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Master's In Space And Aeronautical Engineering

Computational Engineering Assignment 2

Generic Convection- Diffusion Equation

October, 2023

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If it is in the reader's interest, all the code done for this report can be found in the following repository: [Code](#)

1 Introduction

In this work our goal is to solve numerically the generic convection-diffusion equation. It is convenient to start the discussion by stating the Navier-Stokes (N-S) equations. For perfect gases, the N-S equations can be written as [1]:

$$\frac{\partial(\rho\mathbf{v})}{\partial t} + \nabla(\rho\mathbf{v}) = 0 \quad (1)$$

$$\frac{\partial(\rho\mathbf{v})}{\partial t} + \nabla(\rho\mathbf{v}\mathbf{v}) = \nabla(\mu\nabla\mathbf{v}) + [\nabla(\tau - \mu\nabla\mathbf{v}) - \nabla p + \rho\mathbf{g}] \quad (2)$$

$$\frac{\partial(\rho T)}{\partial t} + \nabla(\rho\mathbf{v}T) = \nabla\left(\frac{\lambda}{c_v}\nabla T\right) + \left[\frac{-\nabla\mathbf{q}^R - p\nabla\mathbf{v} + \tau : \nabla\mathbf{v}}{c_v}\right] \quad (3)$$

$$\frac{\partial(\rho Y_k)}{\partial t} + \nabla(\rho\mathbf{v}Y_k) = \nabla(\rho D_{km}\nabla Y_k) + (\dot{\omega}_k) \quad (4)$$

All these transport equations have an unsteady term, a convective term, a diffusion term and other terms. The generic convection diffusion equation for a generic variable ϕ can be expressed as:

$$\frac{\partial(\rho\phi)}{\partial t} + \nabla(\rho\mathbf{v}\phi) = \nabla(\Gamma_\phi\nabla\phi) + \dot{s}_\phi, \quad (5)$$

where Γ_ϕ and \dot{s}_ϕ are the diffusion coefficient and the extra source/sink term, respectively.

One can use the mass conservation equation (Eq. (1)) to write the equivalent convection-diffusion equation:

$$\rho\frac{\partial\phi}{\partial t} + \rho\mathbf{v}\nabla\phi = \nabla(\Gamma_\phi\nabla\phi) + \dot{s}_\phi. \quad (6)$$

To solve this numerically, we will use the finite volume method. This approach involves breaking down our computational domain into smaller Control Volumes (CVs). In Figure 1, we can see a representation of a CV along with its neighboring CVs.

Each CV is enclosed by four walls, each of which contains a node: one to the east (e), one to the north (n), one to the west (w), and one to the south (s). At the center of the CV, we find a node denoted as P. Surrounding this CV, there are four neighboring CVs, each with its central node labeled as East (E), North (N), West (W), and South (S). To specify the dimensions of the CV, we refer to its length in the X direction as Δxp and its height in the Y direction as Δyp .

As it can be noticed, in equation (6) we have a time partial derivative, This implies that we will also need a discretization along the time. This can be achieved by cutting

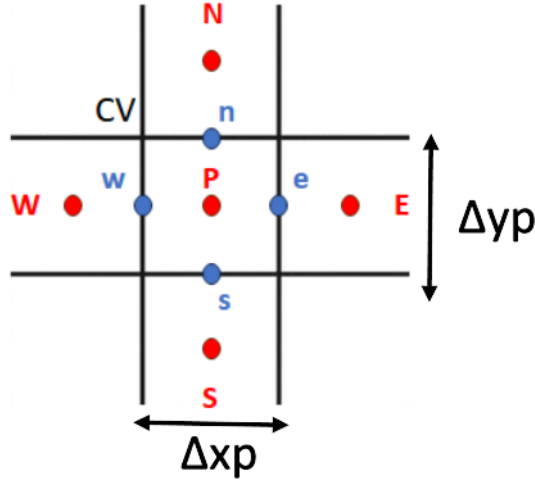


Figure 1: A scheme of a CV and it's neighbours. In red the centred nodes: P , S , N , W and E . In blue the nodes centred in the walls of the CV: s , n , w and e .

the time domain into small Δt time steps, such that the time evolution will be given as $t = t_0 + \Delta t \times i$, where i represents the total time steps.

Let's consider the integral form of the mass conservation (Eq. (1))

$$\frac{\partial}{\partial t} \int_{V_P} \rho dV + \int_{S_f} \rho \mathbf{v} \mathbf{n} dS = 0. \quad (7)$$

where V_P is the CV volume. Considering the centred point P (see Fig. 1) the mass conservation equation can be semi-discretized as:

$$V_P \frac{\partial \bar{\rho}_P}{\partial t} + \dot{m}_e - \dot{m}_w + \dot{m}_n - \dot{m}_s = 0 \quad (8)$$

where $\bar{\rho}_P = 1/V_P \int_{V_P} \rho dV$ and $\dot{m}_i = \int_{S_i} \rho \mathbf{v} \mathbf{n} dS$.

Integrating the above equation over time (between time t^n and t^{n+1}) we obtain:

$$V_P \int_{t^n}^{t^{n+1}} \frac{\partial \rho_P}{\partial t} dt + \int_{t^n}^{t^{n+1}} (\dot{m}_e - \dot{m}_w + \dot{m}_n - \dot{m}_s) dt = 0 \quad (9)$$

which gives the following implicit discretization form:

$$\frac{\rho_P - \rho_P^0}{\Delta t} V_P + \dot{m}_e - \dot{m}_w + \dot{m}_n - \dot{m}_s = 0 \quad (10)$$

where ρ_P and ρ_P^0 are the densities at point P for time steps $n + 1$ and n , respectively.

Going back to Eq. (6), the numerical implicit approximation of the different integral form terms can be written as:

$$\int_{t^n}^{t^{n+1}} \int_{V_P} \frac{\partial(\rho\phi)}{\partial t} dV dt \approx V_P(\rho_P\phi_P - \rho_P^0\phi_P^0) \quad (11)$$

$$\int_{t^n}^{t^{n+1}} \int_{V_P} \nabla(\rho\mathbf{v}\phi) dV dt \approx (\dot{m}_e\phi_e - \dot{m}_w\phi_w + \dot{m}_n\phi_n - \dot{m}_s\phi_s)\Delta t \quad (12)$$

$$\int_{t^n}^{t^{n+1}} \int_{V_P} \nabla(\Gamma_\phi \nabla\phi) dV dt \approx \left(-\Gamma_w \frac{\phi_P - \phi_W}{d_{PW}} S_w + \Gamma_e \frac{\phi_E - \phi_P}{d_{PW}} S_e - \Gamma_s \frac{\phi_P - \phi_s}{d_{PS}} S_s + \Gamma_n \frac{\phi_N - \phi_P}{d_{PN}} S_n \right) \quad (13)$$

$$\int_{t^n}^{t^{n+1}} \int_{V_P} \dot{s}_\phi dV dt \approx (S_C^\phi + S_P^\phi\phi_P) V_P \Delta t \quad (14)$$

Using the above terms and discretized mass conservation equation, Eq. (10), we obtain the following equation:

$$\rho_P^0 \frac{\rho_P - \rho_P^0}{\Delta t} V_P + \dot{m}_e(\phi_e - \phi_P) - \dot{m}_w(\phi_w - \phi_P) + \dot{m}_n(\phi_n - \phi_P) - \dot{m}_s(\phi_s - \phi_P) = \quad (15)$$

$$D_e(\phi_E - \phi_P) - D_w(\phi_P - \phi_W) + D_n(\phi_N - \phi_P) - D_s(\phi_P - \phi_S) + (S_C^\phi + S_P^\phi\phi_P) V_P$$

where $D_e = \Gamma_e S_e / d_{PE}$ and $D_w = \Gamma_w S_w / d_{PW}$. On the one hand, it can be seen that the source and diffusion terms are well defined in the centred nodes (P,E,W,S and N) of the CVs. On the other hand, the convective terms are given in the nodes of the faces (e,w,s and n) of the CVs. At this point, one has to assume some approximations to evaluate the convective terms in the central nodes.

There are different evaluation methods for the convective terms. The simplest one is the central-difference scheme (CDS):

$$\phi_e - \phi_P = f_e(\phi_E - \phi_P) \quad (16)$$

where $f_e = d_{Pe} / d_{PE}$. Even though CDS is a second-order scheme, it yields to stability problems.

For incompressible flows, or gases at low Mach numbers, the upwind-difference scheme (UDS) is a good choice since the convective terms are more influenced by upstream conditions than downstream ones. UDS is more stable than EDS but is a first-order accurate, so we gain stability but we lose accuracy. The UDS approximation can be

written as:

$$\dot{m}_e(\phi_e^{UDS} - \phi_P) = \frac{\dot{m}_e - |\dot{m}_e|}{2}(\phi_E - \phi_P) \quad (17)$$

Other usefull scheme is the exponential-difference scheme (EDS). In this case, f_e function is defined as:

$$f_e = \frac{e^{Pe d_{Pe}/d_{PE}} - 1}{e^{Pe} - 1} \quad (18)$$

where $Pe = \frac{\rho_e v_{xe} d_{PE}}{\Gamma_e}$ is the so-called Peclet number. EDS is still a first-order accurate scheme. More accurate ones (up to second-order) can be found in the literature as second-order upwind linear extrapolation (SUDS) and quadratic upwind interpolation for convective kinematics (QUICK). High-resolution schemes (HRS) can be also applied. The basic idea for the faces, f , can be expressed as:

$$\phi_f^{HRS} - \phi_P = (\phi_f^{UDS} - \phi_P) + (\phi_f^{HRS*} - \phi_f^{UDS*}) \quad (19)$$

In summary, using UDS the final form of the discretized generic convection-diffusion equation can be written as it follows:

$$a_P \phi_P = a_e \phi_E + a_W \phi_W + a_N \phi_N + a_S \phi_S + b_P \quad (20)$$

$$a_E = D_e - \frac{\dot{m}_e - |\dot{m}_e|}{2}; \quad a_W = D_W + \frac{\dot{m}_w + |\dot{m}_w|}{2} \quad (21)$$

$$a_N = D_n - \frac{\dot{m}_n - |\dot{m}_n|}{2}; \quad a_S = D_s + \frac{\dot{m}_s + |\dot{m}_s|}{2} \quad (22)$$

$$a_P = a_E + a_W + a_N + a_S + \frac{\rho_P^0 V_P}{\Delta t} - S_P^\phi V_P \quad (23)$$

$$b_P = \frac{\rho_P^0 V_P}{\Delta t} \phi_P^0 + S_C^\phi V_P - \dot{m}_e(\phi_e^{HRS*} - \phi_e^{UDS*}) + \dot{m}_w(\phi_w^{HRS*} - \phi_w^{UDS*}) - \dot{m}_n(\phi_n^{HRS*} - \phi_n^{UDS*}) + \dot{m}_s(\phi_s^{HRS*} - \phi_s^{UDS*}) \quad (24)$$

where for example, $\dot{m}_e \approx \rho(V_E - V_P)\Delta x \Delta z/2$.

Here we have to remark that in this work we considered no source term, $S_C = 0$ and no HRS correction. This implies that in the resolution of the above equations we will have to consider less terms.

We don't have to forget that the Boundary Conditions (BC) play a key roll in numerical

calculations. A more detailed description around BC will be given in Section 2.

2 Problem Description

In this section, we will outline the key elements of the two problem addressed within this work: the Diagonal Flow problem and the Smith-Hutton problem.

2.1 Diagonal Flow

For the Diagonal Flow scenario, we consider a symmetrical box with dimensions $L \times L$, as illustrated in Figure 2. This box is discretized into CVs, which have the same shape as depicted in Figure 1.

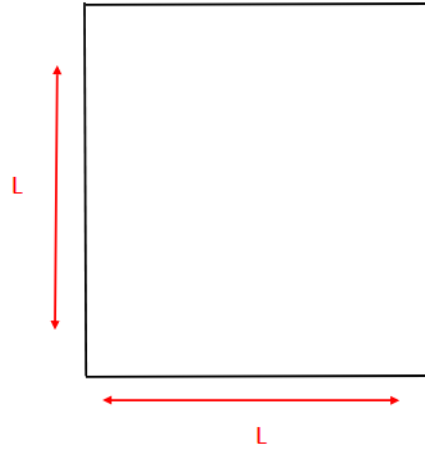


Figure 2: A scheme of the diagonal flow case.

For this problem, we impose Dirichlet boundary conditions on all four walls. Specifically, on the bottom and right walls, we set $\phi = \phi_{low}$ at $(x, y = 0)$ and $(x = L, y)$, while on the top and left walls, we set $\phi = \phi_{high}$ at $(x = 0, y)$ and $(x, y = H)$.

In terms of the coefficients a_i , where $i \in [P, E, W, N, S]$, and b_p , the values at the bottom and right walls are assigned as follows: $a_E = a_W = a_N = a_S = 0$, $a_P = 1$, and $b_P = \phi_{low}$. Similarly, for the top and left walls, the coefficients are set as $a_E = a_W = a_N = a_S = 0$, $a_P = 1$, and $b_P = \phi_{high}$.

The velocity field is described by the following equations:

$$v_x = V_{In} \cos(\alpha) ; \quad v_y = V_{In} \sin(\alpha) \quad (25)$$

Here, V_{In} is a constant value, and $\alpha = \pi/4$. The velocity field represented by Eq. (25) is visualized in Figure 3.

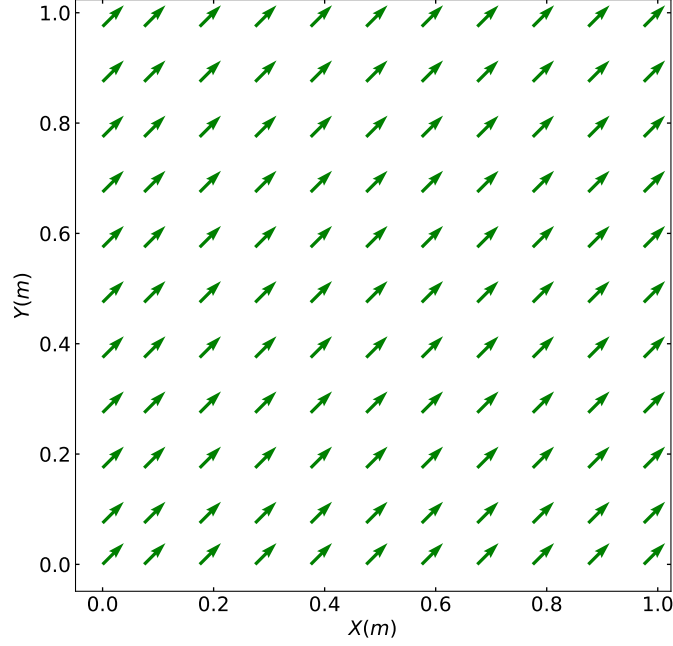


Figure 3: *Velocity field for the diagonal flow problem. Parameters: $N = M = 20$.*

2.2 Smith-Hutton

For the Smith-Hutton scenario, we consider a non-symmetrical box with dimensions $2L \times L$, as illustrated in Figure 4. This box is discretized into CVs, which have the same shape as depicted in Figure 1.

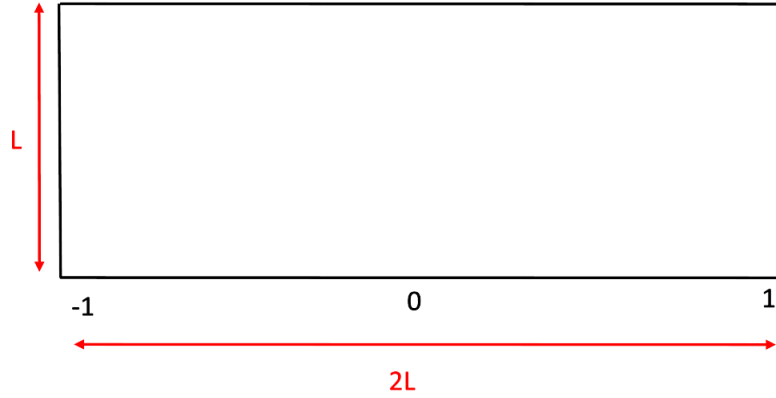


Figure 4: *A scheme of the Smith-Hutton case.*

For this case, the inlet of the problem goes from -1 to 0, while the outlet goes from 0 to 1. We impose Dirichlet boundary conditions on all three walls (left, top and right) and in the inlet. For the outlet we impose Neumann BCs. Specifically, on the left, top and right walls, we set $\phi = 1 - \tanh(10)$ at $(x = -1, y)$, $(-1 \leq x \leq 1, y = 1)$ and $(x = 1, y)$, while for the inlet, we set $\phi = 1 + \tanh[10(2x + 1)]$ at $(-1 \leq x \leq 0, y = 0)$; and for the outlet $\partial\phi/\partial y = 0$ at $(0 < x < 1, y = 0)$.

In terms of the coefficients a_i , where $i \in [P, E, W, N, S]$, and b_p , the values at the left, top and right walls are assigned as follows: $a_E = a_W = a_N = a_S = 0$, $a_P = 1$, and $b_P = 1 - \tanh(10)$. Similarly, for the inlet, the coefficients are set as $a_E = a_W = a_N = a_S = 0$, $a_P = 1$, and $b_P = 1 + \tanh[10(2x + 1)]$, while for the outlet $a_E = a_W = a_S = 0$, $a_N = a_P = 1$, and $b_P = 0$.

The velocity field is described by the following equations:

$$v_x = 2y(1 - x^2) ; \quad v_y = -2x(1 - y^2) \quad (26)$$

The velocity field represented by Eq. (26) is visualized in Figure 5.

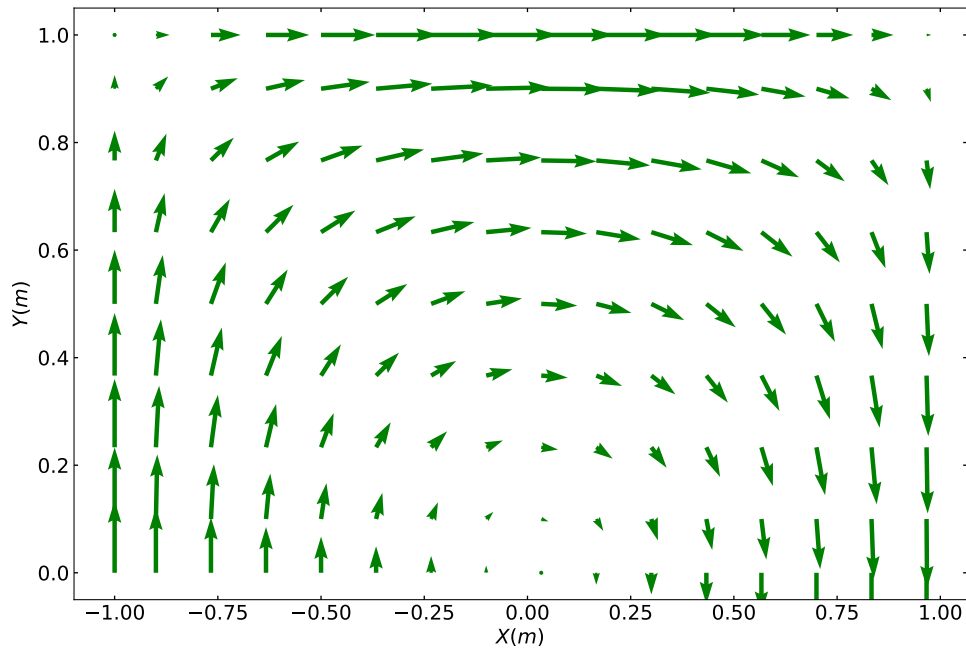


Figure 5: *Velocity field for the Smith-Hutton problem. Parameters: $N = 30$ & $M = 15$.*

3 Code Structure

In this section we introduce the code structure. As it will be shown in Section 7, the code is written for Python3 language. Further discussion on the code can be found in the repository linked below the **Index**.

- Establish the input data, which is segregated into two distinct sections: one dedicated to the physical aspects of the problem encompassing parameters like the cavity's length and height (L and H), the input velocity v_{in} , thermodynamic variables such as initial temperature, pressure, and density (T_{in} , P_{in} , and ρ_{in}) and the ρ/Γ relation. The other is dedicated to numerical aspects as the CVs quantity (N and M) etc.
- Generate the mesh with $N \times M$ CVs over L and H .
- Define all the matrices with dimension of $(N + 2) \times (M + 2)$.
- Compute the velocity fields, which has been explained in Section 2.
- Initialize ϕ and ϕ^0 and evaluate all the internal nodes a_i $i \in [P, E, W, N, S]$ and b_P (here we chose the evaluation scheme). In addition to this we compute the BCs.
- Evaluation of the new time steps: $t = t + \Delta t$.
- Solve the Gauss-Seidel algorithm for all internal nodes.
- Ask for new time step if needed (is $|\phi - \phi^0| < \varepsilon$? if answer is NO we need a new time step).
- Final calculations and create all the plots and save them.

4 Validation and Verification of the Code

Before presenting the results obtained in this work, it is crucial to conduct an analysis to ensure the code's proper functionality.

4.1 Comparison With Reference Solution

The presentation slides in our class included reference values, as presented in Fig. 6, for the outlet solution of ϕ in the Smith-Hutton problem across various ρ/Γ values. Therefore, it is important to conduct a comparative analysis between our solutions for the outlet values using different schemes (UDS and EDS) and the provided reference values.

x -position	$\rho/\Gamma = 10$	$\rho/\Gamma = 10^3$	$\rho/\Gamma = 10^6$
0.0	1.989	2.0000	2.000
0.1	1.402	1.9990	2.000
0.2	1.146	1.9997	2.000
0.3	0.946	1.9850	1.999
0.4	0.775	1.8410	1.964
0.5	0.621	0.9510	1.000
0.6	0.480	0.1540	0.036
0.7	0.349	0.0010	0.001
0.8	0.227	0.0000	0.000
0.9	0.111	0.0000	0.000
1.0	0.000	0.0000	0.000

Figure 6: Reference values of ϕ for the outlet of the Smith-Hutton problem for different ρ/Γ values.

Firstly, in Fig. 7 we show the different values of ϕ in the outlet for $\rho/\Gamma = 10, 10^3$ & 10^6 for UDS. We can observe that the main discrepancies occur at the upper and lower values of $\rho/\Gamma = 10^3$ & 10^6 and at the initial values for $\rho/\Gamma = 10$. This errors might be corrected with a bigger mesh and it will be analysed in Section 4.2.

Next, we replicate the Figure 7 using the EDS approach, as depicted in Fig. 8a. If we compare them, we can observe that there is minimal disparity between the results obtained with UDS and EDS. In Fig. 8b, we extend this comparison to $\rho/\Gamma = 10^3$. Both curves exhibit striking similarity, with the EDS curve consistently being closer to the reference curve. To offer a clearer perspective, we provide zoomed-in views of the upper and lower sections in Fig. 8c and 8d, respectively.

It is useful to put all these solutions into numbers. In Table 6 we compare the errors of ϕ in the outlet for different ρ/Γ values between EDS and UDS. It can be seen that we obtain smaller errors for EDS than UDS for $\rho/\Gamma = 10$ & 10^3 , while we obtain the same error for $\rho/\Gamma = 10^6$ (there is a difference for the fourth decimal, having EDS the smallest error).

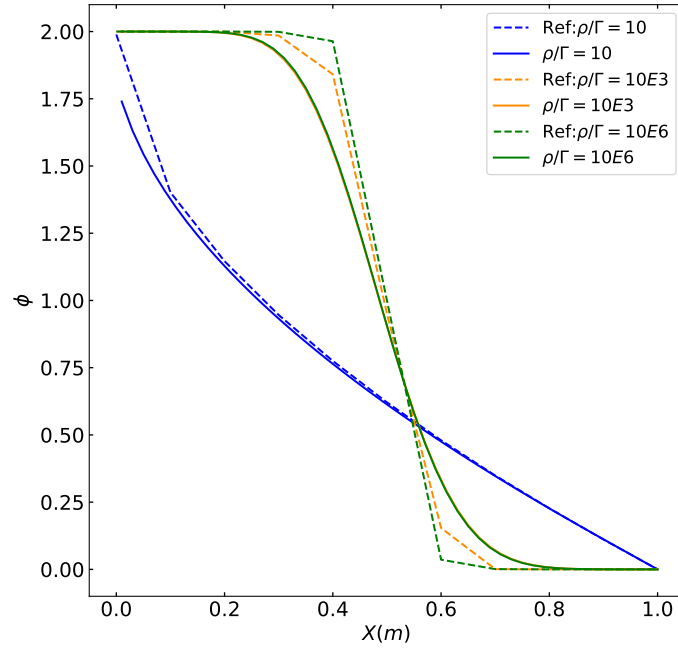


Figure 7: Values of ϕ for the outlet of the Smith-Hutton problem for different ρ/Γ values (solid lines) and the reference values (dashed lines). Here we make use of UDS and $N = 100$ & $M = 50$.

ρ/Γ	Error ^{UDS} (%)	Error ^{EDS} (%)
10	8.46	7.97
10^3	3.48	3.42
10^6	4.02	4.02

Table 1: Relative errors of the outlet values of ϕ respect to the reference for both schemes, UDS and EDS.

4.2 Mesh Refinement

Another method to confirm that the code functions correctly involves observing that simulations with larger meshes should exhibit smaller errors respect to the reference. To demonstrate this, we generated Table 2, 3 and 4 for the three ρ/Γ values and for various mesh configurations: $N = 20$ & $M = 10$, $N = 50$ & $M = 25$, $N = 100$ & $M = 50$, and $N = 200$ & $M = 100$. The results are shown in terms of UDS and EDS.

Mesh	Error ^{UDS} (%)	Error ^{EDS} (%)
$N = 20$ & $M = 10$	20.41	18.37
$N = 50$ & $M = 25$	11.61	10.69
$N = 100$ & $M = 50$	8.46	7.97
$N = 200$ & $M = 100$	6.83	6.31

Table 2: Relative errors of the outlet values of ϕ respect to the reference for both schemes, UDS and EDS. Here we make use of $\rho/\Gamma = 10$.

The tendency is quite clear: for bigger meshes we obtain smaller relative errors respect to the reference values. In addition to this, the errors of EDS are smaller respect to UDS. There

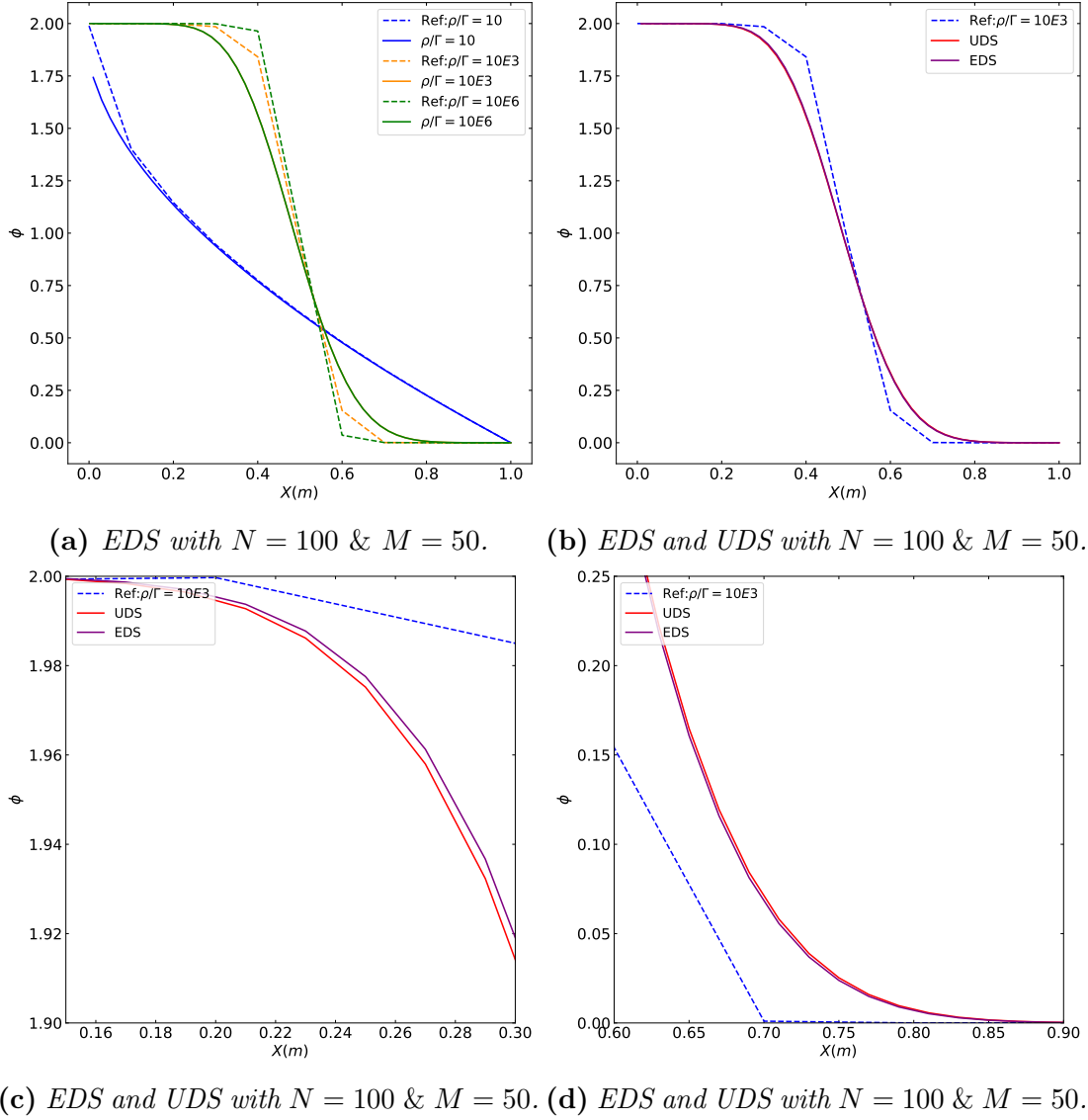


Figure 8: Values of ϕ for the outlet of the Smith-Hutton problem in **a)** for different ρ/Γ values (solid lines) and the reference values (dashed lines), in **b)** the comparison between UDS and EDS for $\rho/\Gamma = 10^3$, in **c)** a zoom into the upper part and in **d)** a zoom into the lower part. Here we make use of EDS and $N = 100$ & $M = 50$.

Mesh	Error ^{UDS} (%)	Error ^{EDS} (%)
$N = 20$ & $M = 10$	18.18%	18.11
$N = 50$ & $M = 25$	7.36	7.30
$N = 100$ & $M = 50$	3.48	3.42
$N = 200$ & $M = 100$	1.47	1.41

Table 3: Relative errors of the outlet values of ϕ respect to the reference for both schemes, UDS and EDS. Here we make use of $\rho/\Gamma = 10^3$.

is an exception, for $\rho/\Gamma = 10^6$ the relative errors are practically the same for all the meshes.

Mesh	Error ^{UDS} (%)	Error ^{EDS} (%)
$N = 20 \ \& \ M = 10$	18.63	18.63
$N = 50 \ \& \ M = 25$	7.88	7.88
$N = 100 \ \& \ M = 50$	4.02	4.02
$N = 200 \ \& \ M = 100$	2.03	2.03

Table 4: Relative errors of the outlet values of ϕ respect to the reference for both schemes, UDS and EDS. Here we make use of $\rho/\Gamma = 10^6$.

It is also noteworthy to discuss the required computational time for each problem to reach convergence. In both cases (Diagonal Flow and Smith-Hutton), we have noticed that a larger mesh size leads to an increased computational time, as it is expected. Furthermore, we have observed that smaller Peclet numbers (as well as lower values of ρ/Γ) requires a longer computational time. Additionally, when comparing both schemes, we have found that the EDS demands slightly more computational time compared to the UDS method.

5 Results

Here we present the results obtained for the Diagonal Flow and Smith-Hutton problems. As it have been shown in the above section, EDS is a little bit more precise. Therefore, the presented results will be computed using EDS.

5.1 Diagonal Flow

Starting with the Diagonal Flow problem, in Table 5 we show the physical and numerical values used to present the results.

The primary results are depicted in Fig. 9. In this figure, we illustrate the variations in ϕ for different Peclet numbers. It is evident that there are two distinct regions: an upper section closely aligned with ϕ_{high} and a lower section closely aligned with ϕ_{low} , with a transitional diffusion zone in between. As we increase the Peclet number, the extent of diffusion decreases. This trend can be attributed to the direct relationship between the Peclet number and the diffusion term. In an ideal scenario (free from numerical errors), as $Pe \rightarrow \infty$, we would expect the elimination of diffusion in our solution, resulting in the upper-mid region being filled with ϕ_{high} values and the lower-mid region populated with ϕ_{low} values. Consequently, our solution aligns with this expectation.

Moreover, we have observed in our results that for large Peclet values, there exists a point beyond which we cannot further reduce the diffusion of the solution (or the change is very small). This limitation arises due to numerical errors, introducing false diffusion.

Name	Symbol	Value	Units
X Direction Mesh	N	140	#
Y Direction Mesh	M	N	#
G-S/Time Convergence Crit.	ε	10^{-6}	#
Time Step	Δt	0.1	s
Cavity Length	L	1	m
Cavity Height	H	1	m
Input Temperature	T_{in}	298	K
Input Pressure	P_{in}	1.013×10^5	N / m ²
Input Velocity	V_{in}	5.15	m / s
Input Density	ρ_{in}	1.18	kg / m ³
Peclet number	Pe	$[1 - 10^7]$	#
ϕ top and left walls	ϕ_{high}	1.0	#
ϕ bottom and right walls	ϕ_{low}	0.0	#

Table 5: *Used numerical and physical input data for simulations.*

Notably, while our results are presented in terms of EDS, the same behavior is observed when employing UDS.

5.2 Smith-Hutton

Continuing with the Smith-Hutton problem, in Table 5 we show the physical and numerical values used to present the results.

Name	Symbol	Value	Units
X Direction Mesh	N	160	#
Y Direction Mesh	M	$\frac{N}{2}$	#
G-S/Time Convergence Crit.	ε	10^{-6}	#
Time Step	Δt	0.1	s
Cavity Length	L	2	m
Cavity Height	H	1	m
Input Temperature	T_{in}	298	K
Input Pressure	P_{in}	1.013×10^5	N / m ²
Input Velocity	V_{in}	5.15	m / s
Input Density	ρ_{in}	1.18	kg / m ³

Table 6: *Used numerical and physical input data for simulations.*

The main results are presented in Fig. 10. In this figure, we depict the variations in ϕ for three different ρ/Γ values. As expected, the variation of ϕ is observed from the inlet to the outlet. Notably, for the smallest ρ/Γ value, we obtain the least symmetrical results, indicating a significant presence of diffusion, meaning that the convective term is more

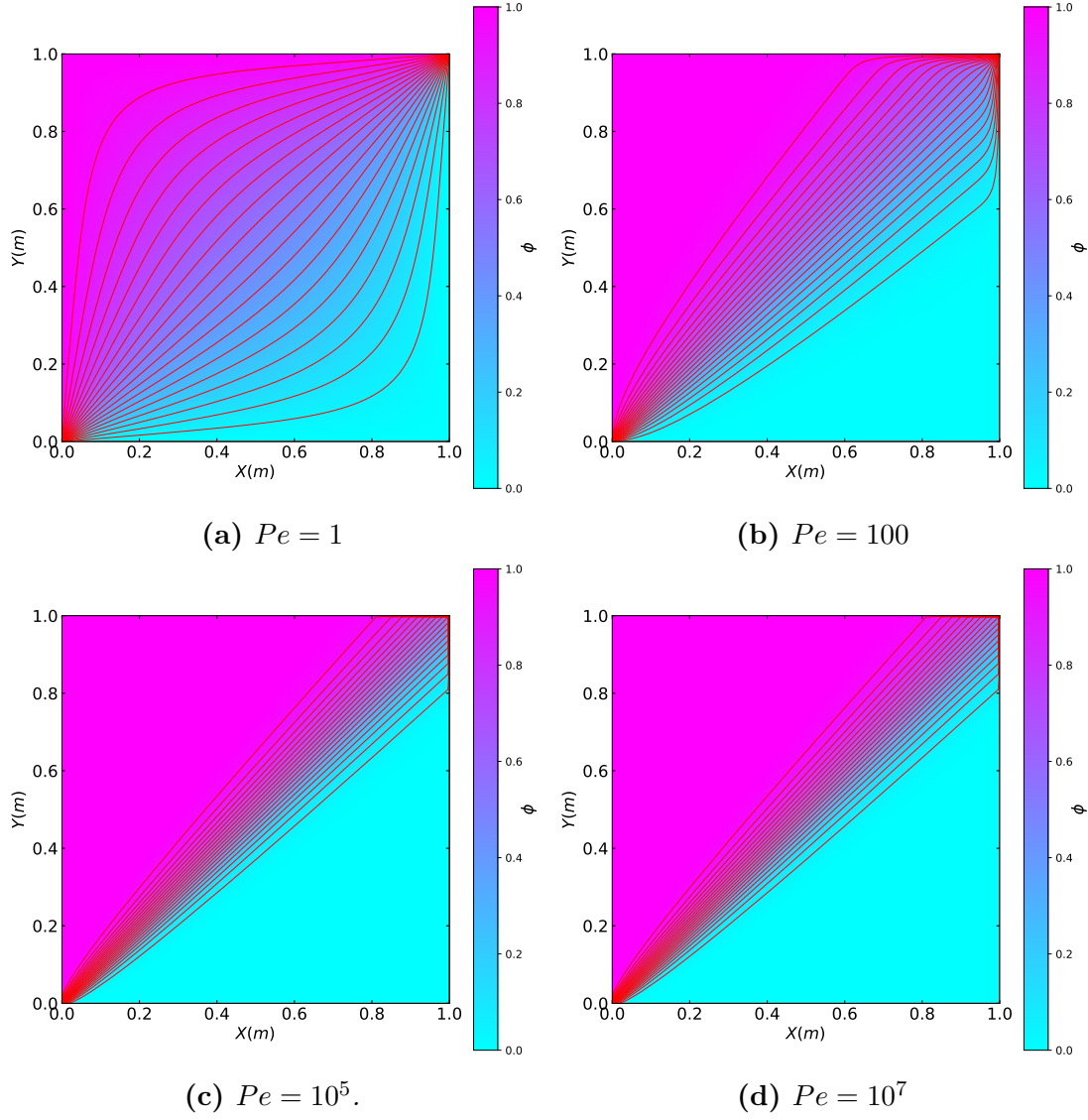


Figure 9: Values of ϕ for the Diagonal problem in **a)** for $Pe = 1$, in **b)** for $Pe = 100$, in **c)** for $Pe = 10^5$ and in **d)** for $Pe = 10^7$. Here we make use of EDS. All the used parameters are shown in Table 5.

dominant for low Peclet values. With an increase in the ρ/Γ value, the solution becomes more symmetrical, although some diffusion is still evident. This means that the diffusive term decreases with increasing ρ/Γ values. Similar to the Diagonal Flow problem, we observe false diffusion caused by numerical errors for the biggest ρ/Γ value. The same behavior is observed when using UDS.

For a better observation of the results we also show them in a 3D plot in Fig. 11.

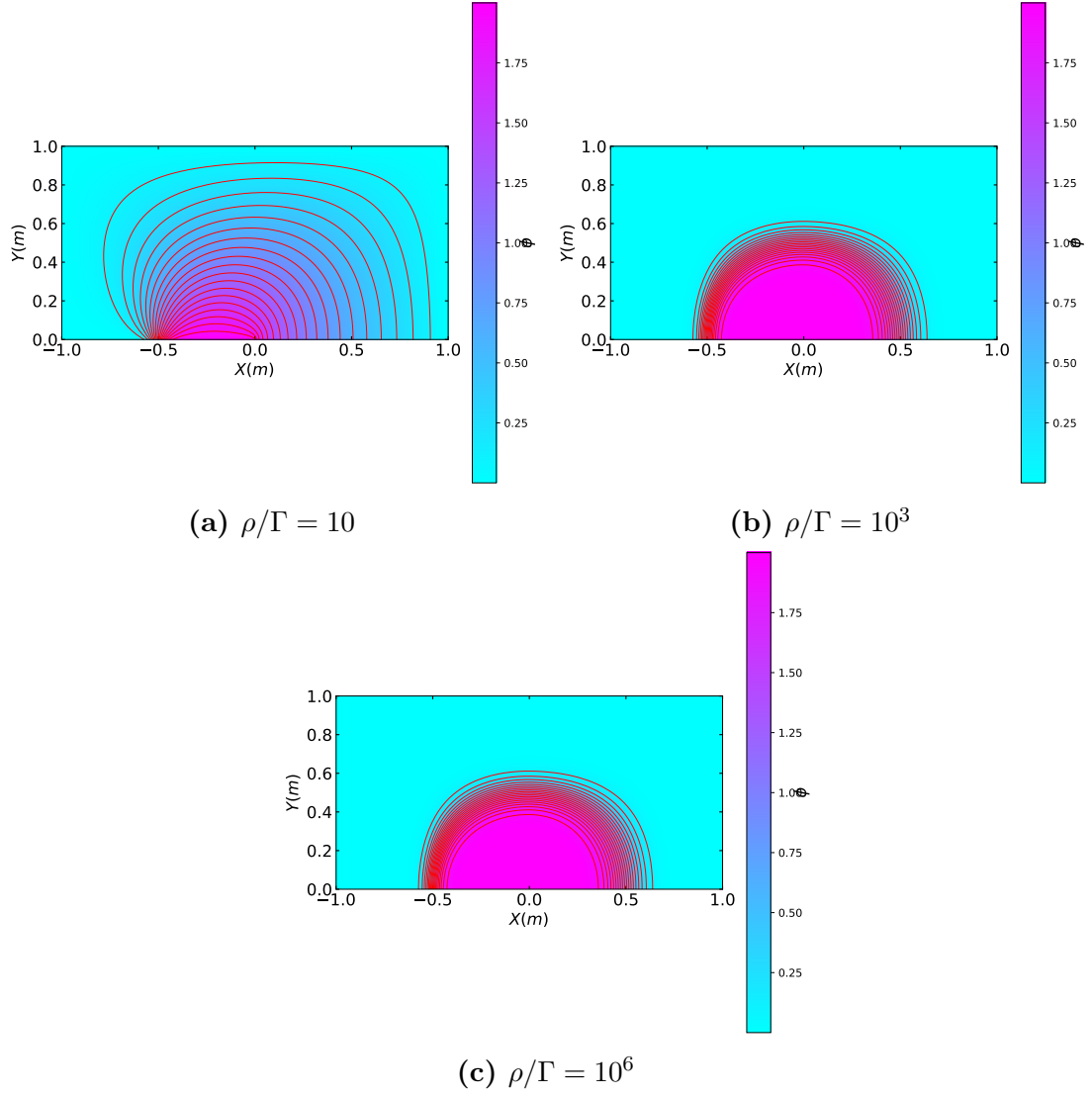


Figure 10: Values of ϕ for the Smith Hutton problem in **a)** for $\rho/\Gamma = 10$, in **b)** for $\rho/\Gamma = 10^3$, and in **c)** for $\rho/\Gamma = 10^6$. Here we make use of EDS. All the used parameters are shown in Table 6.

6 Conclusions

In this work we have analysed the generic convection-diffusion equation for two problems: Diagonal Flow and Smith-Hutton problems, focusing on the comparison with reference solutions and mesh refinement.

In the comparison with the reference solution for the Smith-Hutton problem, it was observed that for $\rho/\Gamma = 10$ and 10^3 , both the UDS and EDS approaches exhibited small errors compared to the reference, with EDS consistently outperforming UDS. For $\rho/\Gamma = 10^6$, the errors were very similar (same up to fourth decimal) between EDS and UDS. The results confirmed that the code could reproduce the reference values.

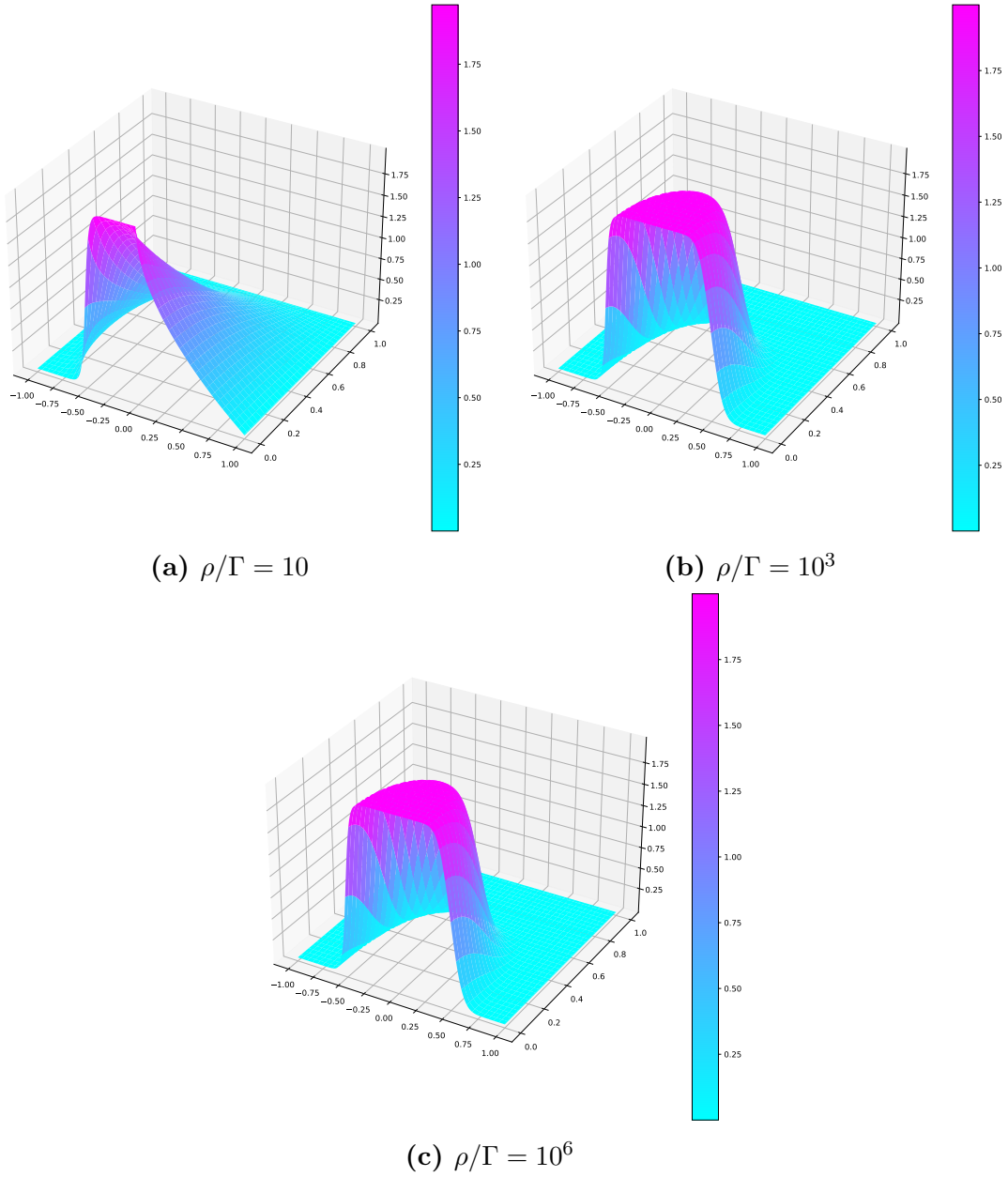


Figure 11: Values of ϕ projected in Z direction for the Smith-Hutton problem in **a)** for $\rho/\Gamma = 10$, in **b)** for $\rho/\Gamma = 10^3$, and in **c)** for $\rho/\Gamma = 10^6$. Here we make use of EDS. All the used parameters are shown in Table 6.

Mesh refinement studies demonstrated that simulations with larger meshes resulted in smaller errors relative to the reference, as expected. Larger meshes reduced the relative errors, and EDS consistently outperformed UDS in terms of accuracy, except the $\rho/\Gamma = 10^6$, which showed same results (again up to fourth decimal).

Additionally, a comment on the computational time for both schemes was performed. It was found that larger meshes required more computational time, while larger ρ/Γ values reduced the computation time. UDS was generally faster than EDS, which could be

advantageous for larger simulations.

In the results section, we presented the results for both the Diagonal Flow and Smith-Hutton problems, focusing on the EDS approach due to its higher precision.

For the Diagonal Flow problem, simulations were conducted across a range of Peclet numbers. The results showed that as the Peclet number increased, the diffusion decreased. However, for very high Peclet numbers, numerical errors introduced a false diffusion component.

In the Smith-Hutton problem, the focus was on different ρ/Γ values. It was observed that for smaller ρ/Γ values, the solution exhibited less symmetry, indicating a significant presence of diffusion. With increasing ρ/Γ , the solution became more symmetrical, but with some remaining diffusion. Numerical errors introduced false diffusion for the largest ρ/Γ values.

References

- [1] S. Patankar, *Numerical heat transfer and fluid flow*. Taylor & Francis, 2018.

7 Appendix: Python3 Code

Here we show the code in Python3 language for the Smith-Hutton problem. The rest of the code will be attached in a .zip in atenea. Also it can be found in the link below the [Index](#) of this report.

```
1 #=====
2 # Code for the Smith-Hutton Case
3 # Master in Space and Aeronautical
4 # Engineering.
5 # Computational Engineering: Assginament
6 # 2.
7 # Author: Inaki Fernandez Tena
8 # email: inakiphy@gmail.com
9 #=====
10
11 #=====
12 # Import modules
13 #=====
14
15 import matplotlib.pyplot as plt
16 import numpy as np
17 from matplotlib.colors import ListedColormap
18 from mpl_toolkits.mplot3d import Axes3D
19 from matplotlib import cm
20
21
22 #=====
23 # Physical input data
24 #=====
25
26 L      = 2.0                # Channel lenght.
27 H      = L/2.0             # Channel height.
28 Z      = 1.0               # Channel depth.
29 V_in   = 1.15              # Air velocity in m/s.
30 T_in   = 298.0             # Normal air temperature in K.
31 P_in   = 1.013E05          # Air pressure at sea level in N/m^2.
32 R      = 287               # Ideal gas constant in J/mol*K
33 rho_in = P_in / (R * T_in) # Air density at sea level in kg/m^3.
34         Incompressible = constant density at all mesh points.
35 delta_t = 0.1
36 rhoGamma = 10E6
37 Gamma    = rho_in / rhoGamma
38 phi_in   = 1.0
39
40 #=====
41 # Numerical input data
42 #=====
43
44 N      = 200                # Control volumes in x direction.
45 M      = int(N / 2)         # Control volumes in y direction.
46 delta_X = L / N             # Control volume lenght in x direction.
47 delta_Y = H / M             # Control volume height in y direction.
48 eps     = 1.0E-06           # Gauss-Seidel method convergence parameter.
49 t_max   = 1000000            # G-S loop steps.
50 time_max = 1000000           # Time loop steps.
51 delta_Z = 1.0               # Just one step into z direction
```



```

52
53 # Mesh generation
54
55 x_cv = np.linspace(-L/2.0, L/2.0, N+1)           # Generate x points with the same
           spacing
56 y_cv = np.linspace(0, H, M+1)                   # Generate y points with the same spacing
57 x_p  = np.zeros(N+2)                             # Vectors for the centered control volumes
           with N+2 elements
58 y_p  = np.zeros(M+2)                             # Vectors for the centered control volumes
           with M+2 elements
59
60 # Fill x_p and y_p
61
62 for i in range(1, N+1):
63     x_p[i] = (x_cv[i] + x_cv[i-1]) / 2.0
64
65 for j in range(1, M+1):
66     y_p[j] = (y_cv[j] + y_cv[j-1]) / 2.0
67
68 # Set boundary points at the ends of the domain
69 x_p[0]  = x_cv[0]                                # Left boundary
70 x_p[-1] = x_cv[-1]                               # Right boundary
71 y_p[0]  = y_cv[0]                                # Bottom boundary
72 y_p[-1] = y_cv[-1]                               # Top boundary
73
74 #=====
75 # Define matrixes
76 #=====
77
78 phi      = np.zeros((M+2,N+2))                   # Phi function
79 phi_zero  = np.zeros((M+2,N+2))                   # Phi at t=0 function
80 phi_ax    = np.zeros((M+2,N+2))                   # Phi auxiliar for G-S
81 rho       = np.zeros((M+2,N+2))                   # Density matrix
82 a_P       = np.zeros((M+2,N+2))                   # Auxiliar matrix at point p
83 a_E       = np.zeros((M+2,N+2))                   # Auxiliar matrix at point east
84 a_S       = np.zeros((M+2,N+2))                   # Auxiliar matrix at point south
85 a_W       = np.zeros((M+2,N+2))                   # Auxiliar matrix at point west
86 a_N       = np.zeros((M+2,N+2))                   # Auxiliar matrix at point north
87 b_P       = np.zeros((M+2,N+2))                   # Generation term
88 v_xP      = np.zeros((M+2,N+2))                   # Velocity in x direction
89 v_yP      = np.zeros((M+2,N+2))                   # Velocity in y direction
90 m_e       = np.zeros((M+2,N+2))                   # east mass flow rate
91 m_s       = np.zeros((M+2,N+2))                   # south mass flow rate
92 m_w       = np.zeros((M+2,N+2))                   # west mass flow rate
93 m_n       = np.zeros((M+2,N+2))                   # north mass flow rate
94 D_e       = np.zeros((M+2,N+2))
95 D_w       = np.zeros((M+2,N+2))
96 D_n       = np.zeros((M+2,N+2))
97 D_s       = np.zeros((M+2,N+2))
98
99
100 #=====
101 # Define velocities
102 #=====
103 for i in range(N+2):
104     for j in range(M+2):
105         v_xP[j,i] = 2.0 * y_p[j] * (1.0 - x_p[i]**2)
106         v_yP[j,i] = -2.0 * x_p[i] * (1.0 - y_p[j]**2)
107
108 #=====

```

```

109 # Initialize the map of \phi_{0} and \phi
110 #=====
111 # Internal nodes
112 for i in range(1, N+1):
113     for j in range(1, M+1):
114         phi_zero[j,i] = phi_in
115         phi[j,i] = phi_in
116
117 # Inlet nodes
118 for i in range(int((N+2) / 2)):
119     for j in range(M+2):
120         phi_zero[0,i] = 1.0 + np.tanh(10.0 * (2.0 * x_p[i] + 1.0))
121         phi[0,i] = 1.0 + np.tanh(10.0 * (2.0 * x_p[i] + 1.0))
122 # Walls
123 for i in range(N+2):
124     for j in range(M+2):
125         phi_zero[j,0] = 1.0 - np.tanh(10.0)
126         phi_zero[-1,i] = 1.0 - np.tanh(10.0)
127         phi_zero[j,-1] = 1.0 - np.tanh(10.0)
128         phi[j,0] = 1.0 - np.tanh(10.0)
129         phi[-1,i] = 1.0 - np.tanh(10.0)
130         phi[j,-1] = 1.0 - np.tanh(10.0)
131
132 #=====
133 # Compute mass flow rates
134 #=====
135 for i in range(1,N+1):
136     for j in range(1,M+1):
137         m_e[j,i] = rho_in * (v_xP[j,i + 1] + v_xP[j,i]) * delta_Y * delta_Z / 2.0
138         m_w[j,i] = rho_in * (v_xP[j,i - 1] + v_xP[j,i]) * delta_Y * delta_Z / 2.0
139         m_n[j,i] = rho_in * (v_yP[j + 1,i] + v_yP[j,i]) * delta_X * delta_Z / 2.0
140         m_s[j,i] = rho_in * (v_yP[j - 1,i] + v_yP[j,i]) * delta_X * delta_Z / 2.0
141
142 #=====
143 # Compute Di
144 #=====
145 # Internal nodes
146 for i in range(1, N+1):
147     for j in range(1, M+1):
148         D_e[j,i] = Gamma * delta_Y * delta_Z / np.abs(x_p[i] - x_p[i + 1])
149         D_w[j,i] = Gamma * delta_Y * delta_Z / np.abs(x_p[i] - x_p[i - 1])
150         D_n[j,i] = Gamma * delta_X * delta_Z / np.abs(y_p[j] - y_p[j + 1])
151         D_s[j,i] = Gamma * delta_X * delta_Z / np.abs(y_p[j] - y_p[j - 1])
152
153 # Internal nodes
154 for i in range(1, N+1):
155     for j in range(1,M+1):
156         #Pe_e = m_e[j,i] / D_e[j,i] # For EDS
157         #Pe_w = m_w[j,i] / D_w[j,i] # For EDS
158         #Pe_n = m_n[j,i] / D_n[j,i] # For EDS
159         #Pe_s = m_s[j,i] / D_s[j,i] # For EDS
160         #a_E[j,i] = D_e[j,i] * (np.abs(Pe_e) / (np.exp(np.abs(Pe_e)) - 1.0)) - ((m_e[j,i]
161         - np.abs(m_e[j,i])) / 2.0) # For EDS
162         #a_W[j,i] = D_w[j,i] * np.abs(Pe_w) / (np.exp(np.abs(Pe_w)) - 1.0) + ((m_w[j,i] +
163         np.abs(m_w[j,i])) / 2.0) # For EDS
164         #a_N[j,i] = D_n[j,i] * np.abs(Pe_n) / (np.exp(np.abs(Pe_n)) - 1.0) - ((m_n[j,i] -
165         np.abs(m_n[j,i])) / 2.0) # For EDS
166         #a_S[j,i] = D_s[j,i] * np.abs(Pe_s) / (np.exp(np.abs(Pe_s)) - 1.0) + ((m_s[j,i] +
167         np.abs(m_s[j,i])) / 2.0) # For EDS
168         a_E[j,i] = D_e[j,i] - ((m_e[j,i] - np.abs(m_e[j,i])) / 2.0)

```

```

165     a_W[j,i] = D_w[j,i] + ((m_w[j,i] + np.abs(m_w[j,i])) / 2.0)
166     a_N[j,i] = D_n[j,i] - ((m_n[j,i] - np.abs(m_n[j,i])) / 2.0)
167     a_S[j,i] = D_s[j,i] + ((m_s[j,i] + np.abs(m_s[j,i])) / 2.0)
168     a_P[j,i] = a_E[j,i] + a_W[j,i] + a_N[j,i] + a_S[j,i] + ((rho_in * delta_X *
        delta_Y * delta_Z) / delta_t)
169
170 # Inlet DIRICHLET
171 for i in range(int((N+2) / 2)):
172     for j in range(M+2):
173         a_E[0,i] = 0.0
174         a_W[0,i] = 0.0
175         a_N[0,i] = 0.0
176         a_S[0,i] = 0.0
177         b_P[0,i] = 1 + np.tanh(10.0 * (2.0 * x_p[i] + 1.0))
178         a_P[0,i] = 1.0
179
180 # Outlet NEUMANN
181 for i in range(int((N+2) / 2), N+2):
182     for j in range(M+2):
183         a_E[0,i] = 0.0
184         a_W[0,i] = 0.0
185         a_N[0,i] = 1.0
186         a_S[0,i] = 0.0
187         b_P[0,i] = 0.0
188         a_P[0,i] = 1.0
189
190 # Walls Dirichlet
191 for i in range(N+2):
192     for j in range(M+2):
193         a_E[j,0] = 0.0 # Left
194         a_W[j,0] = 0.0 # Left
195         a_N[j,0] = 0.0 # Left
196         a_S[j,0] = 0.0 # Left
197         b_P[j,0] = 1.0 - np.tanh(10.0) # Left
198         a_P[j,0] = 1.0 # Left
199         a_E[j,-1] = 0 # Right
200         a_W[j,-1] = 0 # Right
201         a_N[j,-1] = 0 # Right
202         a_S[j,-1] = 0 # Right
203         a_P[j,-1] = 1 # Right
204         b_P[j,-1] = 1.0 - np.tanh(10.0) # Right
205         a_E[-1,i] = 0.0 # Top nodes
206         a_W[-1,i] = 0.0 # Top nodes
207         a_N[-1,i] = 0.0 # Top nodes
208         a_S[-1,i] = 0.0 # Top nodes
209         a_P[-1,i] = 1.0 # Top nodes
210         b_P[-1,i] = 1.0 - np.tanh(10.0) # Top nodes
211
212 print("=====")
213 print("Starting loops")
214 print("=====")
215 # Begin time step (t+delta_t)
216 #=====
217 for time in range(time_max):
218     r2 = np.sum(phi_zero) # Sum the values for t = 0
219     #=====
220     # Begin Gauss-Seidel
221     #=====
222     phi_ax = phi_zero # Set the initial value of the auxiliary
        stream function

```

```

223     for t in range(t_max):
224         # phi_zero outlet NEUMANN
225         for i in range(int((N+2) / 2), N+2):
226             for j in range(M+2):
227                 phi_ax[0,i] = phi_ax[1,i]
228         # Inlet nodes of b_p
229         for i in range(1, N+1):
230             for j in range(1, M+1):
231                 b_P[j,i] = ((rho_in * delta_X * delta_Y * delta_Z) / delta_t) * phi_ax[j
232                     ,i]
233
234         r = np.sum(phi_ax) # Sum the values of the psi auxiliary
235                             matrix
236         for i in range(1, N+1):
237             for j in range(1, M+1):
238                 phi[j, i] = (a_E[j, i] * phi_ax[j, i + 1] + a_W[j, i] * phi_ax[j, i - 1]
239                     + a_N[j, i] * phi_ax[j + 1, i] + a_S[j,i] * phi_ax[j - 1, i] + b_P[j,
240                     i]) / a_P[j, i]
241
242         sum = np.sum(phi) # Sum the values of psi matrix
243         if np.abs(sum-r) <= eps: # Watch if the |psi - psi_aux| <
244             precision # If corveges just break the main G-S
245             print("=====")
246             print("G-S algorithm converged")
247             print("The requiered steps has been:" + " " + str(t))
248             print("=====")
249             break
250         else:
251             phi_ax = phi
252             sum2 = np.sum(phi) # Sum the values for i+1
253             if np.abs(sum2 - r2)<=eps:
254                 print("=====")
255                 print("Time loop converged")
256                 print("The requiered steps has been:" + " " + str(time))
257                 print("=====")
258                 break # If corveges just break the main G-S loop
259             else:
260                 phi_zero = phi
261 # Vectors for outlet plot
262
263 x_plot_exp = [0.0,0.1,0.2,0.3,0.4,0.5,0.6,0.7,0.8,0.9,1.0]
264 phi_10      = [1.989, 1.402, 1.146, 0.946, 0.775, 0.621, 0.480, 0.349, 0.227, 0.111,
265                 0.000]
266 phi_10E3    = [2.0000, 1.9990, 1.9997, 1.9850, 1.8410, 0.9510, 0.1540, 0.0010, 0.0000,
267                 0.0000]
268 phi_10E6    = [2.000, 2.000, 2.000, 1.999, 1.964, 1.000, 0.036, 0.001, 0.000, 0.000,
269                 0.000]
270
271 x_plot      = np.zeros(int((N+2) / 2))
272 phi_plot    = np.zeros(int((N+2) / 2))
273
274 # Write the outlet values
275
276 # Path to output file
277 file_output = f"Results/S-H/Outlet_{N}x{M}_{rhoGamma}.dat"
278
279 # Generate vectors to write
280 for i in range(int((N+2) / 2), N+2):
281     for j in range(M+2):

```

```

274     x_plot[i - int((N+2) / 2)] = x_p[i]
275     phi_plot[i - int((N+2) / 2)] = phi[0,i]
276
277 # Write
278
279 matrix = np.column_stack((x_plot,phi_plot))
280
281 # Save
282
283 np.savetxt(file_output, matrix, fmt='%.10f', delimiter='\t')
284 #=====
285 # Generate the plot
286 #=====
287 output_plot1 = f"Results/S-H/Phi_{N}x{M}_{rhoGamma}.pdf"
288 output_plot2 = f"Results/S-H/Vel_{N}x{M}_{rhoGamma}.pdf"
289 output_plot3 = f"Results/S-H/Phi3D_{N}x{M}_{rhoGamma}.pdf"
290 print(" ")
291 print("=====")
292 print("Starting the plot")
293 print(" ")
294 print("It might take a few seconds")
295 print("=====")
296
297 # Figure specifications
298 fontsize=15
299
300 # Start first plot
301 plt.figure(1)
302 plt.figure(figsize = (8,8))
303 plt.tick_params(axis='both', which='both',length=3, width=1.0,
304 labels=15, right=True, top=True, direction='in') # For ticks in borders
305
306 # Figure labels
307 plt.xlabel(r"$X(m)$", fontsize=fontsize)
308 plt.ylabel(r"$Y(m)$", fontsize=fontsize)
309
310 # Plot
311 plt.contour(x_p, y_p, phi, levels=20, colors='r', linewidths=0.5)
312 plt.imshow(phi, cmap= 'cool', extent=(x_p.min(), x_p.max(), y_p.min(), y_p.max()), origin
313         = 'lower')
314 plt.colorbar()
315
316 # Get the current axis
317 ax = plt.gca()
318
319 # Add text next to the color bar
320 text_x = 1.18 # Adjust the x-coordinate as needed
321 text_y = 0.5 # Adjust the y-coordinate as needed
322 text = r"$\phi$"
323 ax.text(text_x, text_y, text, transform=ax.transAxes, rotation=270, va='center', fontsize
324         =fontsize)
325
326 # Save figure
327 plt.savefig(output_plot1,bbox_inches='tight')
328
329 # Start second plot
330 plt.figure(2)
331 plt.figure(figsize = (12,8))
332 plt.tick_params(axis='both', which='both',length=3, width=1.0,
333 labels=15, right=True, top=True, direction='in') # For ticks in borders

```

```

332
333 # Figure labels
334 plt.xlabel(r"$X(m)$", fontsize=fontsize)
335 plt.ylabel(r"$Y(m)$", fontsize=fontsize)
336
337 # Plot
338 P,Z = np.meshgrid(x_p, y_p)
339 plt.quiver(P[:, :2], Z[:, :2], v_xP[:, :2], v_yP[:, :2], color="green")
340
341
342 # Save figure
343 plt.savefig(output_plot2, bbox_inches='tight')
344
345
346 # Start third plot
347 plt.figure(3)
348 plt.figure(figsize = (8,8))
349 plt.tick_params(axis='both', which='both', length=3, width=1.0,
350 labelsz=15, right=True, top=True, direction='in') # For ticks in borders
351
352 # Figure labels
353 plt.xlabel(r"$X(m)$", fontsize=fontsize)
354 plt.ylabel(r"$Y(m)$", fontsize=fontsize)
355
356 # Plot
357 # Create a 3D figure
358 fig = plt.figure(figsize=(12, 12))
359 ax = fig.add_subplot(111, projection='3d')
360
361 # Create the surface plot
362 surf = ax.plot_surface(P, Z, phi, cmap='cool')
363
364 # Add a colorbar
365 cbar = fig.colorbar(surf)
366
367 # Save figure
368 plt.savefig(output_plot3, bbox_inches='tight')
369
370
371 plt.close(1)
372 plt.close(2)
373 plt.close(3)

```