

# Introduction to Scientific Cameras

Dr Louis Keal  
Teledyne Photometrics

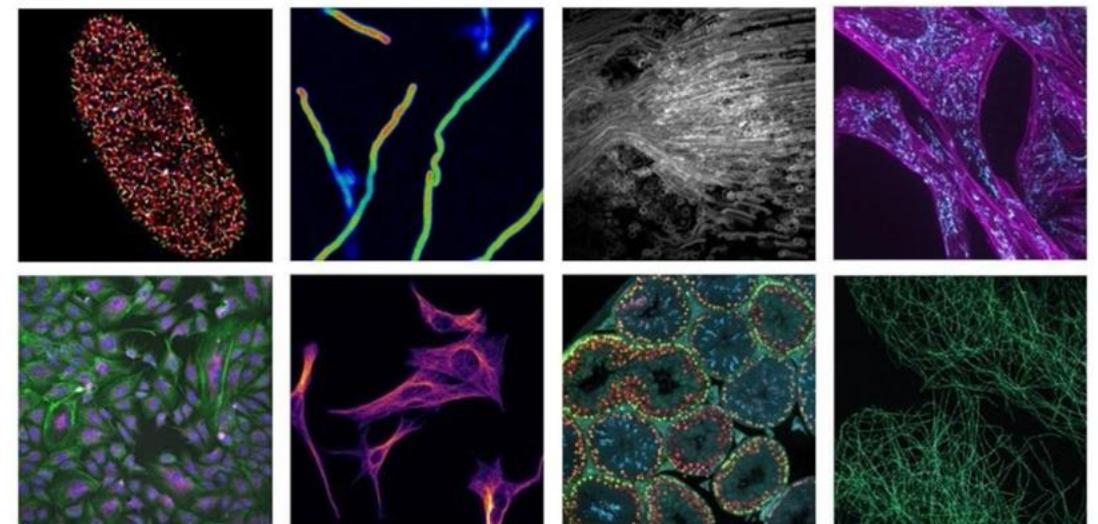
# About Teledyne Photometrics

- Design and manufacture **high end** cameras for **low light imaging**, mostly in **biological applications**
- Market Leader for **Scientific CMOS technology**
- Established in 1978 and part of the **Teledyne Imaging Group** since 2019
- Global Headquarters in Tucson, Arizona, **USA**
- Factory in Surrey, British Columbia, **Canada**
- Sales and Marketing HQ in Birmingham **UK**
- Additional offices around the world



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Part of the Teledyne Imaging Group



# Important Camera Specs & Topics

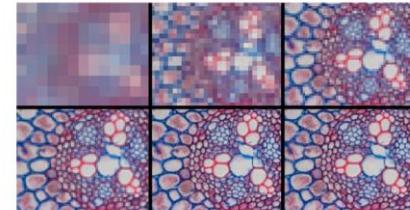
## Sensitivity

- Key question for camera choice:  
Does the camera image with high enough **signal to noise ratio**:  
... at the **experimental conditions** we require (speed, low light dose, etc.)  
... to answer the **scientific questions** we're asking?
- Determined by the balance of the following factors:

**Signal Collection**      vs      **Noise**  
Quantum Efficiency  
(at detection wavelength)  
Image Pixel Size  
  
Read Noise  
Other Noise (Excess Noise Factor)  
Patterns & Artifacts  
Dark Current (only for multi-second exposures)

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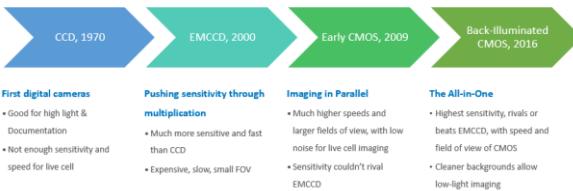
## Resolution



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## Sensitivity

### Timeline of Scientific Imaging



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## Camera Technologies

## Being Quantitative

For **quantitative** intensity analysis, **grey values cannot be used** as they depend on camera-specific settings.  
**Photoelectrons** are our objective, quantitative unit. To calculate back, use:

$$\text{Photoelectrons} = (\text{Grey levels} - \text{Offset}) \times \text{Gain}$$

**Gain:** Ask your camera manufacturer for the gain values for **your specific camera**, or estimate with a **Mean Variance**

**Test:**

**Offset:** Baseline grey level value, **mean image value with no light** (typically 100).

This allows us to:

1. Compare intensity values between **different cameras / camera settings**
2. Estimate **signal to noise ratio**
3. Understand your **light level**

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## Quantitative Imaging



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# Scientific Cameras

Everyone familiar with standard qualitative ‘consumer’ cameras

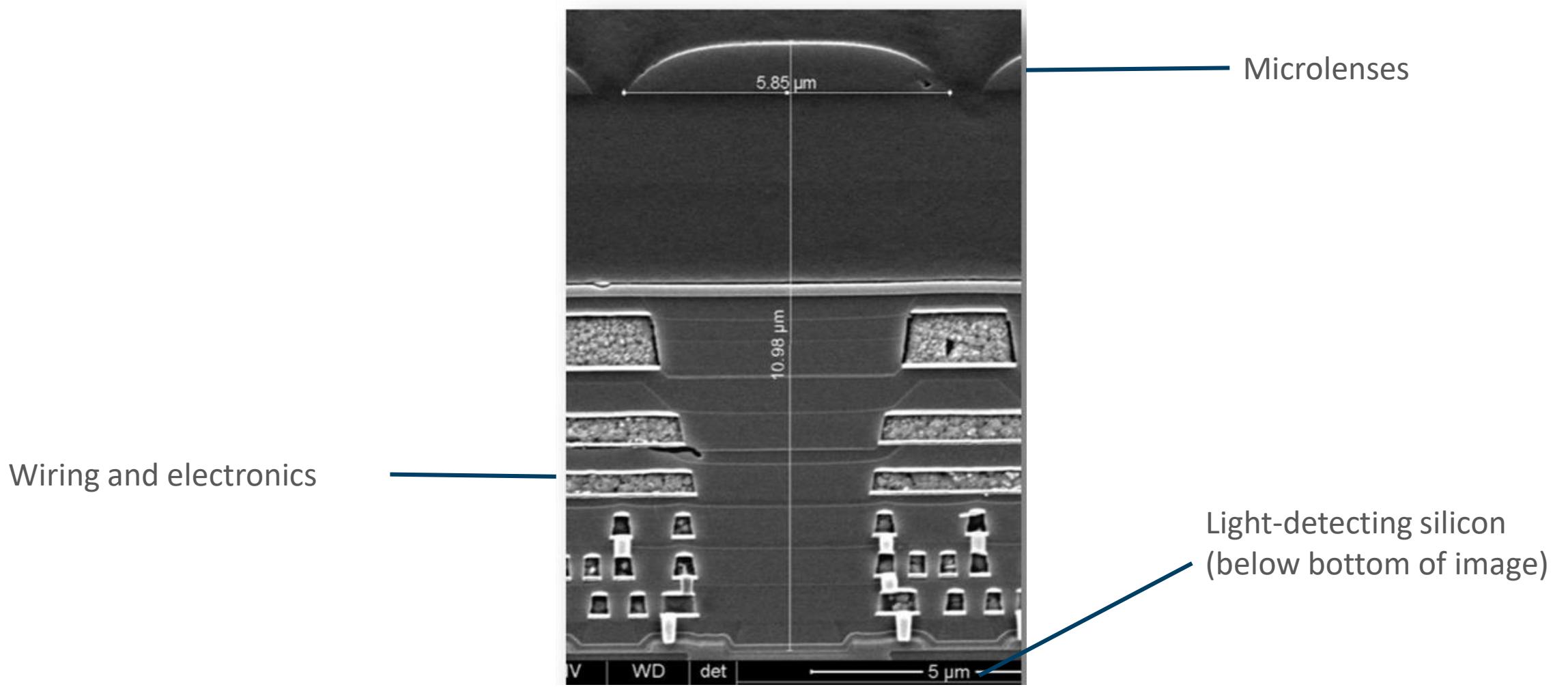
Scientific-grade cameras are **quantitative** and **highly sensitive**

The camera measures specific quantities of light

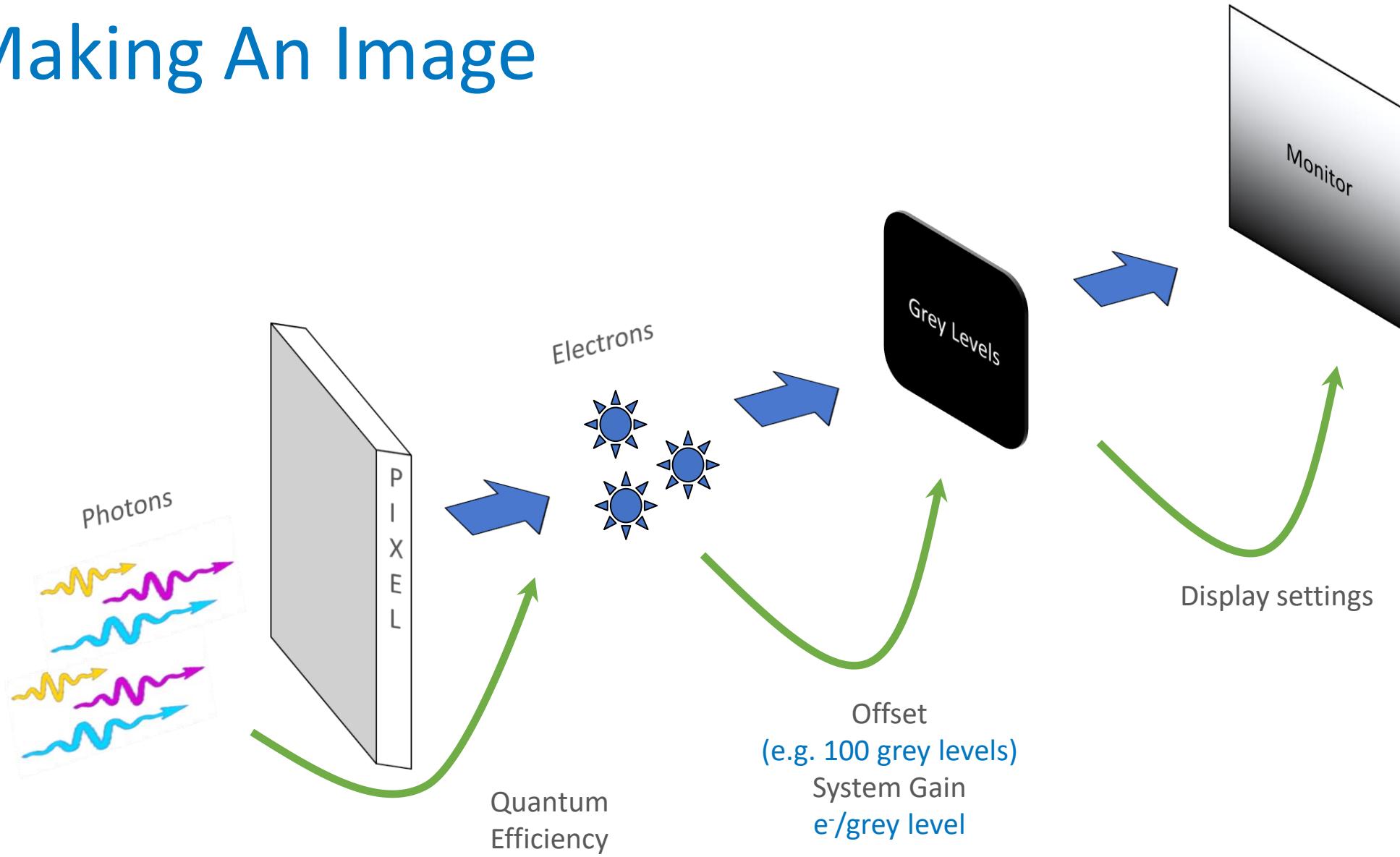
The basic measurable unit of light is the **photon**



# This Is A Pixel



# Making An Image



# Sensitivity

- Key question for camera choice:

Does the camera image with high enough **signal to noise ratio**:

... at the **experimental conditions** we require (speed, low light dose, etc.)

... to answer the **scientific questions** we're asking?

- Determined by the balance of the following factors:

## **Signal Collection**

Quantum Efficiency  
(at detection wavelength)  
Image Pixel Size

**vs**

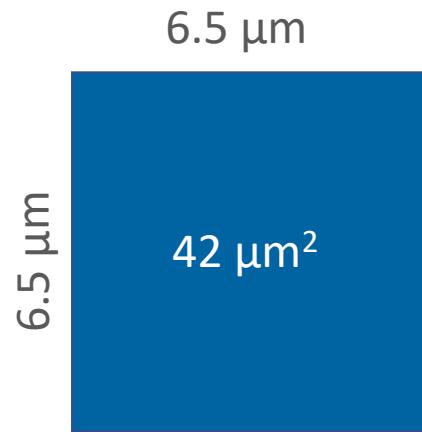
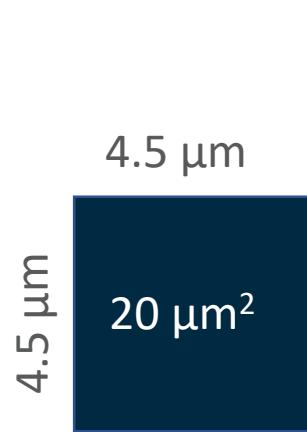
## **Noise**

Read Noise  
Other Noise (Excess Noise Factor)  
Patterns & Artefacts  
Dark Current (only for multi-second exposures)

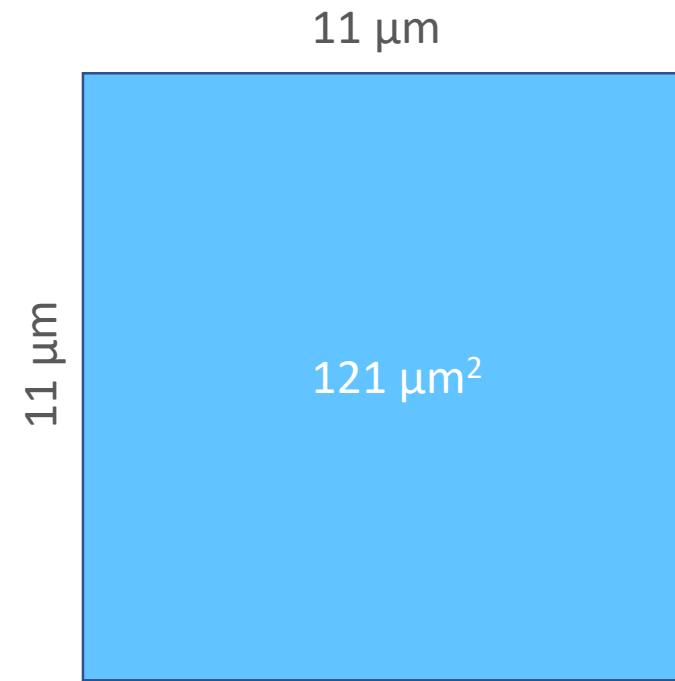


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# Pixel Size Matters



**2x the area**



**6x the area**



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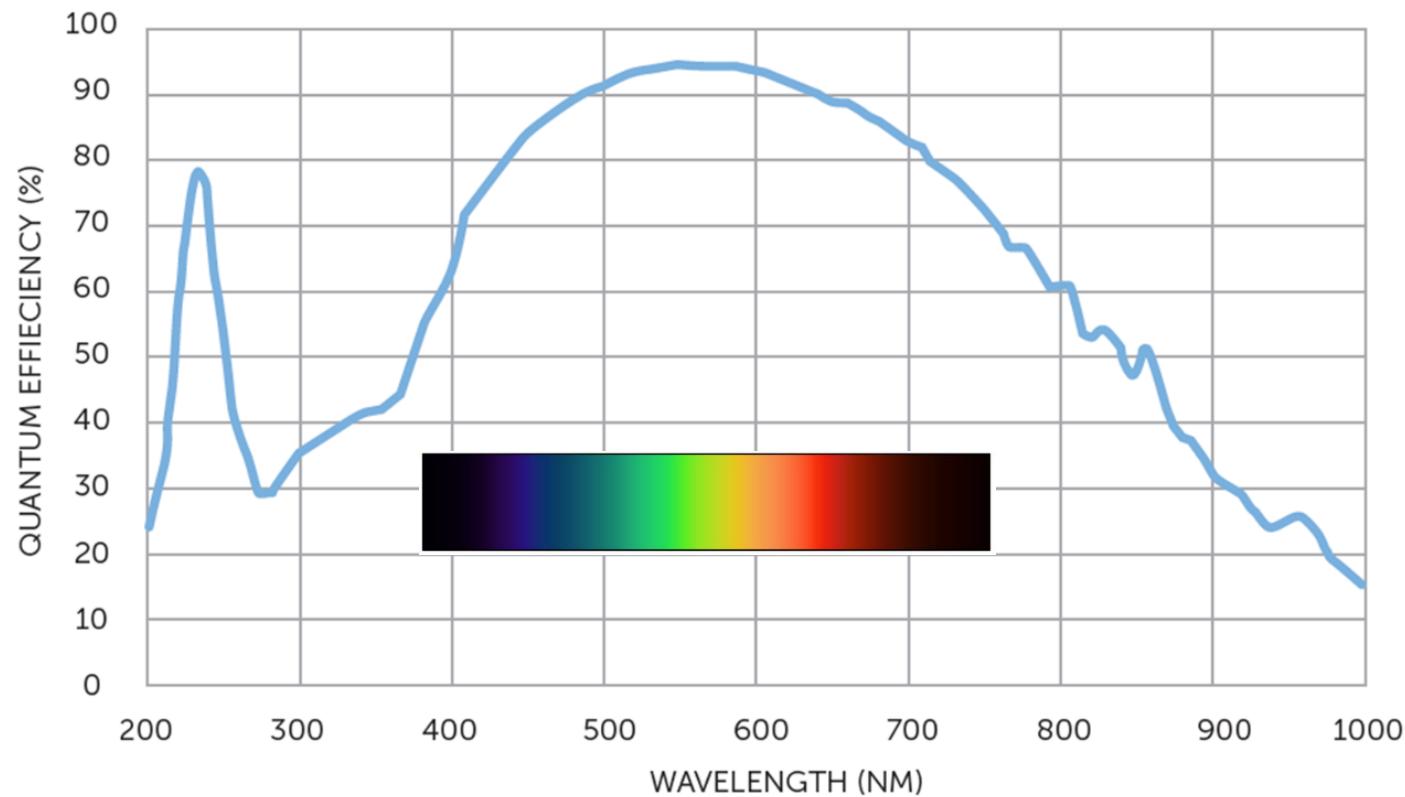
# Quantum Efficiency (QE)

QE is the **percentage** of photons that are converted to photoelectrons

If a sensor has 50% QE and is hit with 500 photons, it will result in a signal of 250 electrons

The lower the QE, the more photons are lost

QE **changes with wavelength**



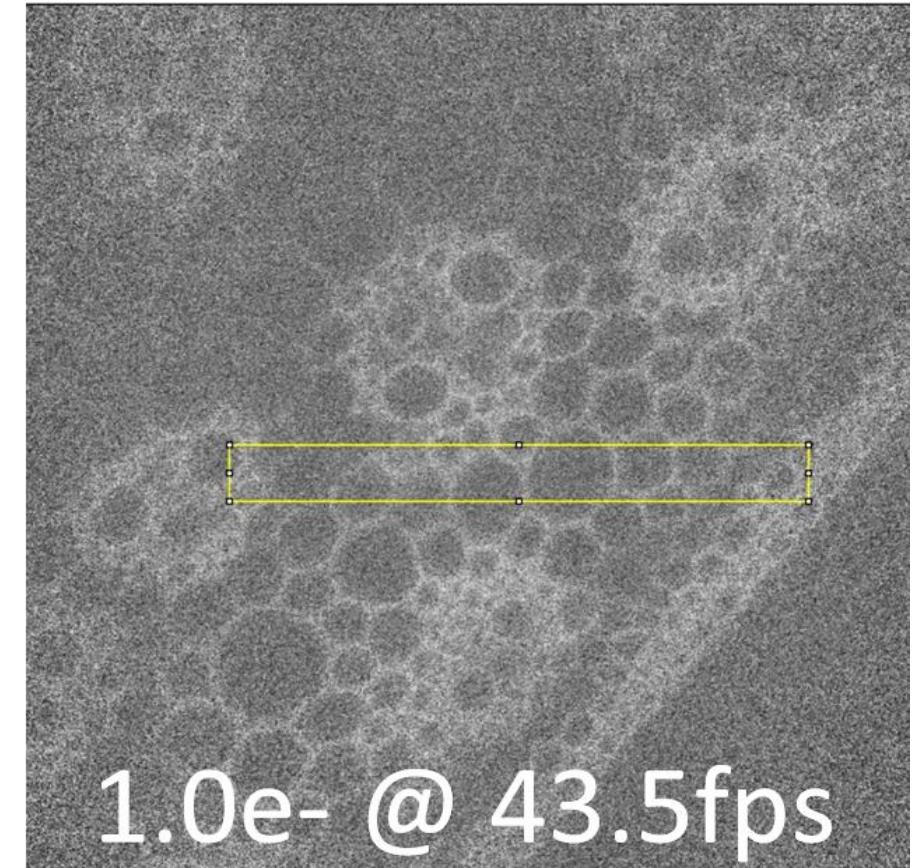
# Read Noise

How accurately can we measure how many photoelectrons we collected?

Read noise is primarily determined by quality of electronic design and speed of readout, and has a **fixed value** for each camera mode.

Prime BSI in CMS mode has a read noise of  **$\pm 1.0$  electrons**.

The lower the light level, the more important it is.



# Photon Shot Noise

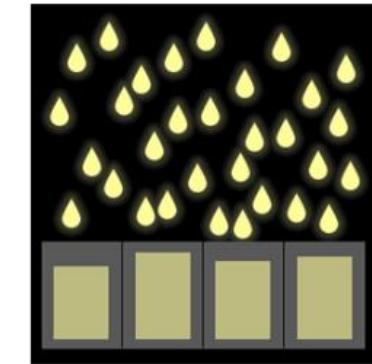
Emission and detection of photons from all sources fluctuates in time.

This is our most commonly visible noise source in imaging.

We may know an **average rate of emission**, but actual photon events are **random** (Poisson Distribution).

If we collect  $N$  photoelectrons of signal,

Our noise will be  $\sqrt{N}$  photoelectrons.



# Dark Current Noise

Dark Current Noise is an **exposure time** and **temperature** dependent noise source.

Thermal motion of electrons can cause them to enter the pixel well **as if they were detected photoelectrons**. This is another random **Poisson behaviour** like Photon Shot Noise.

- **With cooling**, it is typically **negligible for live cell imaging**, and only should be considered for multi-second exposure times.
- **Without cooling**, it must always be considered.

For example, Prime BSI with air cooling achieves **D = 0.5 electrons / pixel / second**.

Our noise contribution is then  $\sqrt{D * t}$ , where  $t$  is our exposure time.

Even at 1 second exposure, this is still well below our read noise.

# Signal To Noise Ratio for CCD & CMOS

$$SNR = \frac{S}{\sqrt{S + D^2 + \sigma_R^2}}$$

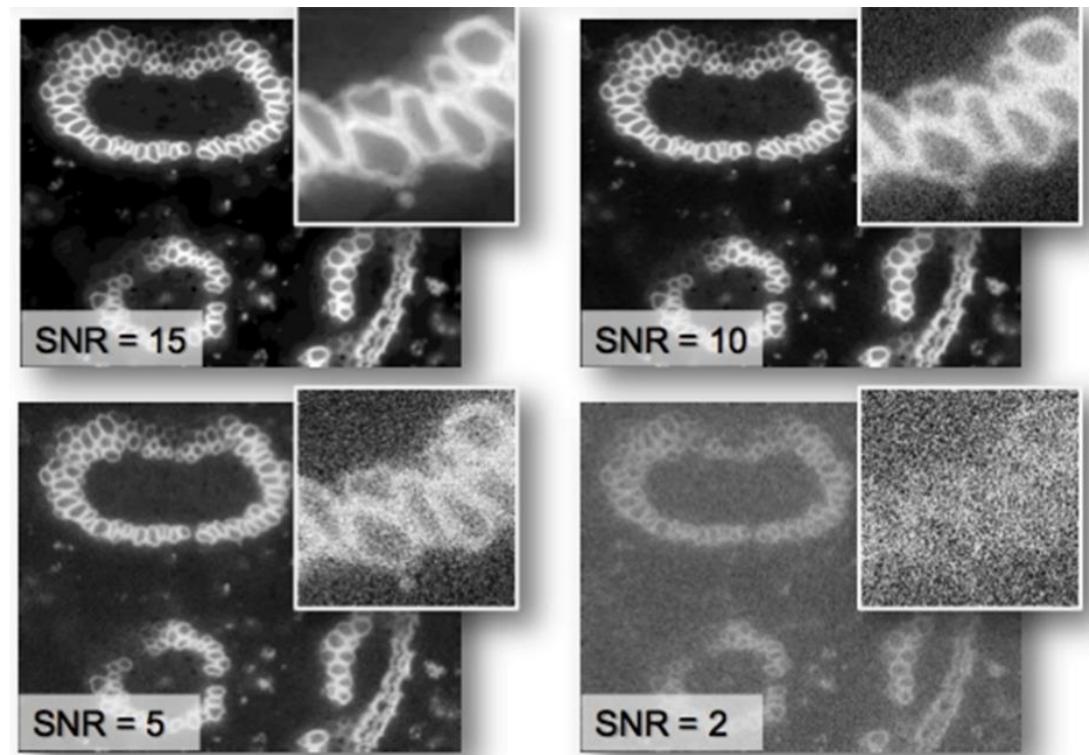
**S:** Signal

**D:** Dark current noise

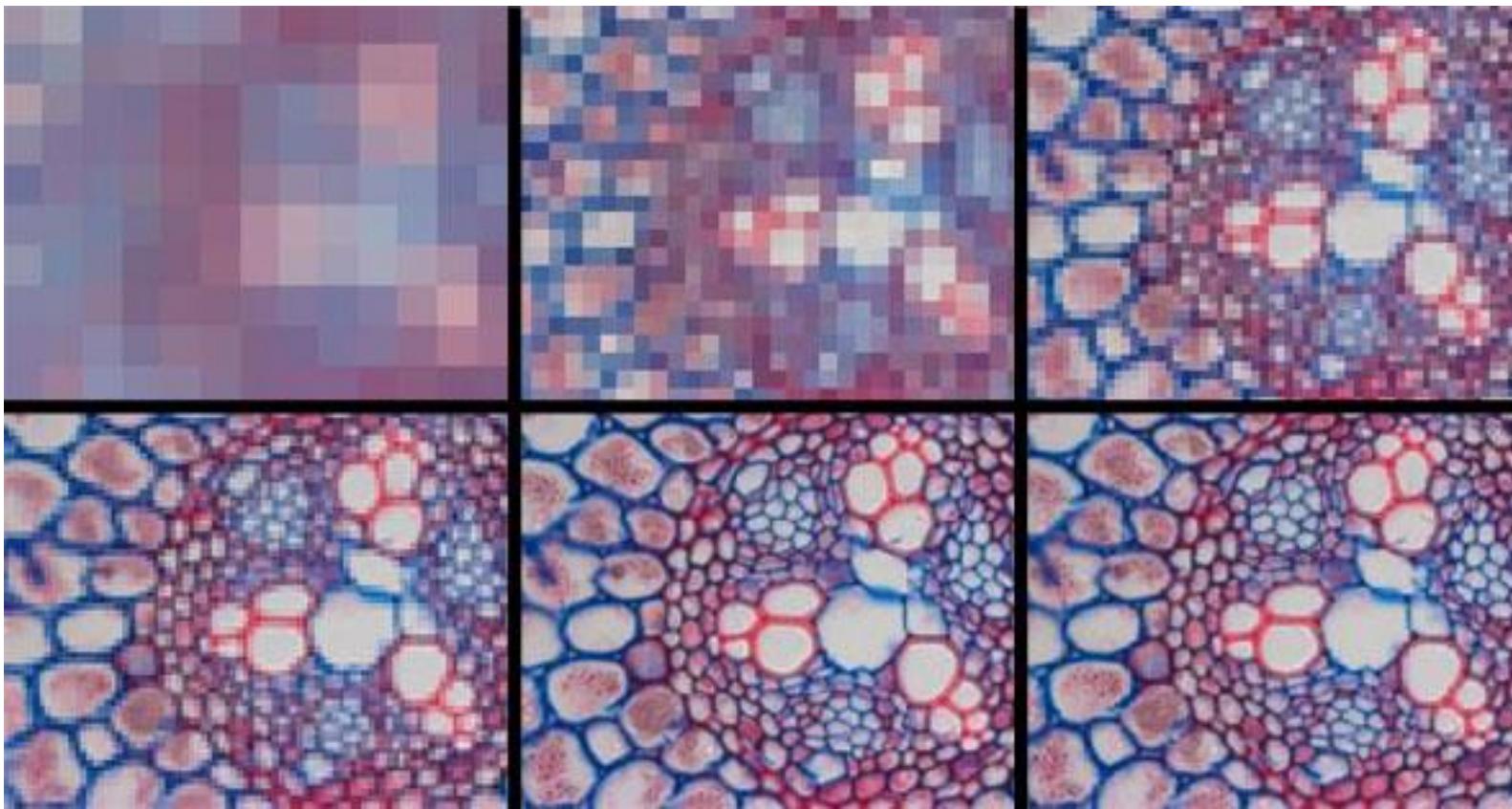
$$D = \sqrt{(\text{Dark current} * \text{exposure time})}$$

**$\sigma_R$ :** Read noise

**$\sqrt{S}$ :** Photon shot noise



# Resolution



# What is Resolution?

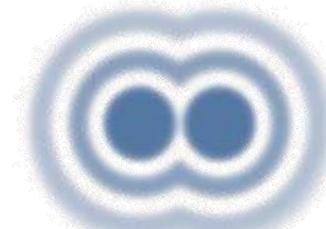
**Spatial Resolution** in fluorescence microscopy is defined as the shortest distance between two points that can still be distinguished.

Determined by two factors:

**Microscope Resolution** – All optics in the light path, numerical aperture of the objective lens, emission wavelength of the sample

**Camera Resolution** – Pixel size after magnification, plus spatial sampling

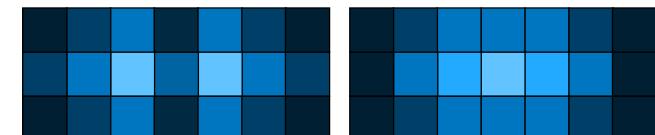
Two Distinct Objects



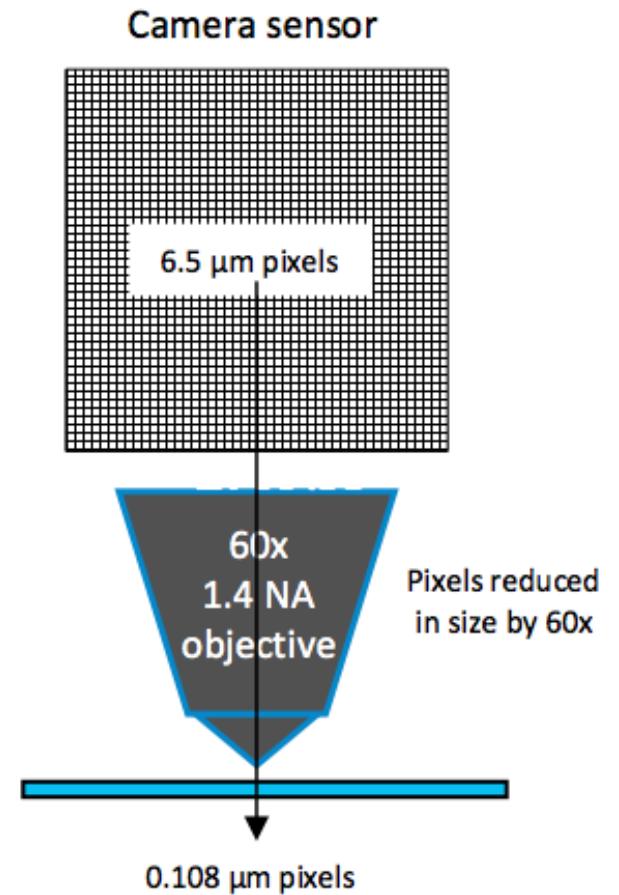
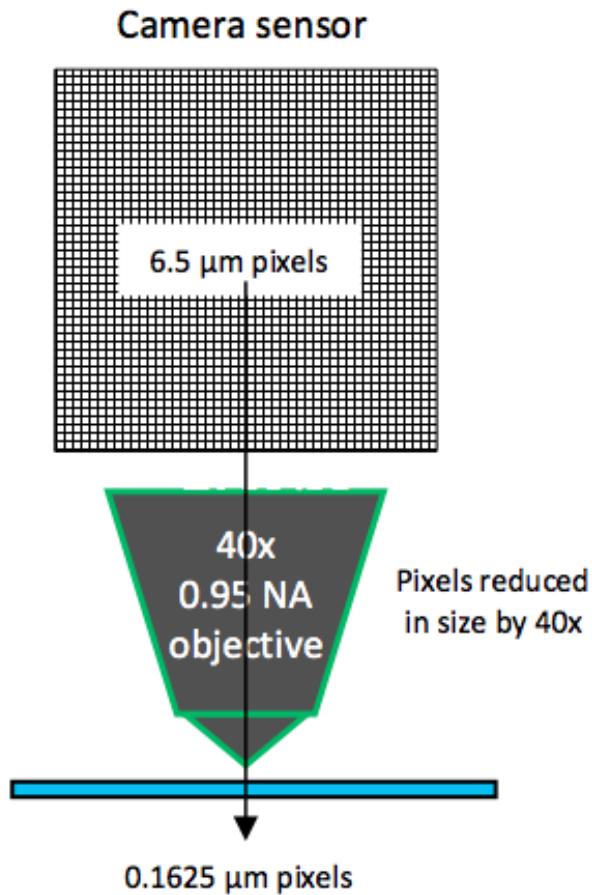
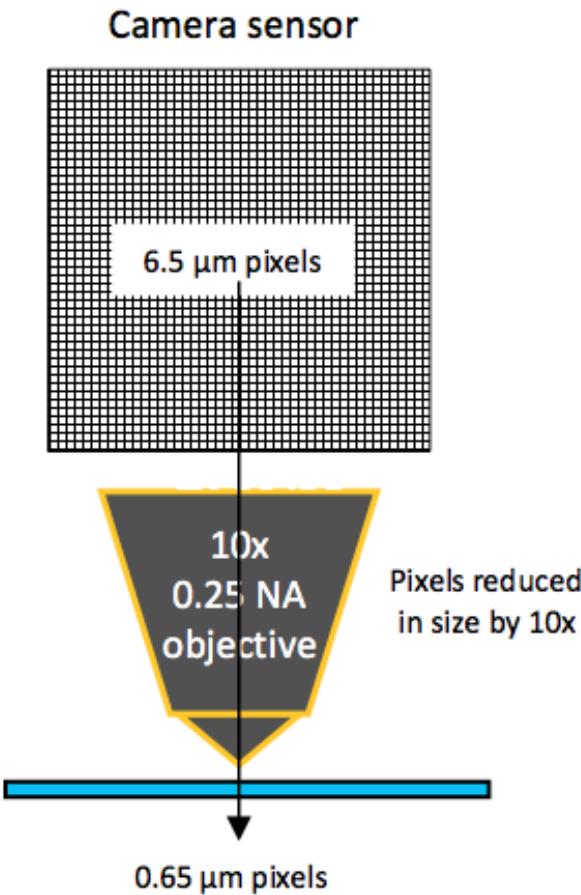
One or Two Objects?



Two Airy disks merge until the two central spots can no longer be differentiated.

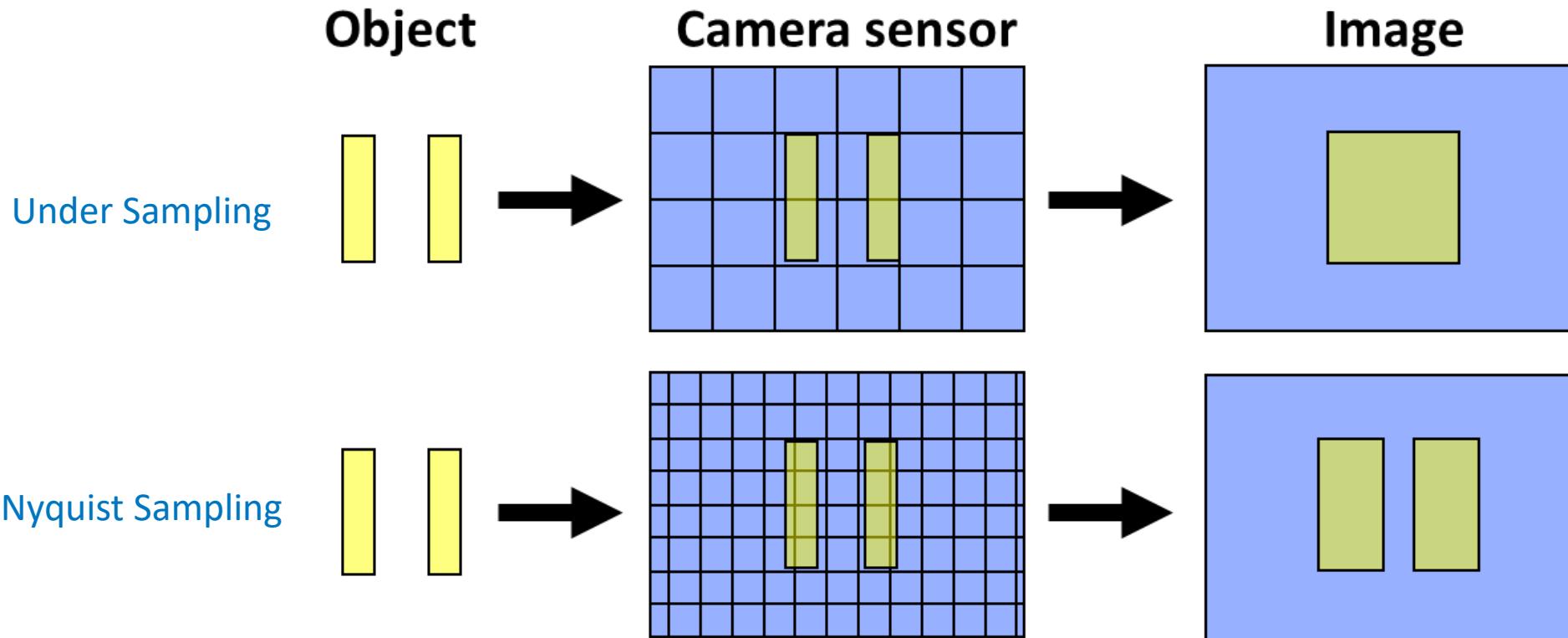


# Pixel Size After Magnification



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# Nyquist Sampling



$$\text{Camera Resolution} = \left( \frac{\text{Pixel size}}{\text{Total magnification}} \right) \times 2.3$$

# When Do I Want To Match Nyquist?

The resolution of your system is limited to the worst component.

When aiming for optimal resolution, match Nyquist.

$$\left( \frac{\text{Pixel size}}{\text{Total magnification}} \right) \times 2.3 > \frac{1.22 \times \text{Wavelength } (\mu\text{m})}{2NA_{obj}}$$

If resolution is less important, use **undersampling**: pixels larger than Nyquist requires. This gives greater sensitivity.

If pixels are smaller than Nyquist requires, this is **oversampling**. This is only used for advanced post-processing techniques such as deconvolution.

# Field Of View

$$FOV = \frac{Sensor\ size\ (diagonal)}{Total\ magnification}$$

## Sensor Sizes:

Blue: 11.6 mm (**EMCCD, 0.25MP**)

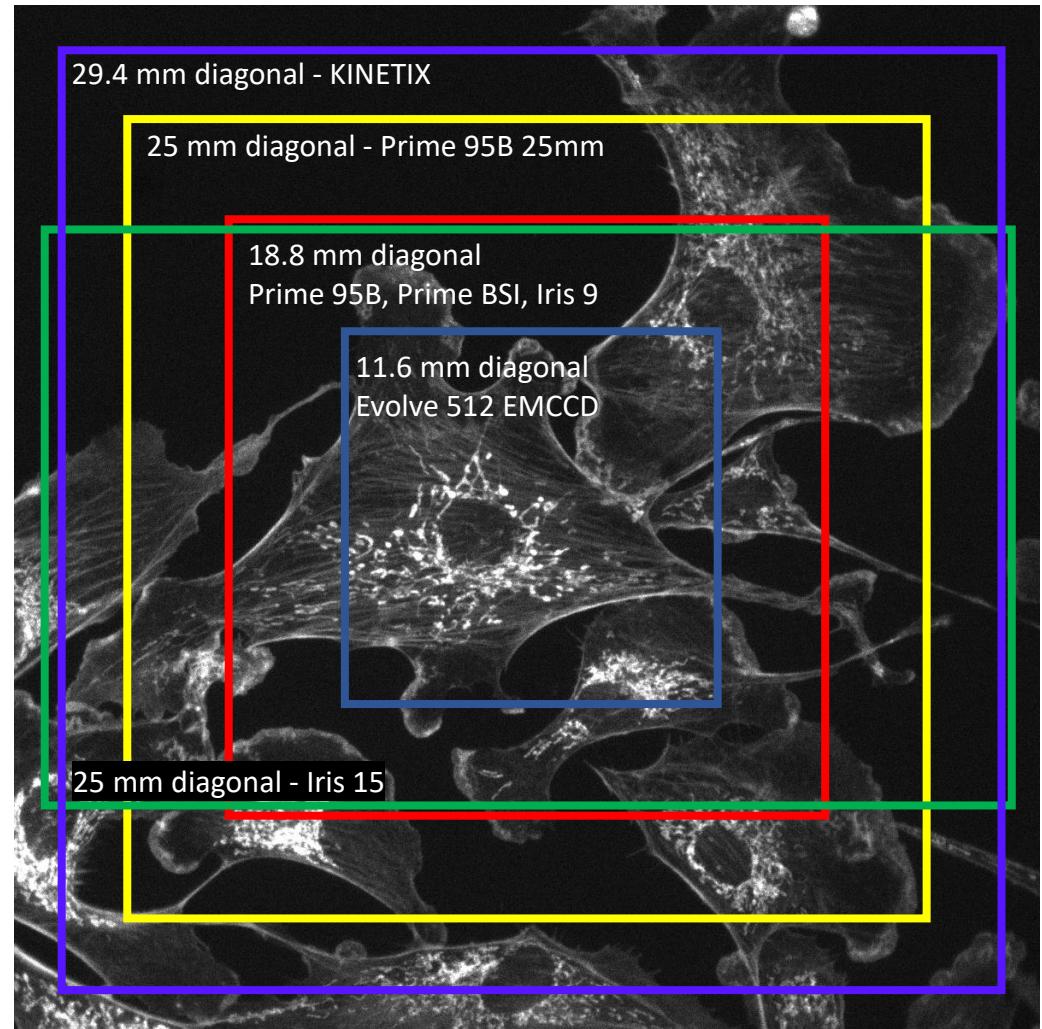
Red: 18.8 mm (**95B 1.4MP, BSI 4.2MP**)

Green: 25 mm (**Iris 15, 15MP**)

Yellow B: 25 mm (**95B 25MM, 2.6MP**)

Purple: 29.4 mm (**Kinetix, 10MP**)

At the same **image pixel size**, FOV is determined by the number of pixels.



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# Speed And Interfaces

**Required exposure time** is the first determining factor for speed → **sensitivity** is a major factor!

Imaging at **50 ms** exposure results in **20 fps**

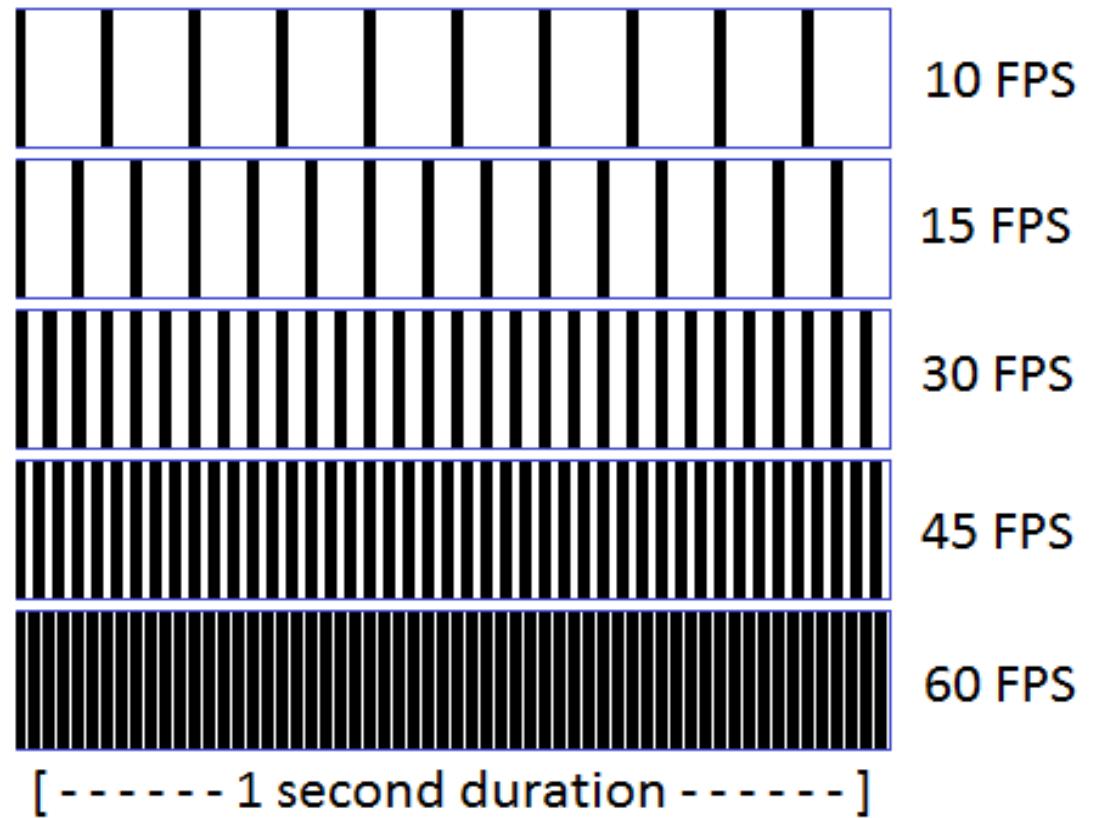
Other speed factors:

Sensor size – row time & number of rows

Digital interface – USB3 vs 3.1g2 vs PCIe

Readout Method – CMS vs High Speed

New CMOS technology is allowing speeds up to **500fps for 10 Megapixels**, maintaining low noise.



# Timeline of Scientific Imaging



## First digital cameras

- Good for high light & Documentation
- Not enough sensitivity and speed for live cell

## Pushing sensitivity through multiplication

- Much more sensitive and fast than CCD
- Expensive, slow, small FOV

## Imaging in Parallel

- Much higher speeds and larger fields of view, with low noise for live cell imaging
- Sensitivity couldn't rival EMCCD

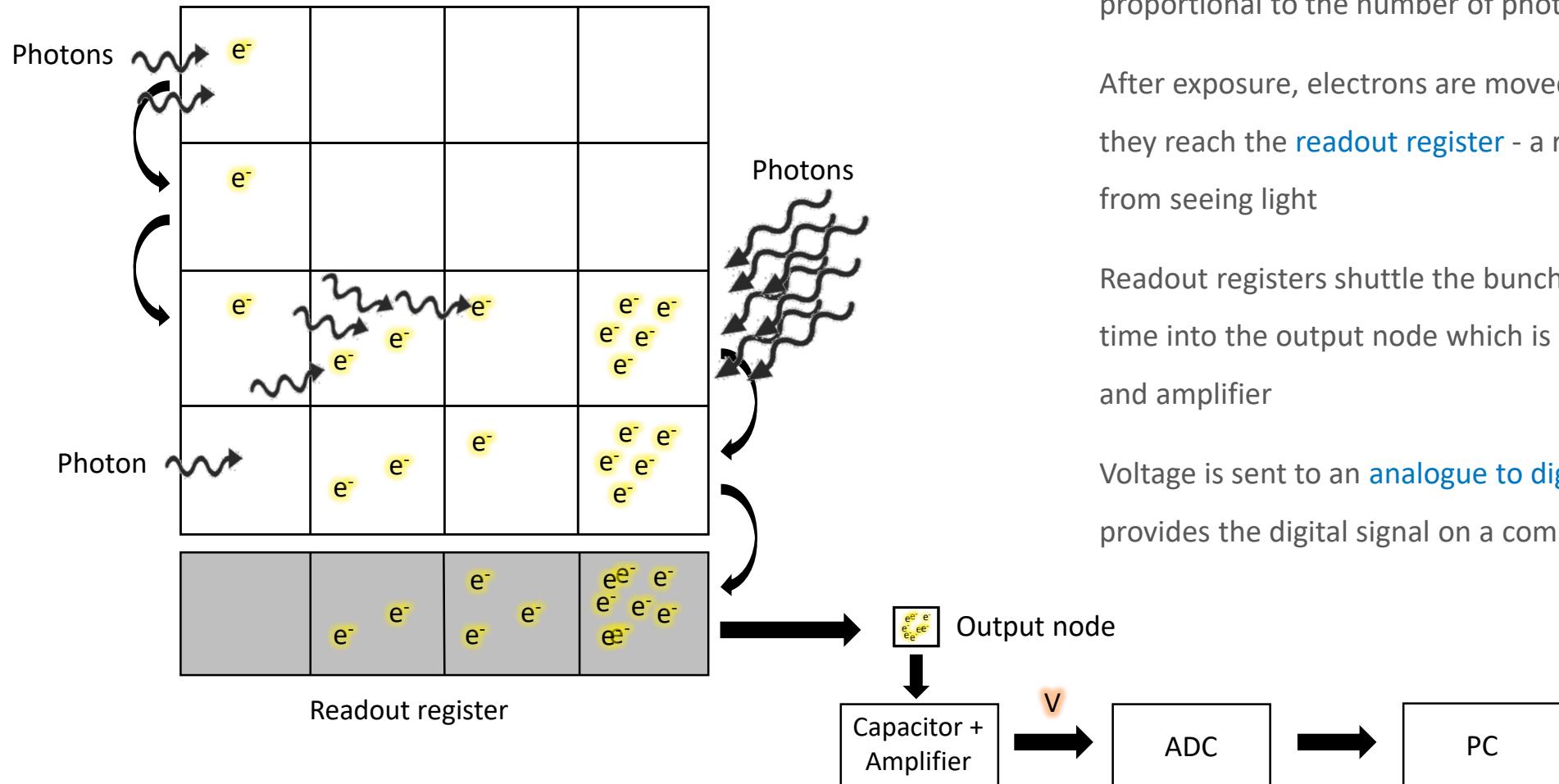
## The All-in-One

- Highest sensitivity, rivals or beats EMCCD, with speed and field of view of CMOS
- Cleaner backgrounds allow low-light imaging



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# CCD Fundamentals



The number of electrons created should be linearly proportional to the number of photons hitting the pixel

After exposure, electrons are moved down, row by row, until they reach the **readout register** - a row that doesn't prevent from seeing light

Readout registers shuttle the bunches of electrons one at a time into the output node which is connected to a capacitor and amplifier

Voltage is sent to an **analogue to digital convertor** which provides the digital signal on a computer

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# EMCCD: Faster and More Sensitive

Used for faster, more sensitive imaging than CCD

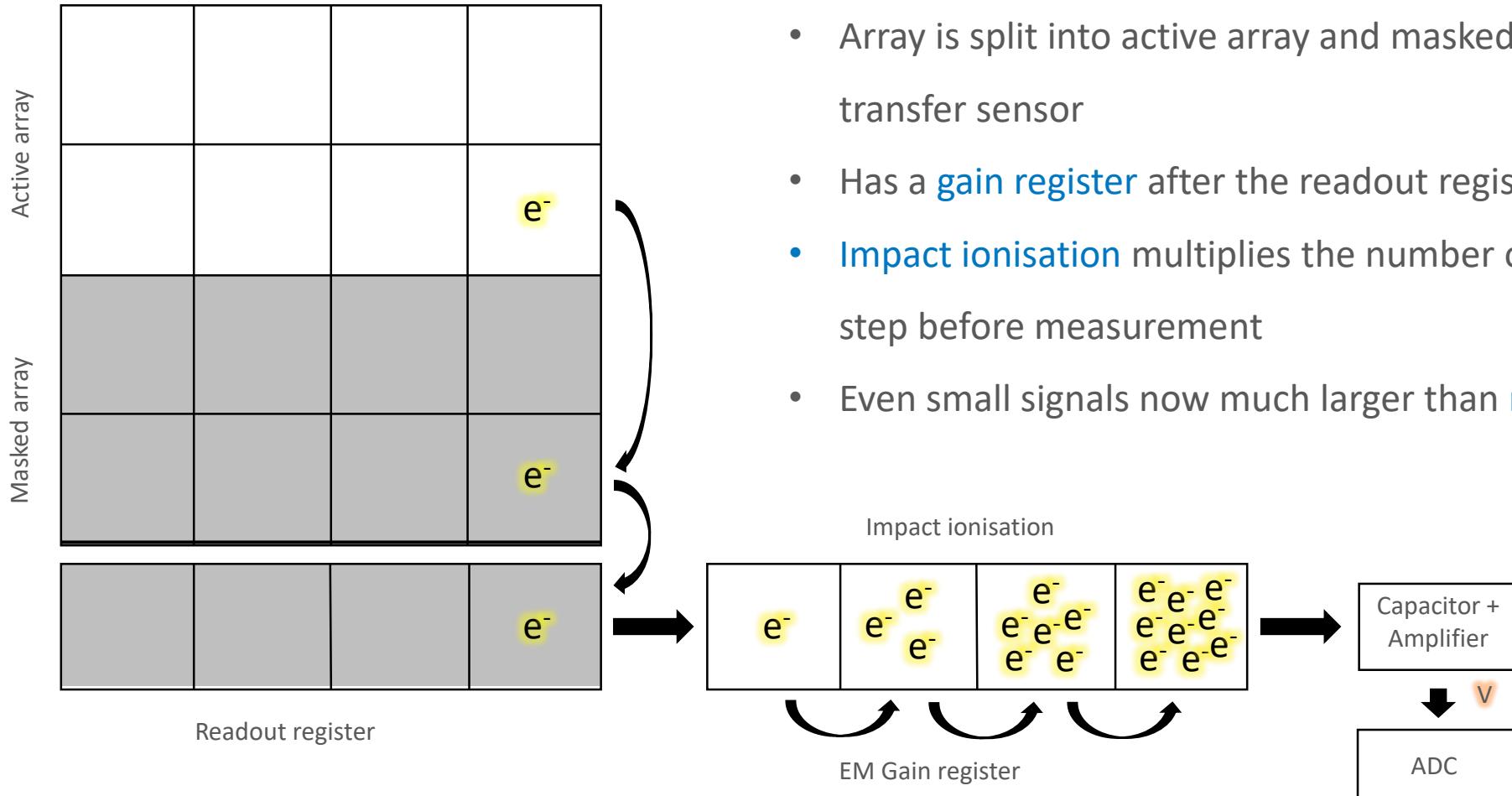
Photometrics introduced the first scientific grade camera (Cascade 650) in 2000

EMCCD sensors are also **back-illuminated** (high QE) and have **large pixels** (up to 24x24  $\mu\text{m}$ )

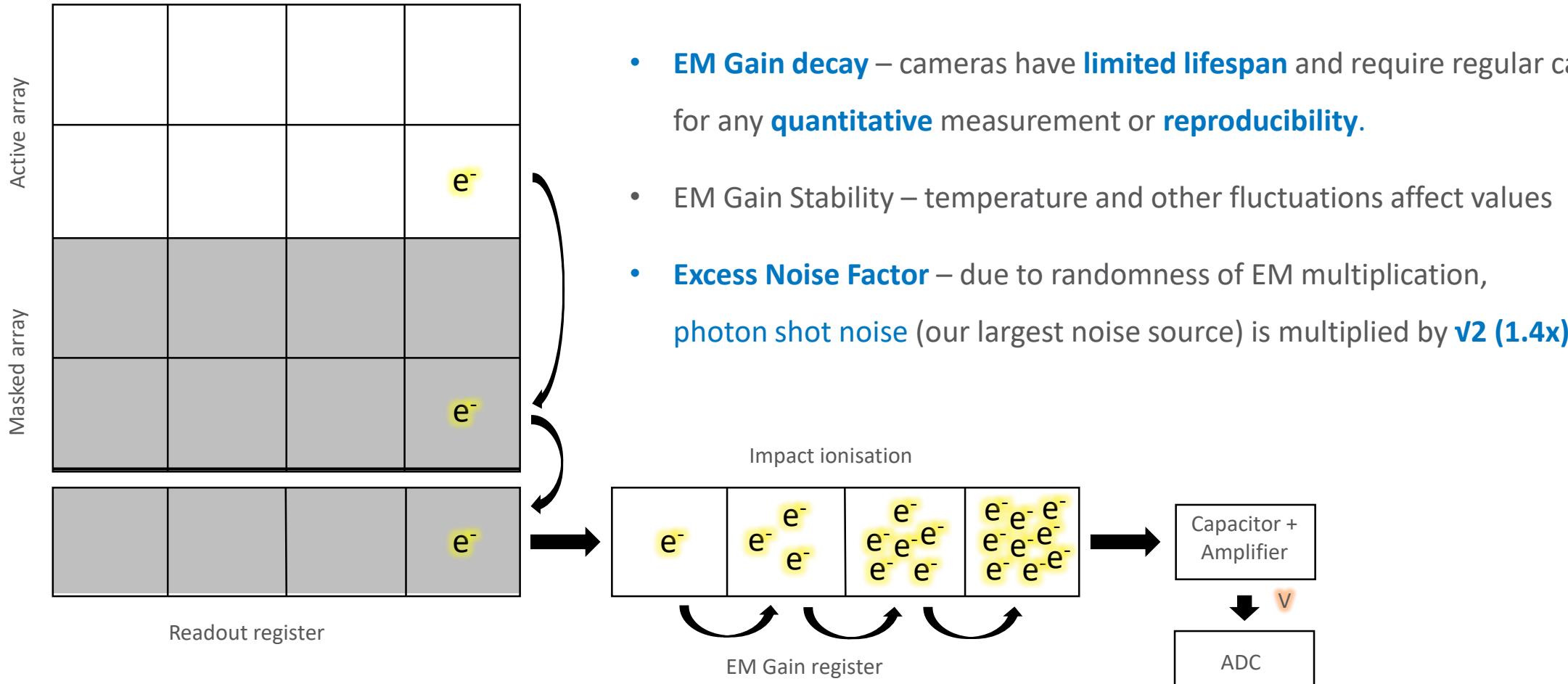
**Electron Multiplication** of detected photoelectrons to overcome read noise, increasing speed and sensitivity above CCDs



# EMCCD Fundamentals



# Issues With EMCCD Cameras



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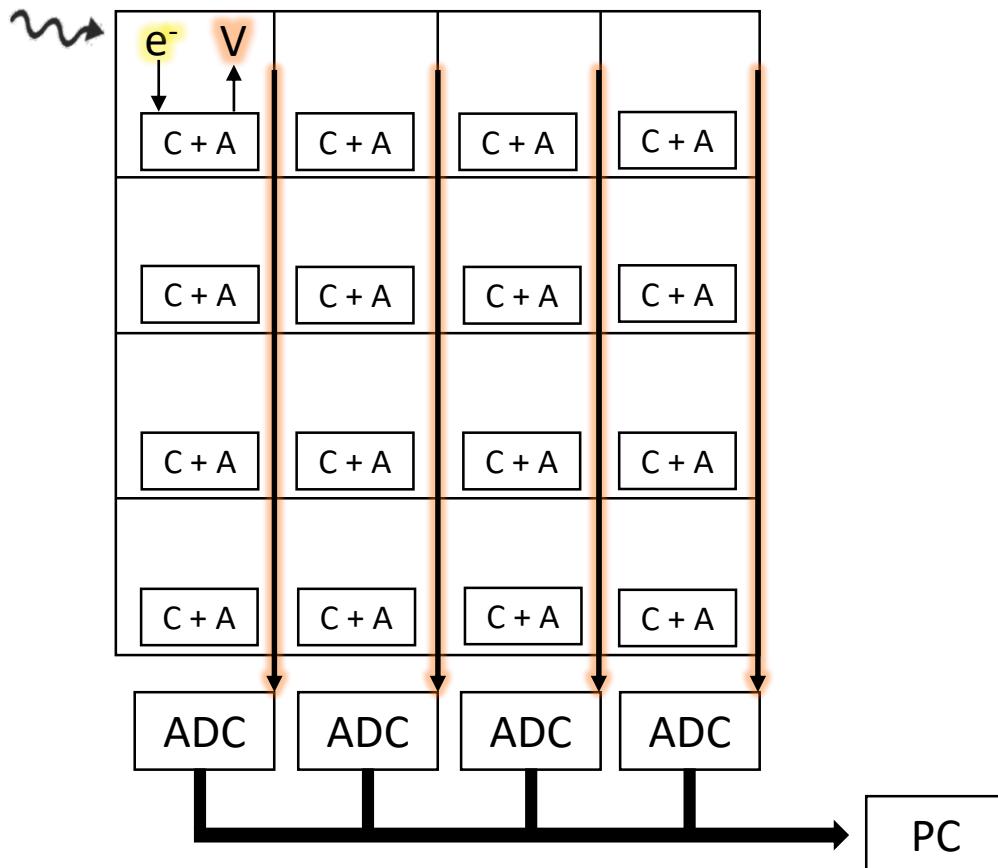
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# CMOS Fundamentals

Photon



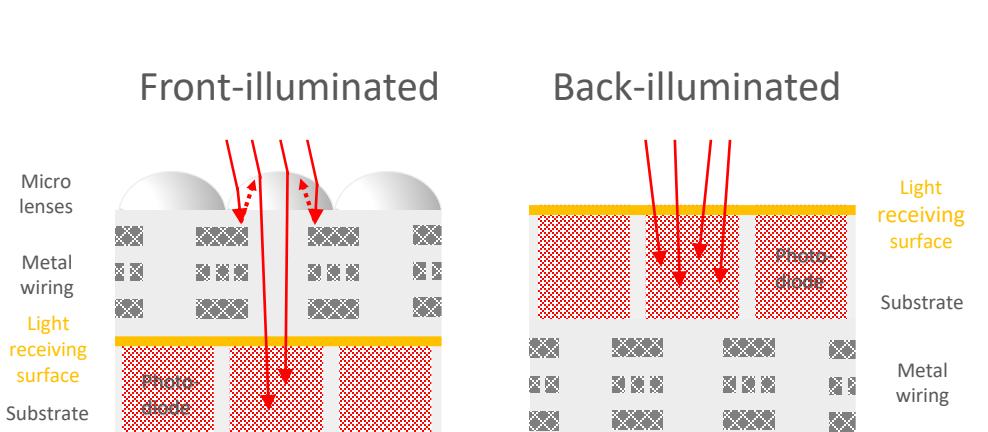
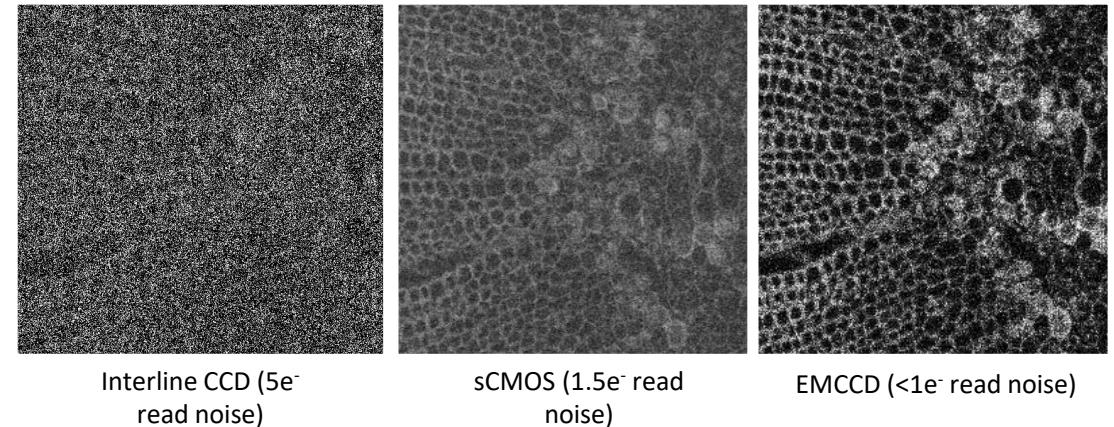
- Capacitor and amplifier on **every pixel**
- An ADC for **every column**
- sCMOS cameras deliver **bigger fields of view** at much **faster speeds**
- Low read noise to detect weak fluorescence and work with live cells
- However early sCMOS sensitivity & image quality **couldn't** rival EMCCD.

# Competing With EMCCD Sensitivity

sCMOS brought many advantages over EMCCD, not least being significantly cheaper!

However, sCMOS couldn't perform the low-light imaging tasks of the more sensitive EMCCD, due to:

- **Front-illumination**, Lower QE
- **Higher noise**
- Problems with **patterns & artefacts**
- Much **smaller pixels**



# Timeline of Scientific Imaging



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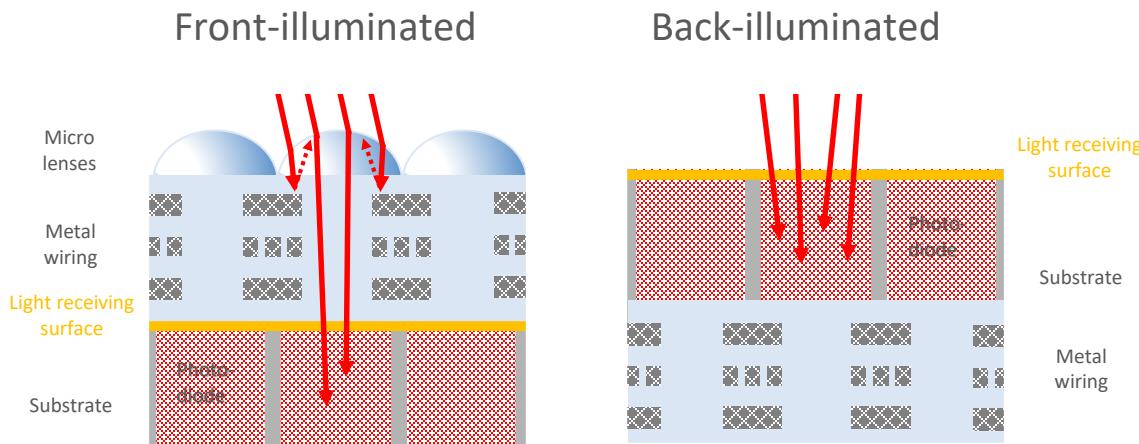
- Highest sensitivity, rivals or beats EMCCD, with speed and field of view of CMOS
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# Back-Illuminated CMOS

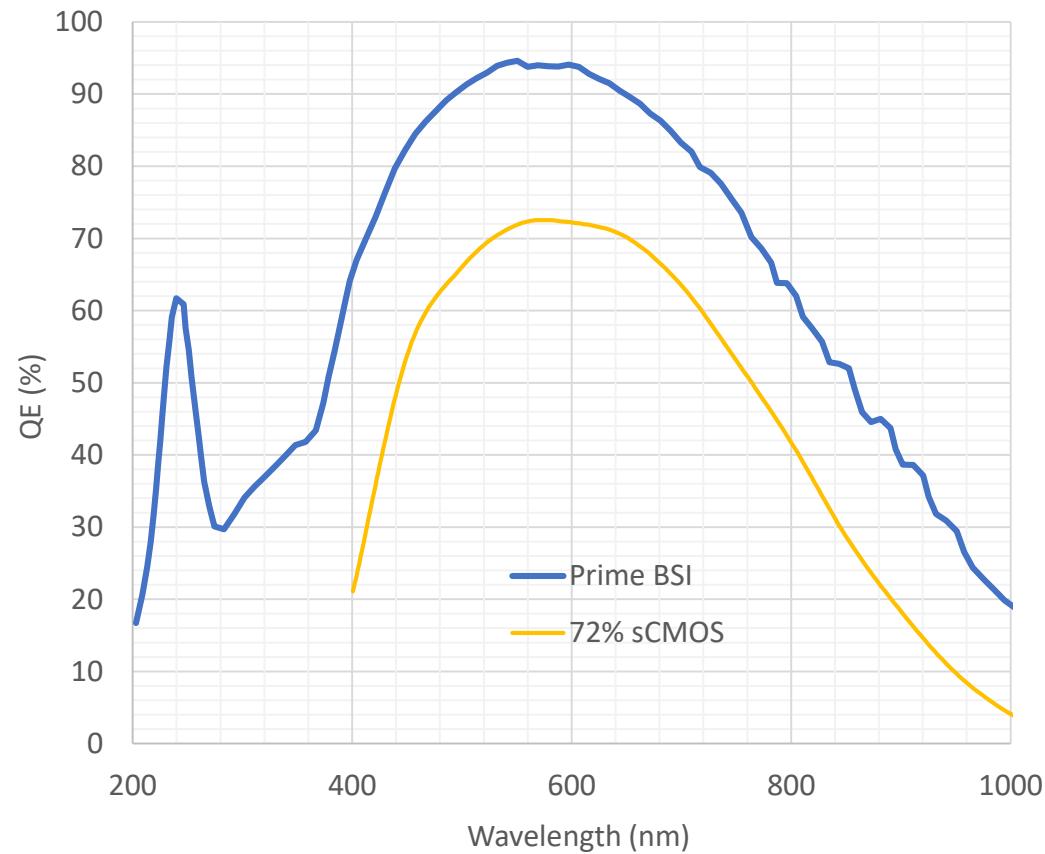
Avoid the reflecting layers completely by bringing the light in from the back of the sensor. Requires thinning of silicon to around 1.1um.



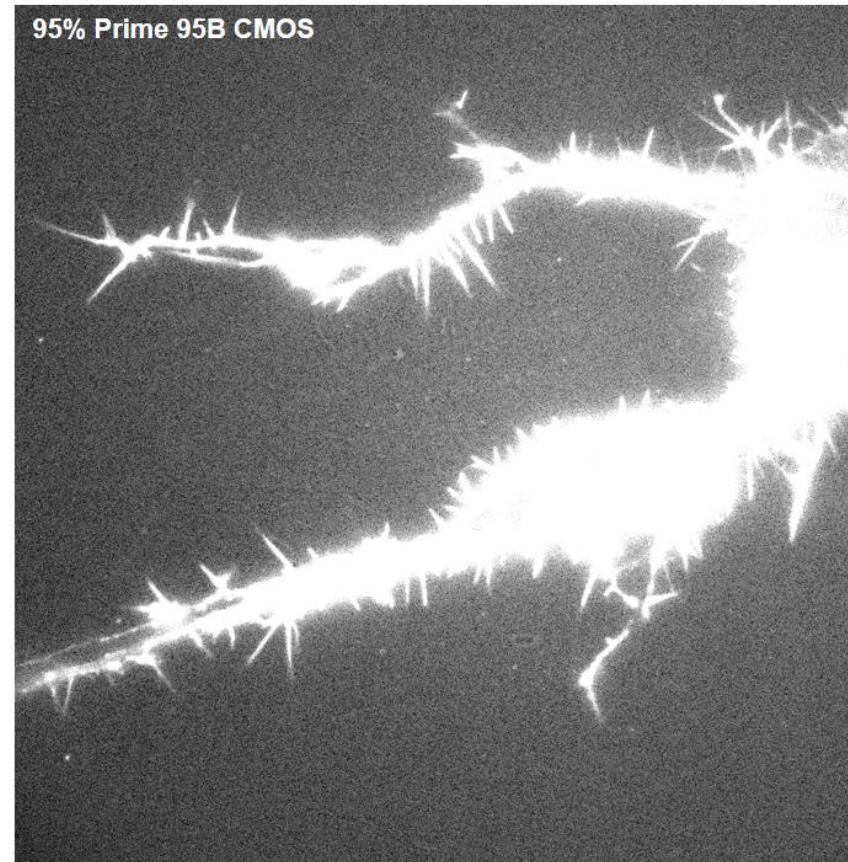
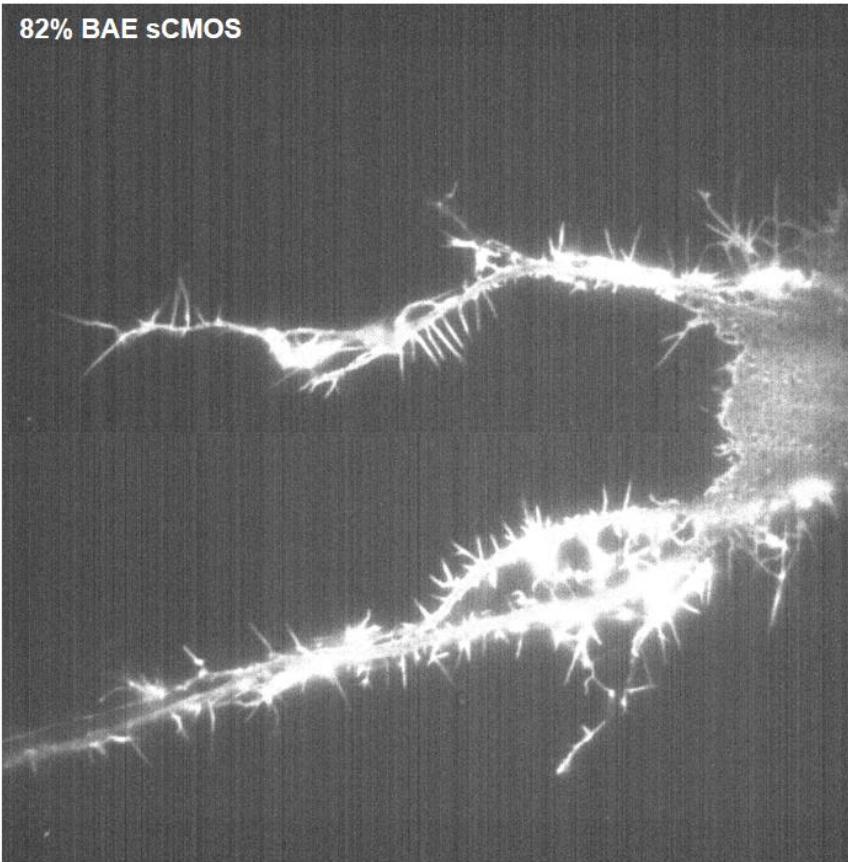
Achieve **95% peak QE**

Maintain 1.0 - 1.6e<sup>-</sup> Read Noise

**Outperforms EMCCD in sensitivity, speed, FOV and more**

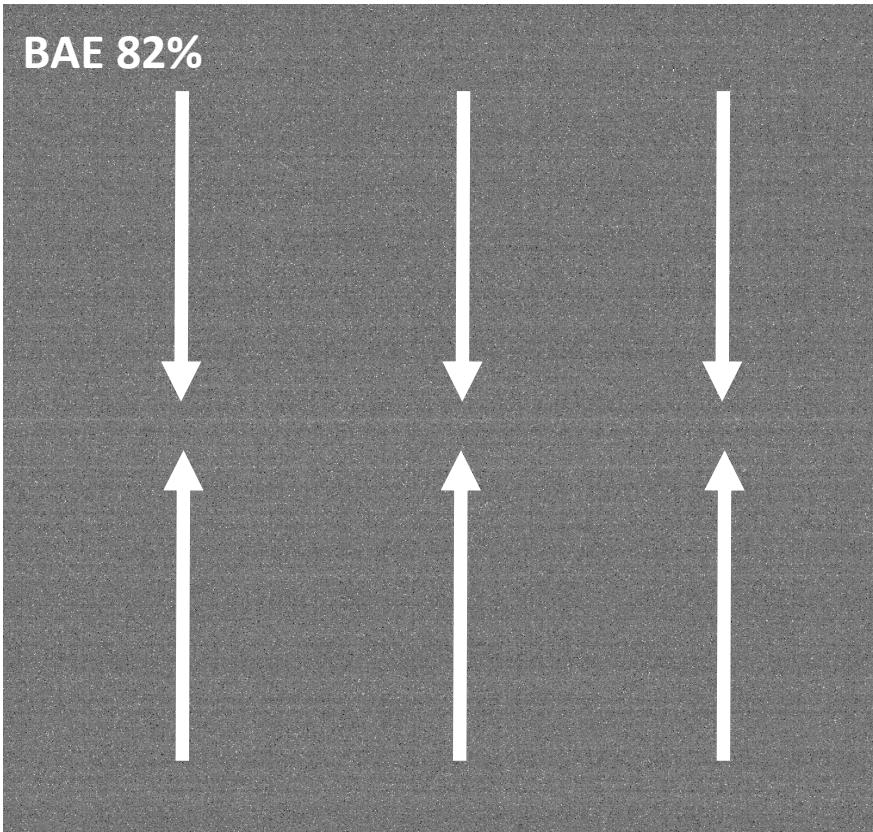


# Fixed pattern column noise



Average of 100 frames, 30 ms exposure time

# Excellent Background Quality



**Prime 95B and BSI**

Single Sensor Design

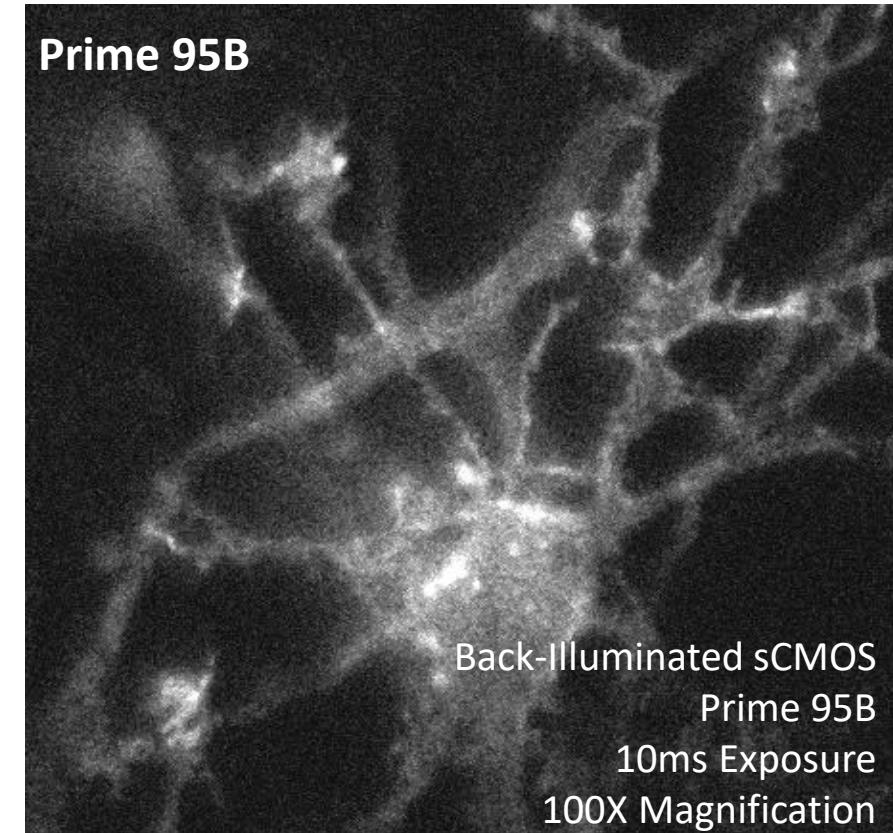
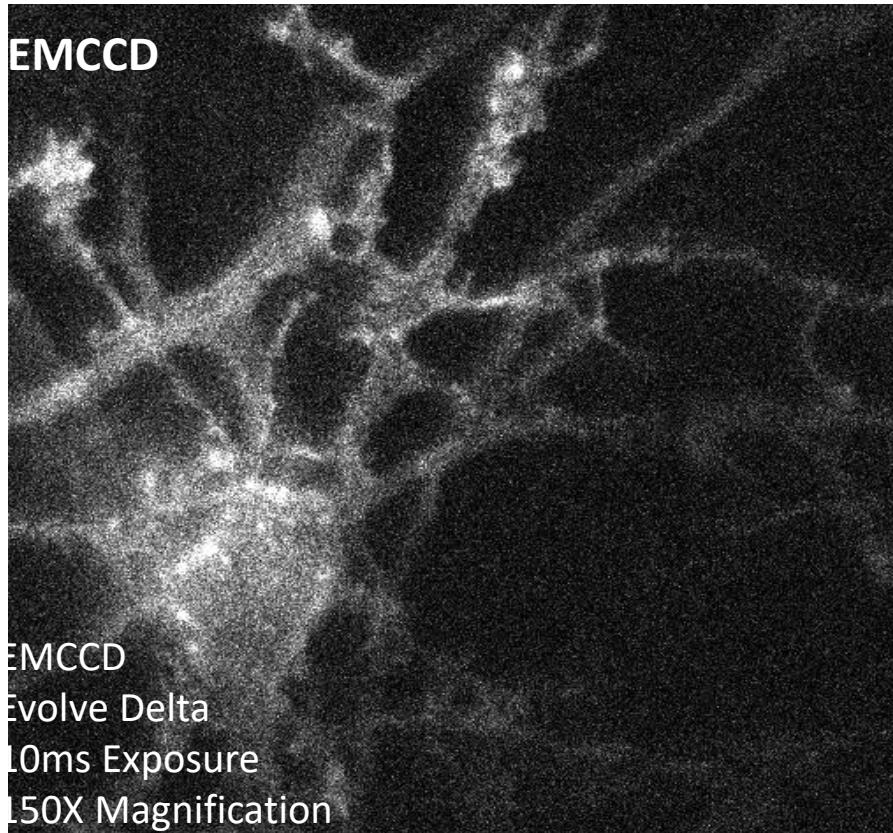
No Split Readout (Centre Line)

No Horizontal or Vertical Pattern

No Moving Pattern

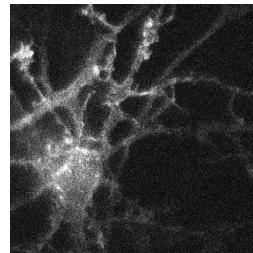
Background quality with the Prime 95B and Prime BSI is excellent, offering the very best **low light imaging performance**

# EMCCD vs Back sCMOS: Sensitivity

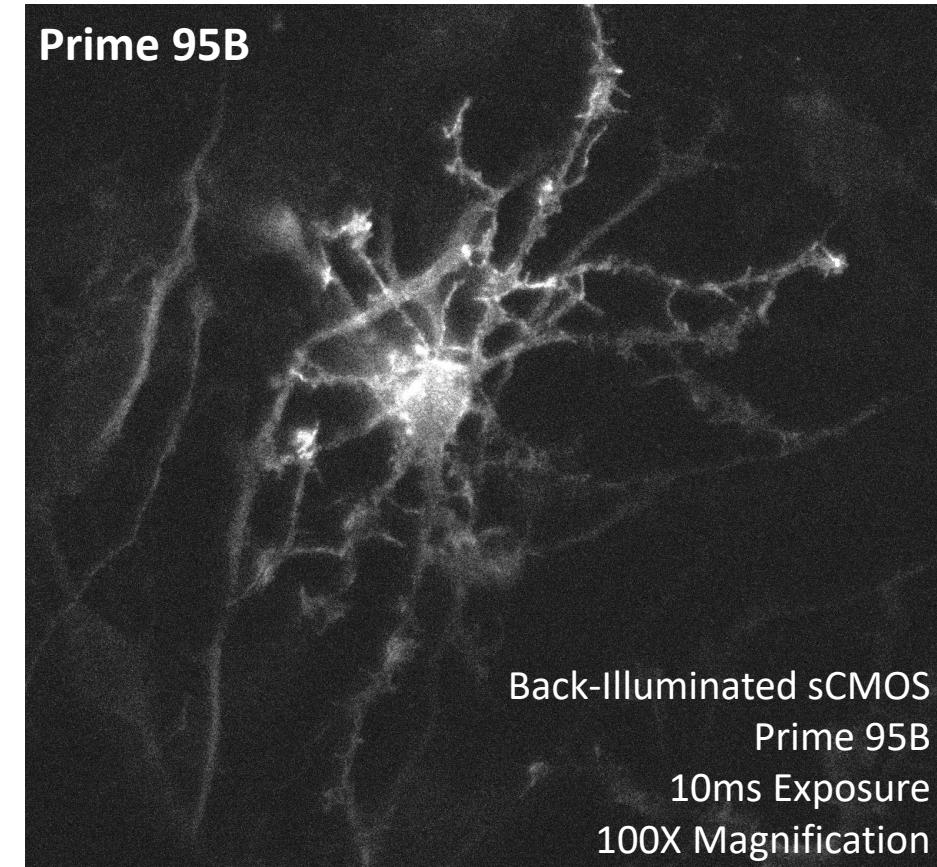


Better SNR than the EMCCD due to absence of **Excess Noise Factor**

# EMCCD vs Back sCMOS: Sensitivity



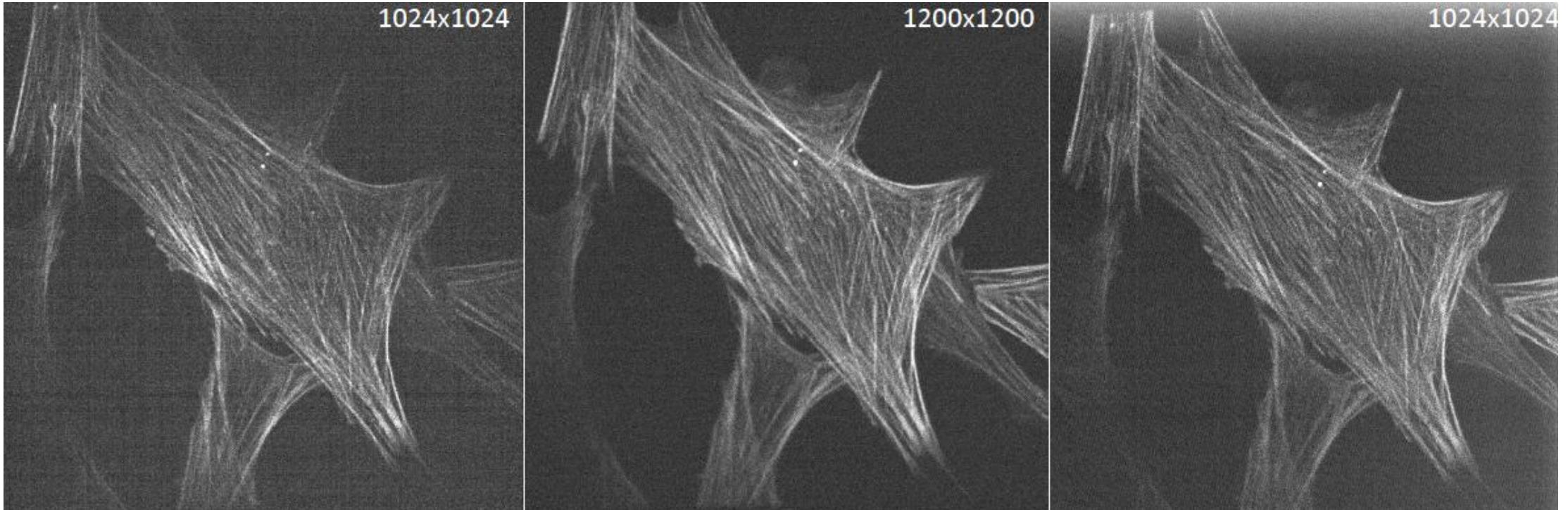
EMCCD  
Evolve Delta  
10ms Exposure  
150X Magnification



Better SNR than EMCCD, with faster frame rate and drastic FOV improvement

# 2x2 Binned CMOS vs 95B vs 1k EMCCD

Fixed sample, Spinning disk confocal, same exposure time & focal plane



sCMOS 2x2Bin - 13 micron  
Pixel Area  $169\mu\text{m}^2$

**Prime 95B - 11 micron**  
**Pixel Area  $121 \mu\text{m}^2$**

1024 EMCCD- 13 micron  
Pixel Area  $169\mu\text{m}^2$

# Being Quantitative

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**Photoelectrons** are our objective, quantitative unit. To calculate back, use:

$$\text{Photoelectrons} = (\text{Grey levels} - \text{Offset}) \times \text{Gain}$$

**Gain:** Ask your camera manufacturer for the gain values for **your specific camera**, or estimate with a **Mean Variance Test**.

**Offset:** Baseline grey level value, **mean image value with no light** (typically 100).

This allows us to:

1. Compare intensity values between **different cameras / camera settings**
2. Estimate **signal to noise ratio**
3. Understand your **light level**

# How is Gain set?

Gain represents the number of **detected photoelectrons** that each **grey level (ADU)** represents. Choosing the value is a **balance** of **precision** and **dynamic range**:

**Dynamic Range:**

**Bit depth x Gain = Full Well Capacity (in e-)**

**Precision & Contrast:**

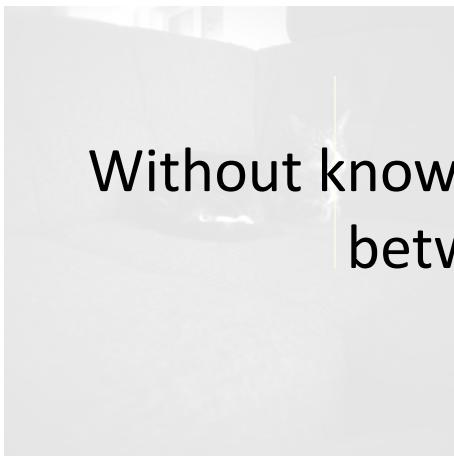
**Gain (in e-/grey) = smallest measurement step**

'Higher' gain means a **smaller number** of **e-/grey** (e.g. 0.25e-/grey), resulting in each detected photoelectron providing more grey levels and **better contrast for low light imaging**.

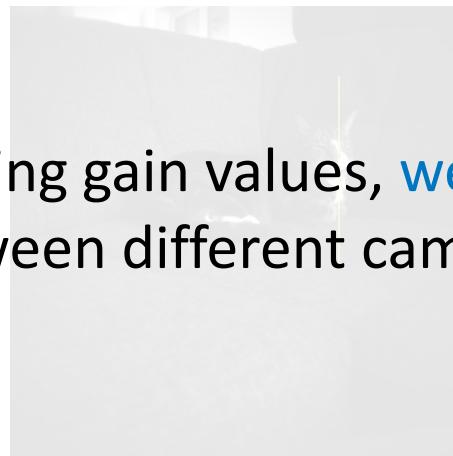
At high light levels, low gain (e.g. 2e-/grey) is necessary to detect strong signals.

# Gain States With Cimba

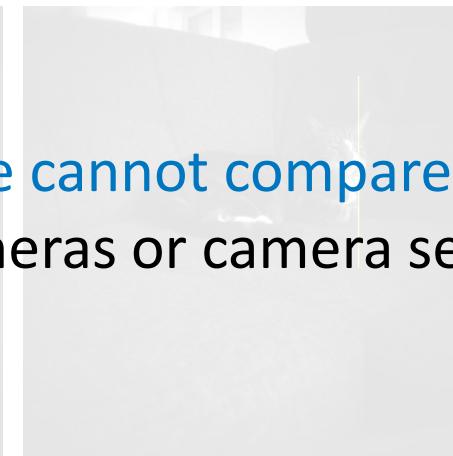
**12-bit sensitivity**  
= 0.6 e<sup>-</sup>/grey level



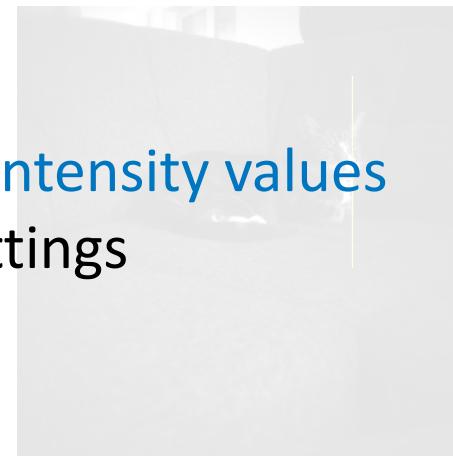
**12-bit balanced**  
= 1.2 e<sup>-</sup>/grey level



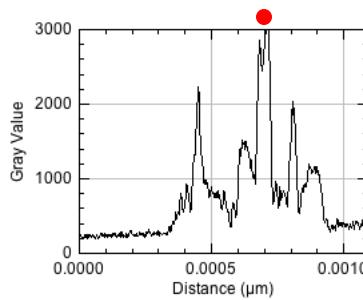
**12-bit full well**  
= 2.4 e<sup>-</sup>/grey level



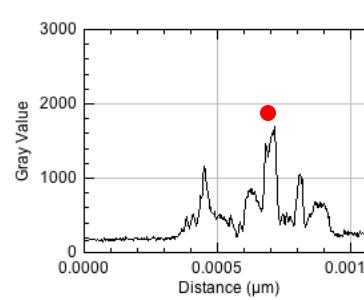
**16-bit HDR**  
= 0.9 e<sup>-</sup>/grey level



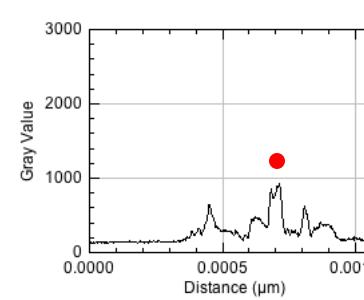
Without knowing gain values, we cannot compare intensity values between different cameras or camera settings



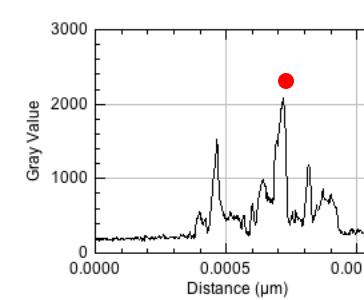
Max grey 2900 = 1700 e<sup>-</sup>



Max grey 1500 = 1700 e<sup>-</sup>



Max grey 800 = 1700 e<sup>-</sup>



Max grey 2100 = 1700 e<sup>-</sup>



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# Thank you for attending! Summary:

**Sensitivity:** **Quantum Efficiency** most important factor,  
**Read Noise** also important at low light levels.

**Pixel Size:** **Larger pixels** are more sensitive  
**Smaller pixels** can offer better resolution unless limited by microscope

**Being Quantitative:**  
**Calculating back to electrons** gives true comparison  
then can **Calculate Signal to Noise Ratio**

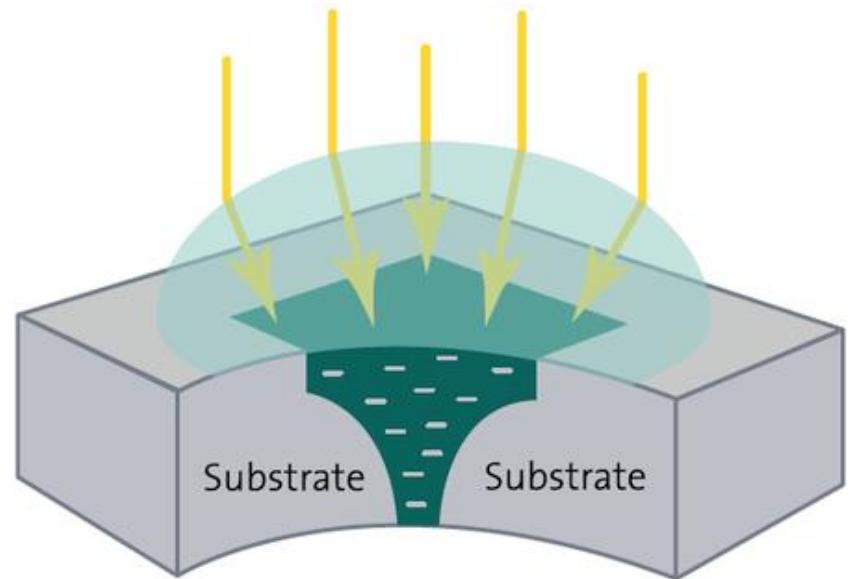
Camera Technologies:  
**Front-Illuminated CMOS:** Fast but not as sensitive as EMCCD  
**EMCCD:** Sensitive but slow & small FOV  
**CCD:** Good for documentation, but not for live cell  
**...Back-Illuminated CMOS:** Fastest, biggest FOV, most sensitive

# Well Depth / Full Well Capacity

How many photoelectrons of signal can a pixel store during an exposure?

After photons are converted to photoelectrons in a pixel, they are stored in the **well** of a pixel.

Too many photoelectrons in a well and it will **saturate** – no additional light can be collected in that pixel.



# Full-well Capacity and Dynamic Range

Pixel Size

4.25 $\mu\text{m}$



Full-Well  
Capacity

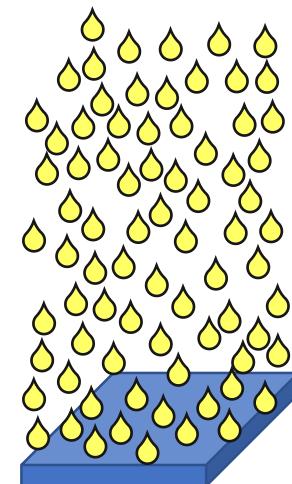
12,000e-

6.5 $\mu\text{m}$



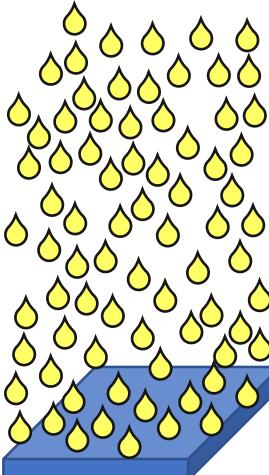
45,000e-  
(~3.8X more e-)

11 $\mu\text{m}$



80,000e-  
(~6.3X more e-)

# Full-well Capacity and Dynamic Range

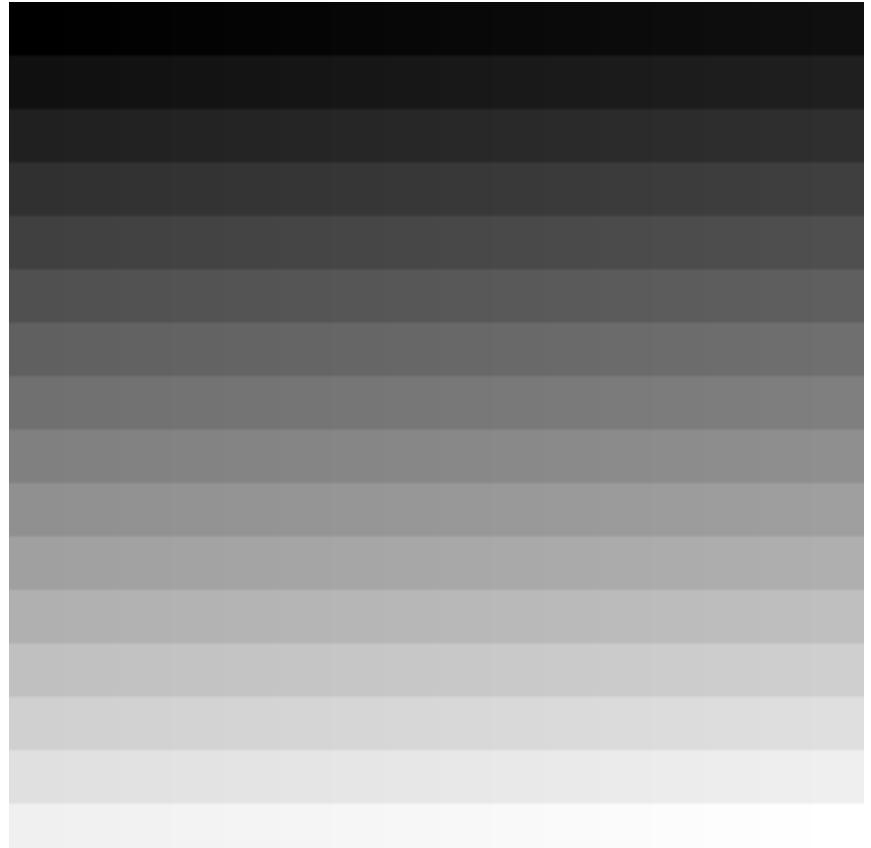
Pixel Size	4.25μm	6.5μm	11μm
Full-Well Capacity	 <b>12,000e-</b>	 <b>45,000e-</b> (~3.8X more e-)	 <b>80,000e-</b> (~6.3X more e-)
Dynamic Range:	8000 : 1	30,000 : 1	53,000 : 1
Full-well Capacity / Read Noise (1.5e-)			

# Grey Levels

The voltage from photoelectrons is converted into an **arbitrary** intensity unit using an **analogue to digital converter (ADC)**

As our cameras are monochrome, these colors are **gray levels**

The maximum number of displayable grey levels depends on the **bit depth**



# Bit Depth



8 bit (**256** grey levels)



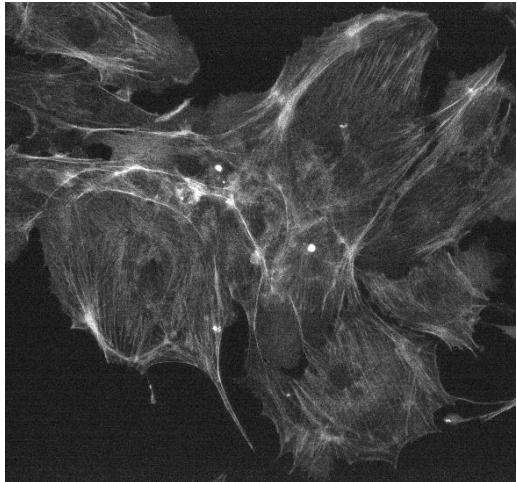
5 bit (**32** grey levels)



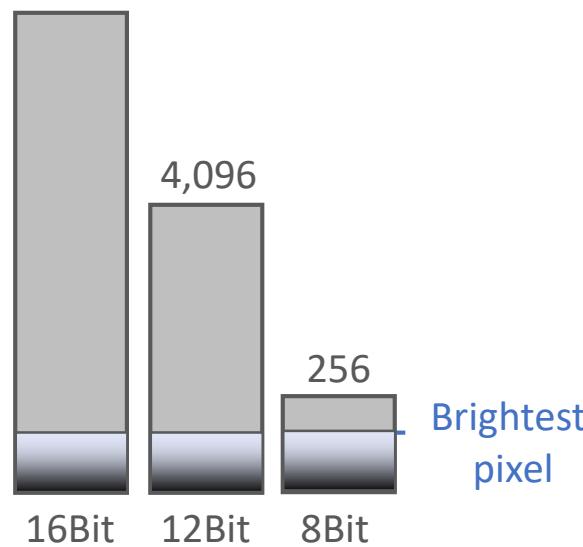
2 bit (**4** grey levels)

# Bit Depth

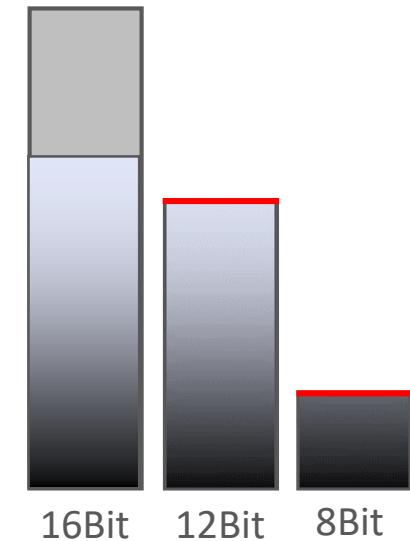
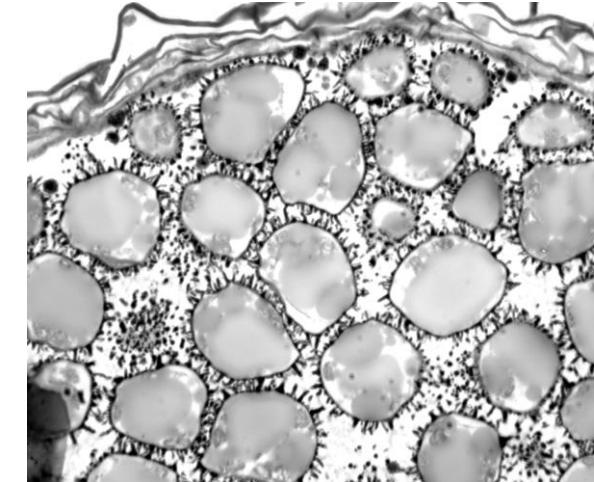
Fluorescence Image



65,536



Brightfield Image



- High bit depth allows precise capture of high contrast samples, such as bright fixed cell fluorescence and brightfield.
- Bit depth is typically not relevant for low light imaging