

DUSTGROWTH IN PHANTOM

Arnaud Vericel, Jean-François Gonzalez

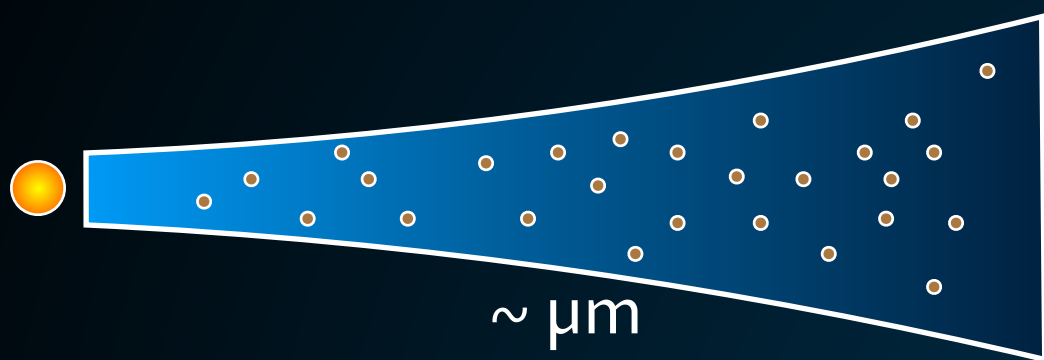


CENTRE DE RECHERCHE ASTROPHYSIQUE DE LYON



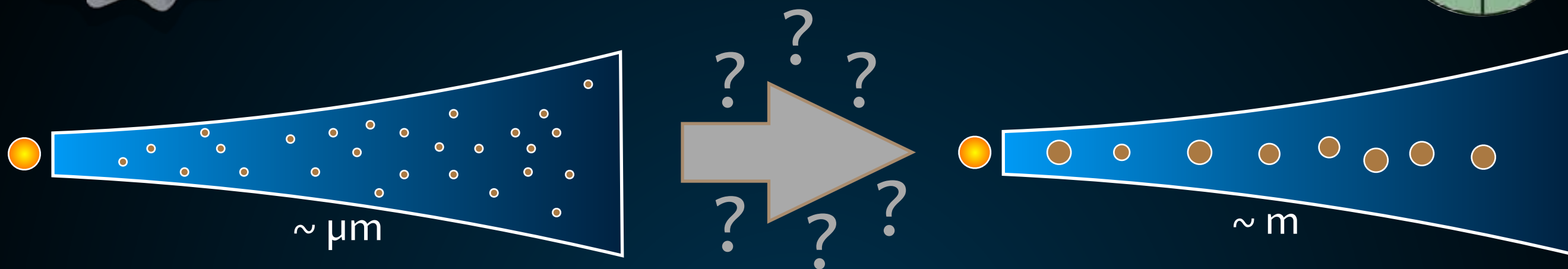


CONTEXT



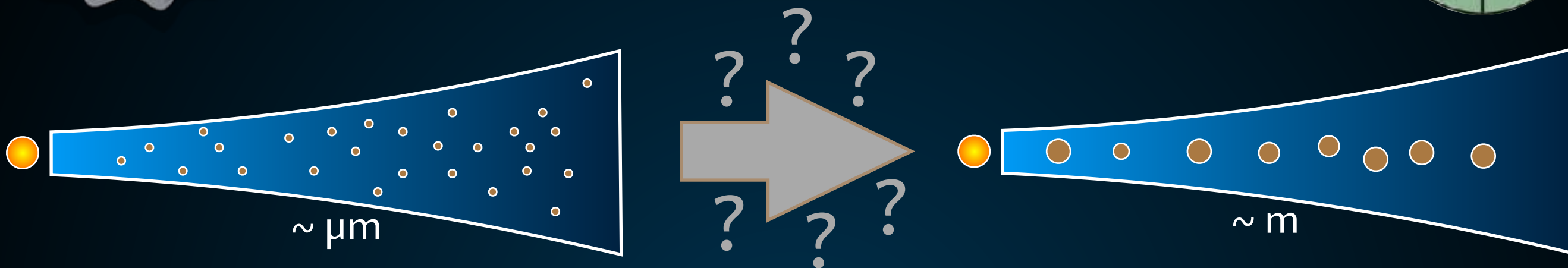


CONTEXT





CONTEXT





★ Smoluchowski equation (1916) (Maxime Lombart's PhD project)

$$\frac{\partial n(m)}{\partial t} = \frac{1}{2} \int_0^m K(m', m - m') n(m') n(m - m') dm' - n(m) \int_0^{+\infty} K(m', m) n(m') dm'$$



MODEL



- ★ Smoluchowski equation (1916) (Maxime Lombart's PhD project)

$$\frac{\partial n(m)}{\partial t} = \frac{1}{2} \int_0^m \text{challenging and expensive}_{m', m) n(m') dm'}$$



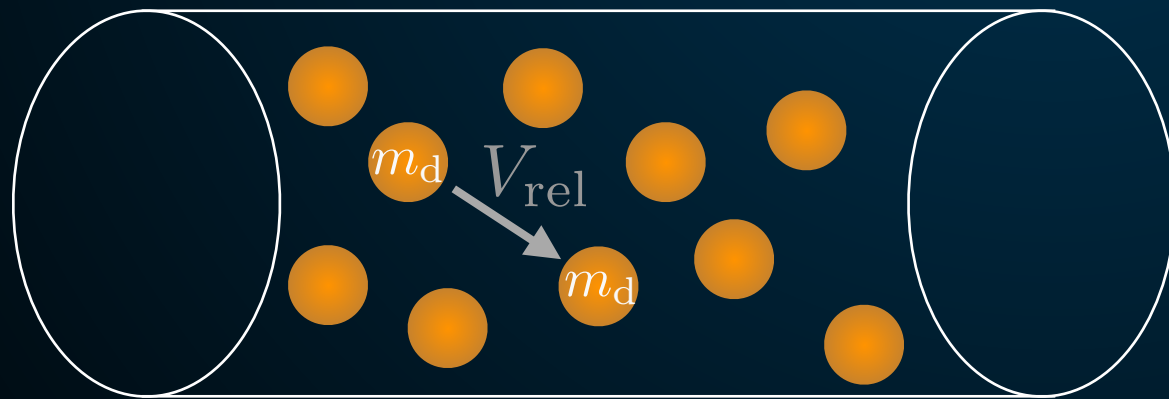
MODEL



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- ★ Stepinski & Valageas (1997) : pure growth





MODEL

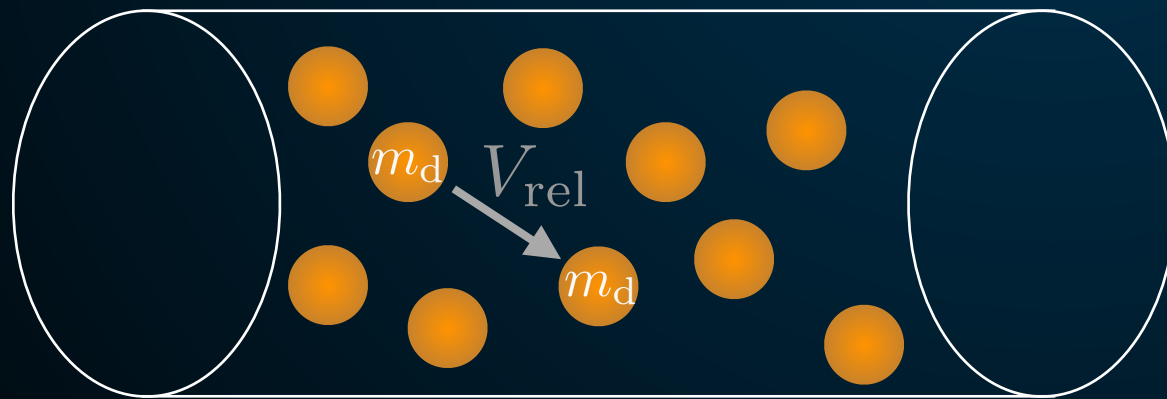


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$$\frac{dm_d}{dt} = \frac{m_d}{t_{\text{col}}}$$





MODEL

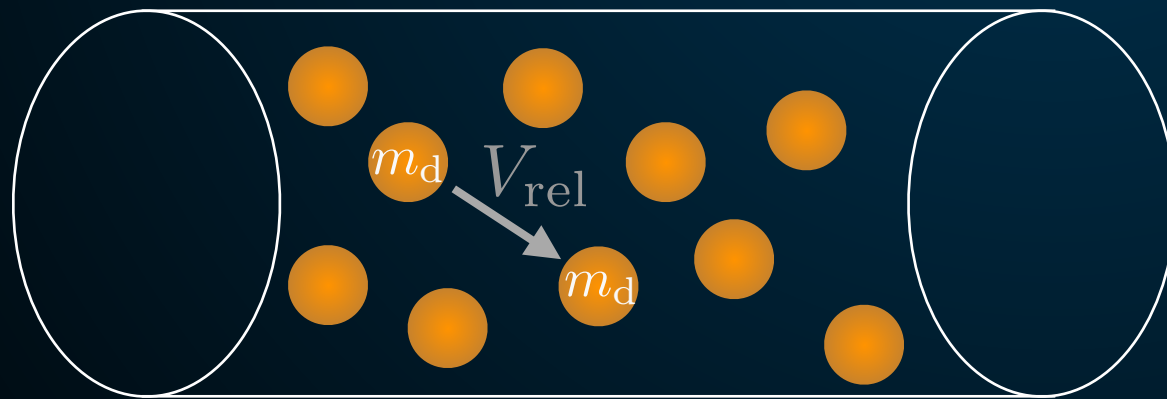


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$$\frac{dm_d}{dt} = \frac{m_d}{t_{\text{col}}}$$



$$t_{\text{col}} = \frac{1}{\sigma V_{\text{rel}} n_d}$$



MODEL

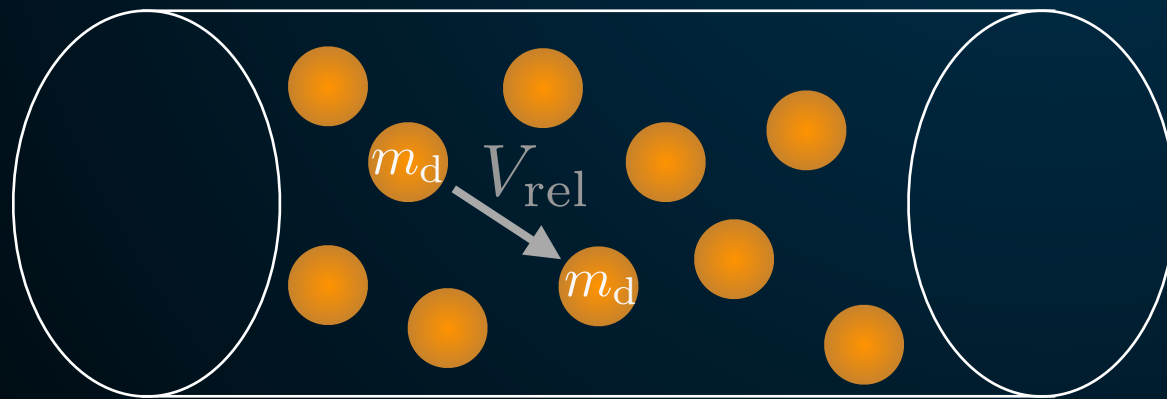


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$$\frac{ds}{dt} = \frac{\rho_d}{\rho_s} V_{\text{rel}}$$



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MODEL

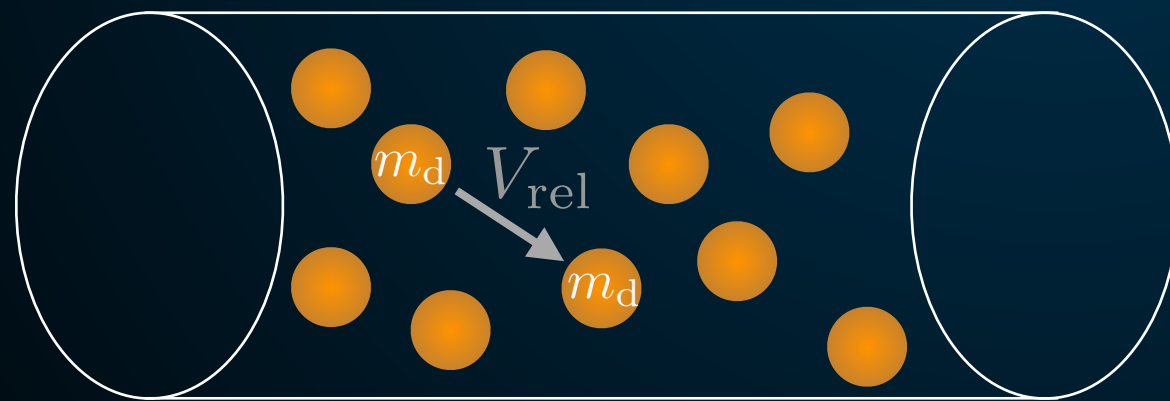


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$$\frac{ds}{dt} = \frac{\rho_d}{\rho_s} V_{\text{rel}}$$



$$t_{\text{col}} = \frac{1}{\sigma V_{\text{rel}} n_d}$$

$$V_{\text{rel}} = \sqrt{2^{3/2} \text{Ro} \alpha c_s} \frac{\sqrt{\text{Sc} - 1}}{\text{Sc}}$$

$$\text{Sc} = 1 + \Omega_K t_s \sqrt{1 + \frac{\Delta v^2}{V_t^2}}$$



MODEL





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★ Fragmentation

We compare V_{rel} with a fragmentation velocity V_{frag}



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$$V_{\text{rel}} < V_{\text{frag}} : \frac{ds}{dt} = \frac{\rho_d}{\rho_s} V_{\text{rel}} : \text{growth}$$

$$V_{\text{rel}} > V_{\text{frag}} : \frac{ds}{dt} = -\frac{\rho_d}{\rho_s} V_{\text{rel}} f : \text{fragmentation}$$



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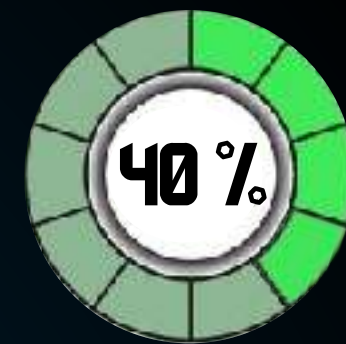
$$V_{\text{rel}} > V_{\text{frag}} : \frac{ds}{dt} = -\frac{\rho_d}{\rho_s} V_{\text{rel}} f : \text{fragmentation}$$

$f = 1$: catastrophic disruption

$$f = \frac{V_{\text{rel}}^2}{V_{\text{rel}}^2 + V_{\text{frag}}^2} : \text{smoother fragmentation derived from Kobayashi \& Tanaka (2009)}$$



MODEL





MODEL



★ Snow lines

We model snow lines by changing V_{frag} between 2 parts of the disk

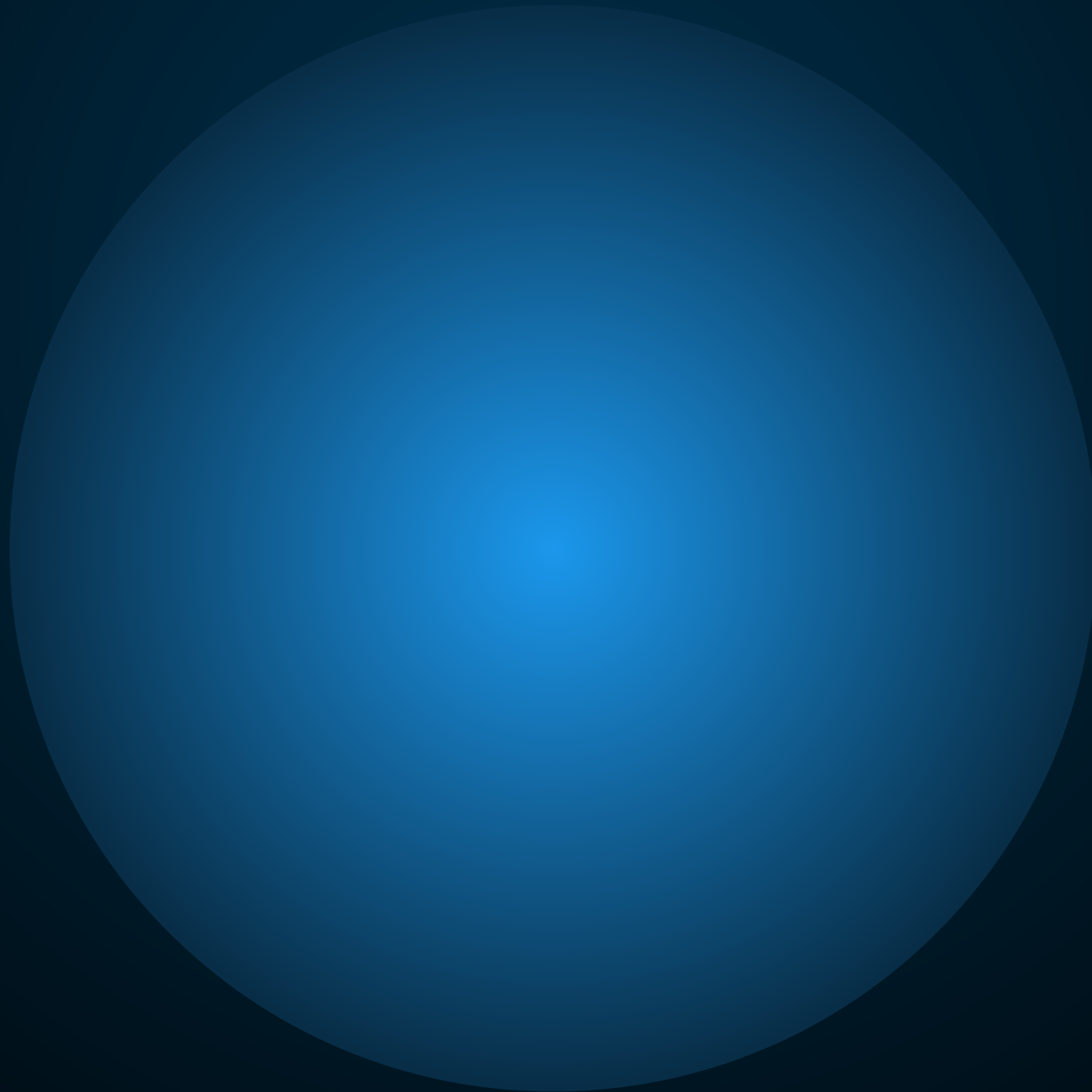


MODEL



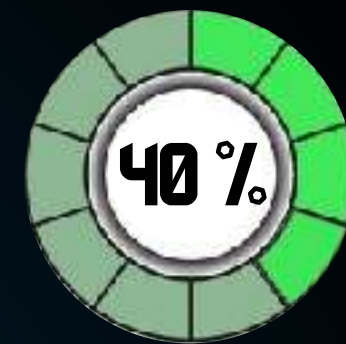
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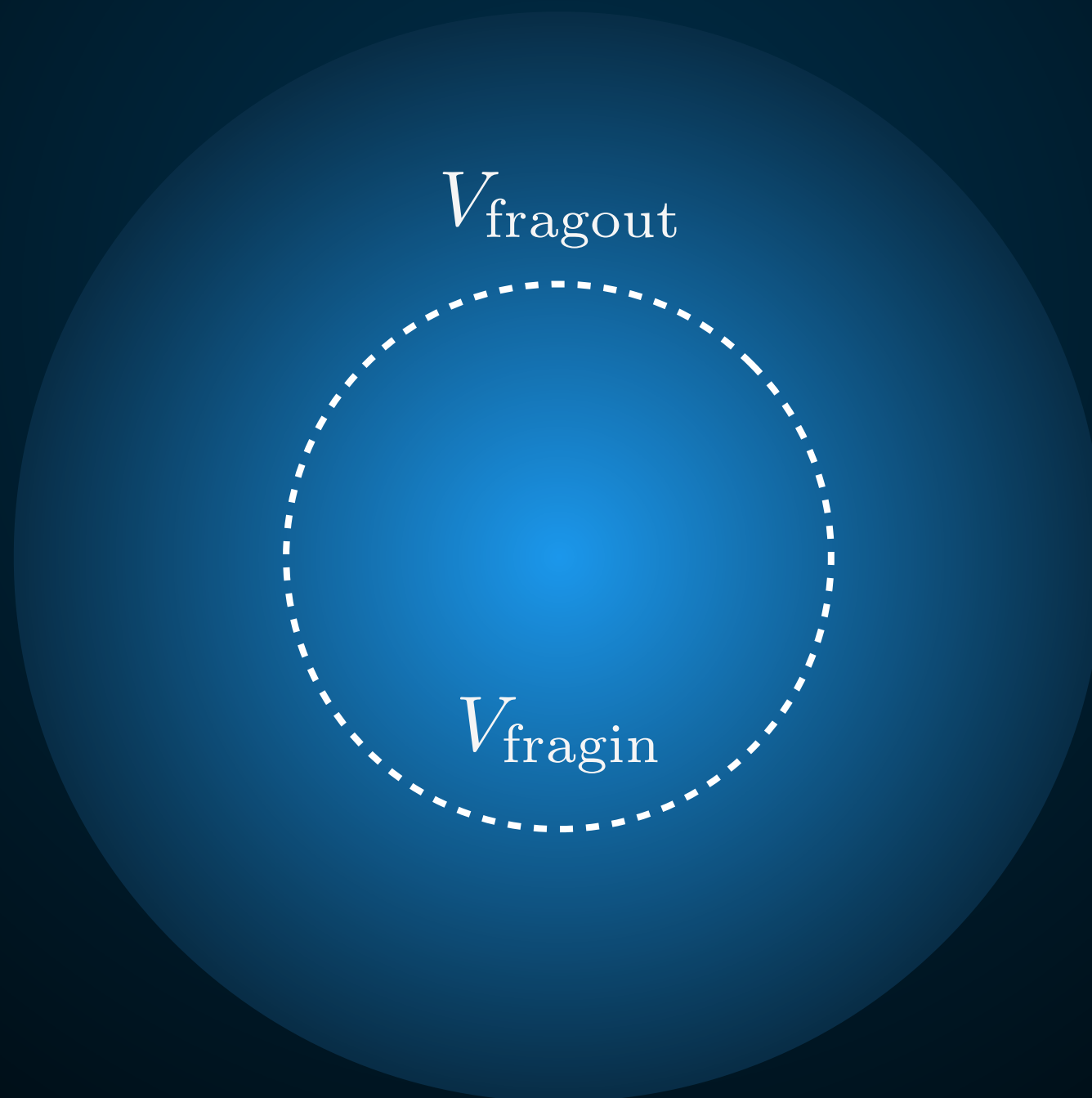


MODEL



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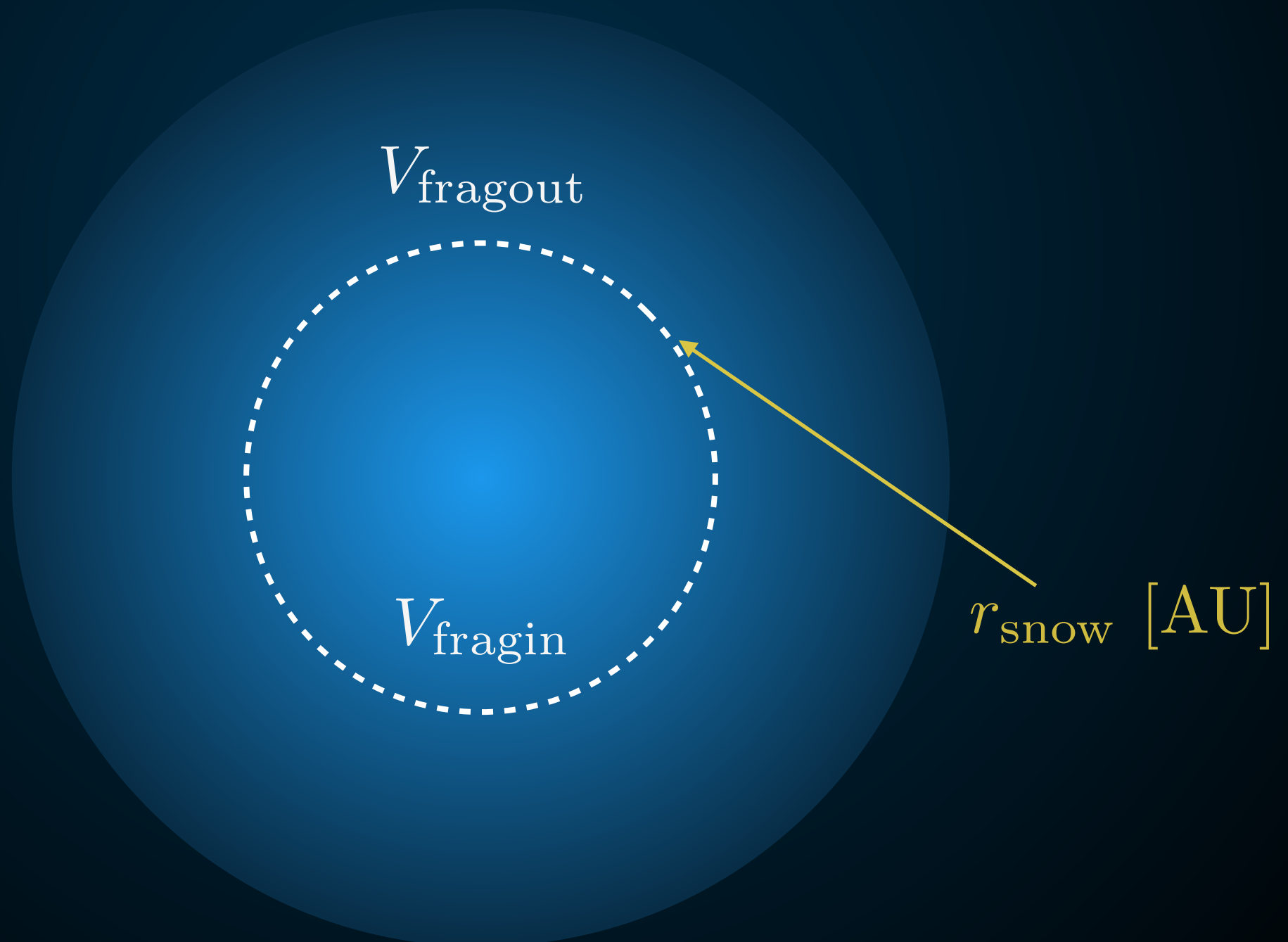
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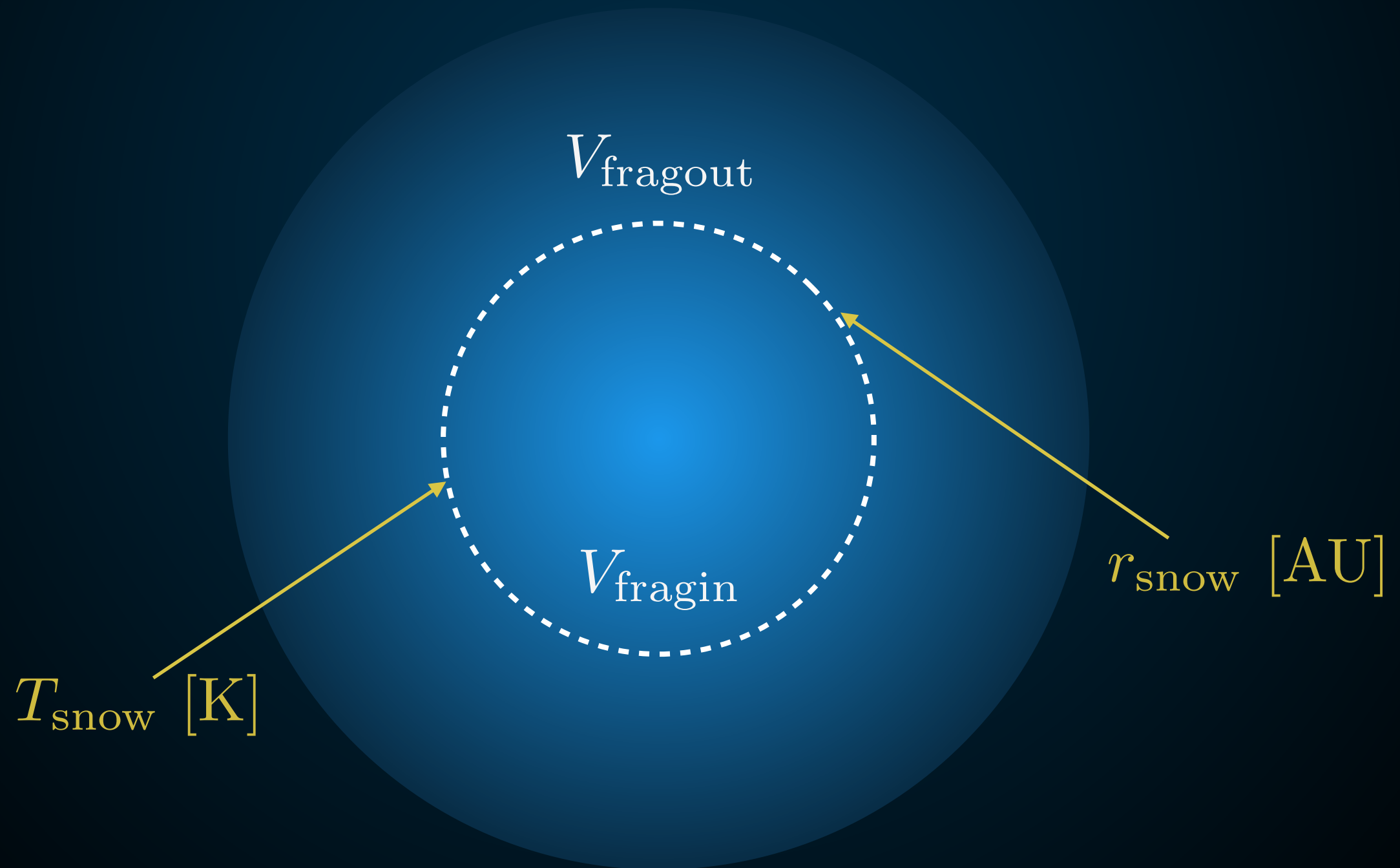
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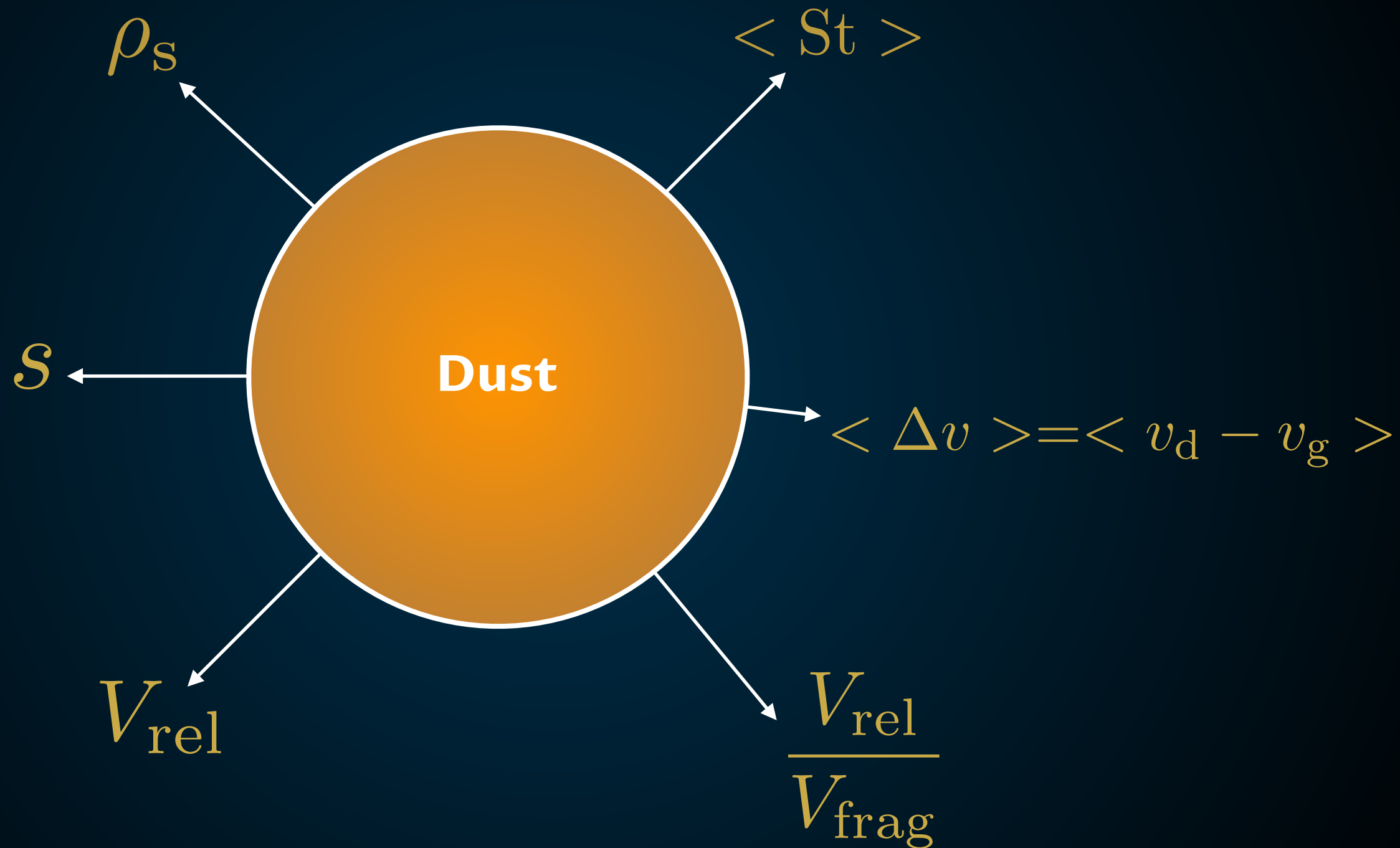
IMPLEMENTATION



Dust

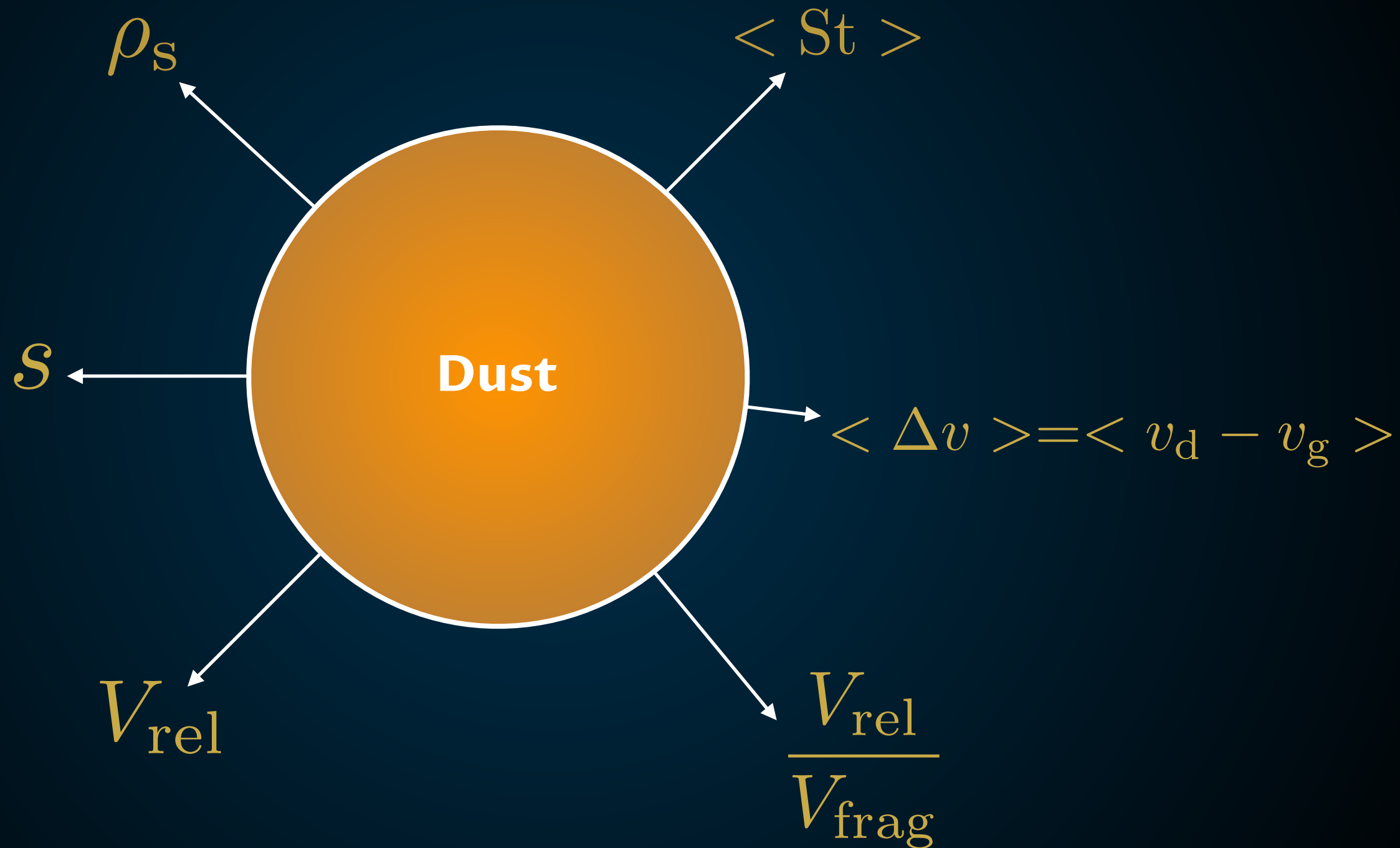


IMPLEMENTATION





IMPLEMENTATION



New array 'dustprop' and new module 'growth'



IMPLEMENTATION





IMPLEMENTATION



interpolate
 Δv & St



IMPLEMENTATION



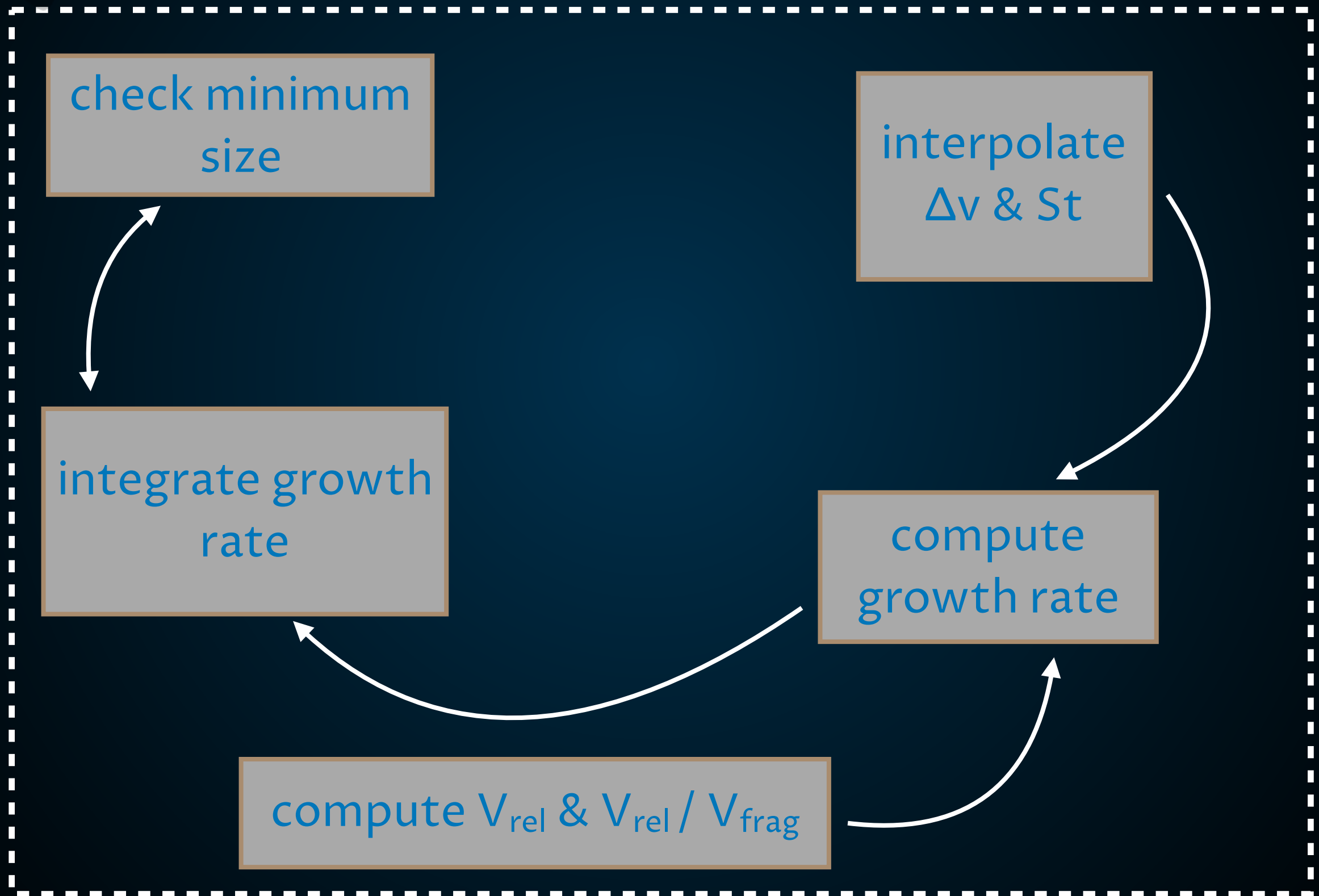
interpolate
 Δv & St

compute
growth rate

compute V_{rel} & V_{rel} / V_{frag}

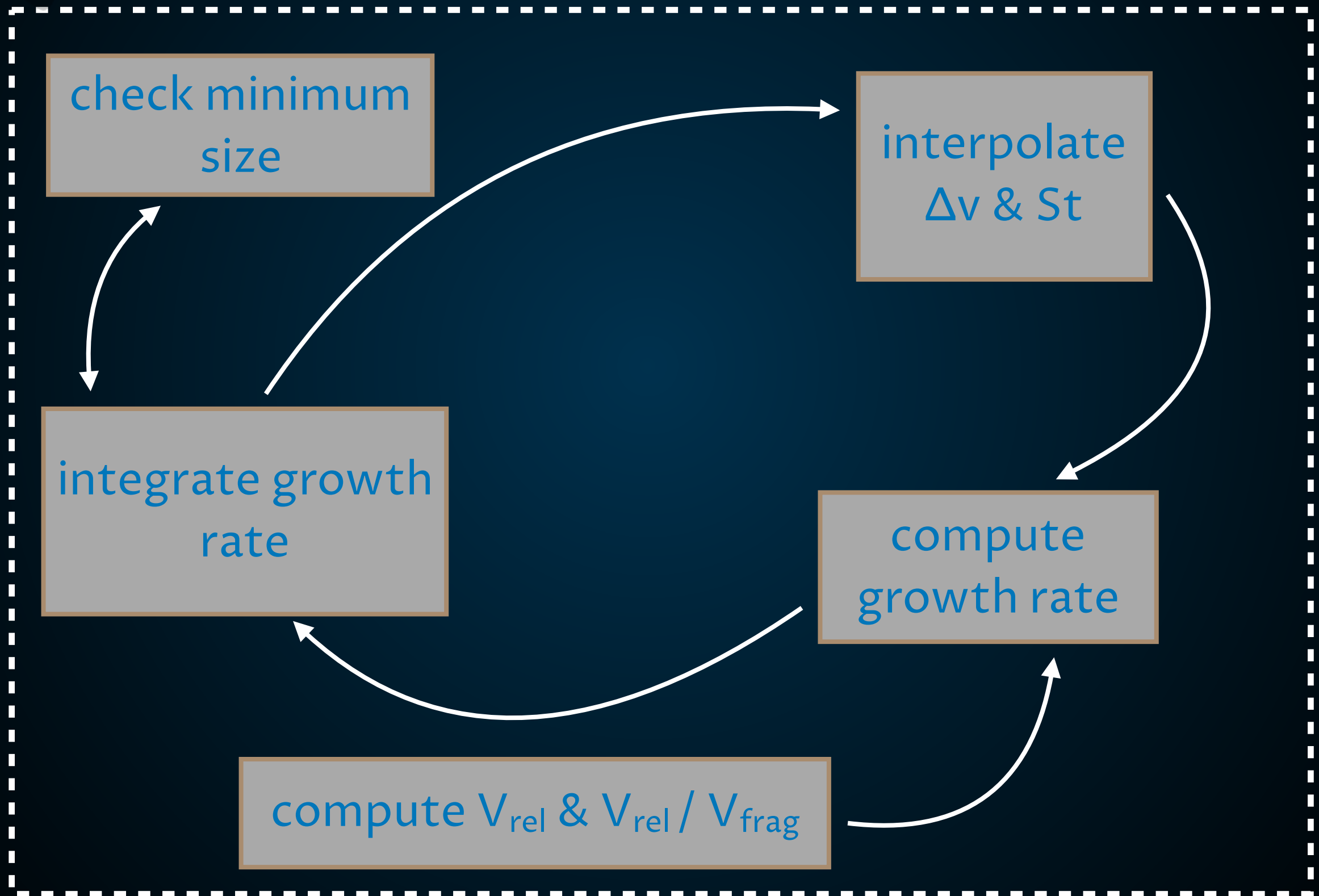


IMPLEMENTATION



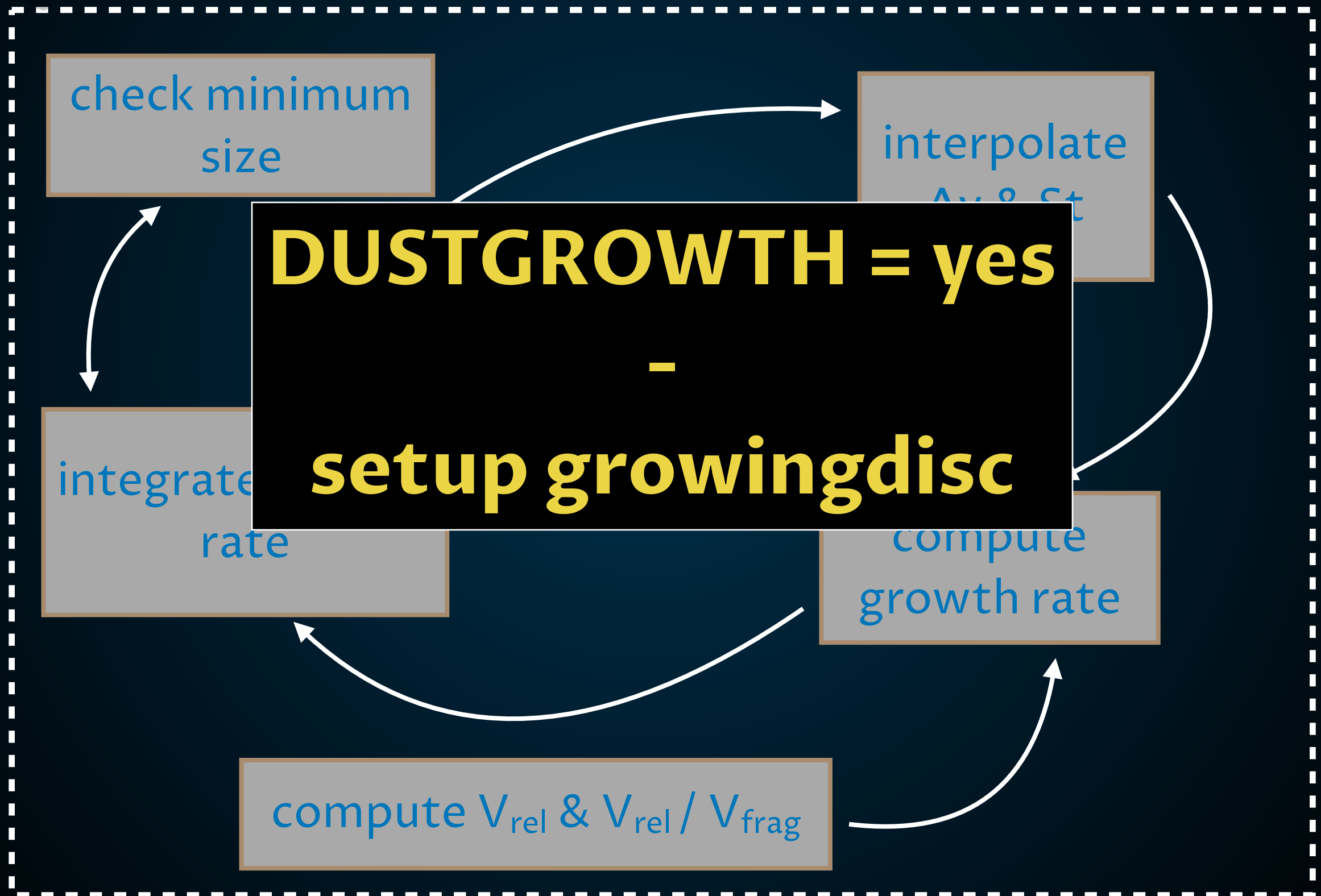


IMPLEMENTATION



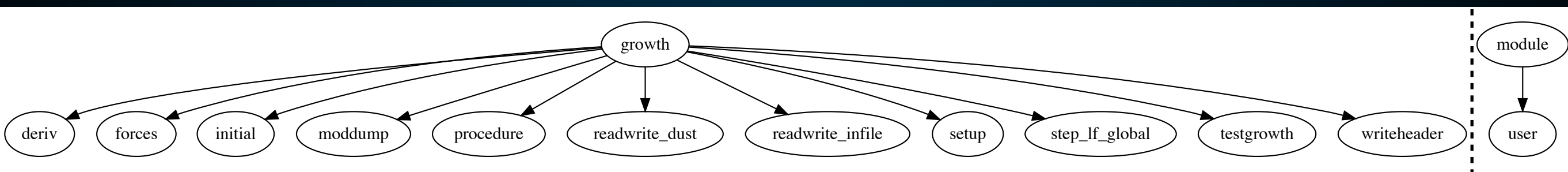
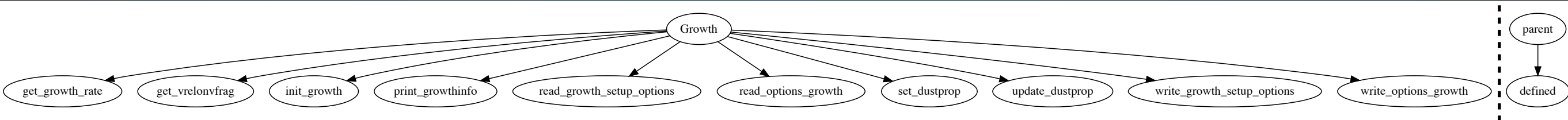
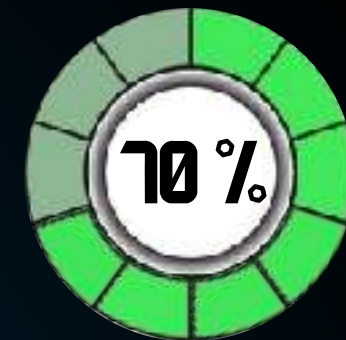


IMPLEMENTATION





FORTREE MAP





GROWINGBOX



The features included in the **growth** module are tested





GROWINGBOX



The features included in the **growth** module are tested

—> check the size integration for several configurations





GROWINGBOX



The features included in the **growth** module are tested

- > check the size integration for several configurations
- > check the interpolation of St



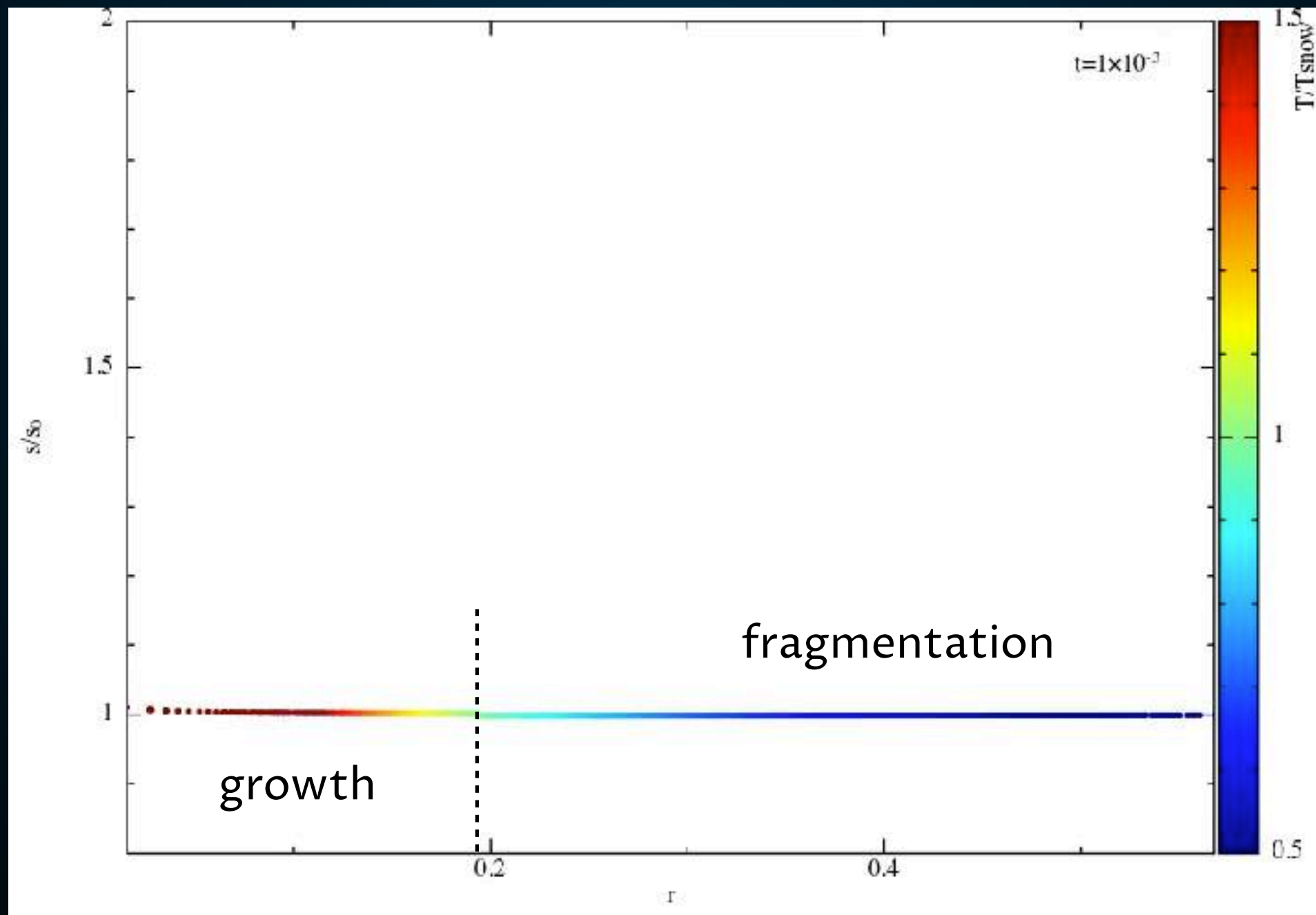


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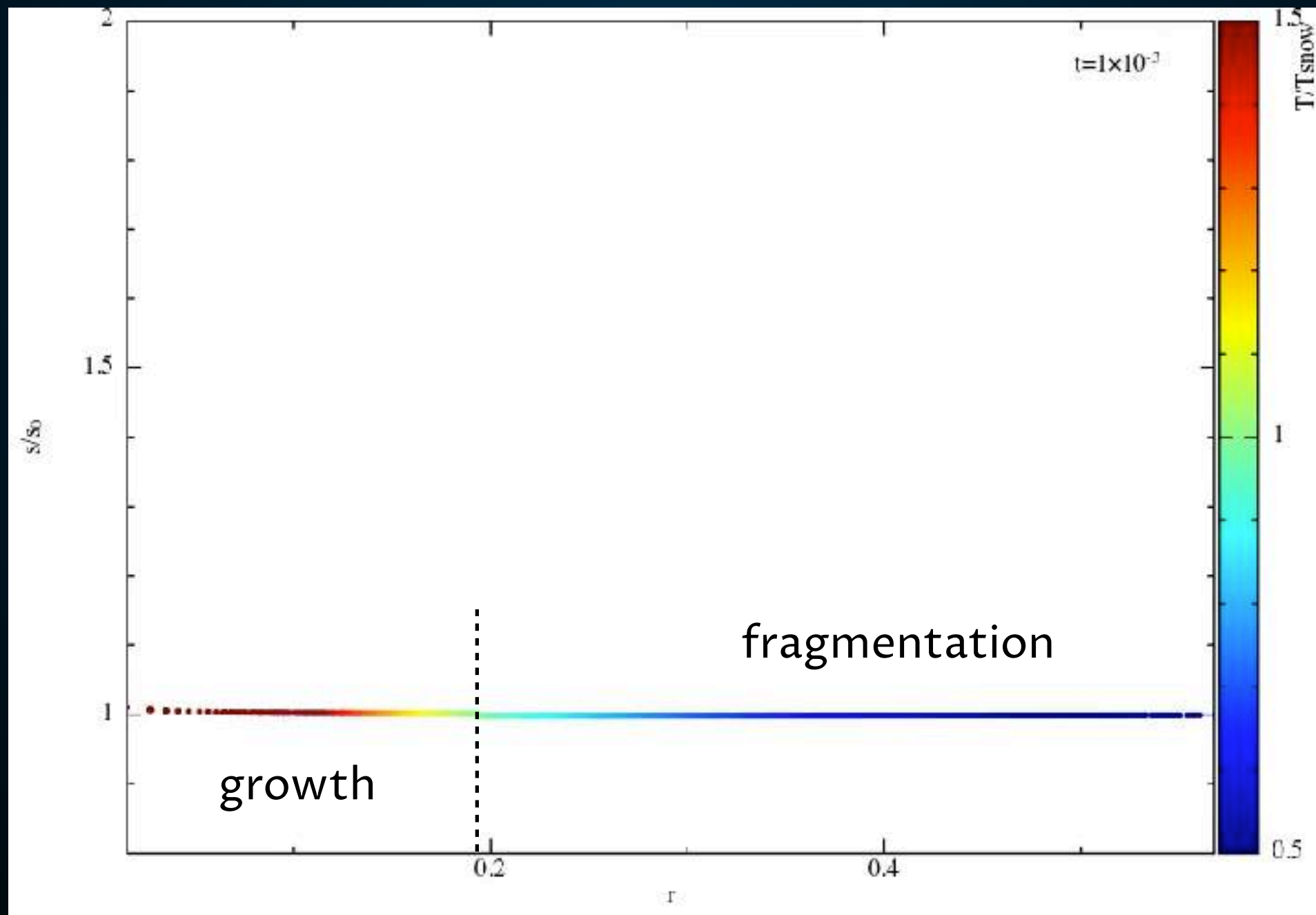


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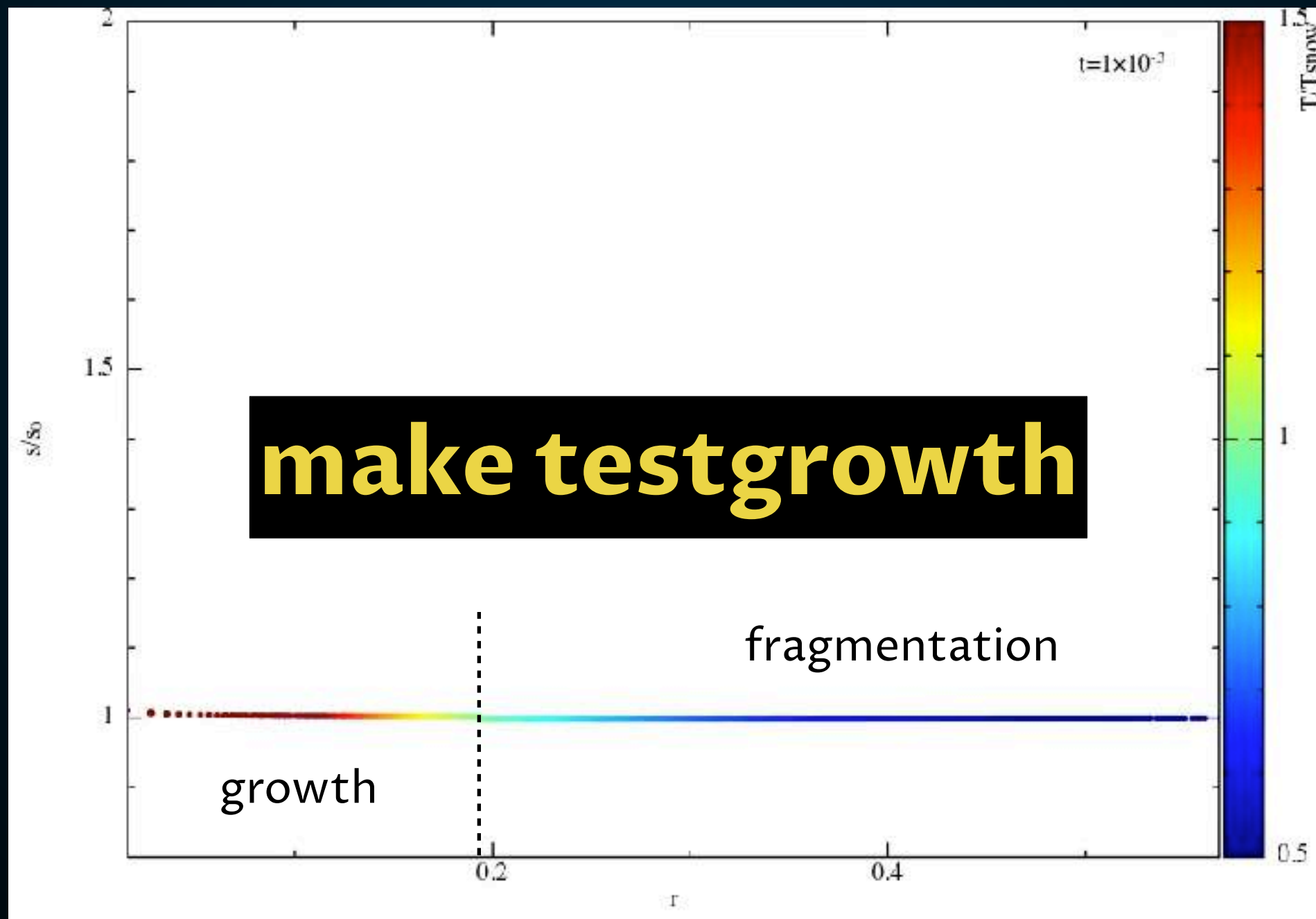


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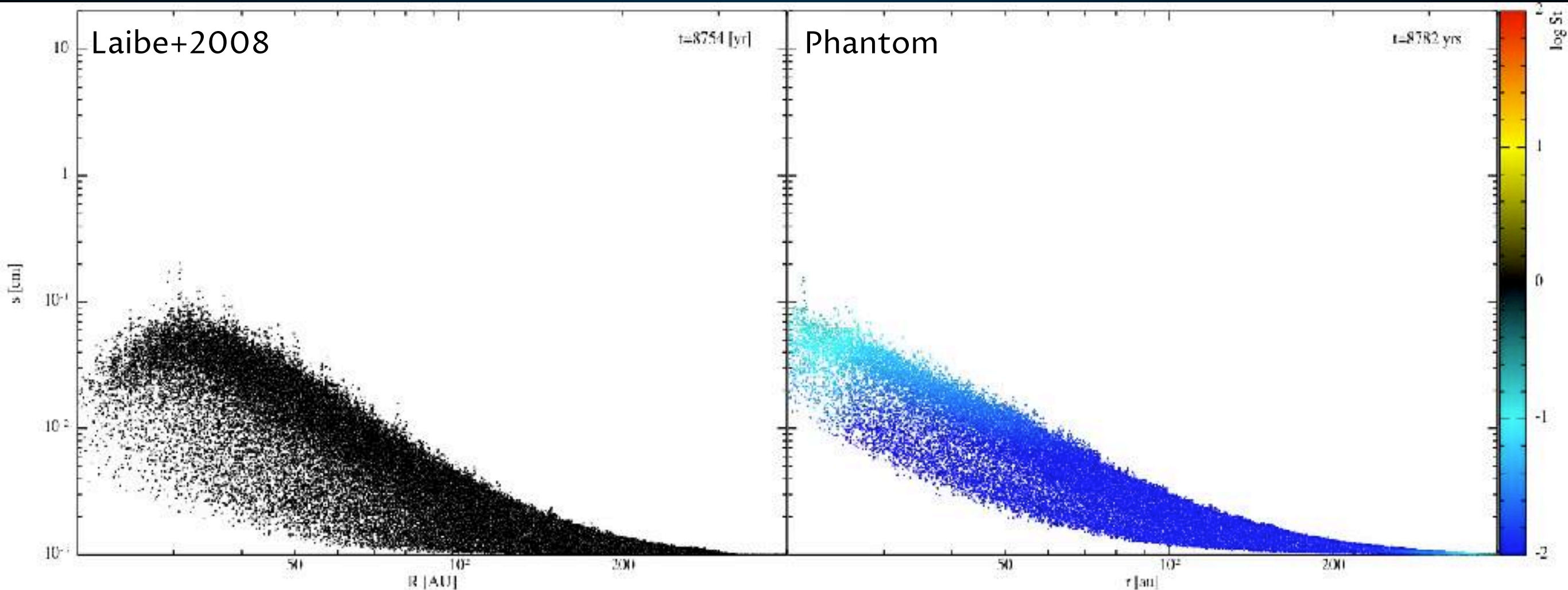


COMPARISON RESULTS



p	q (c_s)	R_{in}	R_{out}	m_{disc}/m_{sun}	ϵ_0	H/R @ 100 AU	s_0
3/2	3/8	20 AU	300 AU	0.01	0.01	0.05	10 μm

Laibe et al. 2008 (pure growth)



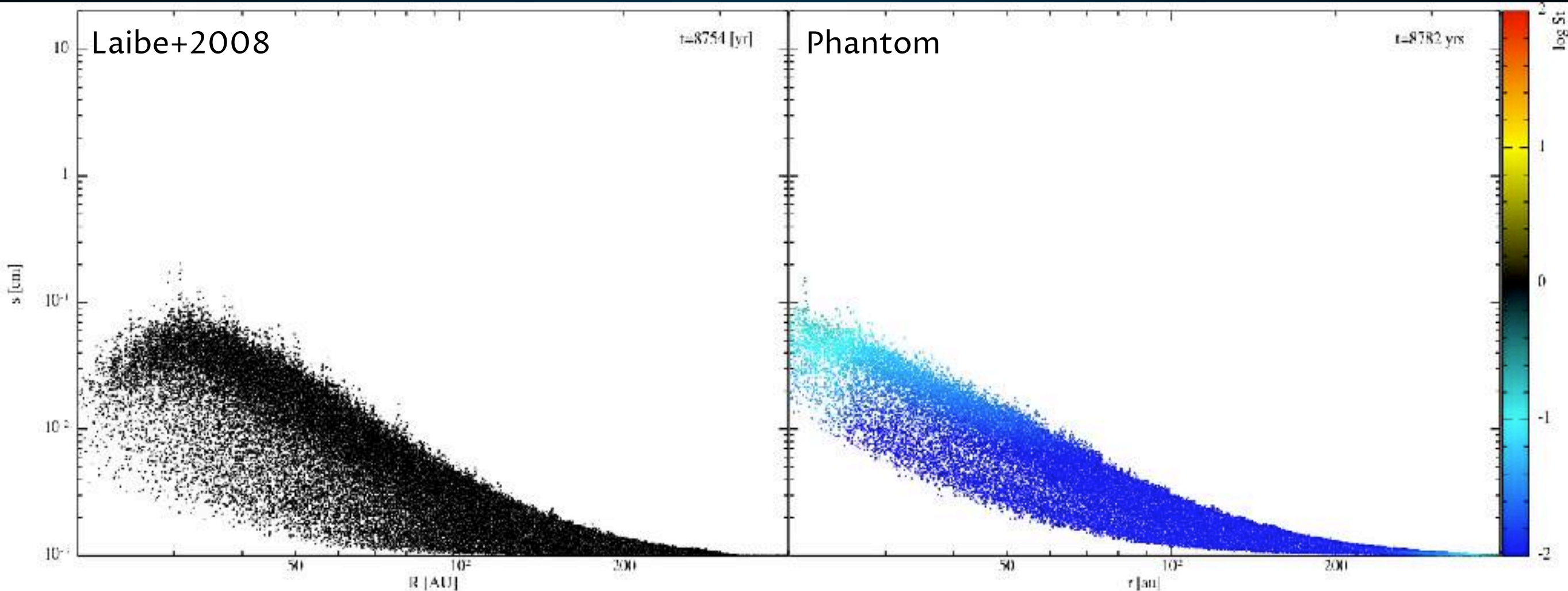


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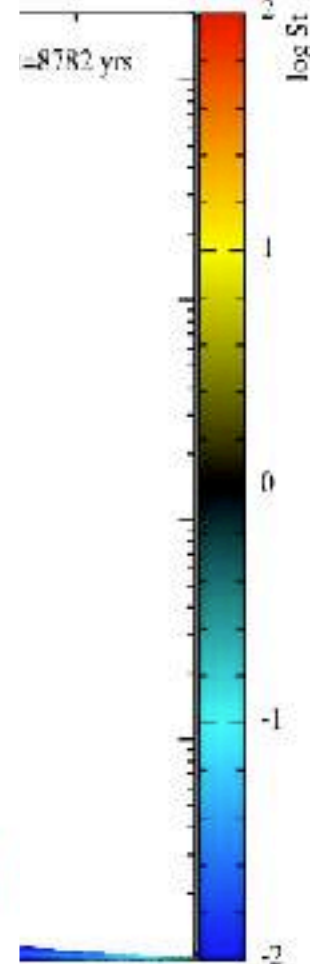
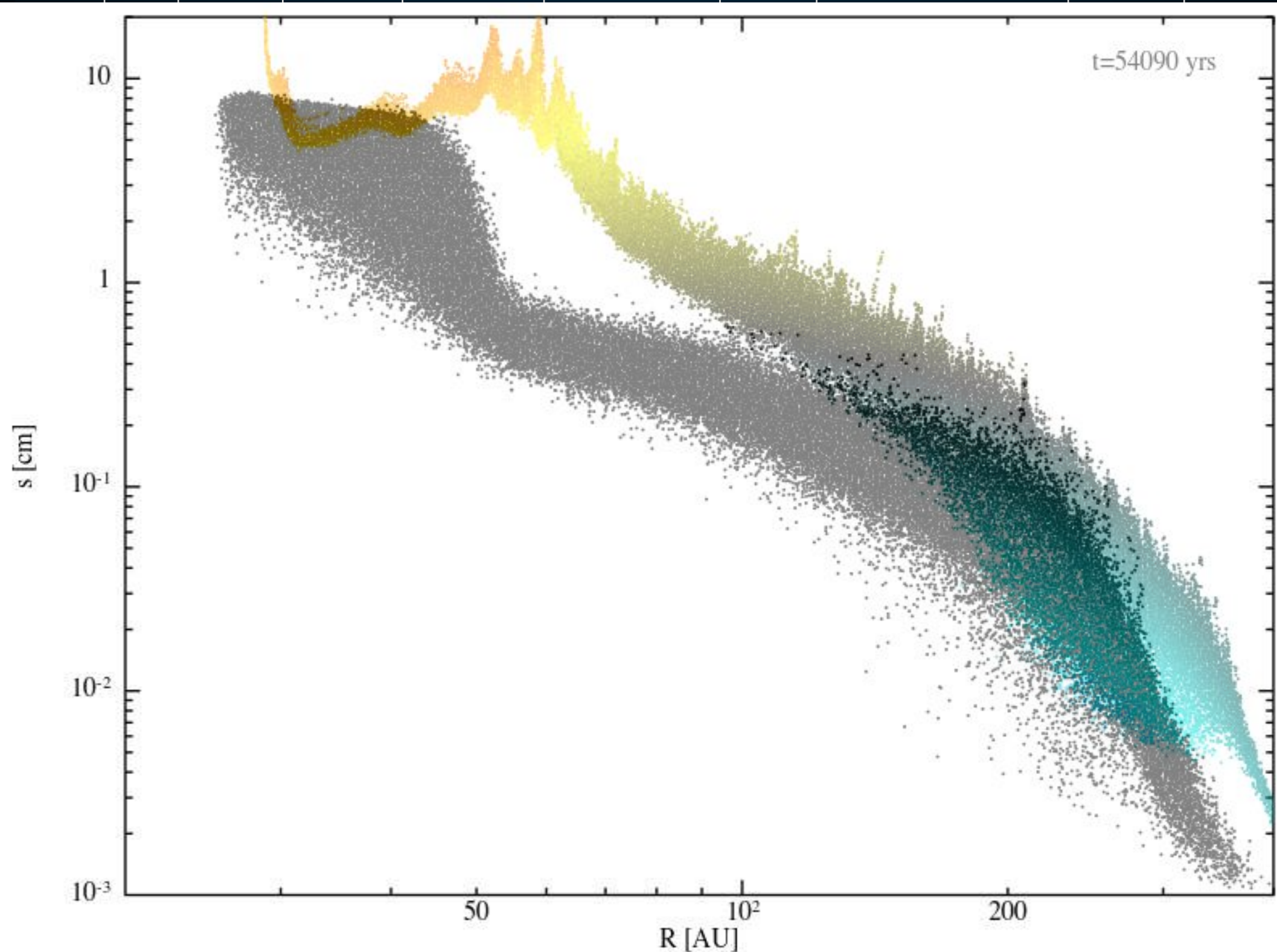
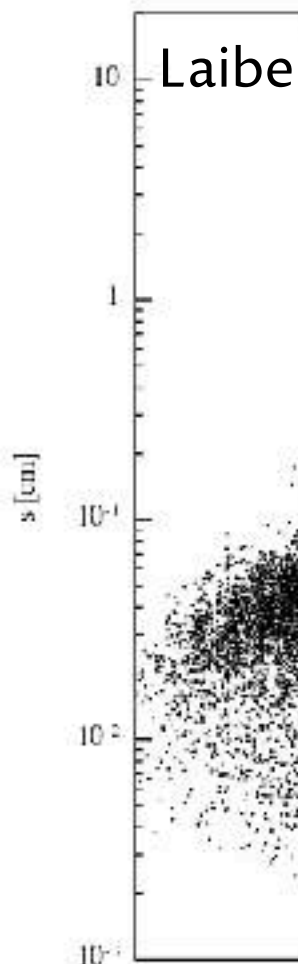




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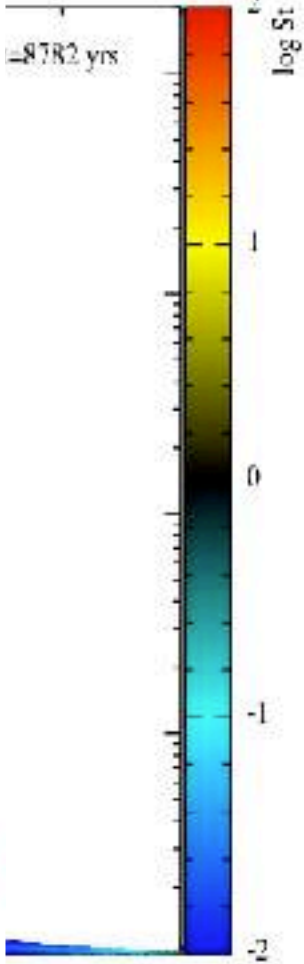
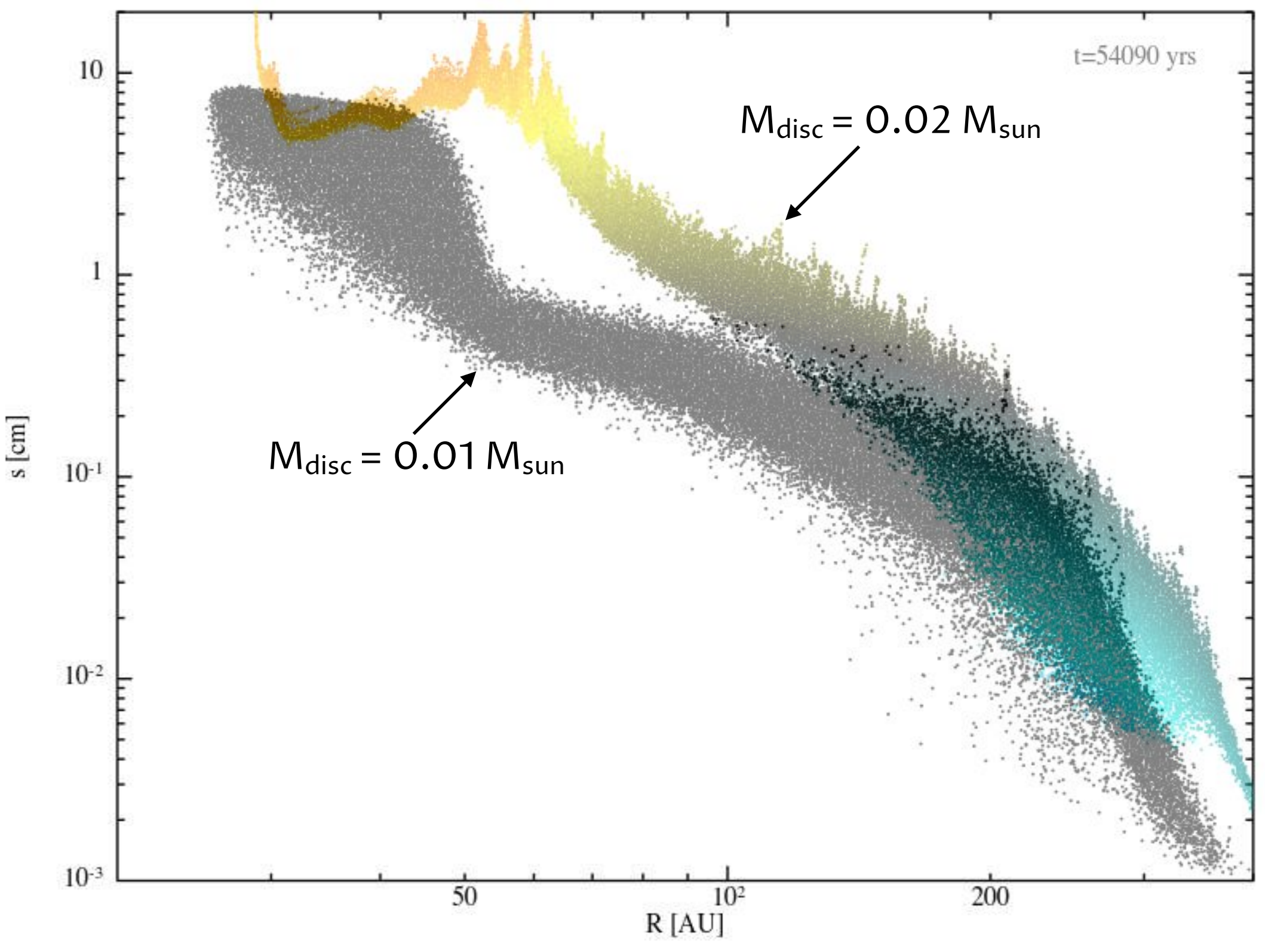
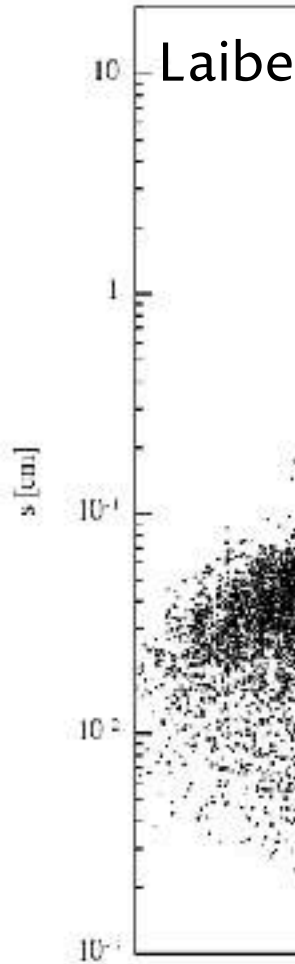




COMPARISON RESULTS



p	q (c _s)	R _{in}	R _{out}	m _{disc} /m _{sun}	ε ₀	H/R @ 100 AU	s ₀
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CONCLUSION





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```
1 |-----|
2 | The Phantom Smoothed Particle Hydrodynamics code, by Daniel Price et al. |
3 | Copyright (c) 2007-2018 The Authors (see AUTHORS) |
4 | See LICENCE file for usage and distribution conditions |
5 | http://users.monash.edu.au/~dprice/phantom |
6 |-----|
7 |-
8 | MODULE: growth
9 |
10 | DESCRIPTION:
11 | Contains routine for dust growth and fragmentation
12 |
13 | REFERENCES:
14 | Slepinski & Valageas (2015)
15 | Kobayashi & Tanaka (2015)
16 |
17 | OWNED BY:
18 | Daniel Price <1700863ef17b1ad96138aeffad86@monash.edu>
19 |
20 | PARAMETERS:
21 | isnow — snow line condensation temperature in K
22 | grainsize_min — minimum allowed grain size in cm
23 | ifrag — fragmentation of dust (0=off, 1=on, 2=Kobayashi)
24 | isnow — snow line (0=off, 1=position based, 2=temperature based)
25 | rsnow — snow line position in AU
26 | vfrag — uniform fragmentation threshold in m/s
27 | vfragin — inward fragmentation threshold in m/s
28 | vfragout — outward fragmentation threshold in m/s
29 |
30 |
31 | DEPENDENCIES: dust, eos, infile_utils, io, options, part, physcon, units
32 |-----|
33 |-----|
34 module growth
```

Dustgrowth is fully implemented



Dustgrowth is tested
(and more importantly it pass the tests)

```

--> TESTING DUSTGROWTH MODULE

--> testing growth initialisation
checking growth initialisation.....OK      lgot      0 should be      0]
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--> testing GROWTHBOX

----- pure growth (ifrag = 0, St = const) -----
checking size match exact solution.....OK      [max err = 1.524E-05, tol = 5.000E-04]

----- 0. St=f(size) inspired from Laibe et al. (2008) -----
checking size match exact solution.....OK      [max err = 1.026E-05, tol = 5.000E-04]

----- growth-fragmentation -----
checking size match exact solution.....OK      [max err = 1.524E-05, tol = 5.000E-04]

----- position based snow line -----
checking size match exact solution (in).....OK      [max err = 1.524E-05, tol = 5.000E-04]
checking size match exact solution (out).....OK      [max err = 1.524E-05, tol = 5.000E-04]

----- temperature based snow line -----
checking size match exact solution (in).....OK      [max err = 1.000E-04, tol = 5.000E-04]
checking size match exact solution (out).....OK      [max err = 1.647E-05, tol = 5.000E-04]

--> testing STOKES NUMBER INTERPOLATION

----- ts = const -----
checking Stokes number interpolation match exact solution..OK      [max err = 4.945E-03, tol = 5.000E-03]

<-- DUSTGROWTH TEST COMPLETE

<-- testing complete
total wall time = 46.12 s
total cpu time = 5 min, 19.05 s (= 3.1905E+02s)

SUMMARY OF ALL TESTS:
PASSED: 3 of 3 100.0%
FAILED: 0 of 3 0.0%

PASS

```


THANK YOU

Run finished on 19/06/2018 at 12:00:00

Total wall time: 20 min, 0.00 s (=1.2000E+03s)

THANKS TO ALL



CENTRE DE RECHERCHE ASTROPHYSIQUE DE LYON



observatoire
de Lyon

