

Task #2: Are We Alone?

1/04/26

Objective: Identify which conditions that make a planet capable of supporting life and determine the most promising exoplanet for future exploration.

The Drake Equation:

- 1) The last four coefficients because as of right now the only basis we can estimate them from is ourselves.
- 2) Distance from our planet to potential other planets.
- 3) We are assuming Earth-like conditions are ideal for all life.
- 4) The pessimism line is essentially the H2O, its probability is so low (statistically significant) that it can't be true and thus there is other life out there.
- 5) If civilizations existence is temporary, we can study their downfall to prevent ours.
- 6) Exoplanets means a possibility of Earth-like life. And if our searches return empty-handed, perhaps the possibility of a hidden factor not yet understood.

$$N = R_* \cdot f_p \cdot N_e \cdot f_i \cdot f_l \cdot f_c \cdot L$$

- N = Number of civilizations we could possibly communicate

- R_* = Rate of star formation in our galaxy

$\approx 1.65 \pm 0.19 M_\odot$ per year (Solar masses/year)

- f_p = fraction of stars with planets
 $\approx 90 - 100\%$.

- N_e = Average Number of planets per star that could support life
 $\approx 0.37 - 0.88$ planets per star

Very rough

estimations

because our

only frame

of reference
is ourselves

- f_i = fraction of habitable planets that actually develop life: $0 - 1$

- f_l = fraction of those with life that evolve intelligent life: $0.003 - 1$

- f_c = Fraction of intelligent civilizations that develop technology to communicate: $10 - 20\%$.

- L = length of time such civilizations release detectable signals: A number from 100

- a few thousand

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- Scenarios:

1) Best-case scenario (optimistic):

$$2 \cdot 10^6 \cdot 1 \cdot 100 \cdot 100 \cdot 100 \cdot 100 = 100,000,000$$

$\approx 200,000,000$ planets with civilizations

2) Worst-case scenario (pessimistic):

$$1 \cdot 10^6 \cdot 1 \cdot 100 \cdot 100 \cdot 100 \cdot 100 = 100$$

≈ 10 civilizations

3) Personal estimate:

$$1.69 \times 10^6 \cdot 1 \cdot 204 \cdot 37 \cdot 104 \cdot 1000 \\ = 3 \text{ civilizations}$$

- Reflection:

- 1) I am surprised by just the sheer impossibility of life on other planets, given the optimistic view.
 - 2) Yes, it has only 0.0005% of stars have planets with civilizations.
 - 3) Yes, because we would understand that we may be a part of a broader civilization.
- A) No, there isn't much we could do about either answer within my life span.
- B) It's so rare! Then why do we exist?

- Habitable Zone:

$$d = \sqrt{\frac{L}{c}}$$

d = distance from star in AU

L = star's luminosity in (solar luminosities)

c = constant based on the habitable zone boundary

C Table:

	<u>C</u> conservative	<u>C</u> optimistic
inner	1.1	1.78
outer	0.53	0.32

* Only applicable to main-sequence stars (Types G, K, M)

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- Factors for Life:

- Star Type: Controls the stability, amount of energy, amount of radiation, and duration of possible life supporting planets.
- Location in the habitable zone: too far to either extreme prevents liquid water, which is key for life.
- Presence of water: A high abundance of liquid water is needed, too much or too little, or the wrong phase could hurt the chances of life.
- Atmosphere: A proper atmosphere traps some (not too much or too little) heat protects from radiation, and is not composed of toxic gases.
- Greenhouse effect: Must be kept in middle, too little means it's too cold, and too much means it's too hot.
- Planetary Mass & Size: Too little gravity cannot sustain an atmosphere, and too much may harm potential life.
- Magnetic Field: Proper field protects against radiation and coronal mass ejections, without this, the atmosphere & possible life can be damaged.
- Geologic Activity: Moderate amounts can recycle carbon & provide heat for liquid water, too much can be deadly.
- Star's Activity: The star must be moderately calm, too much UV, instability & radiation can damage the planet.
- Day length & Temperature: Day length must be moderate for smooth wind speeds & stable circulatory temperatures. Too extreme can cause extreme temperatures on the day & night sides, life could only exist in oases.
- Chemical Building Blocks: Quite literally needed in some high quantity for life to form.
- Energy sources: Life needs constant steady energy, can come in many of ways, but it must be consistent.
- Surface Temperature: Needs to be moderate for liquid water, but not extreme enough that the water changes phase or life is damaged.

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The Hub's Rubric

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~~F₁₀ Planets:~~

- TRAPPIST-1e

- $d_{\text{inner}} = \sqrt{\frac{1}{c}} = \sqrt{0.00052}$
- $d_{\text{outer}} = \frac{d_c + d_o}{2} = \frac{\sqrt{0.00052} + \sqrt{0.00052}}{2} = \frac{1.1}{2} = 0.55 \text{ AU}$
- $d_{\text{inner}} = \sqrt{\frac{1}{c}} = \sqrt{0.00052} = 0.0224 \text{ AU}$
- $d_{\text{outer}} = \frac{d_c + d_o}{2} = \frac{\sqrt{0.00052} + \sqrt{0.00052}}{2} = \frac{0.32}{2} = 0.16 \text{ AU}$
- Location in habitable zone: 4/5
 - Not perfectly in middle but not in either extreme
- Star Type: 4/5
 - M type would be 3/5, but TRAPPIST-1 is relatively quiet
- Mass & size: 3/5
 - Earth-like gravity but likely thinner atmosphere
- Water presence: 2/5
 - Some possibility; NOT known
- Atmosphere: 2/5
 - Unknown composition & likely thin
- Surface temp: 2/5
 - Extreme roughly on the cold side
- Chemical building blocks: 2/5
 - Possible carbons, ONLY 2 points because UNKNOWN

Total: 24/35

- Kepler-70 b:

$$\begin{aligned} d_{\text{inner}} &= 0.006 \text{ AU} \\ d_{\text{outer}} &= 3.19 \text{ AU} = \frac{1}{2} \left(\sqrt{\frac{2.2}{1.01}} + \sqrt{\frac{2.2}{1.78}} \right) \\ d_{\text{inner}} &= 7.367 \text{ AU} = \frac{1}{2} \left(\sqrt{\frac{2.3}{0.53}} + \sqrt{\frac{2.2}{0.32}} \right) \end{aligned}$$

1/5 - Habitability Zone:

- Far too close

1/5 - Star Type:

- Still likely too extreme radiation

2/5 - Mass & size:

- Sub-Earth > than atmosphere

1/5 - H₂O presence

- Impossible

1/5 - Atmosphere:

- Likely stripped away

1/5 - Surface temp:

- Way too extreme

1/5 - Chemical building blocks

- UNKNOWN & unsupported

Total: 8/35

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WASP-A3B

$$\begin{aligned} d_{\text{real}} &= 0.012 \text{ AU} \\ d_{\text{inner}} &= 0.1204 \text{ AU} \\ d_{\text{outer}} &= 0.222 \text{ AU} \end{aligned}$$

$$= \frac{1}{2} \left(\sqrt{\frac{0.02}{1.11}} + \sqrt{\frac{0.02}{1.78}} \right)$$

$$= \frac{1}{2} \left(\sqrt{\frac{0.02}{0.53}} + \sqrt{\frac{0.02}{0.32}} \right)$$

- 1/5 - Habitability zone
- Far too close to star
- 5/5 - Star Type
- Main sequence star type II
- 1/5 - Planet size mass
- Far to large & massive for surface habitability
- 3/5 - Water presence
- Vapor atmosphere
- 3/5 - Atmosphere
- Way too dense
- Not proper elements
- 1/5 - Surface Temp
- Far too hot
- 3/5 - Chemical Building Blocks
C, H, O
- Total: 16/35

Kepler-992b

- $d_{\text{real}} = 0.9 \text{ AU}$
- $d_{\text{inner}} = 0.6650 \text{ AU} = \frac{1}{2} \left(\sqrt{\frac{0.61}{1.11}} + \sqrt{\frac{0.61}{1.78}} \right)$
- $d_{\text{outer}} = 1.2267 \text{ AU} = \frac{1}{2} \left(\sqrt{\frac{0.61}{0.53}} + \sqrt{\frac{0.61}{0.32}} \right)$
- 1/5 - Habitability zone
- Too close to star
- 9/5 - Planet size & Mass
- Super Earth
- 2/5 - Atmosphere
- No water but likely there
- 2/5 - Placeholder
- 5/5 - Star Type
- G-type dwarf, stable, MME radiation, long life
- 0/5 - Water presence
- No water
- 2/5 - Surface Temp
- 900°C
- 3/5 - Chemical Building Blocks
Unknown
- Total: 14

Kepler-182b

$$\begin{aligned} d_{\text{real}} &= 1.096 \text{ AU} \\ d_{\text{inner}} &= 0.9327 \text{ AU} = \frac{1}{2} \left(\sqrt{\frac{1.20}{1.11}} + \sqrt{\frac{1.20}{1.78}} \right) \\ d_{\text{outer}} &= 1.172 \text{ AU} = \frac{1}{2} \left(\sqrt{\frac{1.20}{0.53}} + \sqrt{\frac{1.20}{0.32}} \right) \end{aligned}$$

- 2/5 - H2: Near edge but in 2nd one
- 4/5 - Star Type: G-type Star
- 5/5 - Planet size & Mass: Just big enough without issue
- 0/5 - Water presence: Unknown
- 2/5 - Atmosphere: 2/5 placeholder bc gravity & unknown density
- 3/5 - Surface Temp: -80°C
- 2/5 - Chemical building blocks: Unknown

= 17/35