

Computational Methods for Astrophysical Applications

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References

- **Books**

- *Riemann solvers and numerical methods for fluid dynamics - A practical introduction*, by Toro
- *Numerical methods for conservation laws*, by Leveque
- *Finite volume methods for hyperbolic problems*, by Leveque

- **Courses**

- Rony Keppens & Jon Sundqvist 2016-2017
- *Numerical PDE Techniques for Scientists and Engineers*, by Dinshaw Balsara

- **Schools**

- Les Houches
- Numerical techniques in MHD simulations, Köln University

Lesson 1 : Hyperbolic Partial Differential Equations

- **Conservation laws**

Integral and differential forms

Examples

- **Hyperbolic PDE**

Matrix formulation of conservation laws

Time advance

Linear advective equations

The Riemann problem

- 1 Conservation laws
 - Integral and differential forms
 - Examples

- 2 Hyperbolic PDE
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Integral form

- Continuous medium hypothesis : from actual particles to particles of fluid
 - ⇒ distribution function and integration of Boltzmann equation : the closure problem
 - ⇒ analogy with moments of the specific photon intensity (see Prof. Sundqvist's course)
- Consider **any volume element** dV of a fluid
 - ⇒ XXX bilan XXX over dt of scalar quantity : inflow/outflow, sinks/sources
 - ⇒ for the momentum : tensor formulation
 - ⇒ for energy
- Shocks and jump conditions : Rankine-Hugoniot

Conservation form

- Provided the variables $X(\mathbf{r}, t)$ are XXX differentiable XXX

$$\partial_t X + \nabla \cdot \underbrace{[F(X, \mathbf{v}, t)]}_{\text{fluxes}} = \underbrace{S(\mathbf{r}, \mathbf{v}, t)}_{\text{sources/sinks}}$$

- Properties of this differential form
 - \Rightarrow Coordinate / dimension independent formulation
 - \Rightarrow Given initial and boundary conditions, can provide a general solution
 - \Rightarrow Conservation form : Green-Ostrogradsky (or Gauss) law and Eulerian approach

Primitive form

- Lagrangian approach XXX Picture of a parachute XXX

$$D_t(\cdot) = \partial_t(\cdot) + \mathbf{v} \nabla(\cdot)$$

- Continuity equation

$$D_t(\rho) = -\rho \nabla \mathbf{v}$$

\Rightarrow incompressible fluid \neq flow

- Navier-Stokes equation

$$\rho D_t(\mathbf{v}) = -\nabla P$$

- Energy equation

\Rightarrow internal and mechanical energy

\Rightarrow link w/ 1st principle of Thermodynamics

\Rightarrow entropy formulation

Closure relation

- Equation-of-state of an ideal gas

$$u = \frac{P}{\gamma - 1}$$

\Rightarrow adiabatic index

- A classic way-around : the polytropic assumption

$$P \propto \rho^\alpha$$

$\Rightarrow \alpha$ polytropic index

\Rightarrow the isentropic case

Linear advection equation

- perform von Neumann stability analysis for Lax–Friedrichs

$$e^{\lambda \Delta t} = \cos(k \Delta x) - i \frac{v \Delta t}{\Delta x} \sin(k \Delta x)$$

⇒ **conditional stability** requiring Courant number C

$$C \equiv \frac{|v| \Delta t}{\Delta x} \leq 1$$

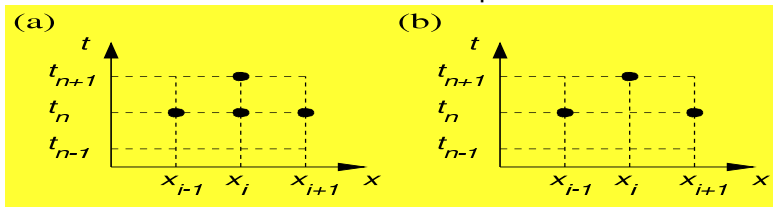
⇒ limitation of the time step Δt for a given resolution Δx

⇒ **Courant–Friedrichs–Lewy** condition (1928)

⇒ necessary (not sufficient) condition for stability!

Euler equations

- in (x, t) space, we identify **stencil** of a method
 \Rightarrow stencils visualizes discrete dependence of method



\Rightarrow stencil for FTCS (a) versus Lax-Friedrichs (b)

- hyperbolic** PDE and **physical characteristics**

\Rightarrow the advection equation is hyperbolic as

$$\frac{\partial u}{\partial t} + \frac{\partial \left(\underbrace{vu}_F \right)}{\partial x} = 0$$

and Flux Jacobian $\frac{\partial F}{\partial u} = v$ is real number 'characteristic speed'

Ideal Magneto-Hydrodynamics equations

- for second order wave equation

$$\frac{\partial^2 u}{\partial t^2} - v^2 \frac{\partial^2 u}{\partial x^2} = 0$$

⇒ factorizes to

$$\left(\frac{\partial}{\partial t} - v \frac{\partial}{\partial x} \right) \left(\frac{\partial}{\partial t} + v \frac{\partial}{\partial x} \right) u = 0$$

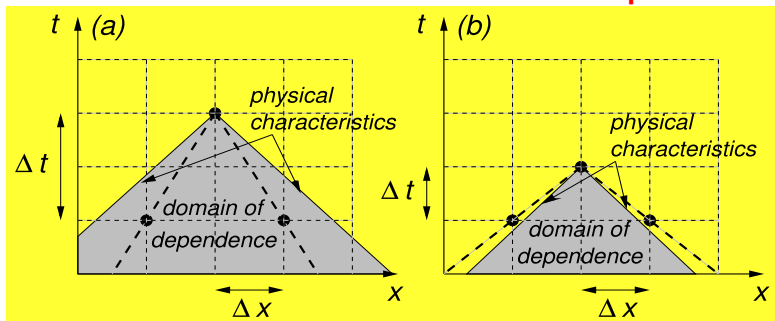
⇒ general solution has left and right going wave with

$$u = f(x - vt) + g(x + vt)$$

⇒ initial shapes $f(x)$, $g(x)$ combine

⇒ 2 characteristics $\frac{dx}{dt} = \pm v$

- illustrate CFL for second order wave equation:
the domain of dependence of the differential equation should be contained in the DOD of the discretised equations



⇒ stability means physical DOD contained in stencil bounds (numerical DOD), hence Δt small enough (right case)

- note: linear advection + wave equation: DOD only involves 1 or 2 points from $t = 0 \leftrightarrow$ HD: DOD bounds set by $v \pm c_s$ with c_s sound speed, delimits $t = 0$ interval

- Second cure: maintain space-time symmetry of the PDE
 - ⇒ use central discretisation for both x and t
 - ⇒ obtain **leapfrog** scheme

$$u_i^{n+1} = u_i^{n-1} - \frac{\Delta t}{\Delta x} (F_{i+1}^n - F_{i-1}^n)$$

- ⇒ numerical flux function for advection is $F_i^n \equiv v u_i^n$
- ⇒ conditionally stable and second-order accurate
- ⇒ multiple time levels involved: $n-1, n, n+1$
- ⇒ potential problem: even/odd time levels may 'decouple'

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Flux Jacobian

- von Neumann stability analysis for BTCS scheme

$$|e^{\lambda \Delta t}| = \frac{1}{|1 + i \frac{v \Delta t}{\Delta x} \sin(k \Delta x)|} < 1 \quad \text{for all } k$$

\Rightarrow **unconditionally stable**, any (large) time step Δt allowed

- note: **stability does not imply accuracy**

\Rightarrow large Δt affects accuracy, defines time resolution:
behavior may involve physical timescale that needs to be resolved!

- implicit backward Euler: first order in time

Eigenvalues & eigenvectors

- von Neumann stability analysis for BTCS scheme

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Hyperbolic, parabolic and elliptic PDE

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Linear, non-linear & quasi-linear

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Explicit

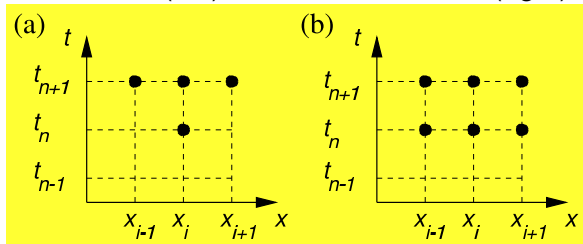
- spatial differences as average of n -th and $(n+1)$ -th time step

$$u_i^{n+1} = u_i^n - \frac{1}{4}v \frac{\Delta t}{\Delta x} (u_{i+1}^{n+1} + u_{i+1}^n - u_{i-1}^{n+1} - u_{i-1}^n)$$

\Rightarrow second order **Crank–Nicolson method**

\Rightarrow **Exercise**: show that this scheme is unconditionally stable,
2nd order accurate

- stencils for BTCS (left) and Crank-Nicolson (right)



Implicit

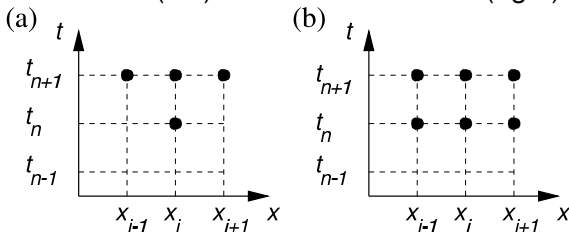
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Properties

- many practical implementations use ‘method of lines’
 - \Rightarrow vector \mathbf{u} of unknowns after first spatial discretization
 - \Rightarrow obtain ODE system

$$\frac{d\mathbf{u}}{dt} = \mathbf{f}(\mathbf{u})$$

\Rightarrow RHS vector function \mathbf{f} could even be nonlinear in \mathbf{u}

- discretize ODE in time using parameter α in

$$\mathbf{u}^{n+1} = \mathbf{u}^n + \Delta t \left[\alpha \mathbf{f}(\mathbf{u}^{n+1}) + (1 - \alpha) \mathbf{f}(\mathbf{u}^n) \right]$$

\Rightarrow note case $\alpha = 0$: explicit (unstable) forward Euler method

Characteristics

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Stencils, domain of dependence, range of influence

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The Riemann problem

$$\mathbf{u}^{n+1} = \mathbf{u}^n + \Delta t \left[\alpha \mathbf{f}(\mathbf{u}^{n+1}) + (1 - \alpha) \mathbf{f}(\mathbf{u}^n) \right]$$

- $\alpha = 1$ is **implicit backward Euler**
- $\alpha = 1/2$ gives second-order accuracy, **trapezoidal method**
 \Rightarrow Crank-Nicolson for central discretization of flux in \mathbf{f}
- when \mathbf{f} nonlinear: linearize using

$$\mathbf{f}(\mathbf{u}^{n+1}) \approx \mathbf{f}(\mathbf{u}^n) + \frac{\partial \mathbf{f}^n}{\partial \mathbf{u}} (\mathbf{u}^{n+1} - \mathbf{u}^n)$$

\Rightarrow introduces matrix $\frac{\partial \mathbf{f}^n}{\partial \mathbf{u}}$ called “**Jacobian matrix**” of \mathbf{f}