LAB-2-CODE

November 16, 2020

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0.1.1 Lab 2 Code

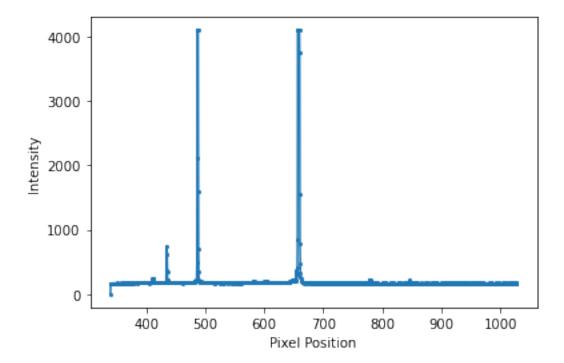
0.1.2 Functions

```
[63]: import matplotlib.pyplot as plt
      import numpy as np
      import astropy.io.fits as fits
      def plotspectrum(name):
          filename = name
          data = np.genfromtxt(filename, skip_header = 17, skip_footer=1)
          pos = data[:,0]
          intensity = data[:,1]
          %matplotlib inline
          plt.figure(1)
          plt.plot(pos, intensity, 'o-', markersize = 2)
          plt.xlabel("Pixel Position")
          plt.ylabel("Intensity")
          plt.show()
      def centroid_finder(x,y,bias): #Take the x and y values of the spectrum.
          s = \prod
          I = \Gamma
          cmean = []
          cstddev = []
          i = 0
          while i < len(x): #Goes through all of the pixels
              if y[i] > bias: #Skip over any pixel lower than the noise
                  check = True
              else:
                  check = False
              if check:
                  s.append(x[i])
                  I.append(y[i])
                  if y[i-1] >= y[i] and y[i] < y[i+1] or y[i+1] <= bias:
                      #If new emission line is found, find standard deviation and
       \rightarrowmean and append it
```

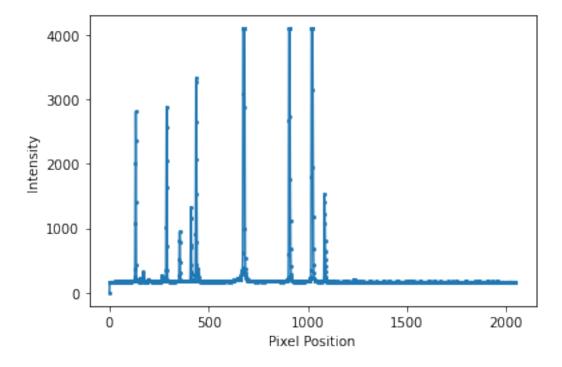
```
m = np.sum((np.multiply(s,I)) / np.sum(I))
                      cmean.append(m)
                      #standard deviations
                      std = np.sum(np.multiply(I,(s-m)**2)) / np.sum(I)
                      cstddev.append(std)
                      #Clear the arrays
                      s = \prod
                      I = []
              i += 1 #if there are more pixels, check again.
          return [cmean,cstddev] #return the arrays of centroids and standard_
       → deviations as a new array
      def linfit(x,y):
          a = np.array([[np.sum(np.power(x,2)),np.sum(x)],[np.sum(x),x.size]])
          b = np.array([[np.sum(np.multiply(x,y))],[np.sum(y)]])
          c = np.matmul(np.linalg.inv(a),b)
          return c
      def linfit error(x,y,m,c):
          N = x.size
          sig2 = 1/(N-2)*(np.sum(np.power(y-(m*x+c),2)))
          sign2=N*sig2/(N*np.sum(np.power(x,2))-np.power(np.sum(x),2))
          sigc2 = sig2*np.sum(np.power(x,2))/(N*np.sum(np.power(x,2))-np.power(np.
       \rightarrowsum(x),2))
          return [np.sqrt(sign2),np.sqrt(sigc2)]
      def reduce(science,flat,bias):
          bias = np.array(bias)/1.00001
          flat = np.array(flat)
          science = np.array(science)
          flat_norm = (flat-bias/np.median(flat-bias))
          science_reduced = science-bias/flat_norm
          return science_reduced
[64]: filenames = ["Hydrogen01870.txt", "groupc-Helium02190.txt", "

¬"groupc-mercury01590.txt",
                   "groupc-neon01250.txt", "groupc-lamp01690.txt", ]
      i = 0
      while i < len(filenames):</pre>
          %matplotlib inline
          print("Data:", filenames[i])
          plotspectrum(filenames[i])
          i+=1
```

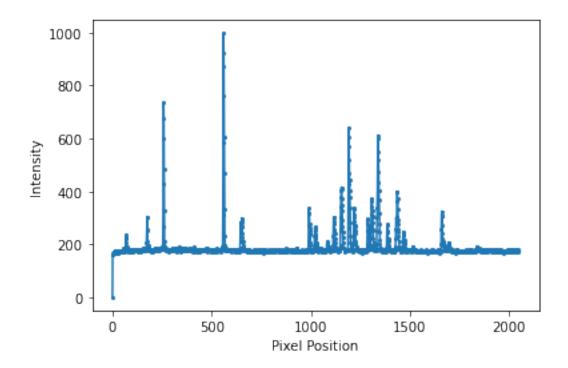
Data: Hydrogen01870.txt



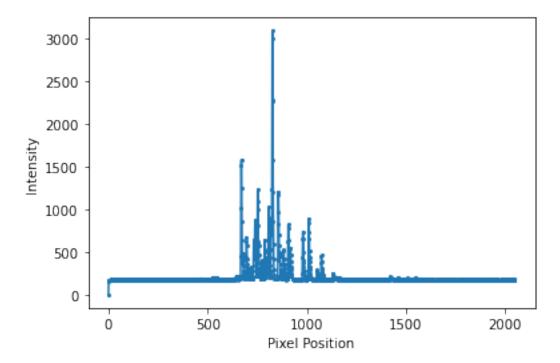
Data: groupc-Helium02190.txt



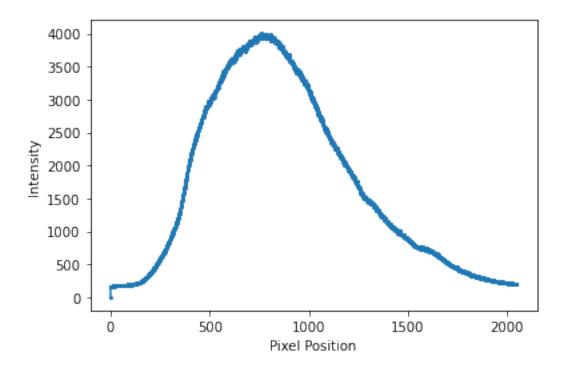
Data: groupc-mercury01590.txt



Data: groupc-neon01250.txt



Data: groupc-lamp01690.txt



0.1.3 Finding Centroids

```
[65]: i = 0
    centroiddata = []
    centroidstds = []
    while i < len(filenames):
        data = np.genfromtxt(filenames[i], skip_header = 17, skip_footer = 1)
        pos = data[:,0]
        intensity = data[:,1]
        c = centroid_finder(pos, intensity, 220)
        print("file :", filenames[i])
        print(c)
        centroiddata.append(c[0])
        centroidstds.append(c[1])
        i += 1</pre>
```

```
file: Hydrogen01870.txt
[[411.1832692307692, 435.0827676322418, 487.36251279960135, 647.99, 654.266424352567, 658.3329715310178, 778.7347128129602], [0.09195277366863903, 0.36206082257952765, 0.4666769172277576, 0.0, 0.4305876236504802, 1.6206449867478276, 0.06923257584465954]]
file: groupc-Helium02190.txt
[[132.02232393050625, 152.495670995671, 168.9281609195402, 265.96989528795814, 288.3707544723879, 353.9975686846584, 411.2439817826936, 436.98832182721463, 445.4630620985011, 646.4967320261437, 665.1218103033221, 677.7912118217544,
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0.0, 10.166391536953068, 7.282890222468051]]
file: groupc-mercury01590.txt
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1.227035490605428, 3.656121498746873]]
file: groupc-neon01250.txt
[[645.0, 670.6976389052038, 679.8787765759046, 696.9963768115942,
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file : groupc-lamp01690.txt
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0.1.4 Wavelength Calibration

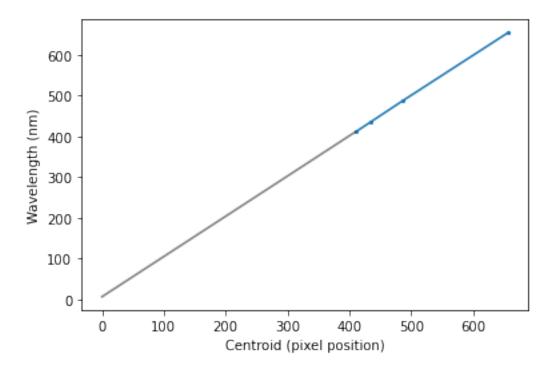
(Wavelength vs Pixel Centroid Plots)

```
[66]: # Wavelengths taken from the webpage recommended by professor
      # Direct Link: https://physics.nist.gov/PhysRefData/ASD/lines_form.html
      #Chosen Wavelengths
      H_{chosen_WL} = np.array([410.1734, 434.0472, 486.135, 656.279])
      HE_chosen_WL = np.array([257.76, 388.8648, 402.61914, 447.14802, 706.5190, 1091.
      →292])
      HG chosen WL = np.array([82.1329, 167.2537, 249.2064, 567.7105, 952.0198, 1128.
      <u>→</u>711)
      NE_chosen_WL = np.array([640.22480, 692.94672, 837.76070, 865.43828, 966.54200, ___
      \rightarrow 1117.75246
      #It is clear that the lamp has a continuous emission spectrum
      #so there is no point in matching wavelengths to centroids for this particular
       \rightarrow object
      #Chosen Centroids
      H_chosen_CDS = np.array([411.1832692307692, 435.0827676322418, 487.
       \rightarrow36251279960135, 654.266424352567])
      HE_chosen_CDS = np.array([265.96989528795814, 353.9975686846584, 411.
       \rightarrow2439817826936, 445.4630620985011, 677.7912118217544,
                        1085.5953213817227])
      HG chosen CDS = np.array([69.48913043478261, 174.50987432675043, 258.
       →10568181818184, 561.0188834154351, 992.7629805786762,
                       1218.9766777724892])
      NE_chosen_CDS = np.array([645.0, 696.9963768115942, 827.6161117961366, 865.
       →846017699115, 981.4306358381504, 1131.9761570827488])
      chosenWL = [H_chosen_WL, HE_chosen_WL, HG_chosen_WL, NE_chosen_WL]
      chosenCDS = [H_chosen_CDS, HE_chosen_CDS, HG_chosen_CDS, NE_chosen_CDS]
```

```
names = ["Hydrogen", "Helium", "Mercury", "Neon"]
i = 0
while i<=3:
   print(names[i], "Fit")
   fit = linfit(chosenWL[i],chosenCDS[i])
   fite = linfit_error(chosenWL[i],chosenCDS[i],fit[0],fit[1])
   print("Slope, Intercept:", np.transpose(fit))
   print("Error:", fite)
   plt.figure()
   %matplotlib inline
   plt.plot(chosenWL[i], chosenCDS[i], 'o-', markersize = 2)
   x = np.array(range(int(np.min(chosenWL[i])),))
   plt.plot(x,fit[0]*x+fit[1],'k',alpha = 0.5)
   plt.xlabel("Centroid (pixel position)")
   plt.ylabel("Wavelength (nm)")
   plt.show()
   i+=1
```

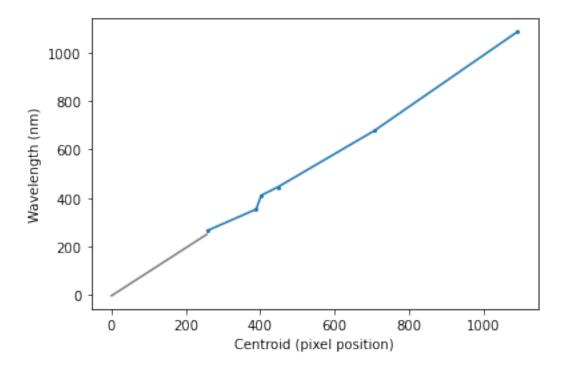
Hydrogen Fit

Slope, Intercept: [[0.98685179 6.84526712]] Error: [0.003405191250559764, 1.72262677954393]



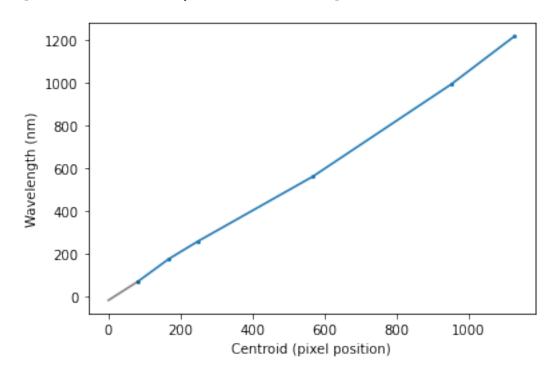
Helium Fit Slope, Intercept: [[0.98804948 -2.46241326]]

Error: [0.030031642999770614, 18.471238251233068]



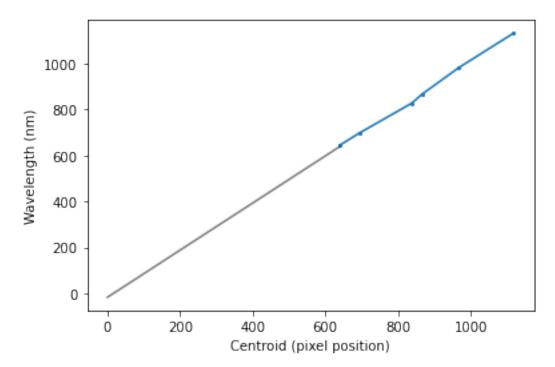
Mercury Fit

Slope, Intercept: [[1.07611609 -18.61832225]] Error: [0.022516000442013254, 14.820268606972371]



Neon Fit

Slope, Intercept: [[1.0255357 -17.09323737]] Error: [0.02323538484490006, 20.177055218310354]



0.1.5 KAST Spectrograph Data Analysis

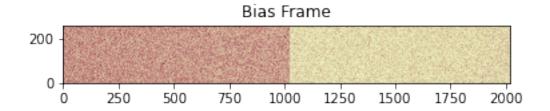
```
[67]: def reduce(science,flat,bias):
    bias = np.array(bias)/1.00001
    flat = np.array(flat)
    science = np.array(science)
    flat_norm = (flat-bias/np.median(flat-bias))
    science_reduced = (science-bias)/flat_norm
    return science_reduced

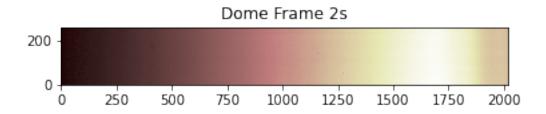
#Will be using these fits files
#b151 Zw 229-015 Seyfert 1 Galaxy
#b158 3C079 Seyfert 2 Galaxy

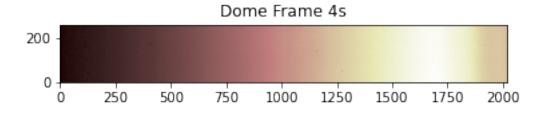
Galaxy1_name = "data-2013-10-26-shane-public/b151.fits"
Galaxy2_name = "data-2013-10-26-shane-public/b158.fits"
Galaxy1 = fits.getdata(Galaxy1_name)
Galaxy2 = fits.getdata(Galaxy2_name)
```

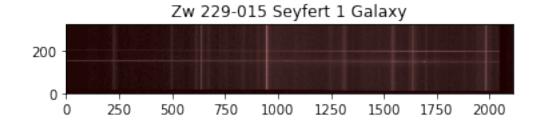
```
bias name = "data-2013-10-26-shane-public/b101.fits"
bias = fits.getdata(bias_name)
domeflat4_name = "data-2013-10-26-shane-public/b112.fits"
domeflat2_name = "data-2013-10-26-shane-public/b121.fits"
dome4s = fits.getdata(domeflat4_name)
dome2s = fits.getdata(domeflat2_name)
w = 2020
h = 20
rGalaxy1 = reduce(Galaxy1[h:279,0:w],dome4s[h:279,0:w],bias[h:279,0:w])
rGalaxy2 = reduce(Galaxy2[h:279,0:w],dome4s[h:279,0:w],bias[h:279,0:w])
%matplotlib inline
plt.figure()
plt.title("Bias Frame")
plt.imshow(bias[h:279,0:w],origin = 'lower', interpolation = 'nearest', cmap = ___
plt.figure()
plt.title("Dome Frame 2s")
plt.imshow(dome2s[h:279,0:w],origin = 'lower', interpolation = 'nearest', cmapu
→= 'pink')
plt.figure()
plt.title("Dome Frame 4s")
plt.imshow(dome4s[h:279,0:w],origin = 'lower', interpolation = 'nearest', cmapu
→= 'pink')
plt.figure()
plt.title("Zw 229-015 Seyfert 1 Galaxy")
plt.imshow(Galaxy1,origin = 'lower', interpolation = 'nearest', cmap = 'pink')
plt.figure()
plt.title("Zw 229-015 Seyfert 1 Galaxy REDUCED")
plt.imshow(rGalaxy1,origin = 'lower', interpolation = 'nearest', cmap = 'pink')
plt.figure()
plt.title("3C079 Seyfert 2 Galaxy")
plt.imshow(Galaxy2,origin = 'lower', interpolation = 'nearest', cmap = 'pink')
plt.figure()
plt.title("3C079 Seyfert 2 Galaxy REDUCED")
plt.imshow(rGalaxy2,origin = 'lower', interpolation = 'nearest', cmap = 'pink')
```

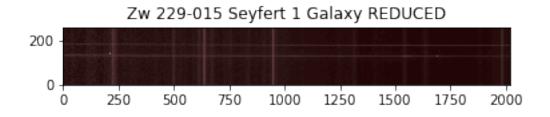
[67]: <matplotlib.image.AxesImage at 0x7f0349bbcb50>

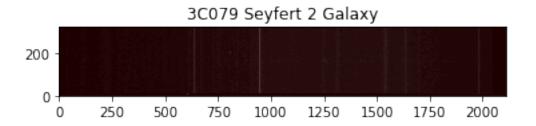




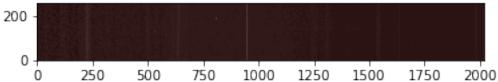






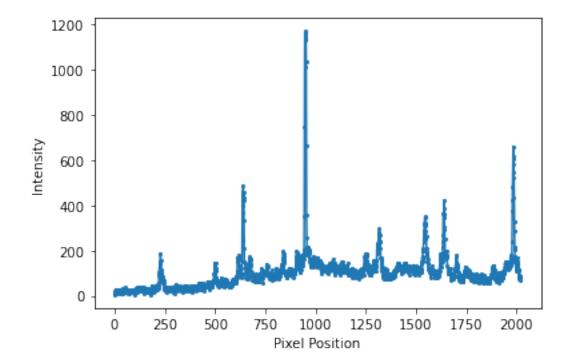


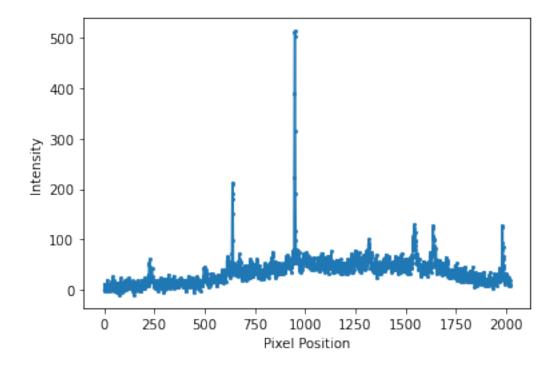




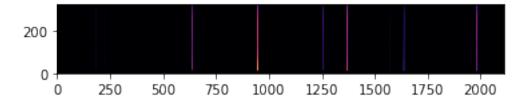
```
[68]: def linearize(science, max1, max2):
          i = 0
          size = science[0].size
          spectrum = []
          m = (max2-max1)/-size
          while i<size:
              spectrum.append(science[int(max1-m*i),i])
              i+=1
          return np.array(spectrum)
      %matplotlib inline
      s1 = linearize(Galaxy1[h:279,0:w]-(bias[h:279,0:w]/1.000001),177,166)
      plt.figure()
      x = np.array(range(0,2020))
      plt.plot(x, s1, 'o-', markersize = 2)
      plt.xlabel("Pixel Position")
      plt.ylabel("Intensity")
      plt.show()
      s2 = linearize(Galaxy2[h:239,0:w]-(bias[h:239,0:w]/1.000001),177,166)
      plt.figure()
      x = np.array(range(0,2020))
      plt.plot(x, s2, 'o-', markersize = 2)
      plt.xlabel("Pixel Position")
```

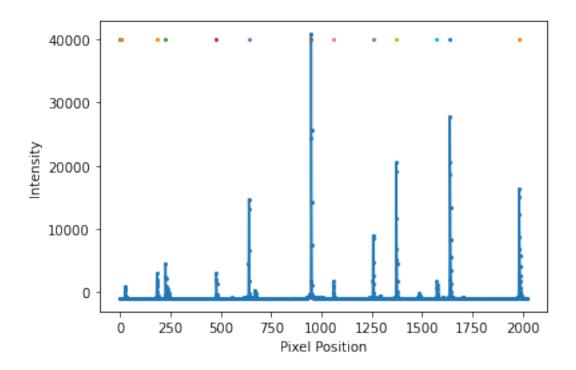
```
plt.ylabel("Intensity")
plt.show()
```





```
[69]: #Need to make this based on wavelength rather than pixe position, so we analyze
      \rightarrow the arc frame.
      arc_name = "data-2013-10-26-shane-public/b100.fits"
      arc = fits.getdata(arc_name)-(bias/1.0000001)
      %matplotlib inline
      plt.figure()
      plt.imshow(arc, origin = 'lower', interpolation = 'nearest', cmap = 'magma')
      larc = linearize(arc[h:279,0:w] - (bias[h:279,0:w]/1.000001),177,166)
      plt.figure()
      x = np.array(range(0,2020))
      arc_centroids = centroid_finder(x, larc, 1000)
      plt.plot(x, larc[0:2020], 'o-', markersize = 2)
      plt.plot(arc_centroids, np.full(len(arc_centroids),40000), 'o', markersize = 2)
      plt.xlabel("Pixel Position")
      plt.ylabel("Intensity")
      plt.show()
      print(arc_centroids)
```

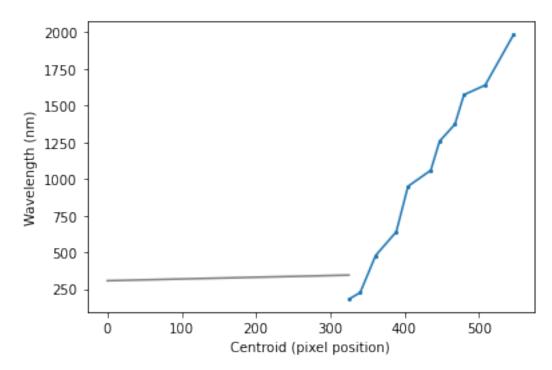




[[183.47805332091258, 225.96198348537462, 477.91213193039266, 639.6690633819476, 949.4593420876913, 1058.6120116976404, 1257.0549606202846, 1372.2137341691166, 1573.3783178601757, 1638.9538631929363, 1982.7861040086375], [0.9913299290853752, 0.4956346976331415, 0.524769672227625, 1.023459406504145, 1.5025611901714209, 0.23745337959171756, 1.5357297773385163, 2.0759824927928427, 0.23519345684777784, 2.998893167204087, 3.90562882050507]]

```
plt.xlabel("Centroid (pixel position)")
plt.ylabel("Wavelength (nm)")
plt.show()
```

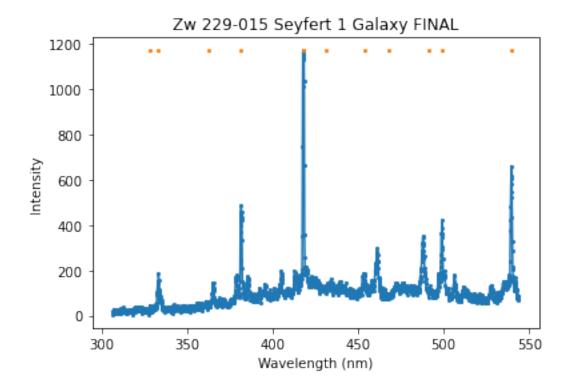
Slope/Intercept: [[1.17516819e-01 3.06513645e+02]] Error: [0.004436790175269142, 5.23273178673198]

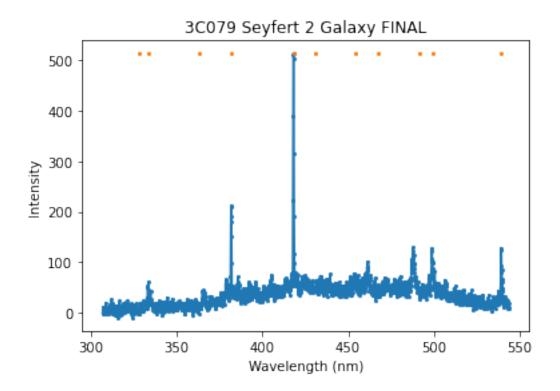


```
[71]: %matplotlib inline
      x = np.array(range(0,2020))
      plt.figure()
      plt.plot(x*cm + ci, s1, 'o-', markersize = 2)
      plt.plot(arc_centroids*cm + ci, np.full(len(arc_centroids),np.max(s1)), 'o',__
      →markersize = 2)
      plt.title("Zw 229-015 Seyfert 1 Galaxy FINAL")
      plt.xlabel("Wavelength (nm)")
      plt.ylabel("Intensity")
      plt.show
      plt.figure()
      plt.plot(x*cm + ci, s2, 'o-', markersize = 2)
      plt.plot(arc_centroids*cm + ci, np.full(len(arc_centroids),np.max(s2)), 'o',__
      →markersize = 2)
      plt.title("3C079 Seyfert 2 Galaxy FINAL")
      plt.xlabel("Wavelength (nm)")
```

```
plt.ylabel("Intensity")
plt.show
```

[71]: <function matplotlib.pyplot.show(*args, **kw)>





[]: