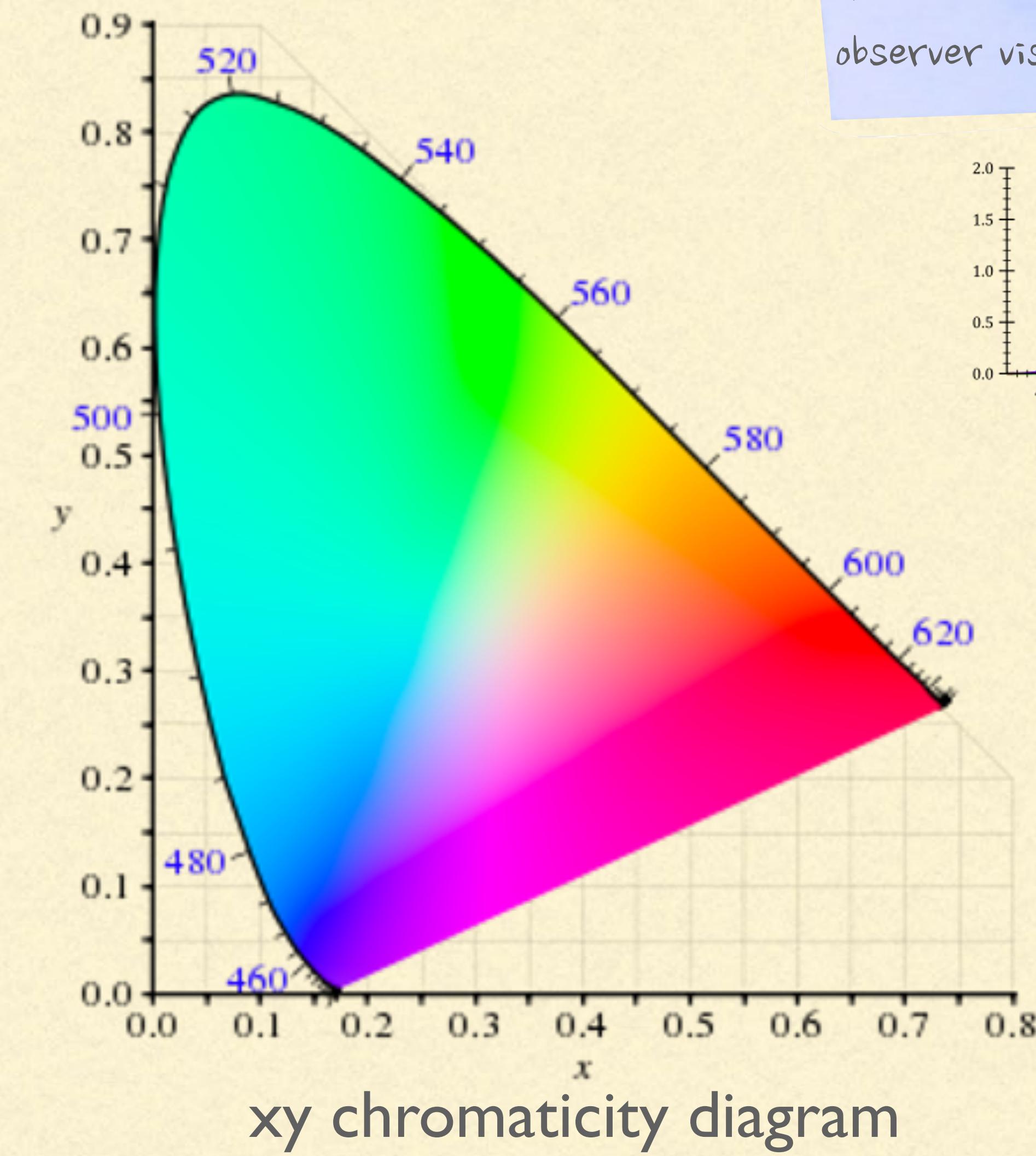
The chromaticity diagram is considered to be a gamut of all colors visible a viewer with the CIE standard observer visual system.



1931 CIE XY CHROMATICITY DIAGRAM

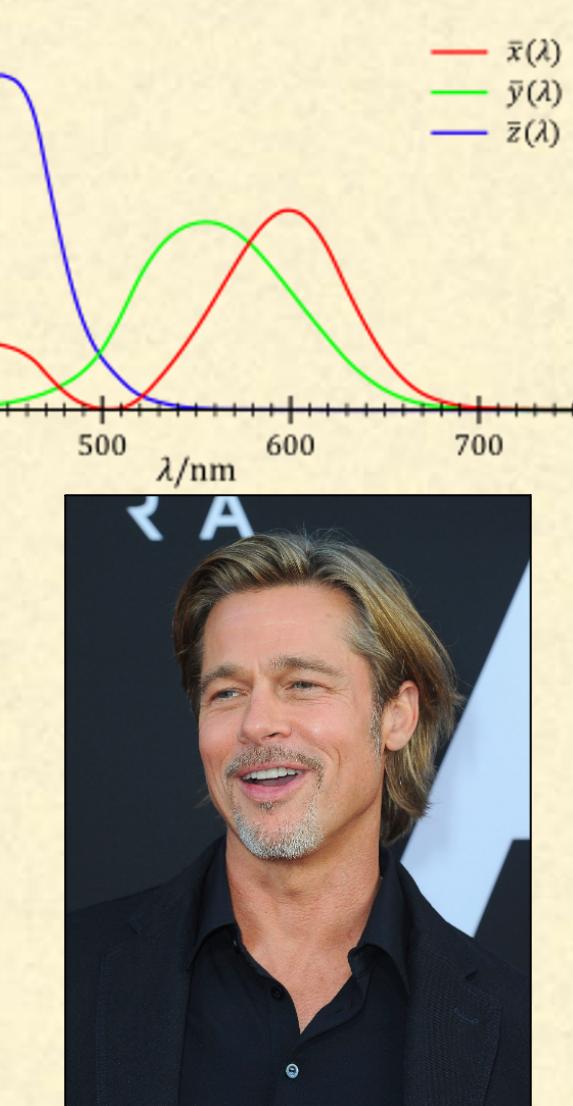


XYZ color solid

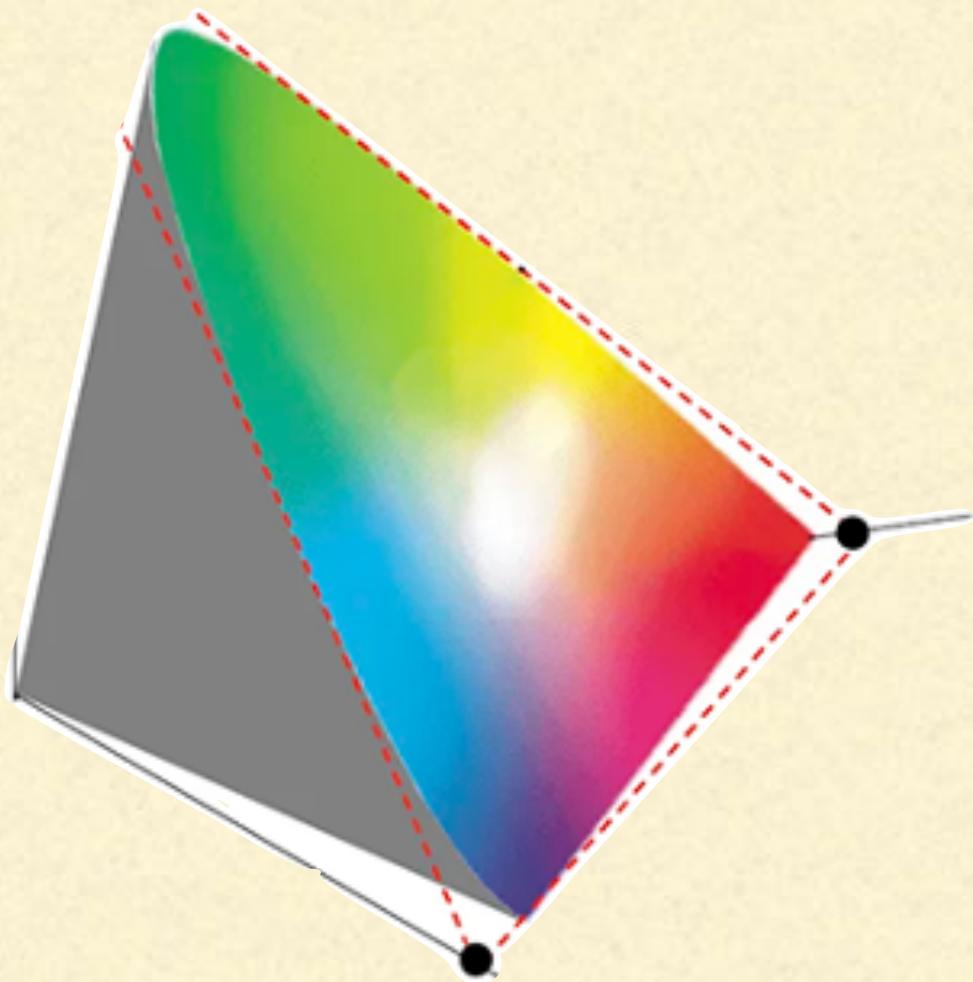


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XYZ TO XY

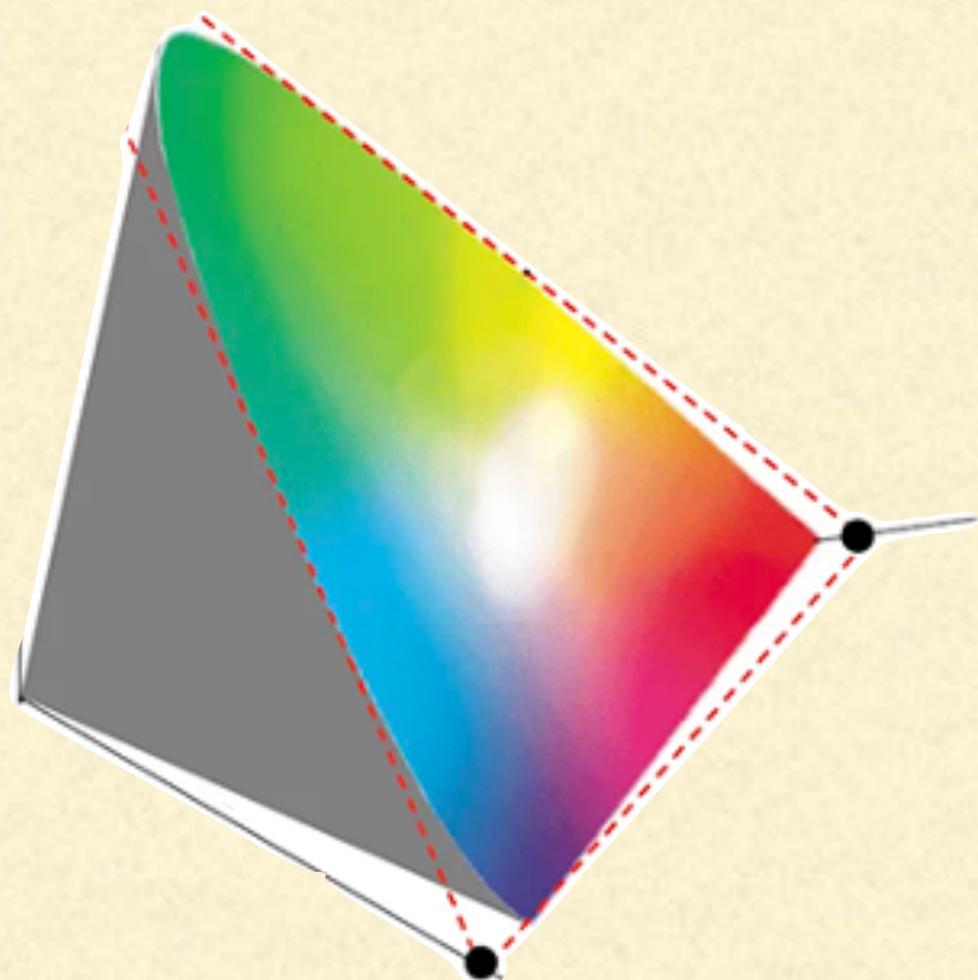


XYZ color solid

$$x = \frac{X}{X + Y + Z}$$
$$y = \frac{Y}{X + Y + Z}$$
$$z = \frac{Z}{X + Y + Z}$$

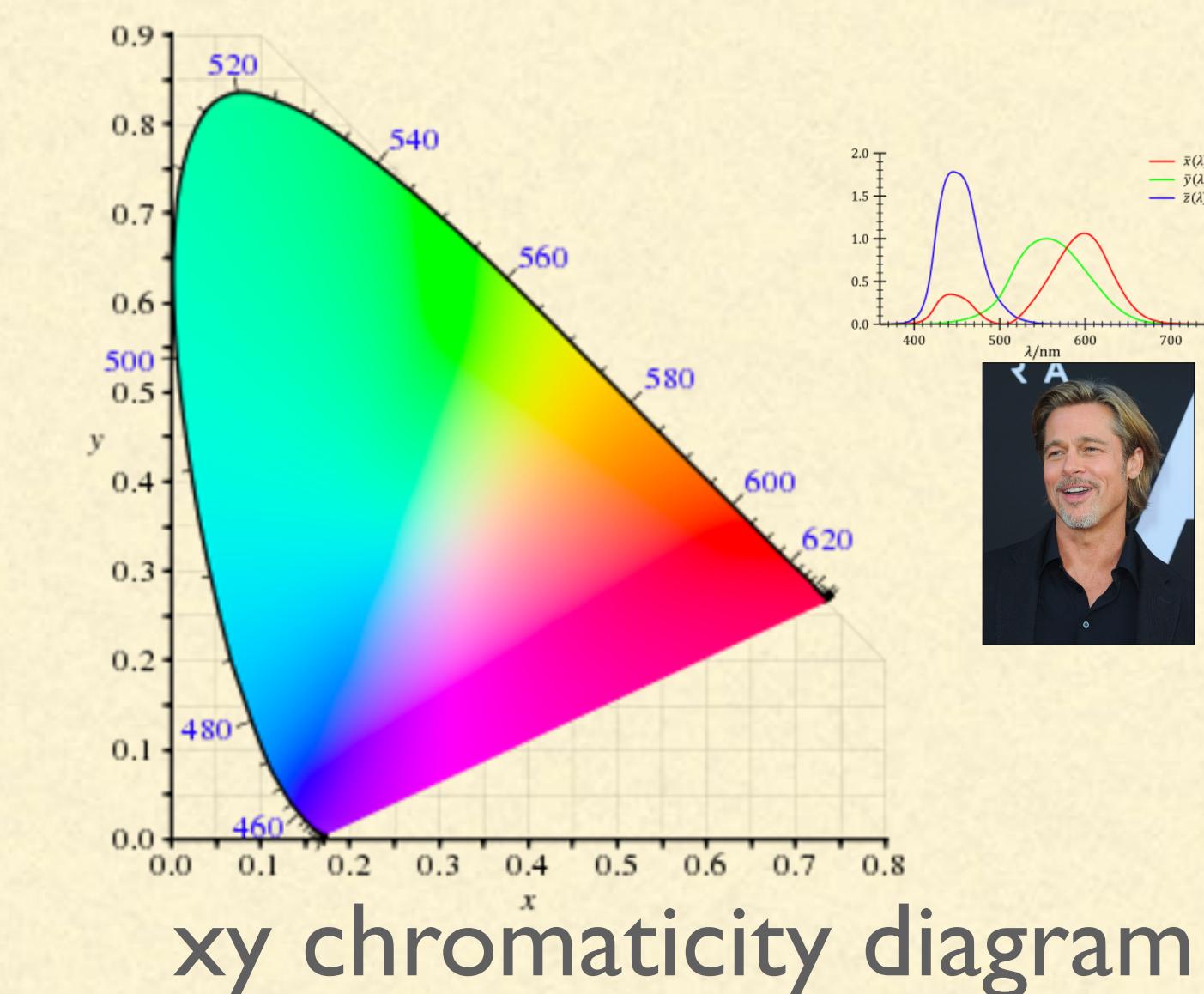
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XYZ TO XY



XYZ color solid

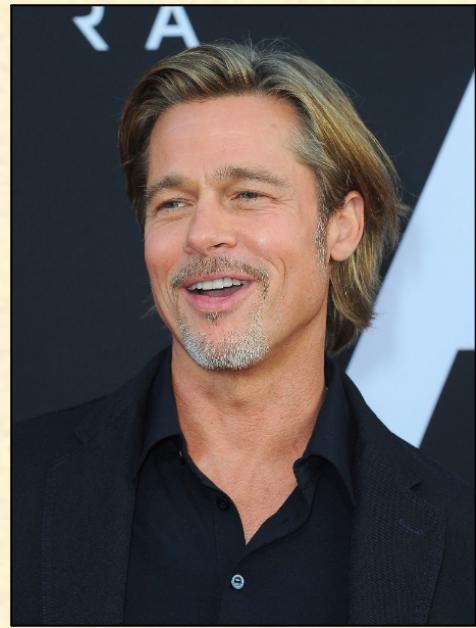
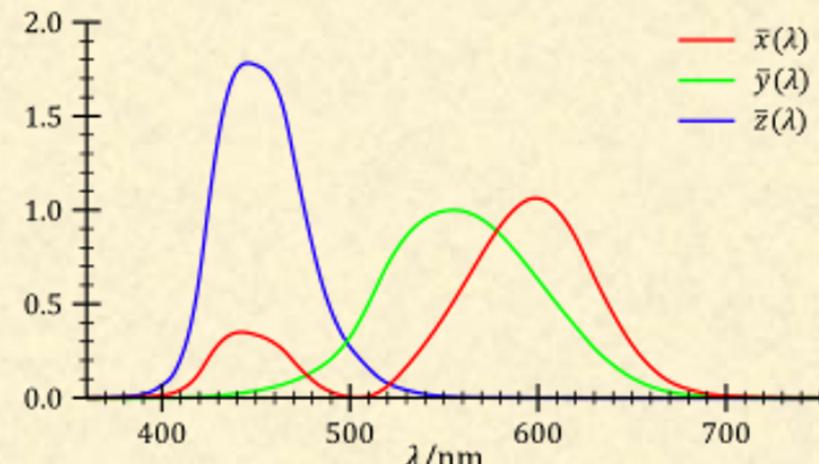
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xy chromaticity diagram

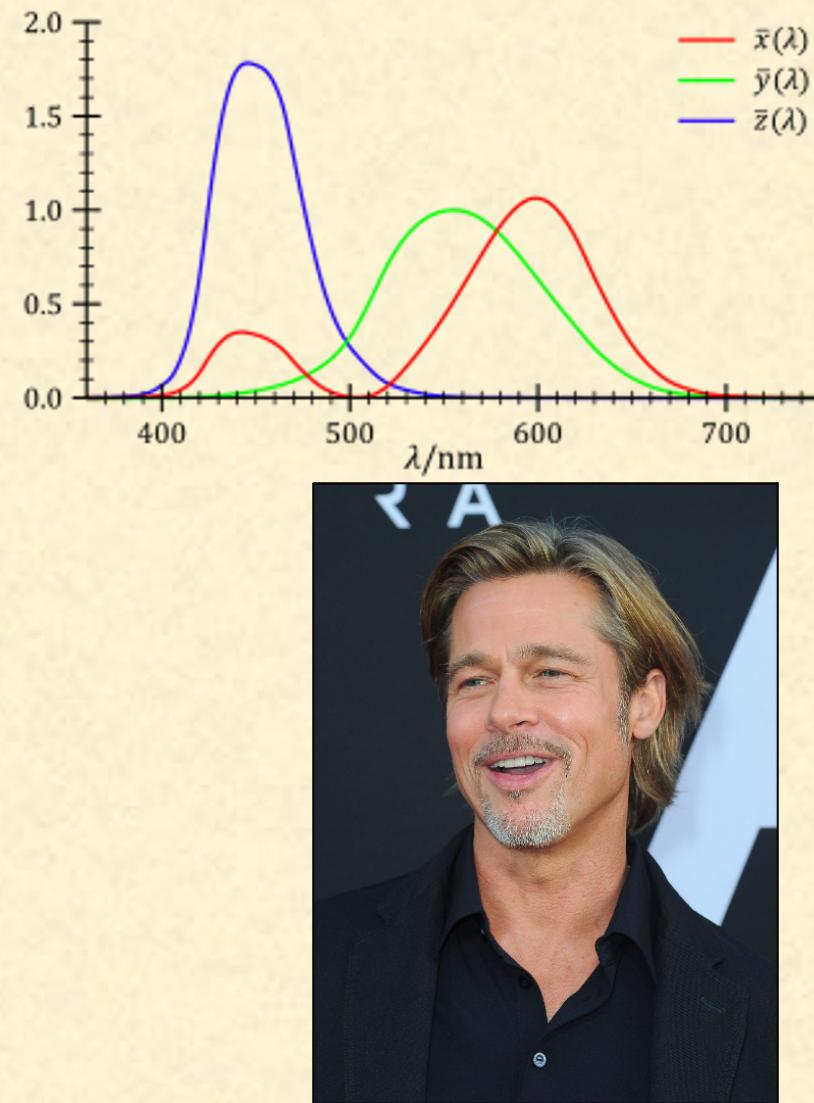
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CIE 1931 XYZ Standard Observer



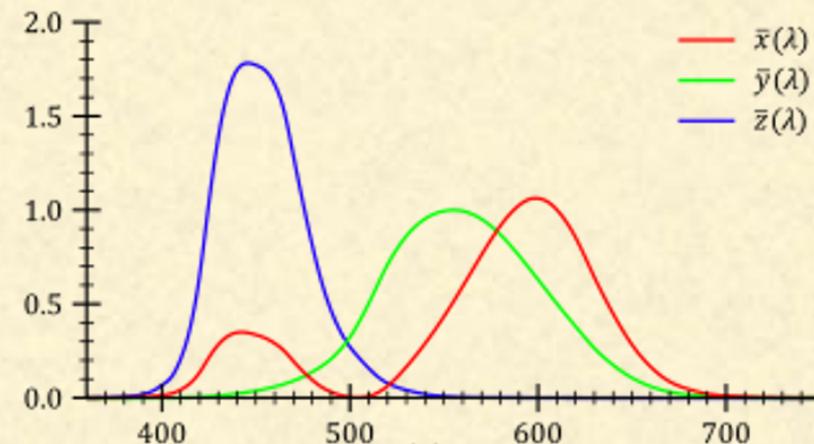
CIE 1931 XYZ Standard Observer

- ▶ “Spectral response of a human with normal color vision”



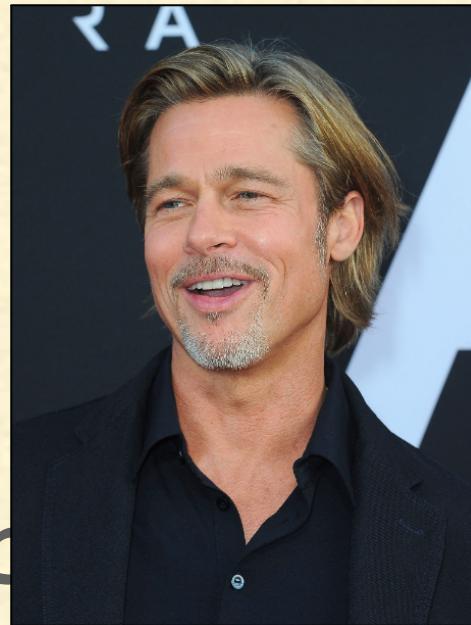
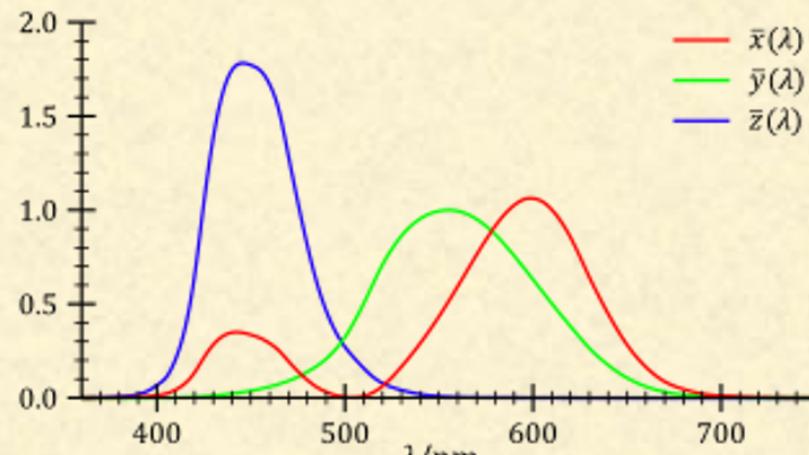
CIE 1931 XYZ Standard Observer

- ▶ “Spectral response of a human with normal color vision”
- ▶ Curves come from handful of subjects tested in 1929 by John Guild, David Wright.



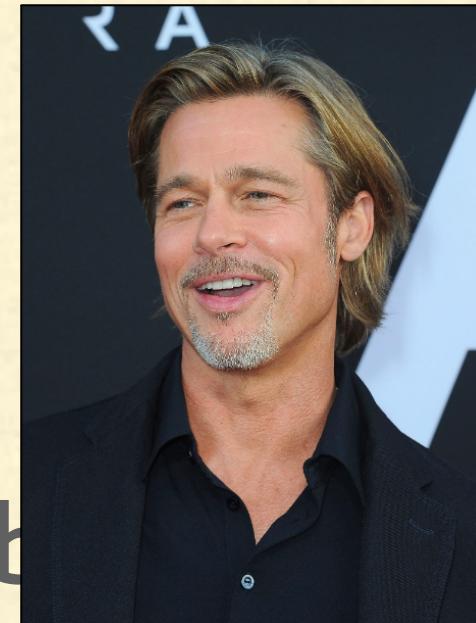
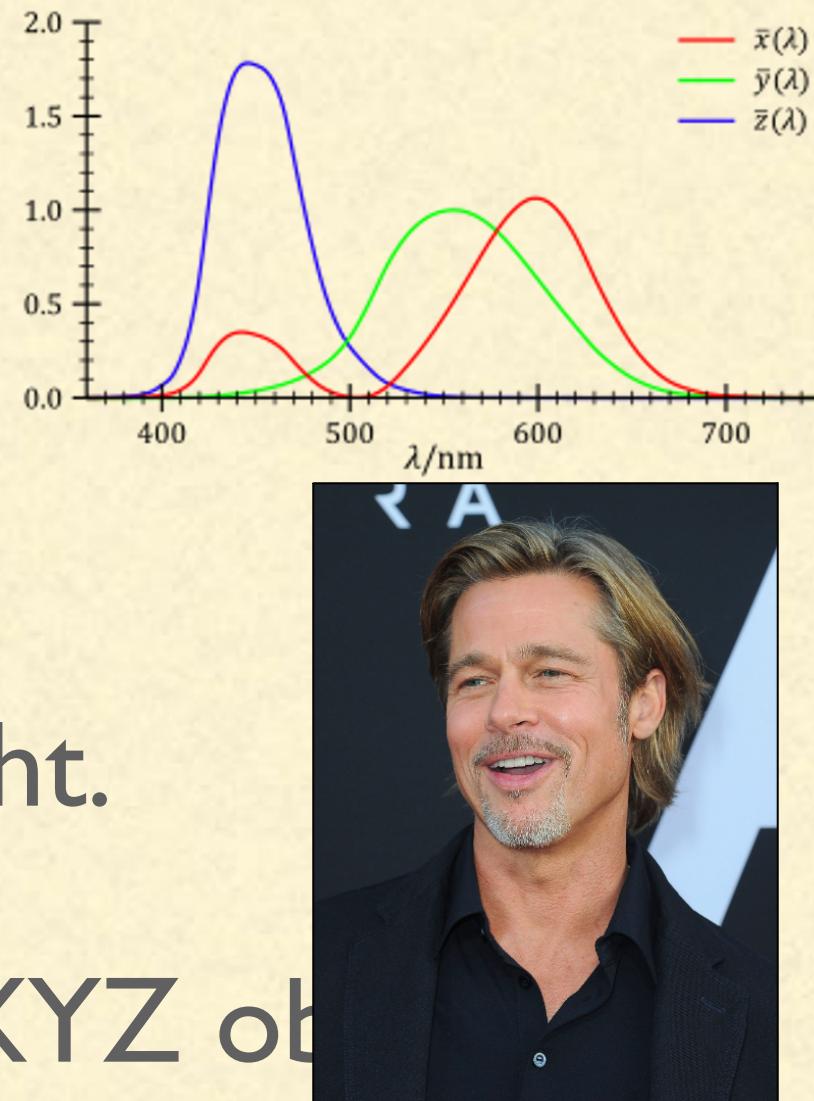
CIE 1931 XYZ Standard Observer

- ▶ “Spectral response of a human with normal color vision”
- ▶ Curves come from handful of subjects tested in 1929 by John Guild, David Wright.
- ▶ Do ALL humans sense colors the same way ? (And is that the same as the CIE XYZ ob



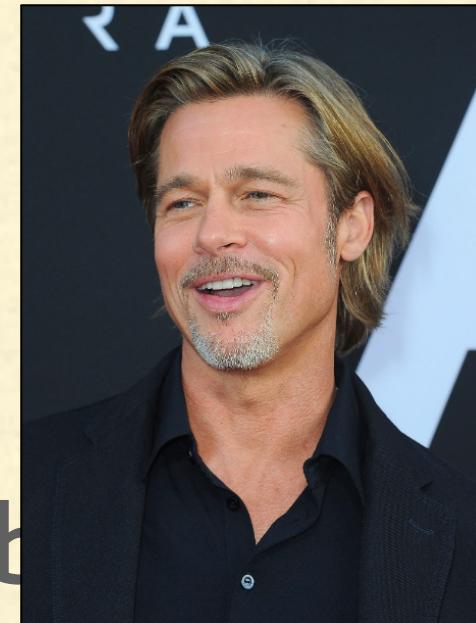
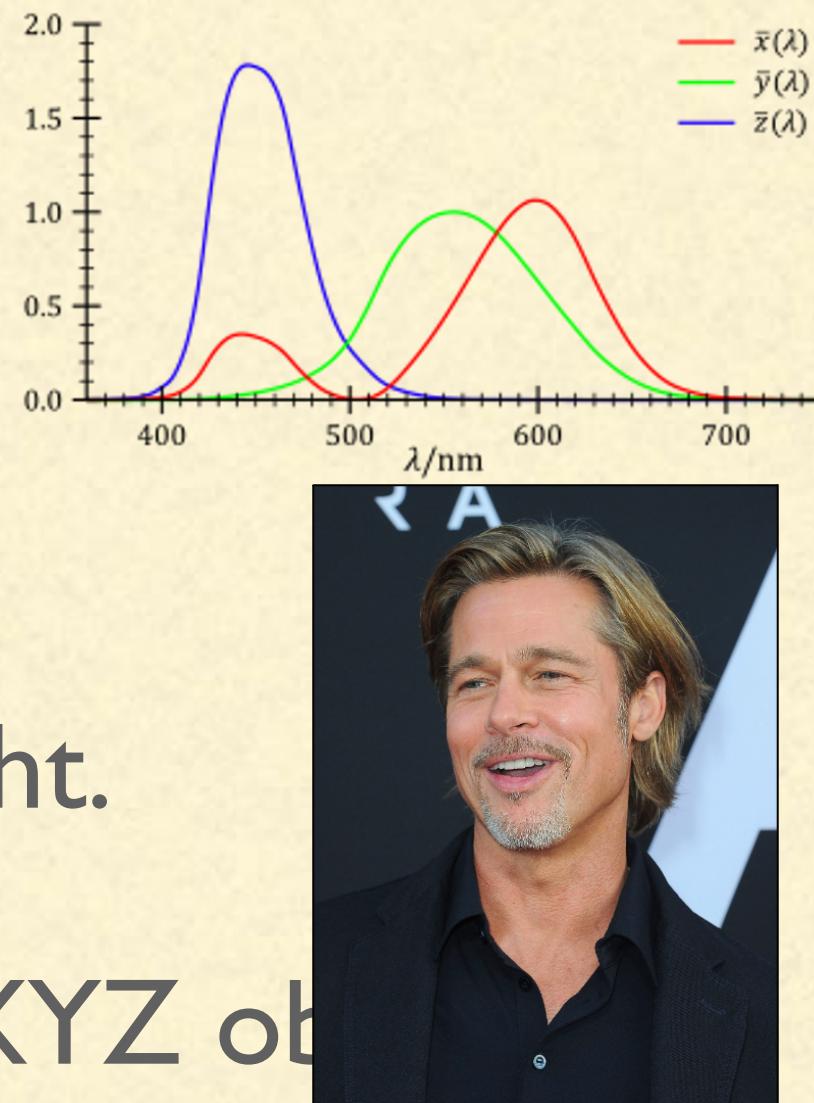
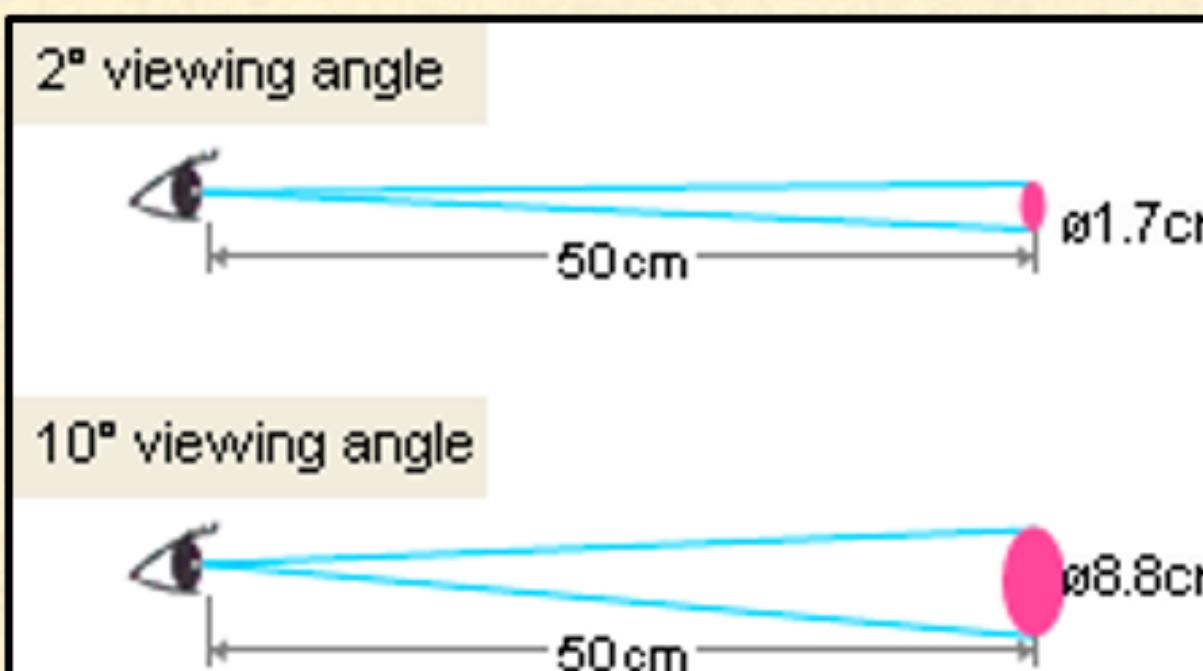
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- ▶ Do ALL humans sense colors the same way ? (And is that the same as the CIE XYZ observer?)
- ▶ Short answer: No, not all humans sense colors the same way, but after decades of repeating color matching experiments, these curves are found to be stable.

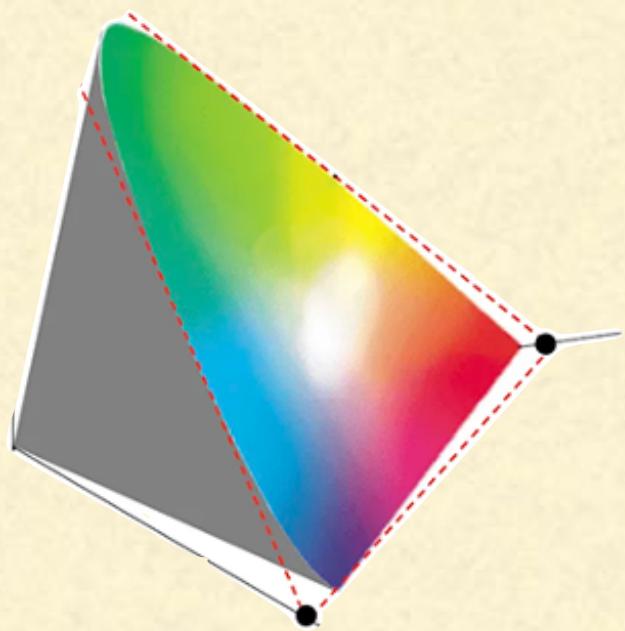


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- ▶ “Spectral response of a human with normal color vision”
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- ▶ Do ALL humans sense colors the same way ? (And is that the same as the CIE XYZ observer?)
- ▶ Short answer: No, not all humans sense colors the same way, but after decades of repeating color matching experiments, these curves are found to be stable.
- ▶ There are slightly different curves for 2 degree vs 10 degree viewing angles.

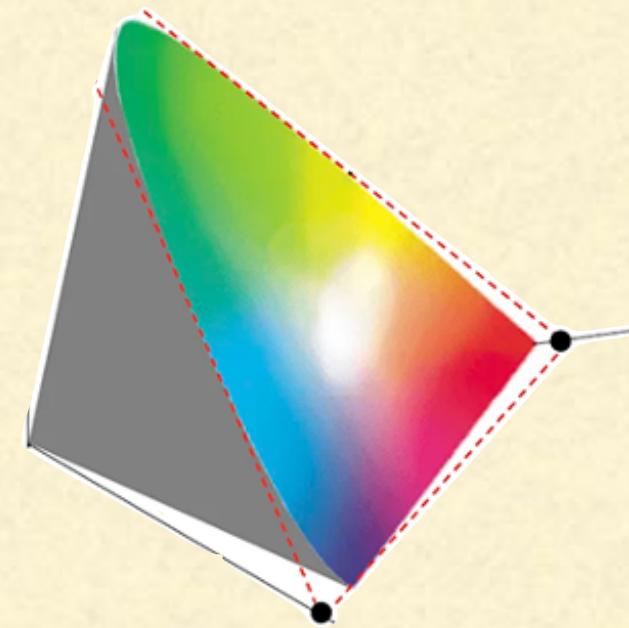


Brief Summary of Key Concepts



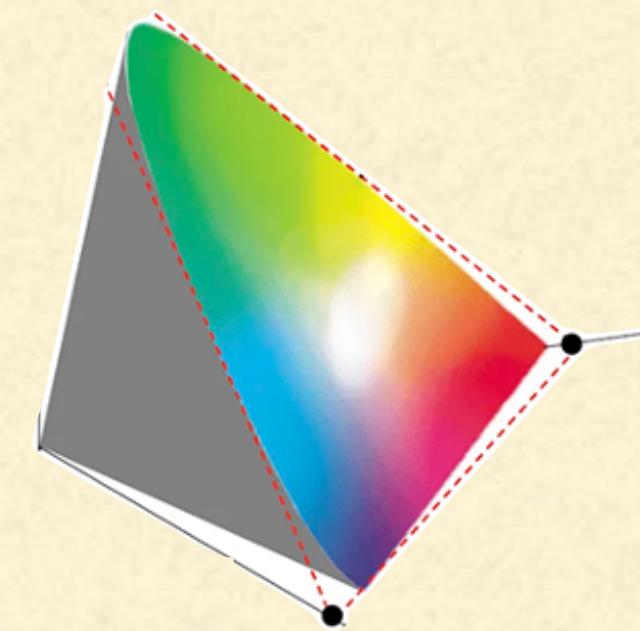
Brief Summary of Key Concepts

- I. The CIE Standard Observer curves represent “Spectral response of a human with normal color vision”

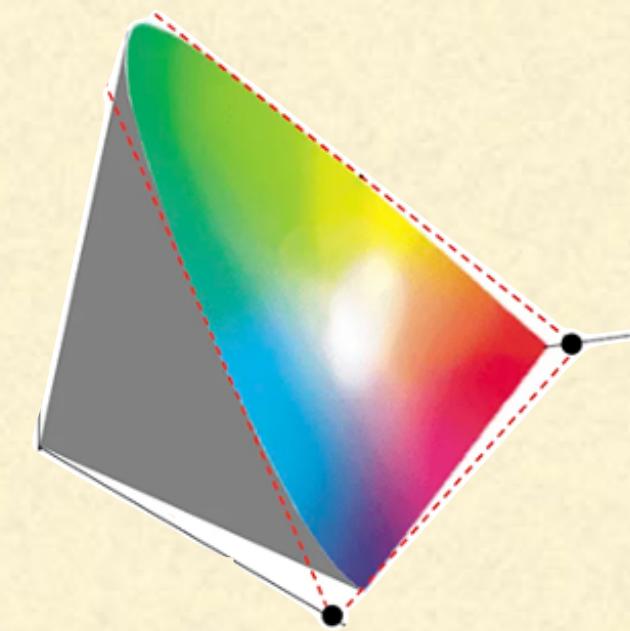


Brief Summary of Key Concepts

1. The CIE Standard Observer curves represent “Spectral response of a human with normal color vision”
2. Given a stimulus, we use the CIE Standard Observer curves to calculate the XYZ values.



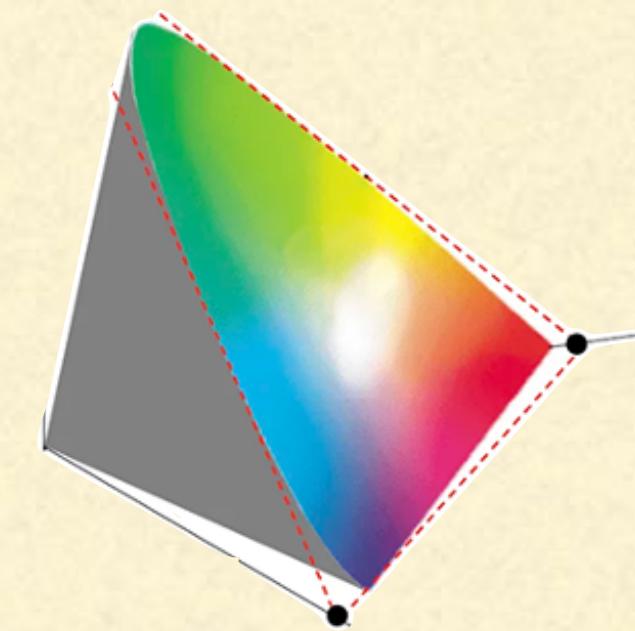
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3. Once in XYZ coordinates, we can convert colors to one of many standard color spaces.



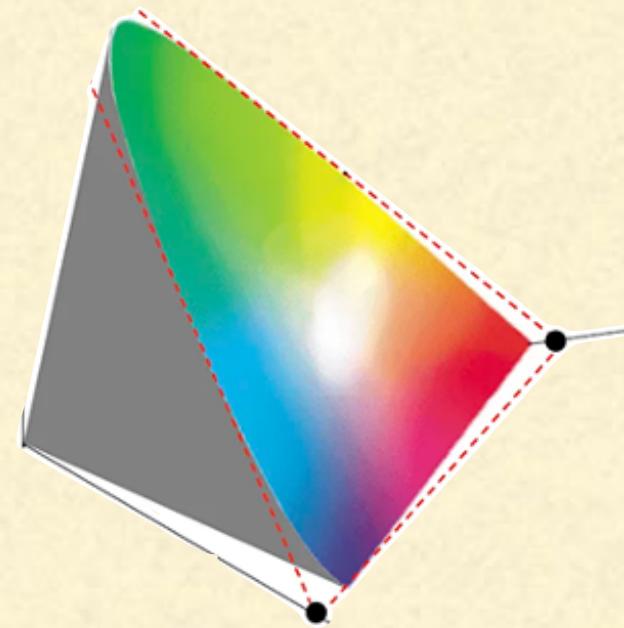
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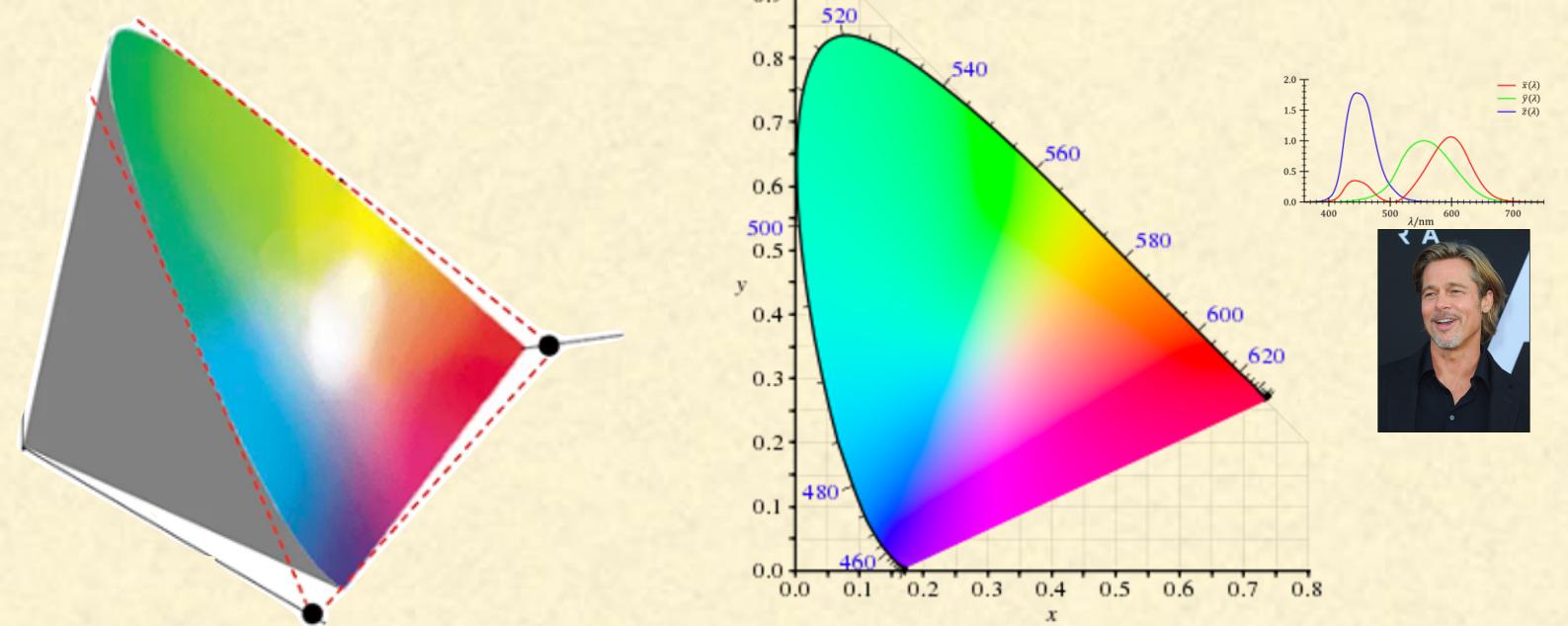
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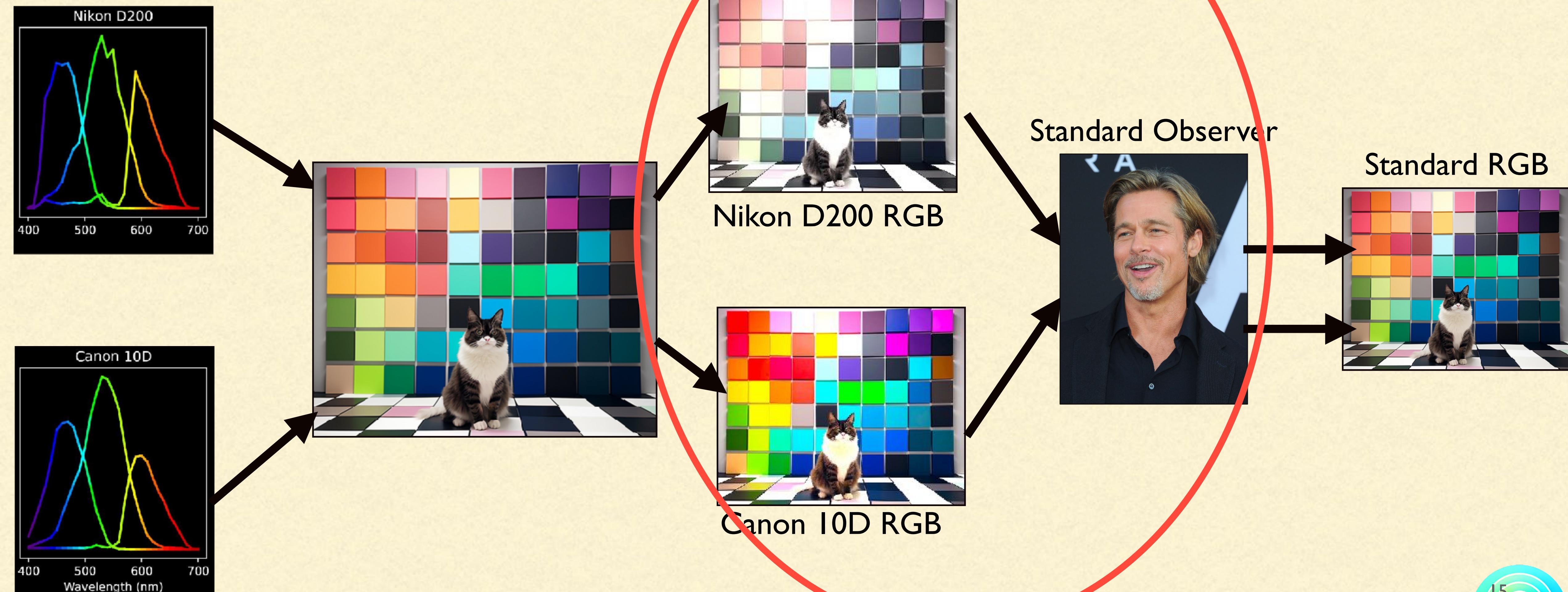


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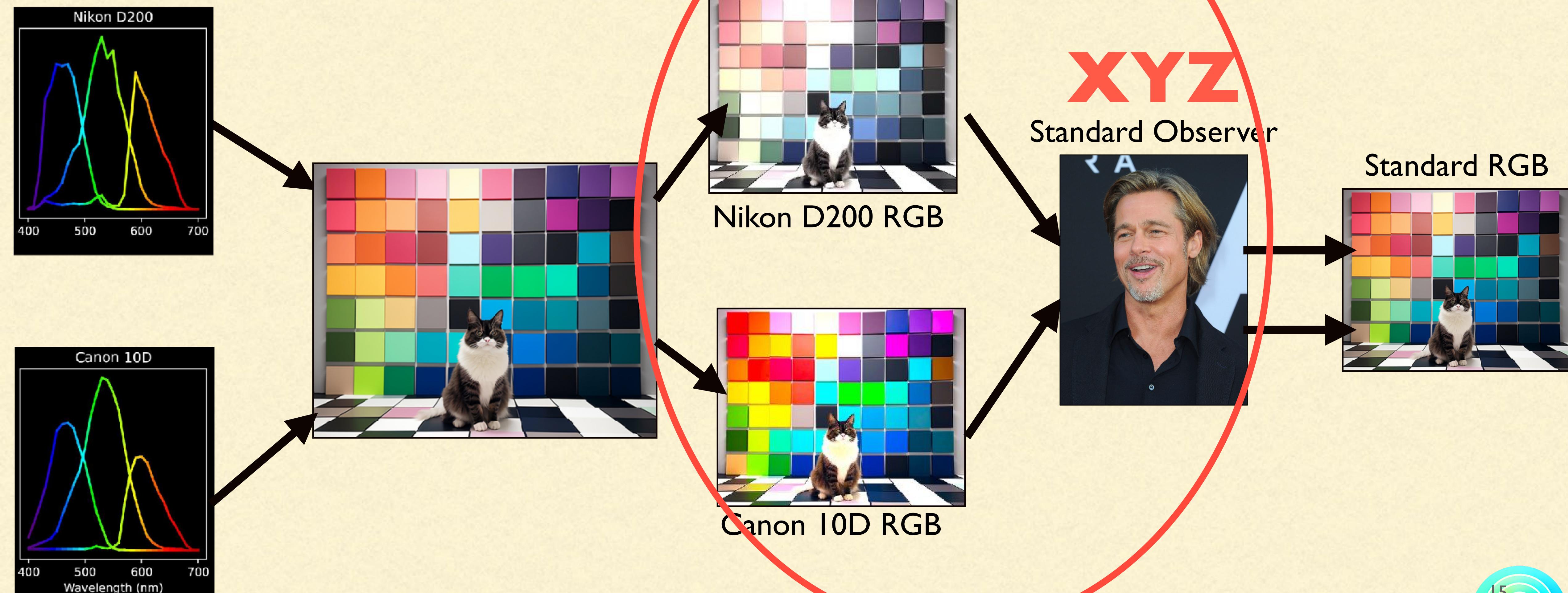


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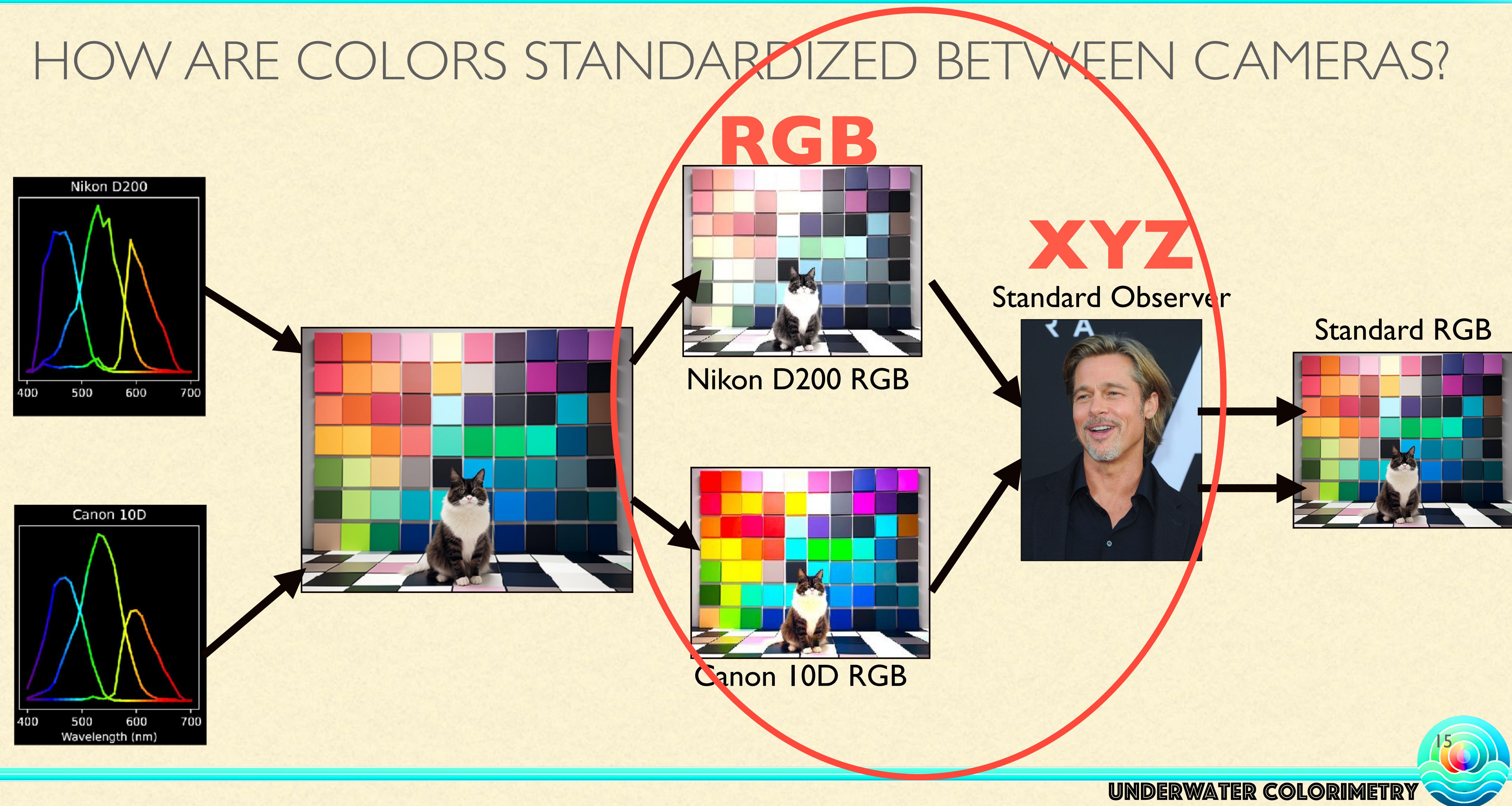
HOW ARE COLORS STANDARDIZED BETWEEN CAMERAS?



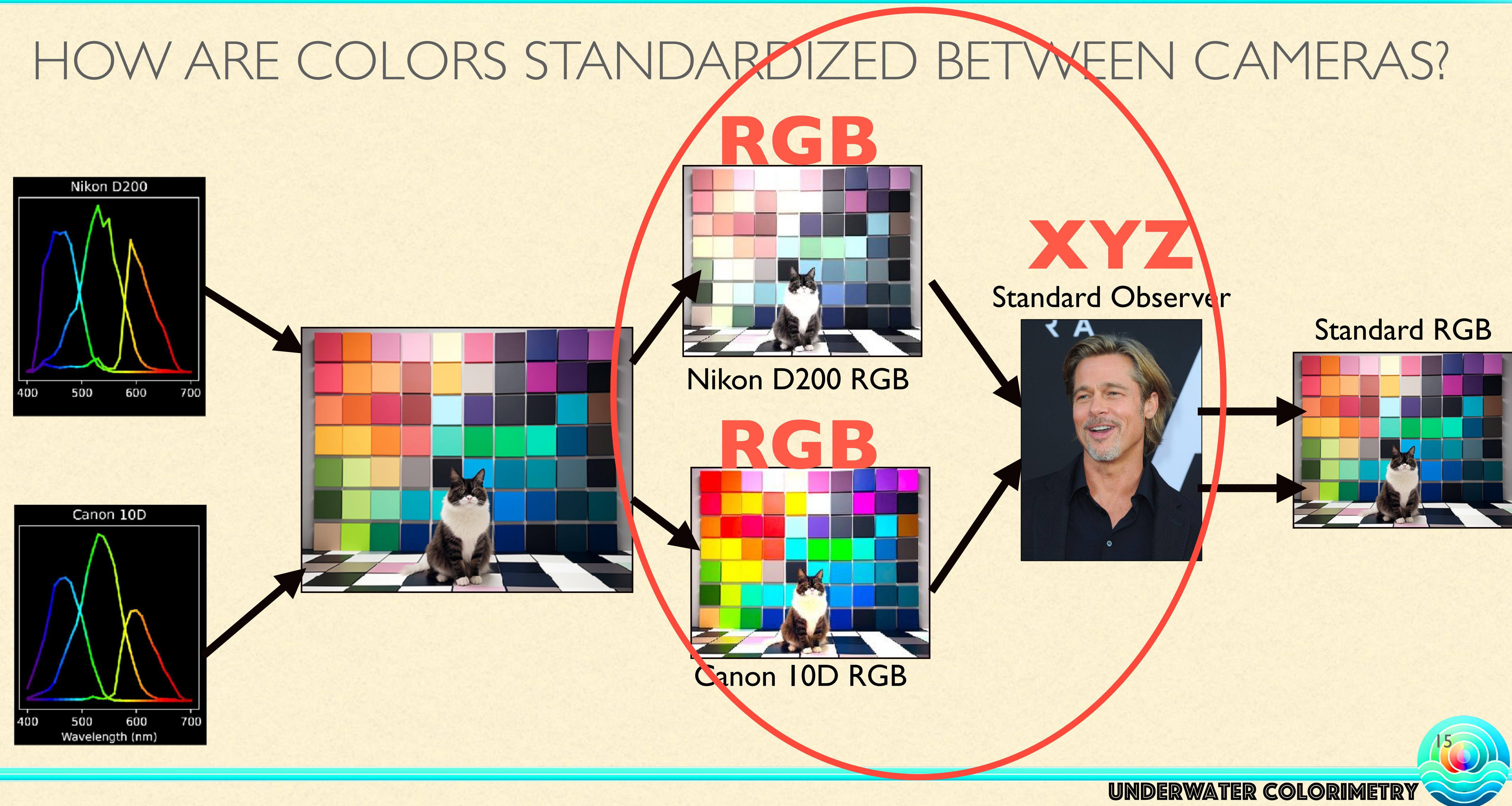
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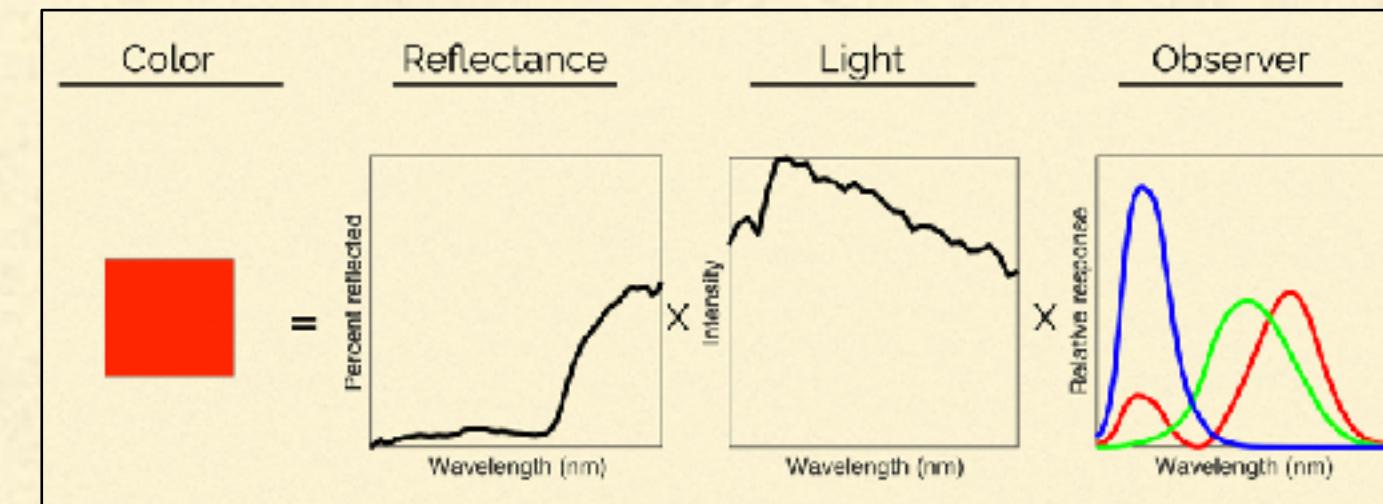
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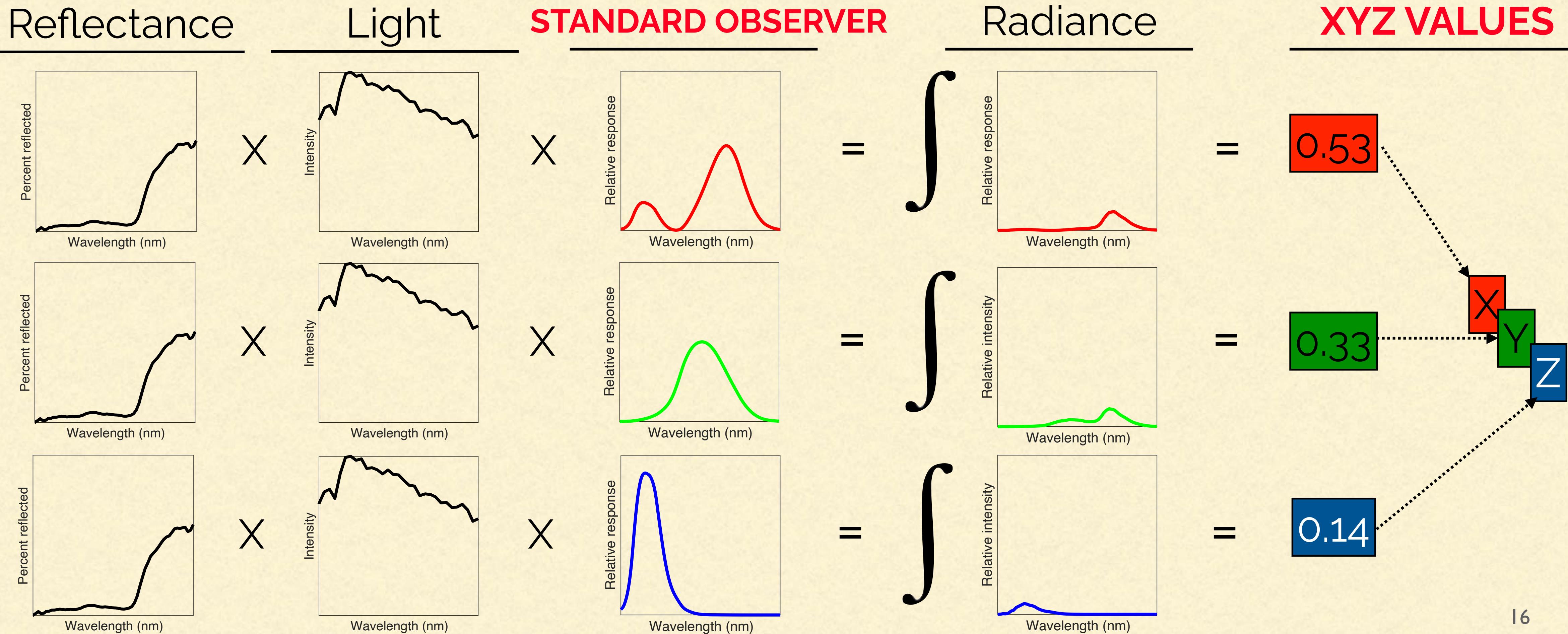


XYZ Values



$$\text{Color} = \frac{1}{\kappa} \int_{\lambda_1}^{\lambda_2} \rho(\lambda) E(\lambda) S(\lambda) d\lambda$$

κ : exposure-related constant



Camera RGB to XYZ transformation

But before that...

Camera RGB to XYZ transformation

But before that...

- I. Will you always use the same camera?

Camera RGB to XYZ transformation

But before that...

1. Will you always use the same camera?
2. Will you compare images from different cameras?

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Camera RGB to XYZ transformation

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Camera RGB to XYZ transformation

But before that...

1. Will you always use the same camera?
2. Will you compare images from different cameras?
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4. Even if you will use the same camera, will you take images under very different light conditions?
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Camera RGB to XYZ transformation

But before that...

1. Will you always use the same camera?
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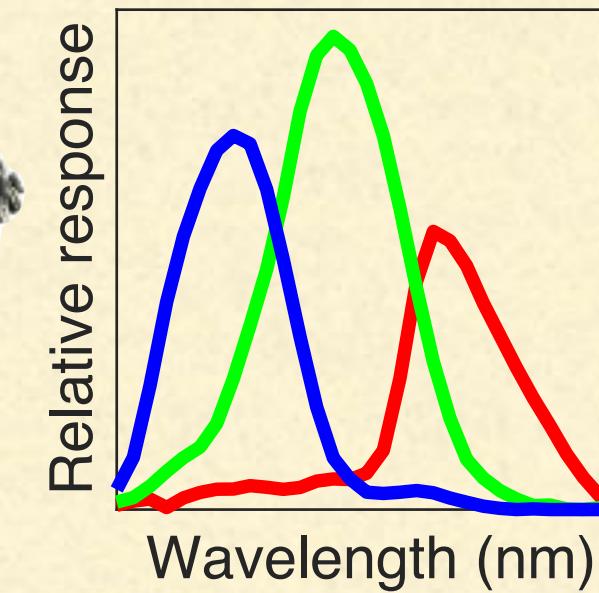
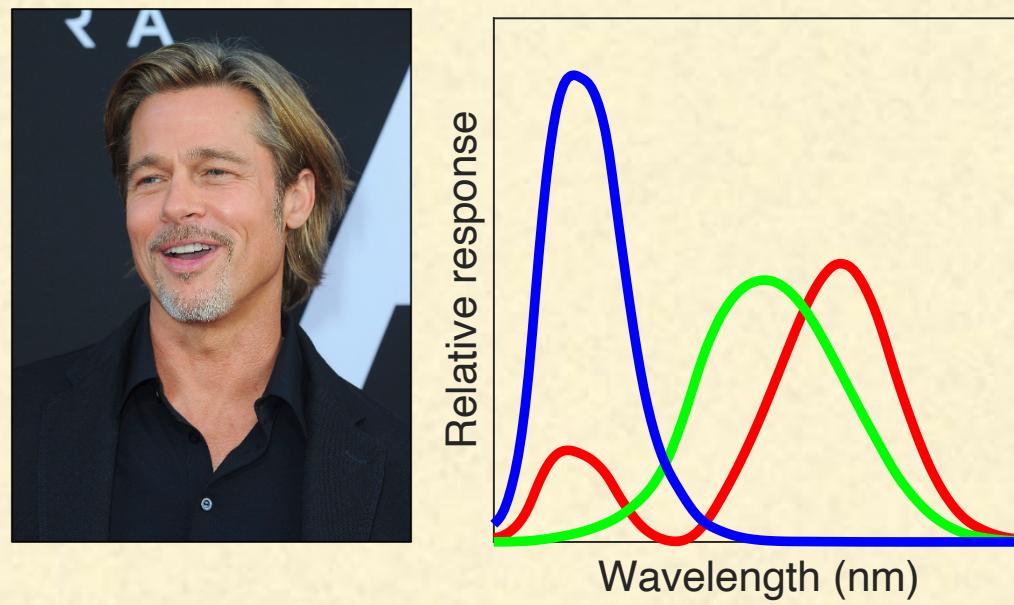
Think about which objects you need to include in the frame to calibrate.



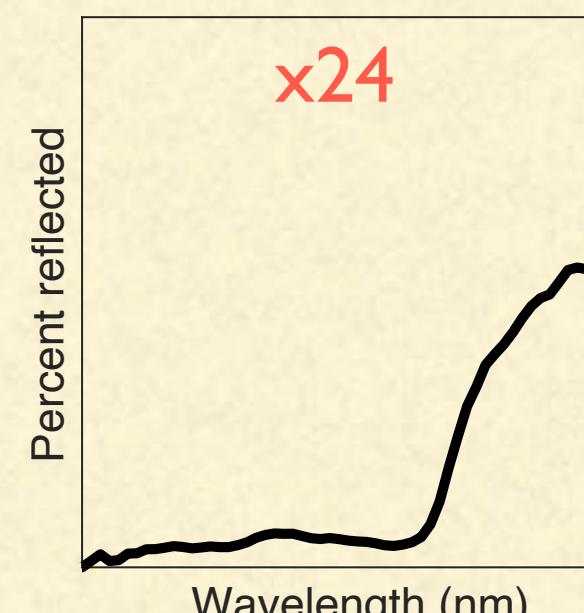
Camera RGB to XYZ transformation

Camera & illuminant specific !!!

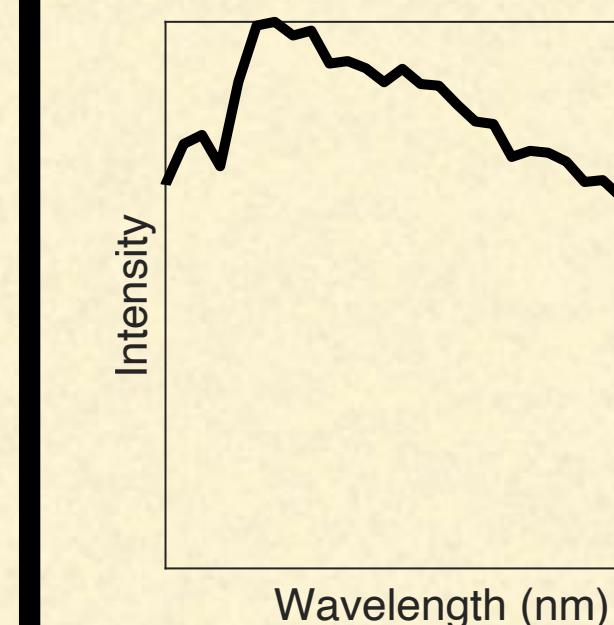
Spectral Response Curves



Reflectances of color chart patches

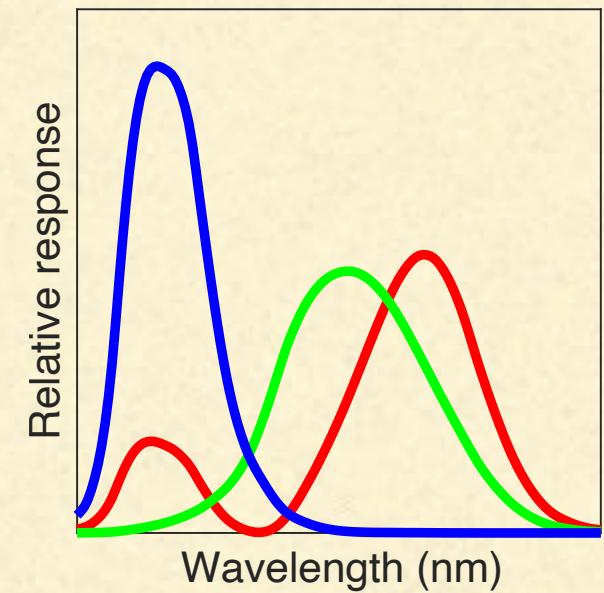


Illuminant Spectrum



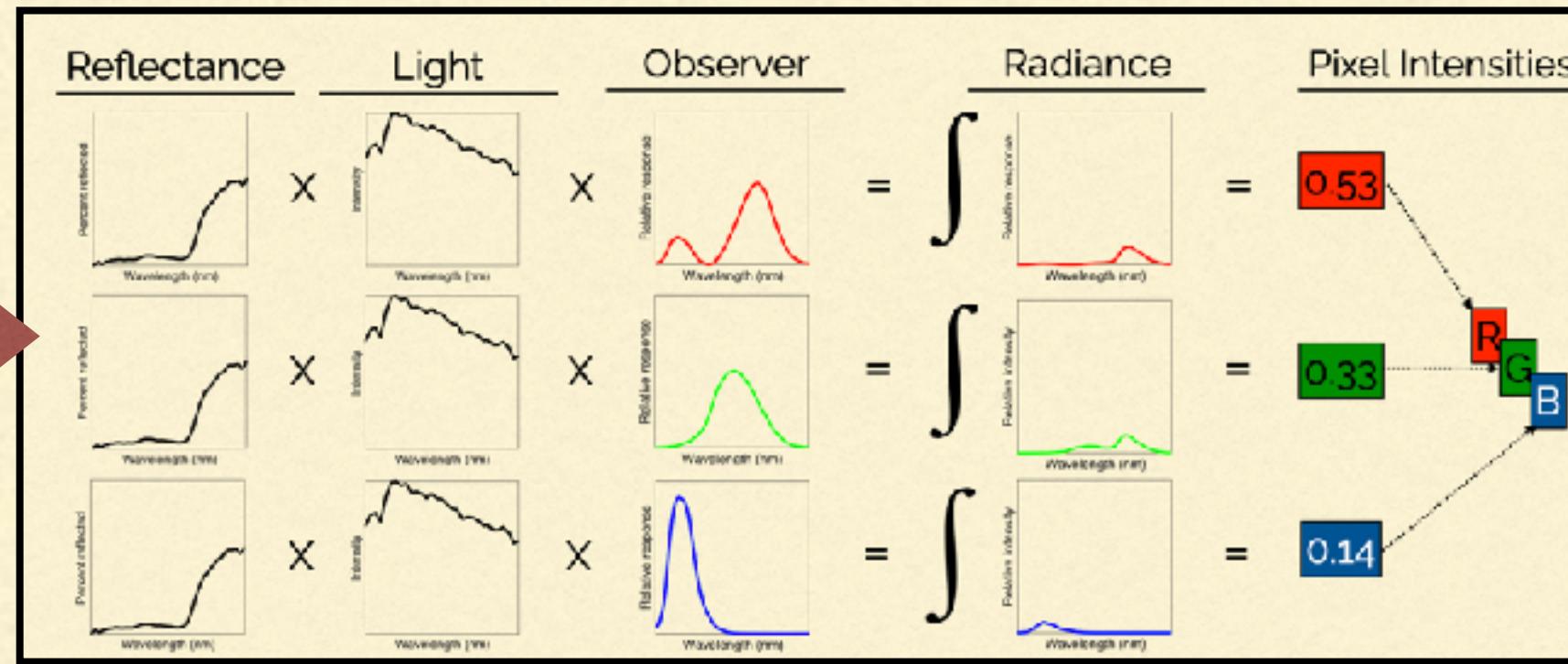
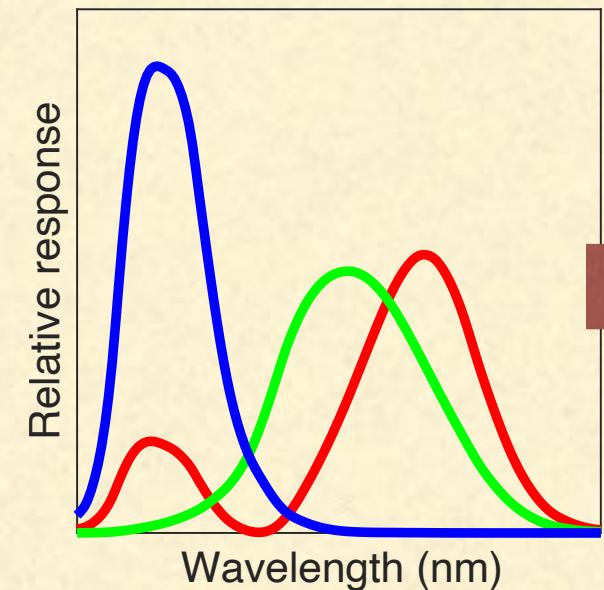
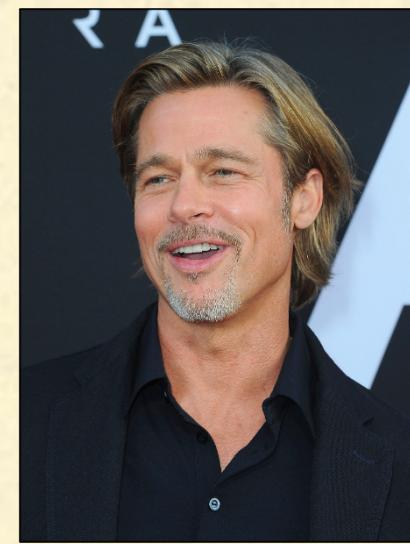
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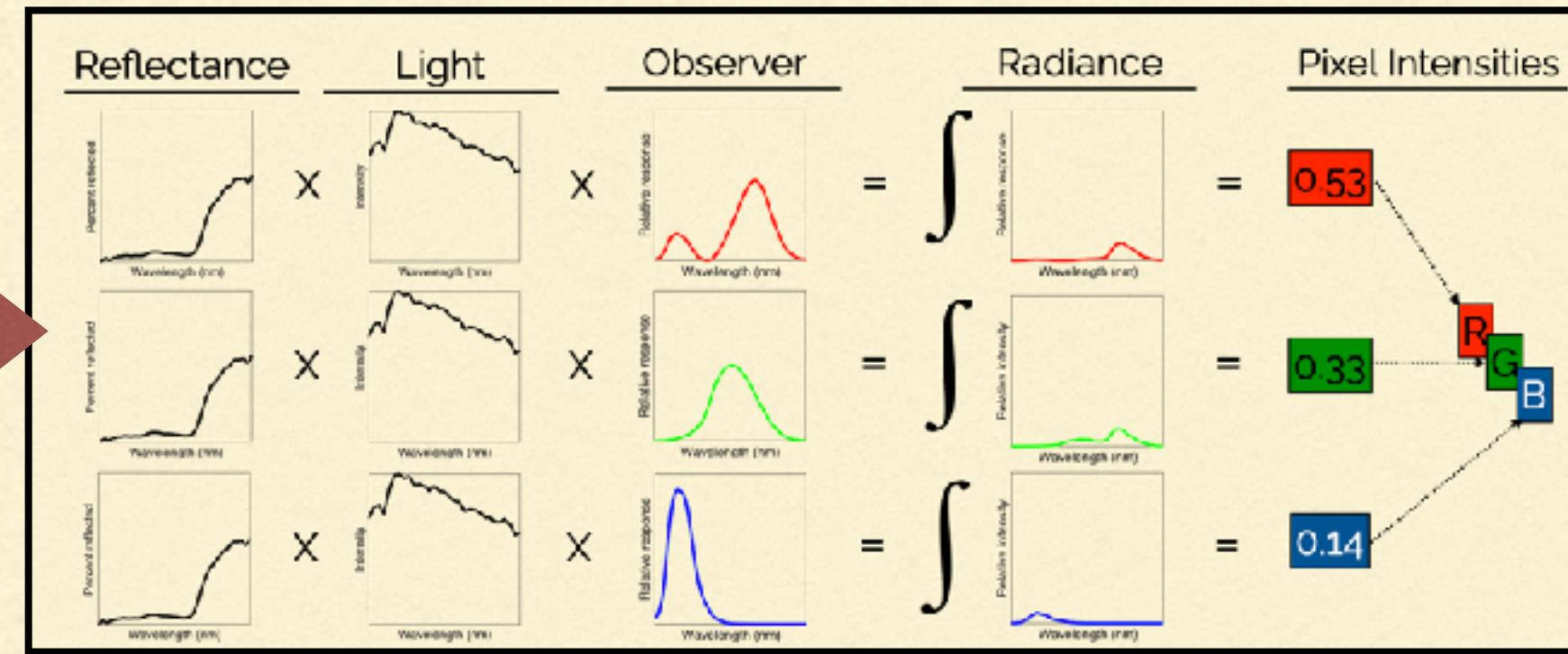
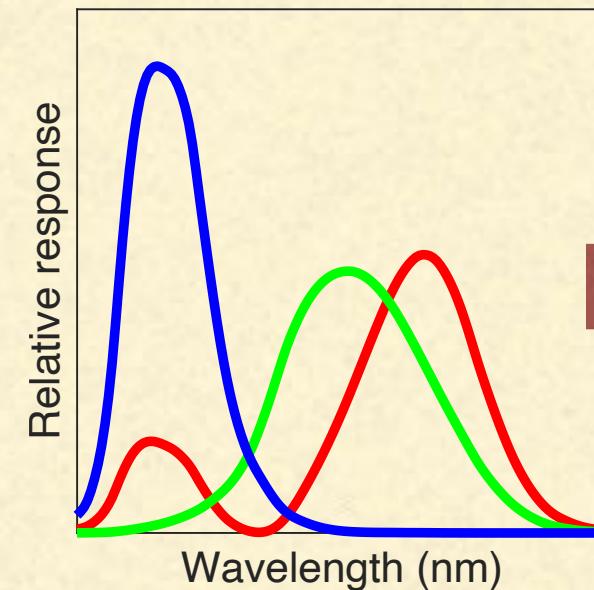
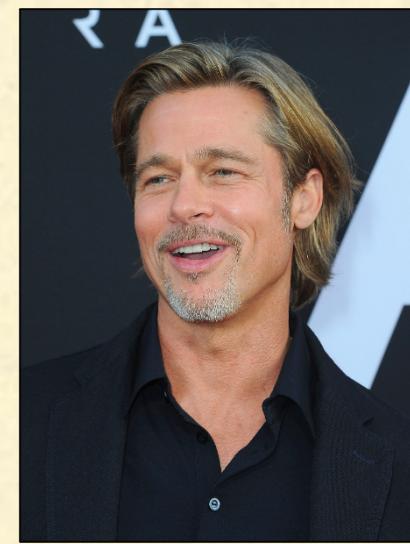
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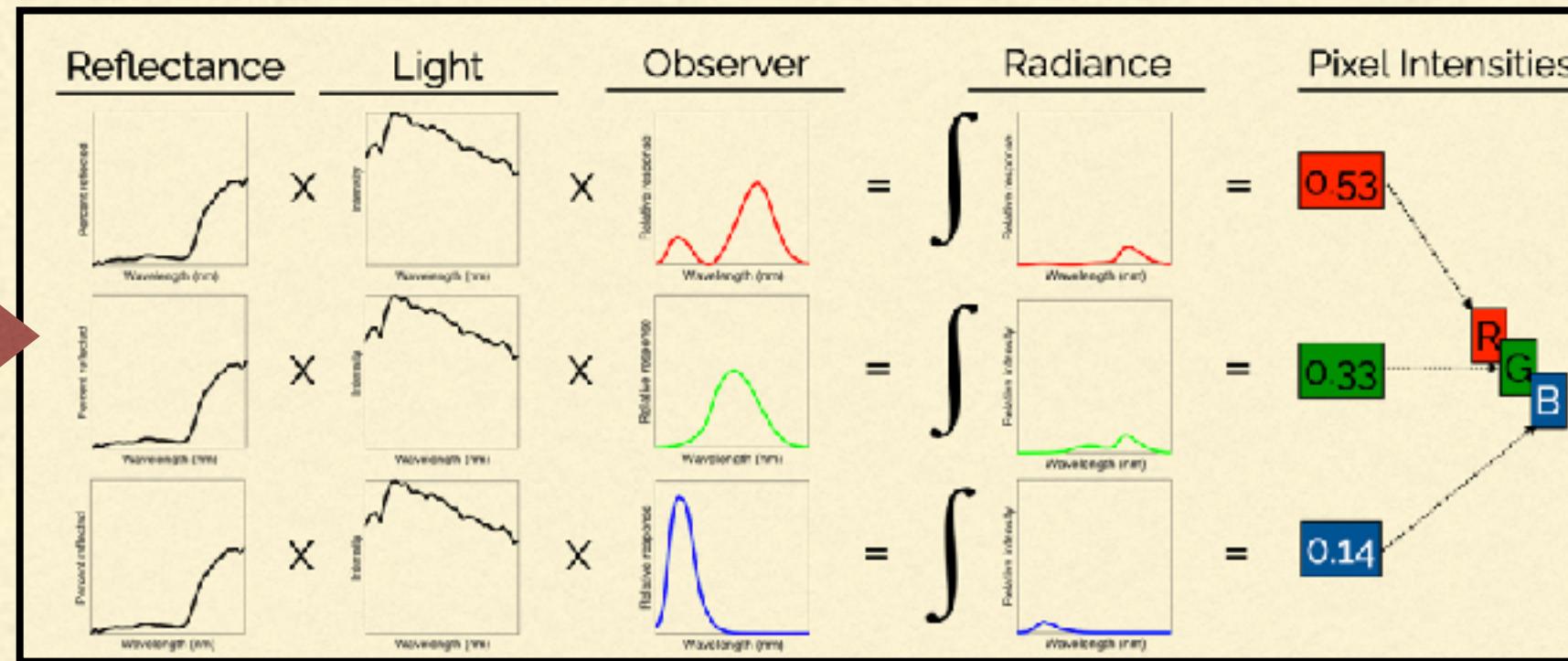
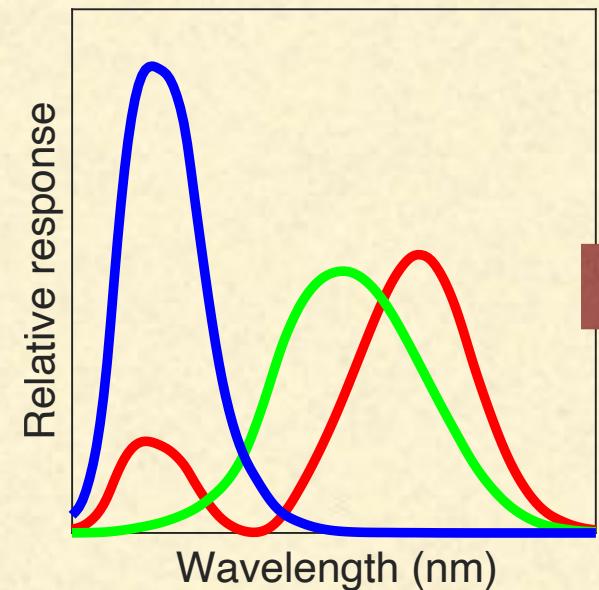
Camera & illuminant specific !!!



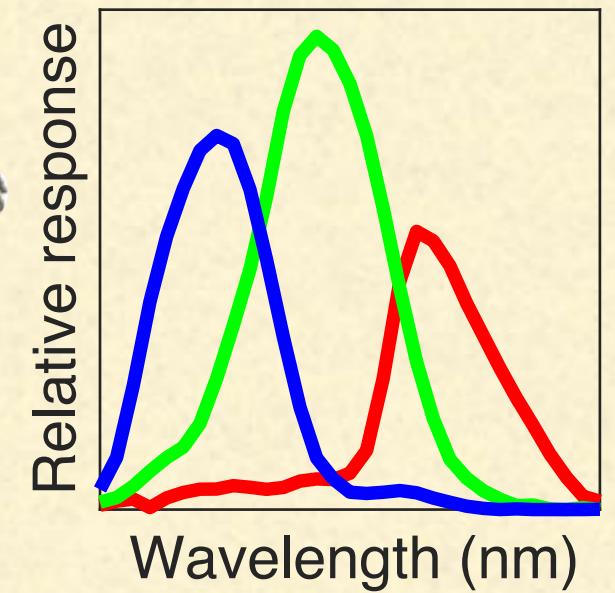
XYZ
[24 x 3]

Camera RGB to XYZ transformation

Camera & illuminant specific !!!

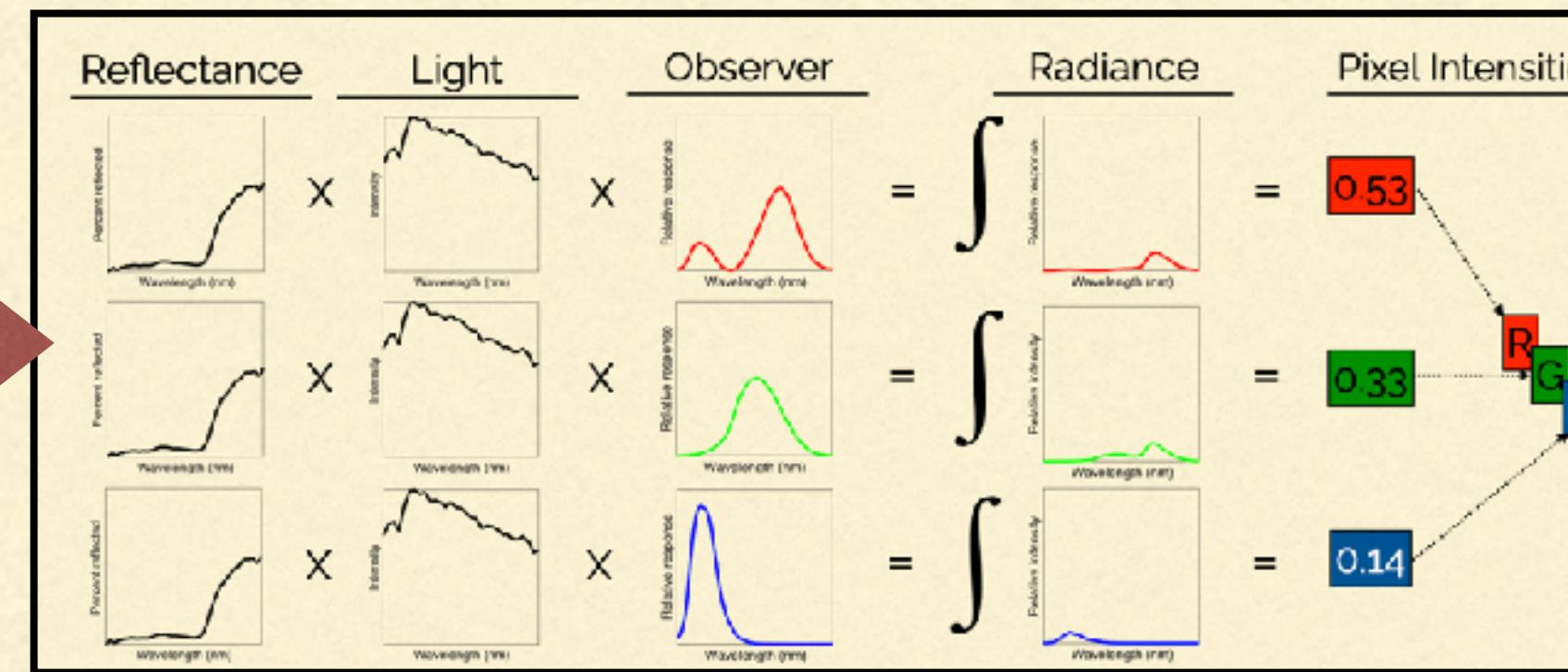
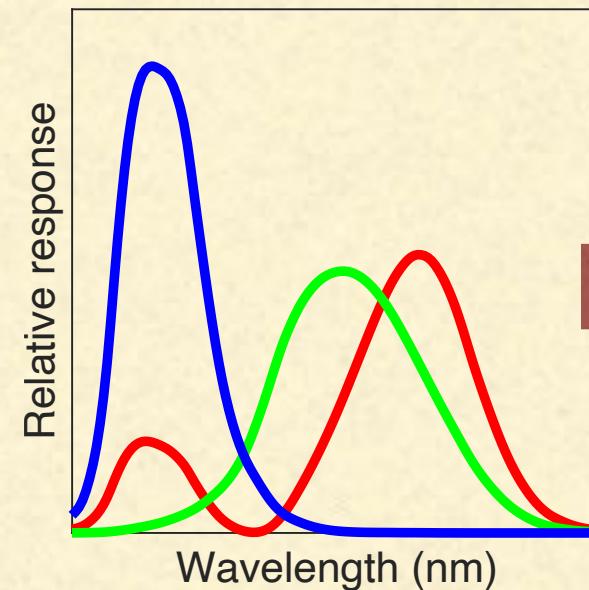


XYZ
[24 x 3]

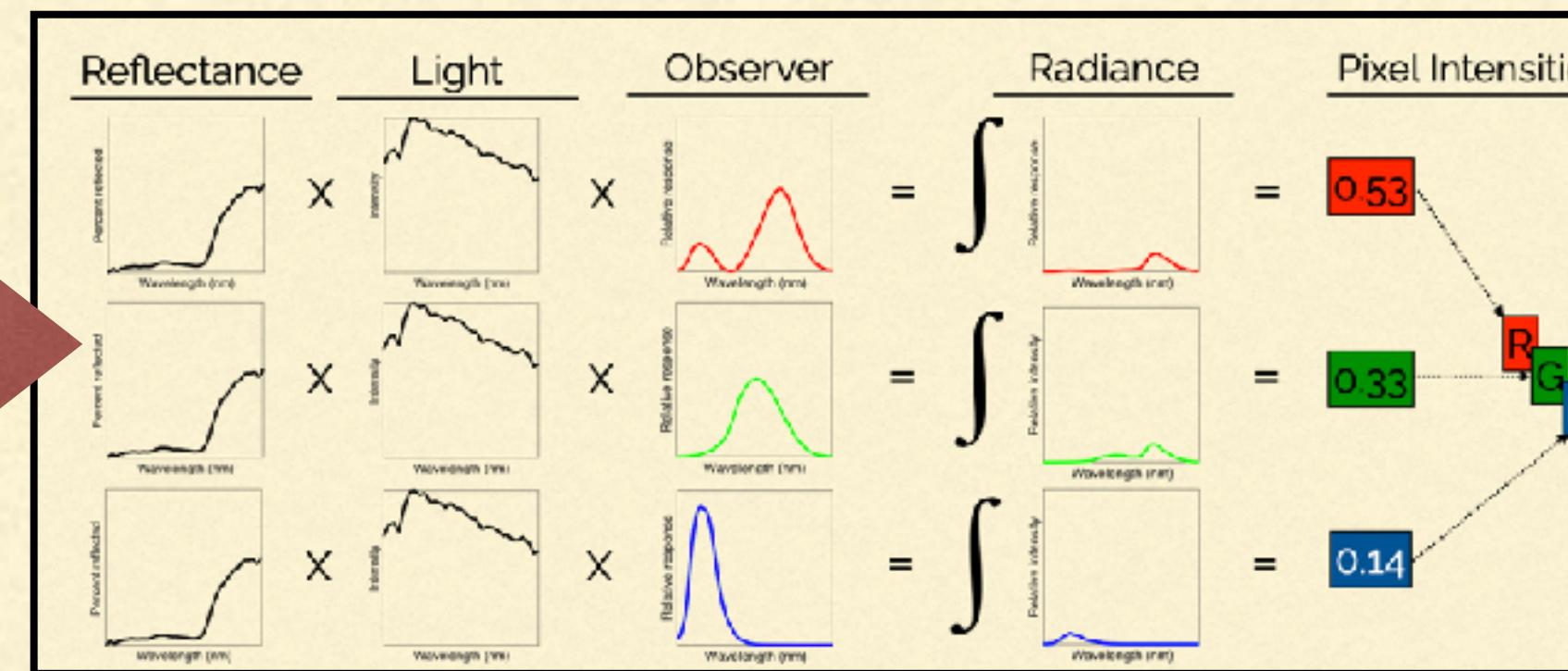
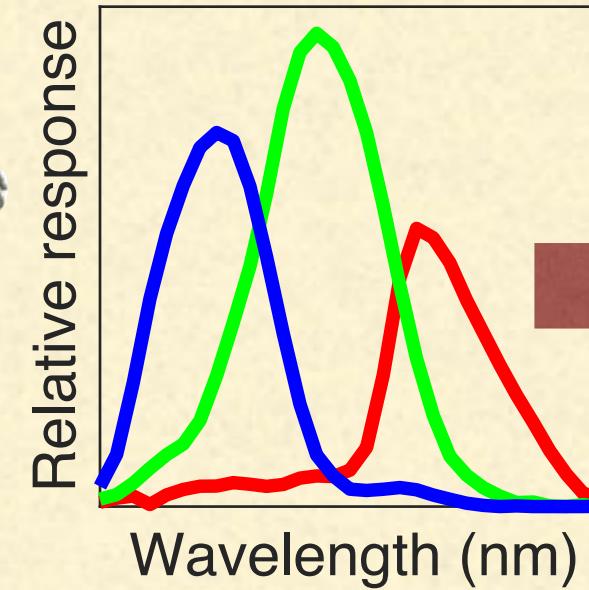


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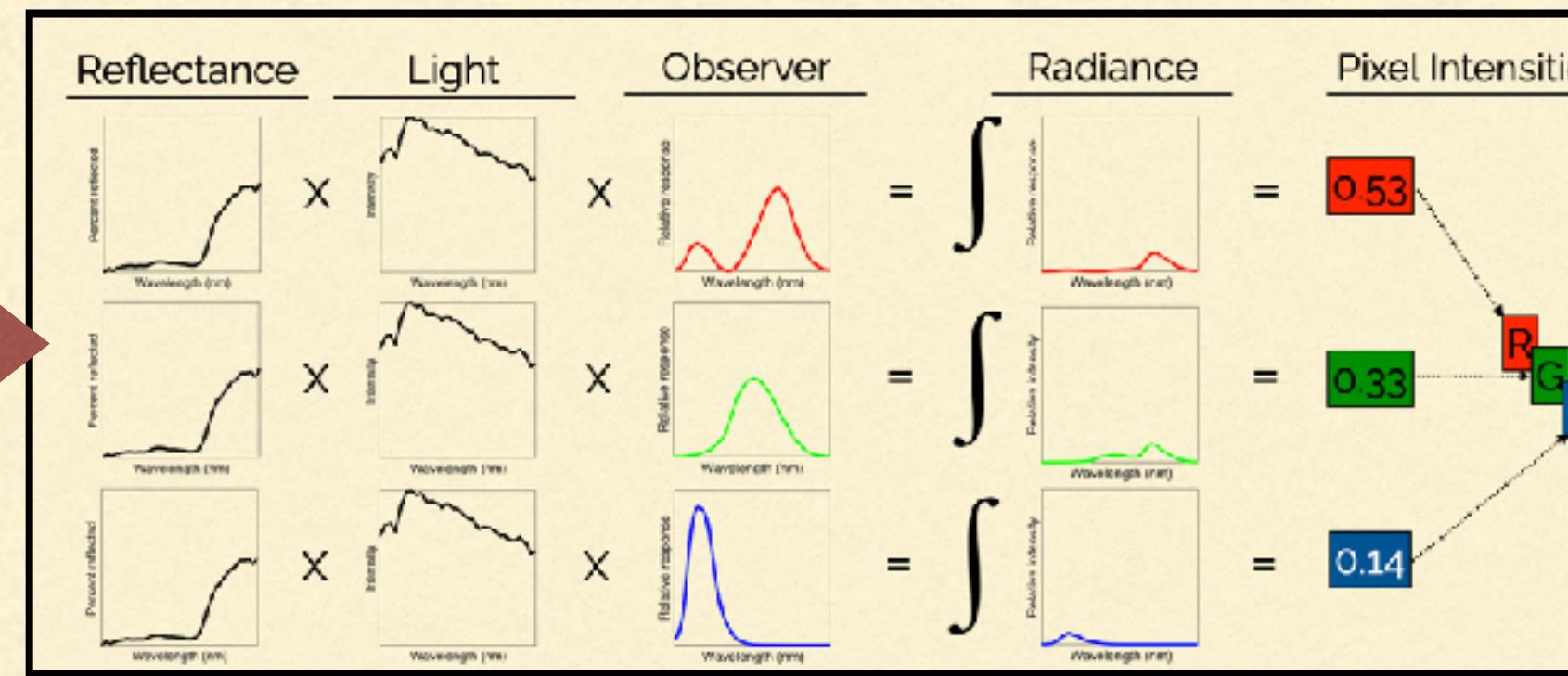
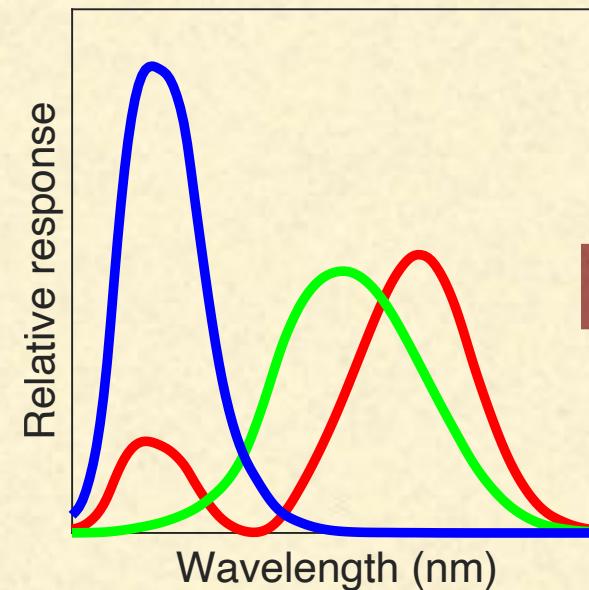


XYZ
[24 x 3]

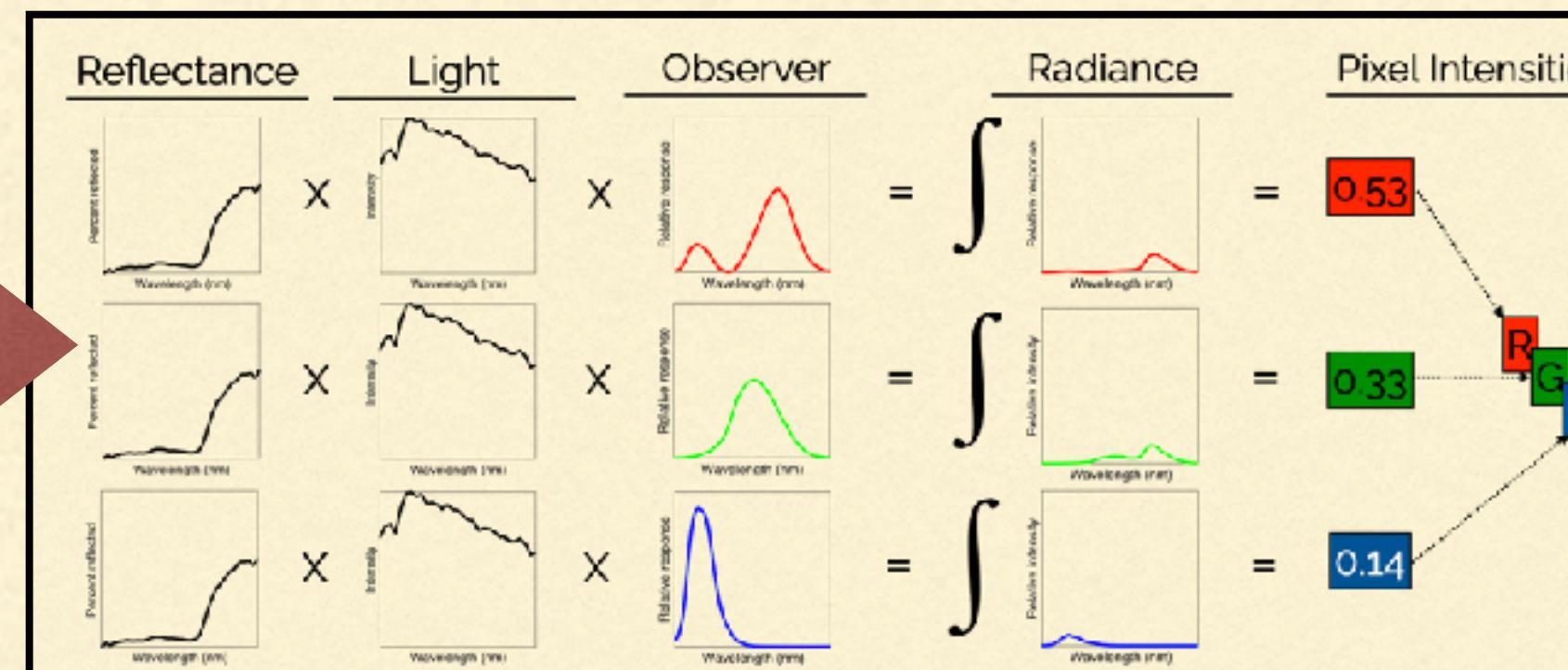
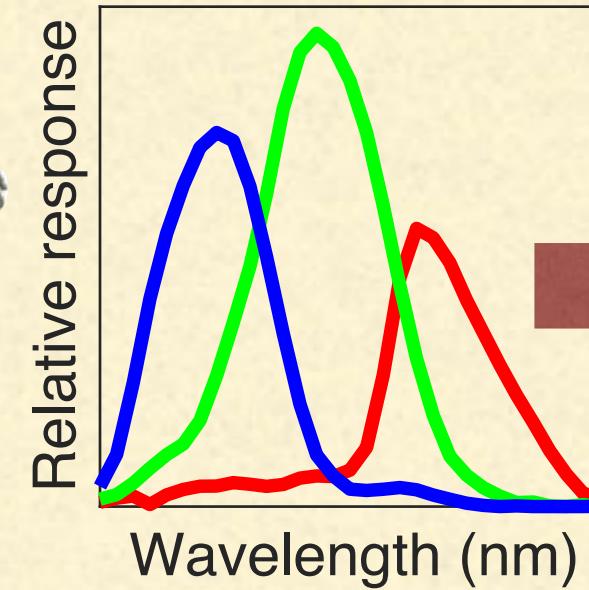


Camera RGB to XYZ transformation

Camera & illuminant specific !!!



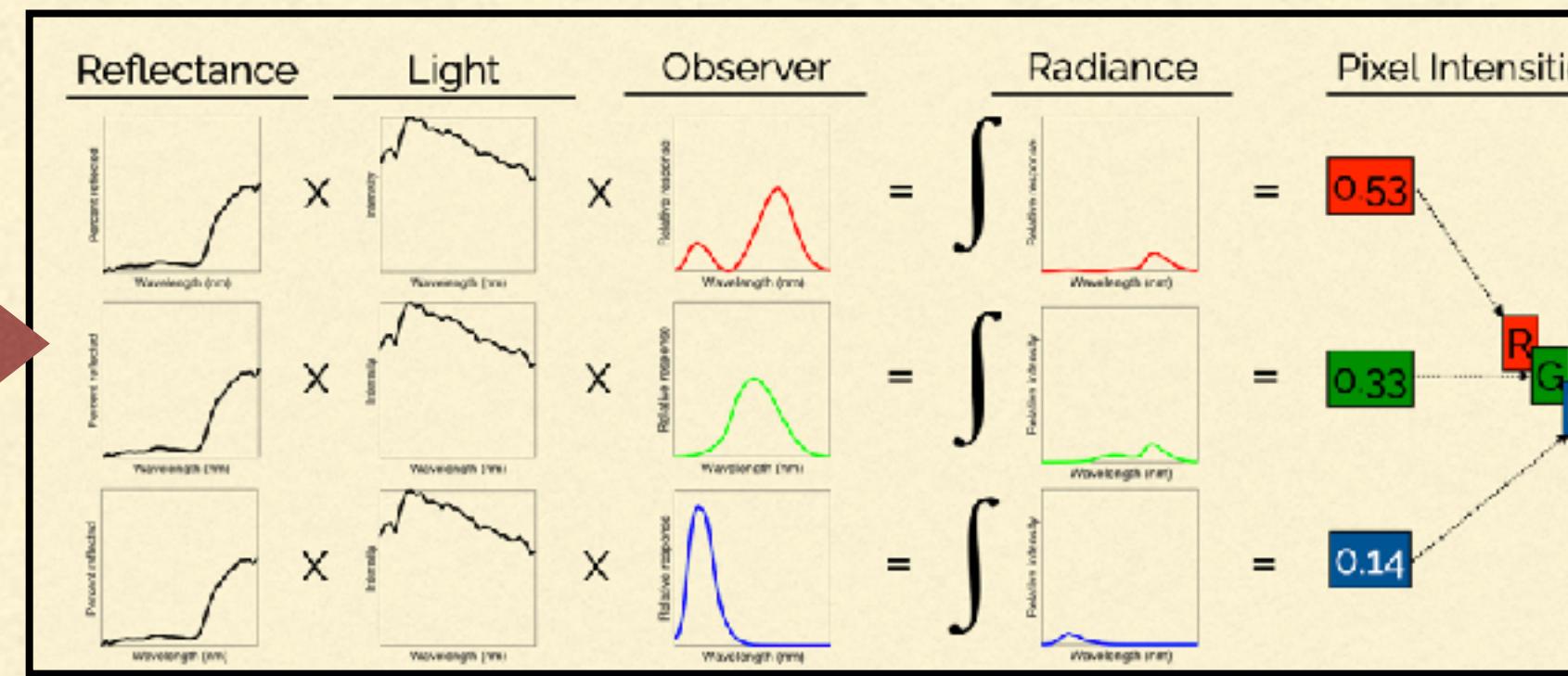
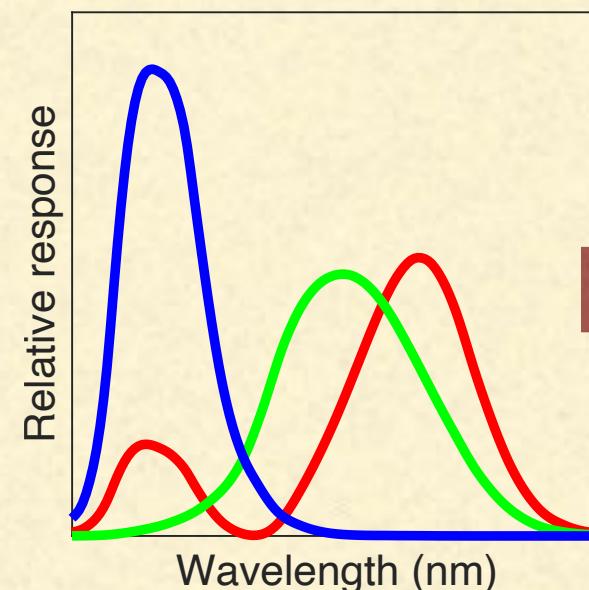
XYZ
[24 x 3]



RGB
[24 x 3]

Camera RGB to XYZ transformation

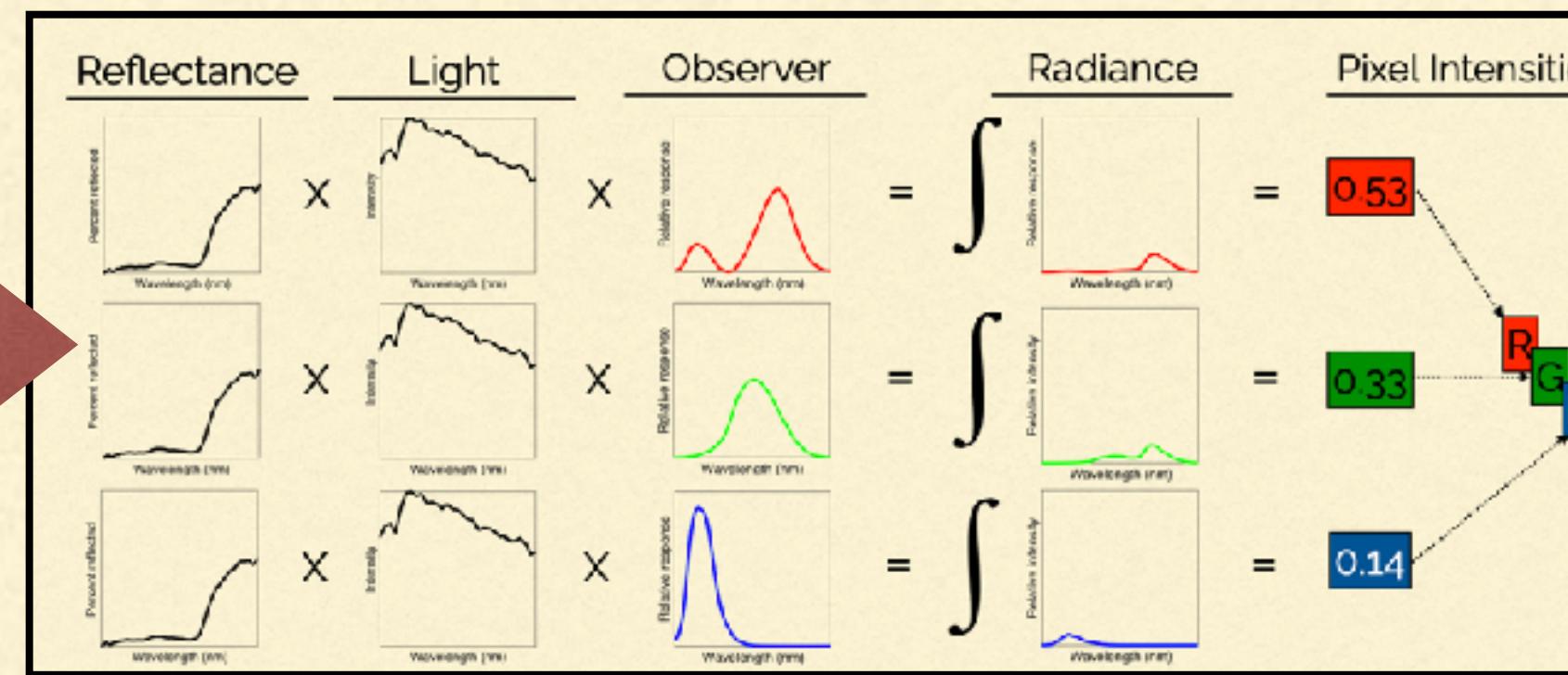
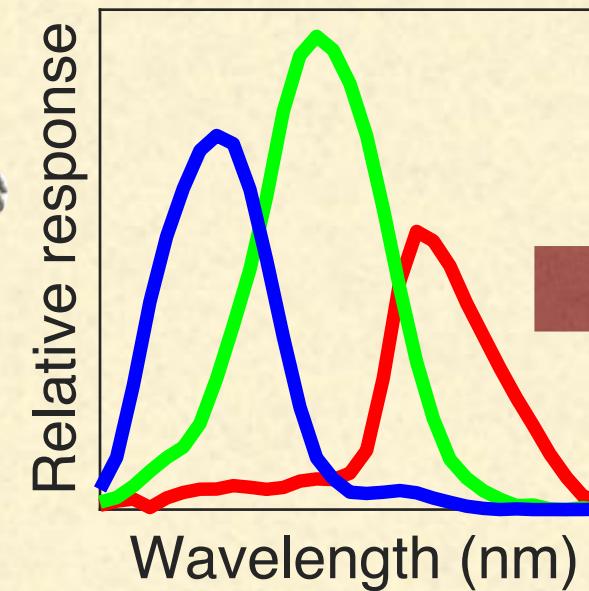
Camera & illuminant specific !!!



XYZ
[24 x 3]

M
[3 x 3]

$$[XYZ] = M[RGB]'$$



RGB
[24 x 3]

Camera RGB to XYZ transformation

Camera & illuminant specific !!!

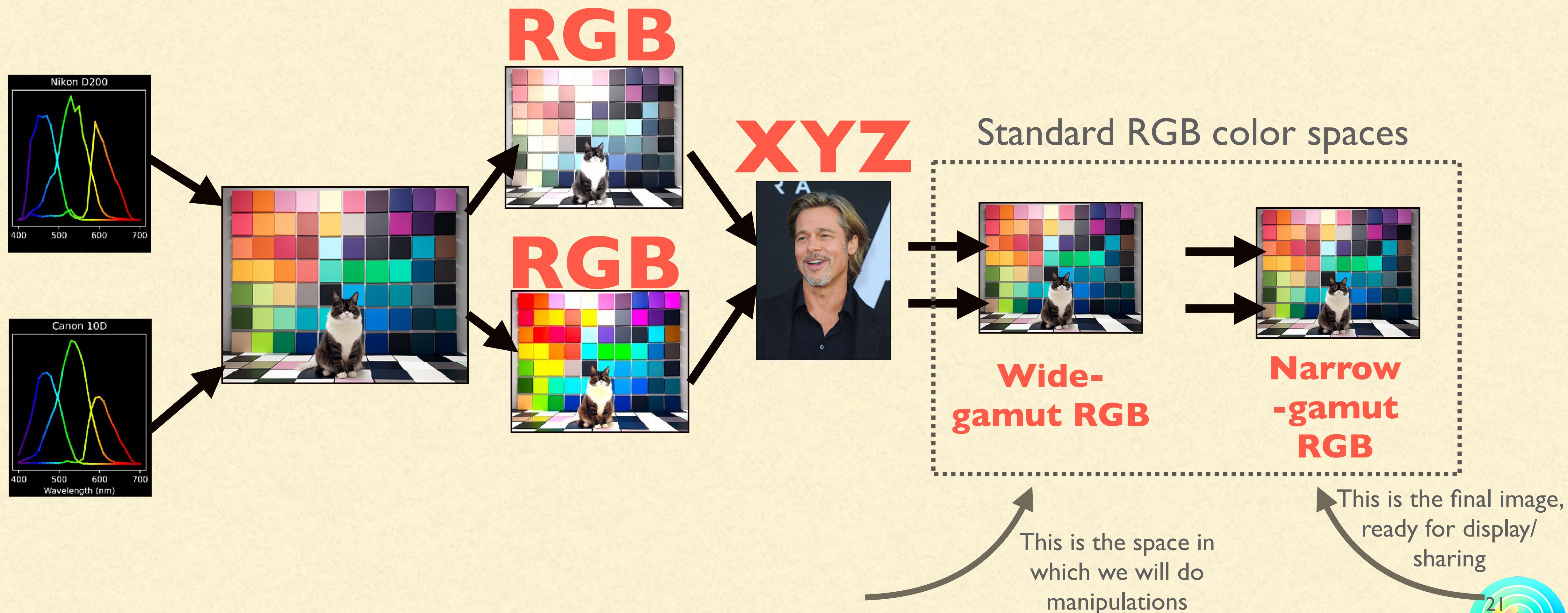


$$\frac{3 \times 3}{[XYZ] = M[RGB]'} \quad \frac{24 \times 3}{(24 \times 3)'} \quad \frac{3 \times 3}{}$$

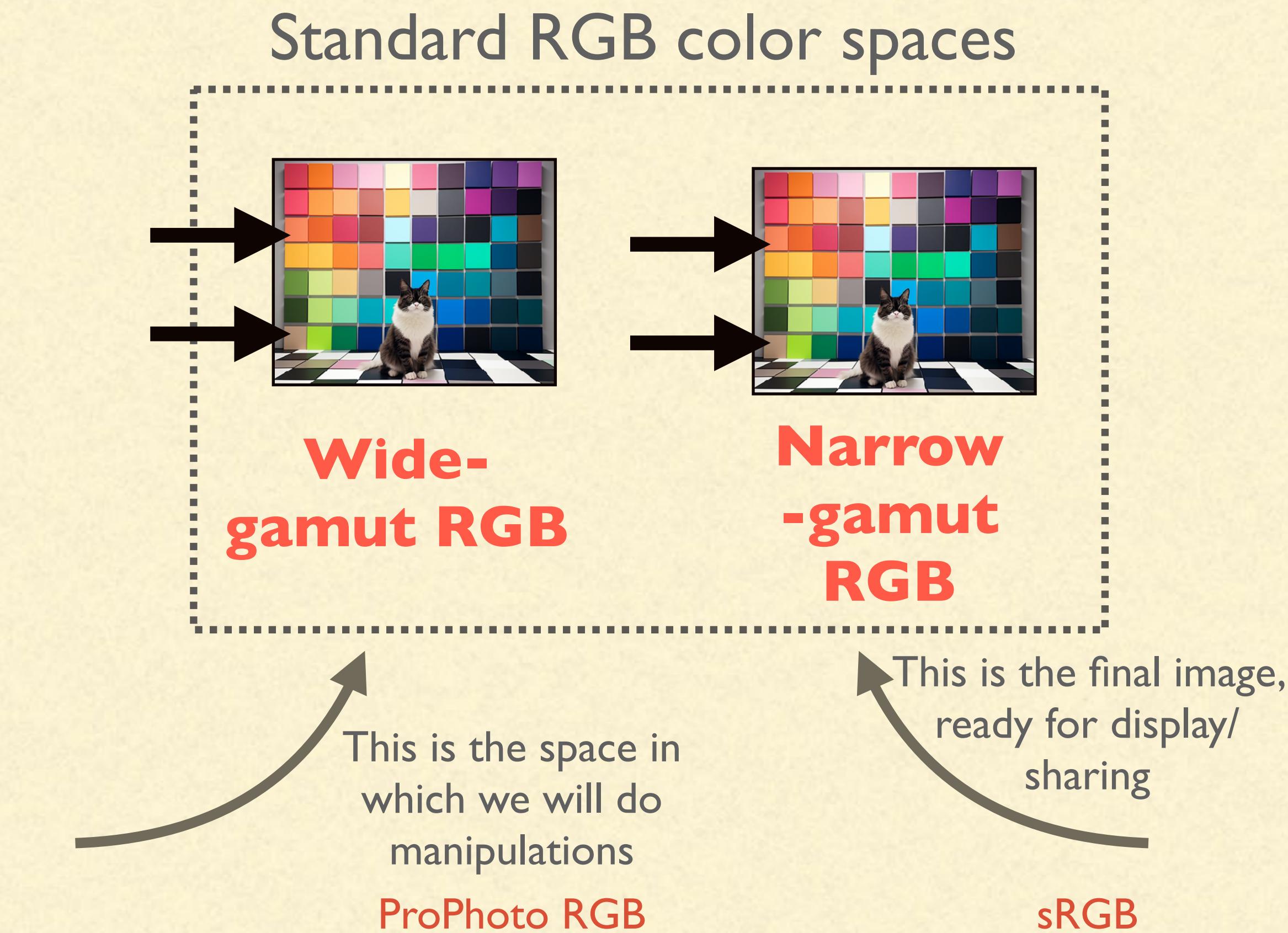
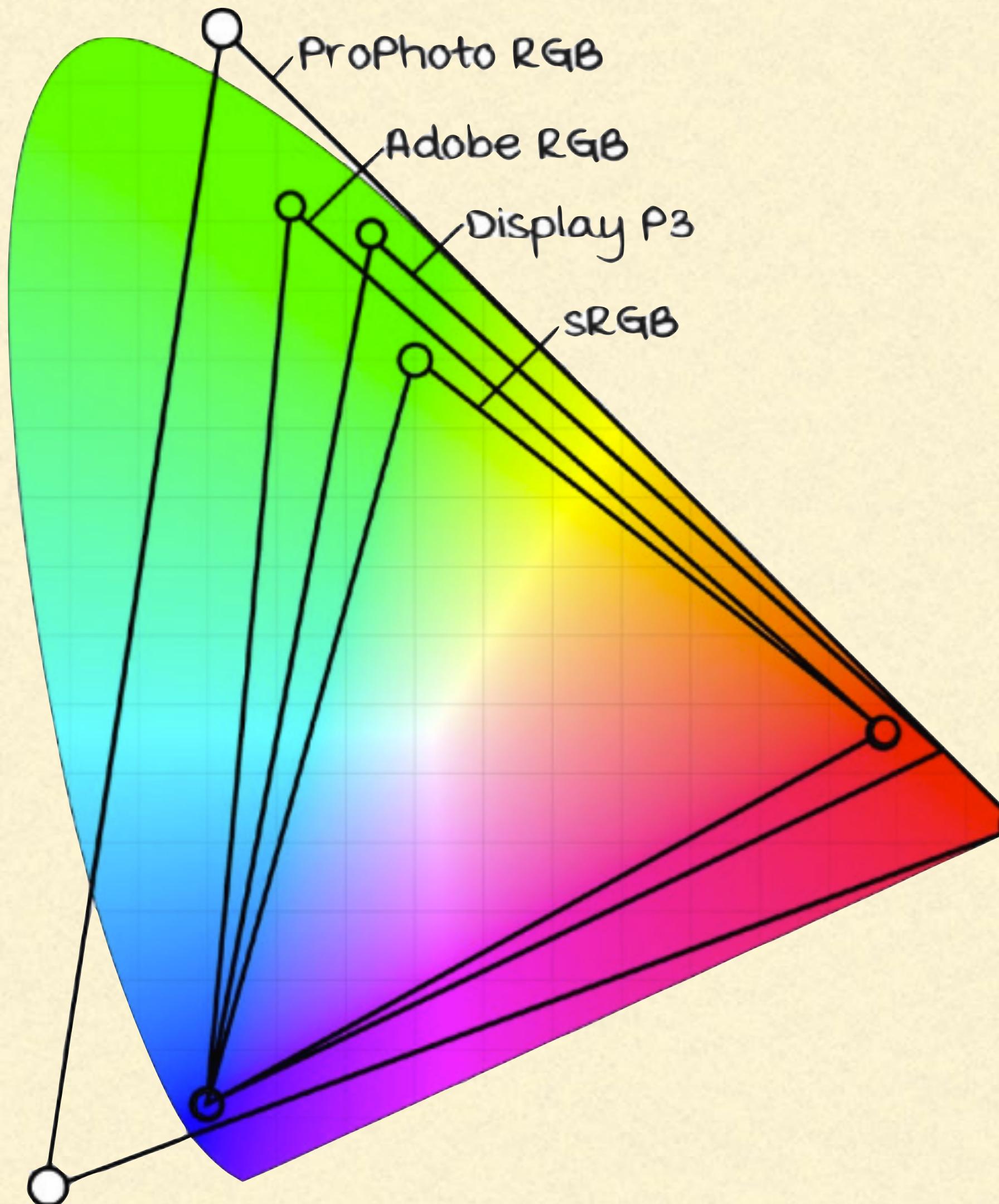
M is a 3×3 linear transformation matrix that maps camera-dependent RGB \rightarrow camera-independent XYZ under a given illuminant.

- M depends on:
- The illuminant
 - the number and type of patches used from the color chart
 - The scaling of the RGB/XYZ matrices

WHAT HAPPENS AFTER XYZ?



STANDARD RGB COLOR SPACES



STANDARD RGB COLOR SPACES

RGB Working Space	Reference White	RGB to XYZ [M]	XYZ to RGB [M] ⁻¹
Adobe RGB (1998)	D65	0.5767309 0.1855540 0.1881852 0.2973769 0.6273491 0.0752741 0.0270343 0.0706872 0.9911085	2.0413690 -0.5649464 -0.3446944 -0.9692660 1.8760108 0.0415560 0.0134474 -0.1183897 1.0154096
AppleRGB	D65	0.4497208 0.3162406 0.1844926 0.2446525 0.6720203 0.0833192 0.0251848 0.1411824 0.9224628	2.9515373 -1.2894116 -0.4730445 -1.0851093 1.9908566 0.0372026 0.0854934 -0.2694964 1.0912975
Best RGB	D50	0.6326696 0.2045558 0.1269946 0.2284569 0.7373523 0.0341908 0.0000000 0.0095142 0.8156958	1.7552599 -0.4836786 -0.2530000 -0.5441336 1.5068789 0.0215528 0.0063467 -0.0175761 1.2256959
Beta RGB	D50	0.6712537 0.1745834 0.1183829 0.3032726 0.6637861 0.0329413 0.0000000 0.0407010 0.7845090	1.6832270 -0.4282363 -0.2360185 -0.7710229 1.7065571 0.0446900 0.0400013 -0.0885376 1.2723640
Bruce RGB	D65	0.4674162 0.2944512 0.1886026 0.2410115 0.6835475 0.0754410 0.0219101 0.0736128 0.9933071	2.7454669 -1.1358136 -0.4350269 -0.9692660 1.8760108 0.0415560 0.0112723 -0.1139754 1.0132541
CIE RGB	E	0.4887180 0.3106803 0.2006017 0.1762044 0.8129847 0.0108109 0.0000000 0.0102048 0.9897952	2.3706743 -0.9000405 -0.4706338 -0.5138850 1.4253036 0.0885814 0.0052982 -0.0146949 1.0093968
ColorMatch RGB	D50	0.5093439 0.3209071 0.1339691 0.2748840 0.6581315 0.0669845 0.0242545 0.1087821 0.6921735	2.6422874 -1.2234270 -0.3930143 -1.1119763 2.0590183 0.0159614 0.0821699 -0.2807254 1.4559877
Don RGB 4	D50	0.6457711 0.1933511 0.1250978 0.2783496 0.6879702 0.0336802 0.0037113 0.0179861 0.8035125	1.7603902 -0.4801198 -0.2536126 -0.7126208 1.6527432 0.0416715 0.0078207 -0.0347411 1.2447743
ECI RGB	D50	0.6502043 0.1780774 0.1359384 0.3202499 0.6020711 0.0776791 0.0000000 0.0678390 0.7573710	1.7827618 -0.4969847 -0.2690101 -0.9593623 1.9477962 -0.0275807 0.0859317 -0.1744674 1.3228273
Ekta Space PS5	D50	0.5938914 0.2729801 0.0973485 0.2606286 0.7349465 0.0044249 0.0000000 0.0419969 0.7832131	2.0043819 -0.7304844 -0.2450052 -0.7110285 1.6202126 0.0792227 0.0381263 -0.0868780 1.2725438
NTSC RGB	C	0.6068909 0.1735011 0.2003480 0.2989164 0.5865990 0.1144845 0.0000000 0.0660957 1.1162243	1.9099961 -0.5324542 -0.2882091 -0.9846663 1.9991710 -0.0283082 0.0583056 -0.1183781 0.8975535
PAL/SECAM RGB	D65	0.4306190 0.3415419 0.1783091 0.2220379 0.7066384 0.0713236 0.0201853 0.1295504 0.9390944	3.0628971 -1.3931791 -0.4757517 -0.9692660 1.8760108 0.0415560 0.0678775 -0.2288548 1.0693490

RGB Working Space	Reference White	RGB to XYZ [M]	XYZ to RGB [M] ⁻¹
ProPhoto RGB	D50	0.7976749 0.1351917 0.0313534 0.2880402 0.7118741 0.0000857 0.0000000 0.0000000 0.8252100	1.3459433 -0.2556075 -0.0511118 -0.5445989 1.5081673 0.0205351 0.0000000 0.0000000 1.2118128
SMPTE-C RGB	D65	0.3935891 0.3652497 0.1916313 0.2124132 0.7010437 0.0865432 0.0187423 0.1119313 0.9581563	3.5053960 -1.7394894 -0.5439640 -1.0690722 1.9778245 0.0351722 0.0563200 -0.1970226 1.0502026
sRGB	D65	0.4124564 0.3575761 0.1804375 0.2126729 0.7151522 0.0721750 0.0193339 0.1191920 0.9503041	3.2404542 -1.5371385 -0.4985314 -0.9692660 1.8760108 0.0415560 0.0556434 -0.2040259 1.0572252
Wide Gamut RGB	D50	0.7161046 0.1009296 0.1471858 0.2581874 0.7249378 0.0168748 0.0000000 0.0517813 0.7734287	1.4628067 -0.1840623 -0.2743606 -0.5217933 1.4472381 0.0677227 0.0349342 -0.0968930 1.2884099

All related to XYZ color space with a 3 x 3 linear transform!

STANDARD RGB COLOR SPACES

$$\underline{[XYZ]} = M \underline{[RGB']}$$

Your image in XYZ
color space

Your image in
camera color space

$$\underline{[RGB_{standard}]} = T \underline{[XYZ]}$$

Your image in a standard RGB
color space, eg ProPhoto

Your image in
XYZ color space

T is a 3x3 linear transformation matrix that maps colors
from the XYZ space to a standard RGB space.

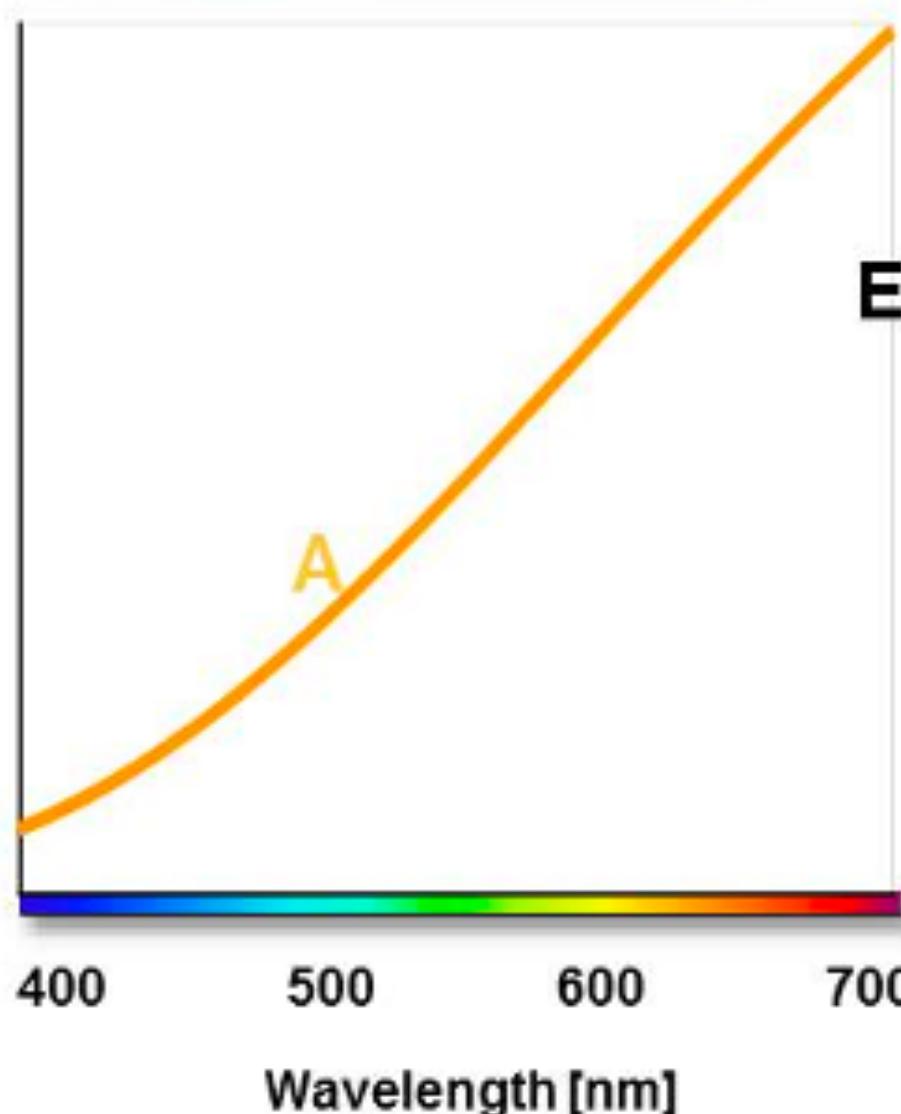
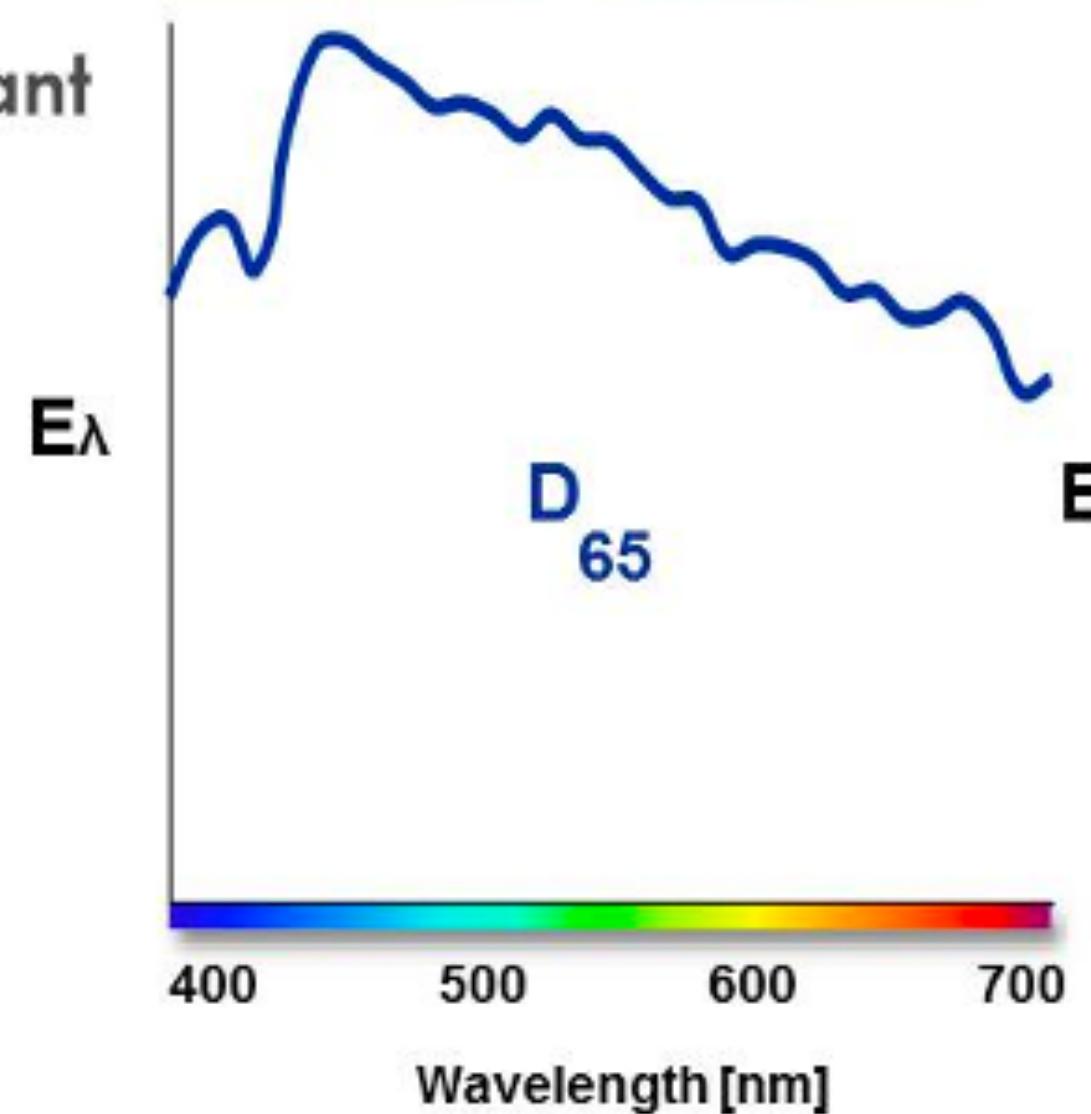


CIE STANDARD ILLUMINANTS : D65, A

Source



Illuminant



From the ISO standards:

a) CIE standard illuminant A

This is intended to represent typical, domestic, tungsten-filament lighting. Its relative spectral power distribution is that of a Planckian radiator at a temperature of approximately 2856 K. CIE standard illuminant A should be used in all applications of colorimetry involving the use of incandescent lighting, unless there are specific reasons for using a different illuminant.

b) CIE standard illuminant D65

This is intended to represent average daylight and has a correlated colour temperature of approximately 6500 K. CIE standard illuminant D65 should be used in all colorimetric calculations requiring representative daylight, unless there are specific reasons for using a different illuminant. Variations in the relative spectral power distribution of daylight are known to occur, particularly in the ultraviolet spectral region, as a function of season, time of day, and geographic location.

Illuminant is a mathematical description of a light source. It has:

- a spectral distribution
- Correlated color temperature

CIE DAYLIGHT MODEL

$$L_T(\lambda) = L_0(\lambda) + M_1(T)L_1(\lambda) + M_2(T)L_2(\lambda)$$

where

$$M_1 = (-1.3515 - 1.7703x_D(T) + 5.9114y_D(T))/M ,$$

$$M_2 = (0.0300 - 31.4424x_D(T) + 30.0717y_D(T))/M ,$$

$$M = 0.0241 + 0.2562x_D(T) - 0.7341y_D(T) ,$$

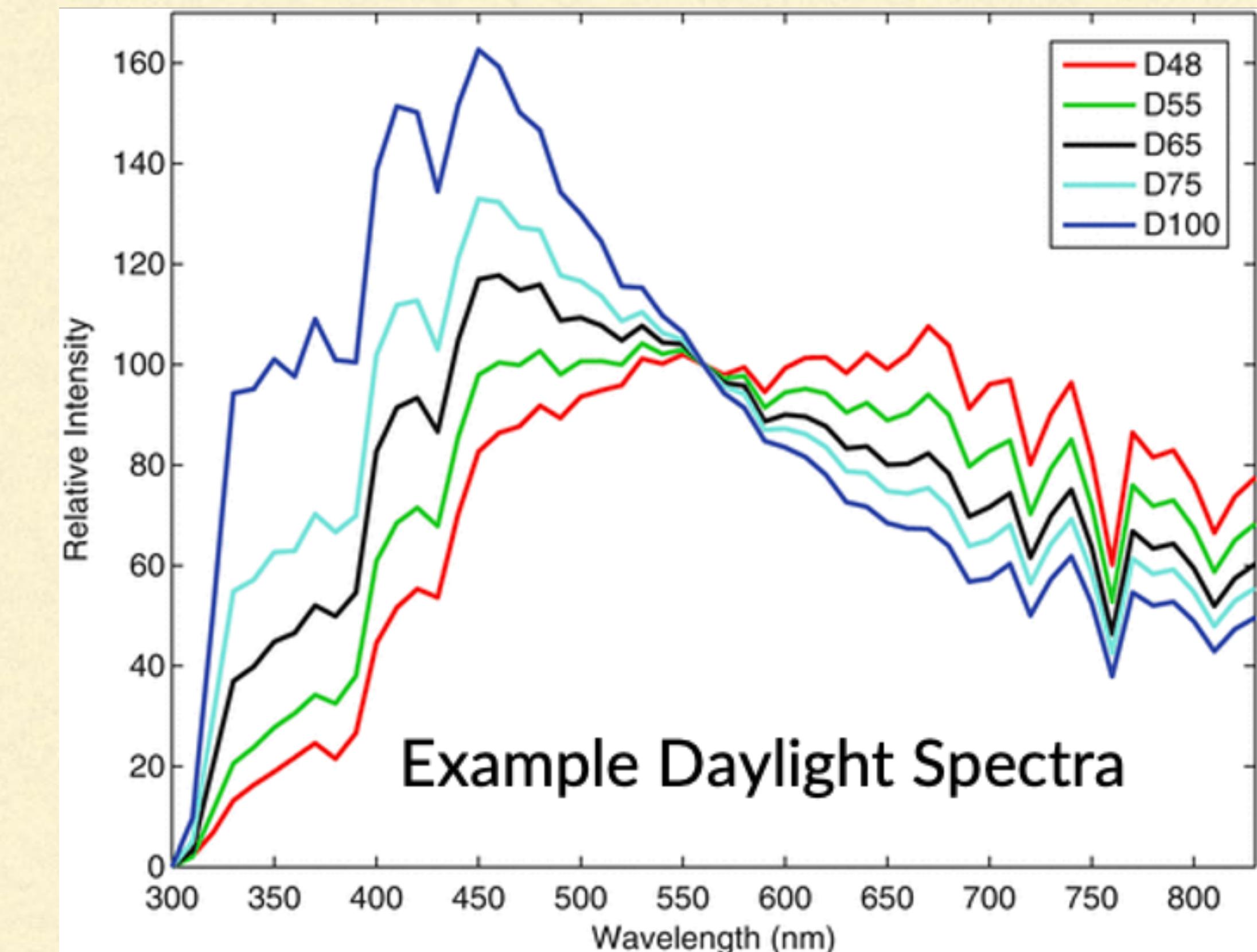
$$y_D(T) = -3.000x_D(T)^2 + 2.879x_D(T) - 0.275,$$

$$x_D(T) = \begin{cases} 0.244063 + 0.09911 \frac{10^3}{T} + 2.9678 \frac{10^6}{T^2} - 4.6070 \frac{10^9}{T^3} & \text{if } 4000\text{K} \leq T \leq 7000\text{K} \\ 0.237040 + 0.24748 \frac{10^3}{T} + 1.9018 \frac{10^6}{T^2} - 2.0064 \frac{10^9}{T^3} & \text{if } 7000\text{K} < T \leq 25000\text{K} \end{cases}$$

T: correlated color temperature: a way to describe the “warmth” or “coolness” of light sources.

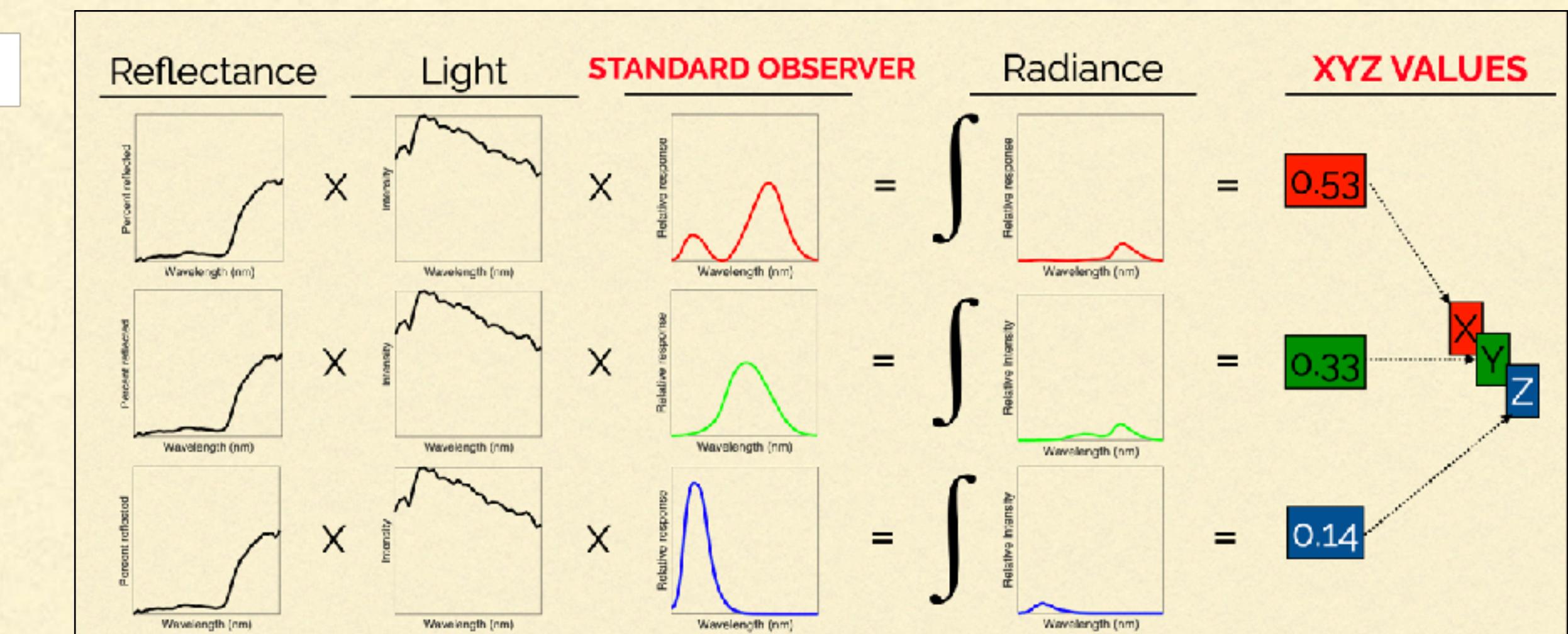
CCT is the temperature at which a theoretical blackbody radiator would emit light of same color.

A blackbody is an “ideal” object that absorbs all electromagnetic radiation and re-emits radiation based on its temperature.



WHITE POINT OF AN ILLUMINANT

RGB Working Space	Reference White	RGB to XYZ [M]	XYZ to RGB [M^{-1}]
ProPhoto RGB	D50	0.7976749 0.1351917 0.0313534 0.2880402 0.7118741 0.0000857 0.0000000 0.0000000 0.8252100	1.3459433 -0.2556075 -0.0511118 -0.5445989 1.5081673 0.0205351 0.0000000 0.0000000 1.2118128
SMPTE-C RGB	D65	0.3935891 0.3652497 0.1916313 0.2124132 0.7010437 0.0865432 0.0187423 0.1119313 0.9581563	3.5053960 -1.7394894 -0.5439640 -1.0690722 1.9778245 0.0351722 0.0563200 -0.1970226 1.0502026
sRGB	D65	0.4124564 0.3575761 0.1804375 0.2126729 0.7151522 0.0721750 0.0193339 0.1191920 0.9503041	3.2404542 -1.5371385 -0.4985314 -0.9692660 1.8760108 0.0415560 0.0556434 -0.2040259 1.0572252
Wide Gamut RGB	D50	0.7161046 0.1009296 0.1471858 0.2581874 0.7249378 0.0168748 0.0000000 0.0517813 0.7734287	1.4628067 -0.1840623 -0.2743606 -0.5217933 1.4472381 0.0677227 0.0349342 -0.0968930 1.2884099



Calculate the XYZ values just the same way you would, but assume reflectance is 100% uniform (i.e., a perfect white surface).

CHROMATIC ADAPTATION TRANSFORM (CAT)

- ▶ Converting from the white point of one illuminant to another (source → destination)
- ▶ A kind of white-balancing but more elaborate as it relates to color constancy
- ▶ Once again, a sequence of simple 3×3 linear transforms

$$\begin{bmatrix} X_D \\ Y_D \\ Z_D \end{bmatrix} = [M] \begin{bmatrix} X_S \\ Y_S \\ Z_S \end{bmatrix}$$

**XYZ white point
of your
destination color
space (eg, D50
for prophoto)**

**XYZ white point
of your source
illuminant**

$$[M] = [M_A]^{-1} \begin{bmatrix} \rho_D/\rho_S & 0 & 0 \\ 0 & \gamma_D/\gamma_S & 0 \\ 0 & 0 & \beta_D/\beta_S \end{bmatrix} [M_A]$$

$$\begin{bmatrix} \rho_S \\ \gamma_S \\ \beta_S \end{bmatrix} = [M_A] \begin{bmatrix} X_{WS} \\ Y_{WS} \\ Z_{WS} \end{bmatrix}$$

$$\begin{bmatrix} \rho_D \\ \gamma_D \\ \beta_D \end{bmatrix} = [M_A] \begin{bmatrix} X_{WD} \\ Y_{WD} \\ Z_{WD} \end{bmatrix}$$

Method	[M _A]	[M _A] ⁻¹
XYZ Scaling	$\begin{bmatrix} 1.0000000 & 0.0000000 & 0.0000000 \\ 0.0000000 & 1.0000000 & 0.0000000 \\ 0.0000000 & 0.0000000 & 1.0000000 \end{bmatrix}$	$\begin{bmatrix} 1.0000000 & 0.0000000 & 0.0000000 \\ 0.0000000 & 1.0000000 & 0.0000000 \\ 0.0000000 & 0.0000000 & 1.0000000 \end{bmatrix}$
Bradford	$\begin{bmatrix} 0.8951000 & 0.2664000 & -0.1614000 \\ -0.7502000 & 1.7135000 & 0.0367000 \\ 0.0389000 & -0.0685000 & 1.0296000 \end{bmatrix}$	$\begin{bmatrix} 0.9869929 & -0.1470543 & 0.1599627 \\ 0.4323053 & 0.5183603 & 0.0492912 \\ -0.0085287 & 0.0400428 & 0.9684867 \end{bmatrix}$
Von Kries	$\begin{bmatrix} 0.4002400 & 0.7076000 & -0.0808100 \\ -0.2263000 & 1.1653200 & 0.0457000 \\ 0.0000000 & 0.0000000 & 0.9182200 \end{bmatrix}$	$\begin{bmatrix} 1.8599364 & -1.1293816 & 0.2198974 \\ 0.3611914 & 0.6388125 & -0.0000064 \\ 0.0000000 & 0.0000000 & 1.0890636 \end{bmatrix}$

CHROMATIC ADAPTATION TRANSFORM (CAT)

- ▶ Converting from the white point of one illuminant to another (source → destination)
- ▶ A kind of white-balancing but more elaborate as it relates to color constancy
- ▶ Once again, a sequence of simple 3×3 linear transforms

$$\begin{bmatrix} X_D \\ Y_D \\ Z_D \end{bmatrix} = [M] \begin{bmatrix} X_S \\ Y_S \\ Z_S \end{bmatrix}$$

**XYZ white point
of your
destination color
space (eg, D50
for prophoto)**

**XYZ white point
of your source
illuminant**

$$[M] = [M_A]^{-1} \begin{bmatrix} \rho_D/\rho_S & 0 & 0 \\ 0 & \gamma_D/\gamma_S & 0 \\ 0 & 0 & \beta_D/\beta_S \end{bmatrix} [M_A]$$

$$\begin{bmatrix} \rho_S \\ \gamma_S \\ \beta_S \end{bmatrix} = [M_A] \begin{bmatrix} X_{WS} \\ Y_{WS} \\ Z_{WS} \end{bmatrix}$$

$$\begin{bmatrix} \rho_D \\ \gamma_D \\ \beta_D \end{bmatrix} = [M_A] \begin{bmatrix} X_{WD} \\ Y_{WD} \\ Z_{WD} \end{bmatrix}$$

Method	[M _A]	[M _A] ⁻¹
XYZ Scaling	$\begin{bmatrix} 1.0000000 & 0.0000000 & 0.0000000 \\ 0.0000000 & 1.0000000 & 0.0000000 \\ 0.0000000 & 0.0000000 & 1.0000000 \end{bmatrix}$	$\begin{bmatrix} 1.0000000 & 0.0000000 & 0.0000000 \\ 0.0000000 & 1.0000000 & 0.0000000 \\ 0.0000000 & 0.0000000 & 1.0000000 \end{bmatrix}$
Bradford	$\begin{bmatrix} 0.8951000 & 0.2664000 & -0.1614000 \\ -0.7502000 & 1.7135000 & 0.0367000 \\ 0.0389000 & -0.0685000 & 1.0296000 \end{bmatrix}$	$\begin{bmatrix} 0.9869929 & -0.1470543 & 0.1599627 \\ 0.4323053 & 0.5183603 & 0.0492912 \\ -0.0085287 & 0.0400428 & 0.9684867 \end{bmatrix}$
Von Kries	$\begin{bmatrix} 0.4002400 & 0.7076000 & -0.0808100 \\ -0.2263000 & 1.1653200 & 0.0457000 \\ 0.0000000 & 0.0000000 & 0.9182200 \end{bmatrix}$	$\begin{bmatrix} 1.8599364 & -1.1293816 & 0.2198974 \\ 0.3611914 & 0.6388125 & -0.0000064 \\ 0.0000000 & 0.0000000 & 1.0890636 \end{bmatrix}$

SUMMARY

- If using “colors” as data, you must specify in which color space your data are in.
- Values in your camera’s color space will not be reproducible by anyone else (that might be ok depending on your question/application)
- You always need a calibrated color chart, ideally in every frame
- You must have proper grays (ie with uniform reflectance) to capture the color of light
- Wait till you hear how all of this gets harder underwater!



Break

- Thank you for not using your cell phones during the lectures.
- **Brain rot:** A condition of mental foginess, lethargy, reduced attention span, and cognitive decline that results from an overabundance of screen time.

