#flo #ret #ref #disorganized #incomplete #hw

## 1 | Les go.

- we need to figure out:
  - what is the curve of generation over a day
  - how does this curve shift over the seasons?
- · our inputs
  - location
  - time of year (season)
- output
  - generation curve over a day

look into: Global Tilted Irradience.

all we care about is the relative shape and how the relative shape changes! this is because the other stuff will be consistent, and we aren't recommending a solar system

## 1.0.1 | terms:

solar irradiance: power per unit area  $(W/m^2)$  integrated over time gives us: insolation  $(j/m^2)$  solar irradiance aka solar flux: power per unit area!

TSI: total solar irradiance. when the sun is perpendicular! over a sqaure meter. this is just a constant zenith angle: angle between suns rays and vertical direction (of earth). "local normal to earths surface" and sun rays (line between point on earth surface and sun)

declination angle: lattitude of point directly under the sun at **noon** complement of solar zenith angle subsolar point: point that is closest to the sun on a planet

hour angle *h*: defined as the longitude of the subsolar point relative to its position at noon. AKA how far it moves in an hour!

A cos zenith angle is the area of sunlight recieved per area on earth AKA how much sunlight area ur actully getting for an area on earth.

## 1.0.2 | helpful relations

spherical law of cosines!

Q day= $S0\pi d d^2[h0 \sin\phi \sin\delta + \cos\phi \cos\delta \sin h0]$ 

## 1.0.3 | Vars!

Assume circular orbit?

$$\text{charge } Q = S_0 \left(\frac{\overline{R_0}}{R_e}\right)^2 \cos \theta_s \text{ or } Q = \begin{cases} S_o \frac{R_o^2}{R_E^2} \cos(\Theta) & \cos(\Theta) > 0 \\ 0 & \cos(\Theta) \leq 0 \end{cases} \\ \text{can be aproximated as } Q \approx S_0 \cos \theta_s \\ \text{cos}(\Theta) \leq 0 \\$$

declination angle  $\delta = -0.409 \cdot \cos\left(\frac{2\pi}{365} \cdot (d+10)\right)$ 

spherical law of cosines  $\cos(c) = \cos(a)\cos(b) + \sin(a)\sin(b)\cos(C)$  and derivation C = h  $c = \Theta$   $a = \frac{1}{2}\pi - \phi$   $b = \frac{1}{2}\pi - \delta$ 

to calculation of  $\cos(\text{zenith})\cos(\Theta) = \sin(\phi)\sin(\delta) + \cos(\phi)\cos(\delta)\cos(\delta)$ 

substituting back in  $Q=S_0\left(\frac{\overline{R_0}}{R_e}\right)^2\left(\sin\left(\phi\right)\sin\left(\delta\right)+\cos\left(\phi\right)\cos\left(\delta\right)\cos\left(h\right)\right)$ 

we can get the delta with  $\delta=-0.409\cdot\cos\left(\frac{2\pi}{365}\cdot(d+10)\right)$  where 23.45deg in radians in 0.409

integrating over a day, h goes from pi to negative pi  $\overline{Q}^{\rm day} = -\frac{1}{2\pi} \int_{\pi}^{-\pi} Q \, dh$ 

 $\frac{R_o^2}{R_F^2}$  is constant, so the integral becomes

$$\begin{split} \int_{\pi}^{-\pi} Q \, dh &= \int_{h_o}^{-h_o} Q \, dh \\ &= S_o \frac{R_o^2}{R_E^2} \int_{h_o}^{-h_o} \cos(\Theta) \, dh \\ &= S_o \frac{R_o^2}{R_E^2} \left[ h \sin(\phi) \sin(\delta) + \cos(\phi) \cos(\delta) \sin(h) \right]_{h=h_o}^{h=-h_o} \\ &= -2S_o \frac{R_o^2}{R_E^2} \left[ h_o \sin(\phi) \sin(\delta) + \cos(\phi) \cos(\delta) \sin(h_o) \right] \end{split}$$
 factoring in the -1/2pi,

we get the:

final

$$\overline{Q}^{\mathrm{day}} = \frac{S_0}{\pi} \frac{R_0^2}{R_e^2} \, \left[ h_0 \sin \left( \phi \right) \sin \left( \delta \right) + \cos \left( \phi \right) \cos \left( \delta \right) \sin \left( h_0 \right) \right]$$

wiki: Let  $\_h_0$  be the hour angle when Q becomes positive. This could occur at sunrise when  $\Theta = \frac{1}{2}\pi$ , or for  $\_h_0$  as a solution of

$$\sin(\phi)\sin(\delta) + \cos(\phi)\cos(\delta)\cos(h_o) = 0$$

or

$$\cos(h_o) = -\tan(\phi)\tan(\delta)$$
 end wiki therefore,  $h_0 = \cos^{-1}(-\tan(\phi)\tan(\delta))$ 

theoretical daily average insolation at the top of the atmosphere as a function of lattitude and time of year Pasted image 20211110162119.png

Pasted image 20211110172859.png equator, summer solstice