

Computational Methods for Astrophysical Applications

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References

- **Books**

- *Riemann solvers and numerical methods for fluid dynamics - A practical introduction*, by Toro
- *Numerical methods for conservation laws*, by Leveque (Leveque1)
- *Finite volume methods for hyperbolic problems*, by Leveque (Leveque2)

- **Courses**

- Rony Keppens & Jon Sundqvist 2016-2017
- *Numerical PDE Techniques for Scientists and Engineers*, by Dinshaw Balsara

- **Schools**

- Les Houches
- *Numerical techniques in MHD simulations*, Köln University

Lesson 1 : Hyperbolic Partial Differential Equations

- **Conservation laws**

Integral and differential forms

Examples

- **Hyperbolic PDE**

Matrix formulation of conservation laws

Time advance

Linear advective equations

The Riemann problem

- 1 Conservation laws
 - Integral form
 - Differential forms
 - Examples
- 2 Hyperbolic PDE
 - Classifications of PDE
 - Matrix formulation of conservation laws
 - The Riemann problem

Integral forms

- Continuous medium hypothesis : from actual particles to particles of fluid
 - ⇒ distribution function and integration of Boltzmann equation : the closure problem
 - ⇒ analogy with moments of the specific photon intensity (see Prof. Sundqvist's course)
- Consider **any volume element** dV of a fluid
 - ⇒ balance over dt of scalar quantity : inflow/outflow, sinks/sources
 - ⇒ for the momentum : tensor formulation and dyadic product
 - ⇒ for energy
- Shocks and jump conditions : Rankine-Hugoniot
- See Toro 2.4.1

Conservative form

- Provided the variables $U(\mathbf{r}, t)$ are differentiable

$$\partial_t U + \nabla \cdot \underbrace{[F(U)]}_{\text{fluxes}} = \underbrace{S(\mathbf{r}, \mathbf{v}, t)}_{\text{sources/sinks}}$$

⇒ Matrix formulation

- Properties of this differential form

⇒ Coordinate / dimension independent formulation

⇒ Given initial and boundary conditions, can provide a general solution

⇒ Conservative form : Green-Ostrogradsky (or Gauss) law and Eulerian approach

⇒ Conservative variables : ρ , $\rho \mathbf{v}$ and e (*total energy*)

Primitive form

- Lagrangian approach

$$D_t(\cdot) = \partial_t(\cdot) + \mathbf{v} \nabla(\cdot)$$

- Continuity equation : $D_t(\rho) = -\rho \nabla \mathbf{v}$
 \Rightarrow incompressible fluid \neq flow
- Navier-Stokes equation

$$\rho D_t(\mathbf{v}) = -\nabla P$$

- Energy equation
 \Rightarrow internal and mechanical energy
 \Rightarrow link w/ 1st principle of Thermodynamics
 \Rightarrow entropy formulation
- Primitive variables : ρ, \mathbf{v}, P

Closure relation

- Equation-of-state of an ideal gas

$$u = \frac{P}{\gamma - 1}$$

\Rightarrow adiabatic index γ

- A classic way-around : the polytropic assumption

$$P \propto \rho^\alpha$$

\Rightarrow polytropic index α

\Rightarrow the isentropic case

The linear advection equation

- Consider the 1D continuity equation w/ constant & uniform speed v

$$\partial_t \rho + v \partial_x \rho = 0$$

\Rightarrow initial conditions $\rho(x, t = 0) = \rho_0(x)$

- Analytic solution

$$\rho(x, t) = \rho_0(x - vt)$$

- Fourier analysis

$$\rho(x, t) = r(k, t) e^{ikx}$$

\Rightarrow non-dissipative

\Rightarrow non-dispersive

\Rightarrow see also Balsara 1 (slides#32-33)

- Analytic preliminary analysis \Rightarrow mathematical properties of the PDE \Rightarrow appropriate numerical scheme
- Numerical solution

Euler equations

- 2 scalar variables (ρ & e) + 1 vector ($\rho \mathbf{v}$) = 5 variables
- Conservative form $\partial_t \mathbf{U} + \nabla \cdot \bar{\bar{\mathbf{F}}} = \mathbf{0}$ with, in 3D :

$$\mathbf{U} = \begin{bmatrix} \rho \\ \rho \mathbf{v} \\ e = u + \rho \mathbf{v}^2 / 2 \end{bmatrix} \quad \text{and} \quad \bar{\bar{\mathbf{F}}} = \begin{bmatrix} \rho \mathbf{v} \\ \rho \mathbf{v} \otimes \mathbf{v}^T + P \bar{\mathbb{1}} \\ (e + P) \mathbf{v} \end{bmatrix} \quad (1)$$

with $\bar{\bar{\mathbf{F}}}$ a 5 rows \times 3 columns matrix

\Rightarrow beware the divergence operator! Retrieve primitive form

- in Cartesian? in cylindrical? in spherical?

Navier-Stokes equations

- Consider the viscous tensor $\overline{\overline{\tau}}$

$$\overline{\overline{\tau}} = \mu \left[\left(\overline{\overline{\nabla \mathbf{v}}} + \overline{\overline{\nabla \mathbf{v}^t}} \right) - \frac{2}{3} (\nabla \cdot \mathbf{v}) \overline{\overline{\mathbb{1}}} \right]$$

- \Rightarrow see [Landau Lifschitz](#) 2.15 for hypothesis & shape
- \Rightarrow μ dynamic viscosity in Pa·s, intrinsic
- \Rightarrow conservative formulation
- \Rightarrow Dimensionless form, Reynolds number and singular limit (D'Alembert's paradox)

Ideal Magneto-Hydrodynamics equations

- Ideal MHD fundamental hypothesis

$$\mathbf{E} + \mathbf{j} \wedge \mathbf{B} = \mathbf{0}$$

\Rightarrow \mathbf{E} secondary variable

- Maxwell-Faraday and Maxwell-Ampère give the induction equation :

$$\partial_t \mathbf{B} = \nabla \wedge (\mathbf{v} \wedge \mathbf{B})$$

\Rightarrow magnetic pressure and magnetic β number

\Rightarrow magnetic tension

- Frozen-flux theorem

Ideal Magneto-Hydrodynamics equations

- 2 scalar variables (ρ & e) + 2 vectors ($\rho\mathbf{v}$ and \mathbf{B}) = 8 variables
- Conservative form $\partial_t \mathbf{U} + \nabla \cdot \overline{\overline{\mathbf{F}}} = \mathbf{0}$ with, in 3D :

$$\mathbf{U} = \begin{bmatrix} \rho \\ \rho\mathbf{v} \\ e = u + \rho\mathbf{v}^2/2 + \mathbf{B}^2/2\mathbf{B} \end{bmatrix} \quad (2)$$

$$\overline{\overline{\mathbf{F}}} = \begin{bmatrix} \rho\mathbf{v} \\ \rho\mathbf{v} \otimes \mathbf{v}^T + (P + \mathbf{B}^2/2) \overline{\overline{\mathbf{1}}} - \mathbf{B} \otimes \mathbf{B}^T \\ (e + P + \mathbf{B}^2/2) \mathbf{v} - (\mathbf{v} \cdot \mathbf{B}) \mathbf{B} \\ \mathbf{v}\mathbf{B}^T - \mathbf{B}\mathbf{v}^T \end{bmatrix} \quad (3)$$

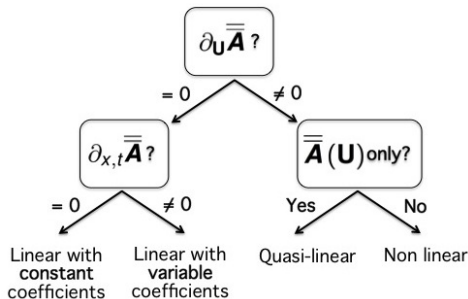
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Linear, non-linear, quasi-linear

- Homogeneous

$$\partial_t \mathbf{U} + \overline{\overline{\mathbf{A}}}(\mathbf{U}, x, t) \partial_x \mathbf{U} = \mathbf{0} \quad (4)$$

- Nature of the PDE



Linear, non-linear, quasi-linear

- Non-homogeneous : $\dots + \mathbf{B}(\mathbf{U}, x, t) = \mathbf{0}$
 \Rightarrow still linear if $\mathbf{B} \propto \mathbf{U}$
- Exemples :
 - \Rightarrow Time and space dependent linear advection equation?
 - \Rightarrow Burgers equation?
 - \Rightarrow Conservation equations?
- Systematic way to solve quasi-linear PDE?

Eigenvalues & eigenvectors

- Consider the following system of linear constant coeff. PDE :

$$\begin{cases} \partial_t \rho_1 + \partial (\rho_1 a_{11} + \rho_2 a_{12}) = 0 \\ \partial_t \rho_2 + \partial (\rho_1 a_{21} + \rho_2 a_{22}) = 0 \end{cases}^{(5)}$$

⇒ Linear? Homogeneous?

⇒ Matrix form $\partial_t \mathbf{U} + \mathbf{A} \partial_x \mathbf{U}$?

⇒ \mathbf{A} diagonalisable?

⇒ Diagonal matrix \mathbf{D} ?

⇒ Eigenvector matrix \mathbf{K} such as $\mathbf{A} = \mathbf{K} \mathbf{D} \mathbf{K}^{-1}$?

Hyperbolic and elliptic 1st order PDE

- Hyperbolic $\Leftrightarrow \mathbf{A}$ diagonalisable and real eigenvalues
 - \Rightarrow eigenvalues λ_i as wave speeds
 - \Rightarrow Was (5) hyperbolic?
- Systematic way to solve *hyperbolic* quasi-linear PDE?
- Rk : Elliptic $\Leftrightarrow \mathbf{A}$ has non-real eigenvalues
 - \Rightarrow ex: Cauchy-Riemann equations (2D steady)

$$\begin{cases} \partial_x u - \partial_y v = 0 \\ \partial_x v + \partial_y u = 0 \end{cases}$$

General solutions

- Toro 2.3.1-2 and Balsara 1.5.3
- If *constant coefficient linear* system, the procedure is the following :
 $\Rightarrow \mathbf{K}^{-1} \times (4) \Rightarrow$ set of independent linear advection solutions, the canonical form :

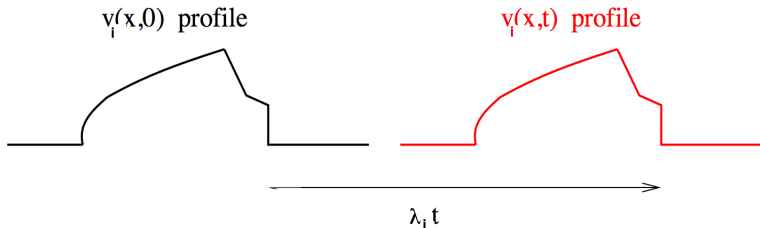
$$\partial_t \mathbf{V} + \mathbf{D} \partial_x \mathbf{V} = 0$$

with the characteristic variables or *waves* :

$$\mathbf{V} = \mathbf{K}^{-1} \mathbf{U} \tag{6}$$

General solutions

\Rightarrow Given initial conditions $\mathbf{U}(x, 0)$, \mathbf{V} basis of decoupled linear advection solutions : $V_i(x, t) = V_i(x - \lambda_i t, 0)$



whose linear superposition via (6) provides, the general solution

\Rightarrow In (5), what if a_{11} and $a_{22} = 0$? What if $a_{12} = a_{21}$?

General solutions

- Examples
 - ⇒ the 1D second order wave equation (Leveque1 6.1)
 - ⇒ the 2D second order wave equation (Leveque1 6.2)
 - ⇒ the **linearized shallow water** equations (Leveque1 6.3, Balsara 1.7)
 - ⇒ the linearized isothermal HD equations (Toro 2.5, 2.10)

Flux Jacobian matrix

- In matrix form, the 1D conservation laws without source terms are written :

$$\partial_t \mathbf{U} + \partial_x [\mathbf{X} \mathbf{X} \mathbf{F}(\mathbf{U})] = 0$$

where

$$\mathbf{U} = (\rho, \rho v, e)$$

and the vector of fluxes is :

$$\mathbf{F}(\mathbf{U}) = \begin{pmatrix} \rho v \\ \rho v^2 + P \\ (e + P) v \end{pmatrix}$$

Flux Jacobian matrix

- Chain rule => quasi-linear form

$$\partial_t \mathbf{U} + \mathbf{A}(\mathbf{U}) \partial_x \mathbf{U} = 0 \quad (7)$$

with $\mathbf{A}(\mathbf{U})$ the Jacobian matrix of the flux :

$$\mathbf{A} = \partial \mathbf{F} / \partial \mathbf{U}$$

⇒ conservation equations are quasi-linear

- The Jacobian matrices of the flux for conservative and primitive variables are *similar*

⇒ which eigenvalues for the ideal 2D MHD equations?

⇒ hint: the Jacobian matrix of the flux for primitive variables is, in Cartesian :

$$\mathbf{A} = \begin{pmatrix} u & v & w & 0 & 0 & 0 \\ 0 & u & 0 & 0 & 0 & 0 \\ 0 & 0 & u & 0 & 0 & 0 \\ 0 & 0 & 0 & u & 0 & 0 \\ 0 & 0 & 0 & 0 & u & 0 \\ 0 & 0 & 0 & 0 & 0 & u \end{pmatrix}$$

Linearization of non-linear PDE

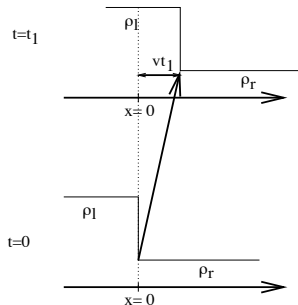
Good news : \exists Taylor expansion to linearize Leveque1 6.4

Linearized Euler equations

- Leveque1 6.4.1
- Solve for wave-like perturbations (Balsara 1.5.2, 1st part)

The Riemann problem

- Step initial state : $\mathbf{U}(x, t = 0) = \begin{cases} \mathbf{U}_L & \text{if } x < 0 \\ \mathbf{U}_R & \text{if } x > 0 \end{cases}$
- Applied to the *linear advection* equation, \exists analytic solution (Toro 2.2.2, Leveque1 6.5)



The Riemann problem

- And for a 1D hyperbolic system of 2 linear PDE with constant coefficients (e.g. linearized isothermal gas dynamics, Toro 2.3.4)?
 \Rightarrow Solve the Riemann problem for (5) with initial state :

$$\begin{cases} \rho_1(x, 0) = \rho_{1,L} & \text{and} & \rho_2(x, 0) = \rho_{2,L} & \text{for } x < 0 \\ \rho_1(x, 0) = \rho_{1,R} & \text{and} & \rho_2(x, 0) = \rho_{2,R} & \text{for } x > 0 \end{cases}$$

- Draw explicitly the solution at $t=1$ for the following numerical values :

$$\begin{cases} \rho_{1,L} = 4 \\ \rho_{1,R} = 1 \\ \rho_{2,L} = 2 \\ \rho_{2,R} = 100 \end{cases} \quad \text{and} \quad \begin{cases} a_{11} = 0.5 \\ a_{12} = 1 \\ a_{21} = -1.25 \\ a_{22} = 3.5 \end{cases}$$

The Riemann problem

- For 3 equations, see linearized Euler (Balsara 1.5.2, 2nd part)
- In the general 1D case of N equations, see Toro 2.3.3 and Balsara 3.4.1 and 3.4.2
- A 1st test for a num. solver (e.g. the finite difference method)
- At the basis of any finite volume method (i.e. 'on a grid'), where data can be represented as a piecewise constant function
- Connection with Rankine-Hugoniot jump conditions at shock

- ⇒ 5 first chapters of Leveque1
- ⇒ Next course → Chapter 10 of Leveque1