

Ray Tracing In Functional Programming Language

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Abstract—to do

Index Terms—ray tracing, three-dimensional graphics, functional programming.

I. INTRODUCTION

There are two approaches to rendering a three-dimensional scene: object-order rendering and image-order rendering. [7]

Object-order rendering involves iterating over each object and figuring out which set of pixels it maps to and coloring those pixels. In image-order rendering, for each pixel, the nearest visible object is found, and that pixel is colored using the given shading model. The former is classified into rasterization algorithms (or projective algorithms), and the latter is termed the ray tracing algorithm (image-space algorithm). Same effects, like reflections and shadows, are easy to implement in ray tracing compared to rasterization [4].

In ray tracing, image synthesis happens by shooting rays that go through the projection plane, and the nearest intersection with an object is found. If the object is reflective, rays bounce off it and hit other objects, which also influences color according to the shading model [4]. In our implementation [15], we have utilized the simplest shading model called Phong shading [7] and used the simplest three-dimensional geometrical object, a sphere, to illustrate ray tracing in the functional programming paradigm. To this end, we decided to use OCaml for its strong support for functional programming and its mechanisms to write imperative code as well, which will enable us to demonstrate alternative ways in the same programming language.

Historically, ray tracers have been developed using object-oriented (OO) techniques for several reasons, primarily due to their size and complexity. However, many students learning computer graphics are unlikely to write a production-grade ray tracer from scratch. Therefore, for a simple ray tracer, we recommend using a functional programming style for the following reasons.

First, ray tracing is inherently mathematical in nature, making it straightforward to convert concepts into code. Second, due to immutability, code is easier to isolate and debug. Most importantly, the code is succinct and readable, due to the type inference feature of functional programming.

Algorithm 1 Ray Tracing Algorithm

- 1: Initialize scene with objects and lights
- 2: Set canvas resolution (width, height)
- 3: For each pixel (x, y) in the canvas:
- 4: Generate ray from camera through pixel (x, y)
- 5: Initialize color to background color
- 6: Initialize closest intersection distance to infinity
- 7: For each object in the scene:
- 8: If ray intersects object:
- 9: Calculate intersection point
- 10: Calculate normal at intersection point
- 11: Calculate distance to intersection
- 12: If distance is less than closest intersection distance:
- 13: Update closest intersection distance
- 14: Calculate color at intersection point
- 15: Update color based on lighting
- 16: Set pixel color to calculated color
- 17: Render the image

II. ARCHITECTURE

We defined following functionalities for our ray tracer as illustrated through Fig. 1, Fig. 2, Fig. 3, Fig. 4 and Fig. 5.

```
Core
+ diffuse_i (normal : Point, i : Point, i : float, c_intensity : float) : float
+ specular_i (o : Point, p : Point, normal : Point, i : Point, i : float, s : int) : float
+ til_inner (normal : Point, p : Point, o : Point, s : int, light_l : light list, sphere_l : sphere list)
+ rtx_inner (o : Point, d : Point, tmin : float, tmax : float, sl : sphere list, ll : light list, limit : int) : sphere*float
```

Fig. 1. Core functions

```
Utils
+ quad(a : float, b : float, c : float) : float*float
+ clamp(x : float) : int
+ myrgb(tuple : int*int*int, intensity : float) : rgb
+ sphere_color(sl : sphere, intensity : float) : rgb
+ plotc(x : int, y : int, color : rgb)
+ g_to_viewport(gx : int, gy : int) : Point
+ create_tuple(min_x : int, max_x : int, min_y : int, max_y : int) : (int*int)list
+ shader(o : Point, tmin : float, tmax : float, ls : sphere list, ll : light list, pair : int*int)
```

Fig. 2. Utility functions

III. COORDINATE SYSTEM

In the OCaml standard graphics library, the origin of the drawing window is situated at the bottom left corner. For

Point
- x : float
- y : float
- z : float
+ sub3(p1 : Point, p2 : Point) : Point
+ add3(p1 : Point, p2 : Point) : Point
+ scale(k : float, p : Point) : Point
+ sproduct(p1 : Point, p2 : Point) : float
+ norm(p : Point) : float
+ reflected_ray(ncap : Point, l : Point) : Point

Fig. 3. Point type

light
- k : char
- i : float
- v : Point

Fig. 4. Light type

convenience, we want it in the middle of the window. Thus, so we wrote wrapper function *plotc*.

This function translates the origin to the middle of the graphics window.

IV. COLORS

We use the additive color model (RGB model), where each color channel is represented by an 8-bit positive integer. Therefore, the range is $[0, 255]$. We can treat the triplet of RGB values as a vector. Adding two colors together yields a new color, and multiplying a color by a scalar increases its brightness, i.e.,

$$k(r, g, b) = (kr, kg, kb)$$

There is a chance that any channel value may go out of the range $[0, 255]$ while manipulating colors. We can handle this as follows:

Suppose there is a channel value x :

- If $x > 255$, then we set $x = 255$.
- If $x < 0$, then we set $x = 0$.

This process is called *clamping*.

V. SCENE

The scene is simply a set of objects that we want to render on the screen.

To represent objects in a scene, we need to use a 3D coordinate system that employs real numbers to represent continuous values, which are then mapped to a discrete 2D graphics window while drawing.

sphere
- r : float
- c : Point
- color : int*int*int*
- s : int
- rf : float
+ intersect_sphere(o : Point, d : Point, s : sphere) : float*float
+ closest_sphere_inner(o : Point, d : Point, tmin : float, tmax : float, sl : sphere list, closest_s : sphere, closest_t : float) : sphere*float

Fig. 5. Sphere type

The viewing position from which we look at the scene is called the **camera position**. We assume that the camera is fixed and occupies a single point in space, which will often be the origin $O(0, 0, 0)$.

The *camera orientation* [7] is the direction in which the camera points, or from where rays enter the camera. We will assume the camera orientation to be Z_+ , the positive z-axis.

The frame has dimensions V_w and V_h and is frontal to the camera (perpendicular to the camera orientation, irrespective of camera position, in our case perpendicular to Z_+) and at a distance d away from the camera. Technically, this is called the *viewport* [7].

The angle visible from the camera is called the *field of view* (FOV). [7] It depends on the distance d from the camera to the viewport and the dimensions of the viewport V_w and V_h .

In our case, we assume

$$V_w = V_h = d = 1 \implies FOV \approx 53^\circ$$

We will represent the coordinates of the viewport as (V_x, V_y) in worldly units and the graphics window as (G_x, G_y) in pixels. Thus, the conversion is given by:

$$V_x = G_x \times \frac{V_w}{G_w}$$

$$V_y = G_y \times \frac{V_h}{G_h}$$

where G_w and G_h are the maximum width and height of the graphics window, respectively.

However, we know that the viewport is in 3D space, so it also has $V_z = d$ for every point on this viewport (in mathematical terms, called the *projection plane*).

Thus, for every pixel (G_x, G_y) on the graphics window, we can calculate the corresponding (V_x, V_y, V_z) of the viewport.

VI. RAYS

The rays are simply straight lines emanating from the origin O and passing through various points of the viewport, ultimately intersecting the objects in the scene.

We will represent a straight line using the parametric form:

$$P = O + t(V - O)$$

Here:

- P is any point on the line.
- O is the origin (where the camera is positioned).
- V is any point on the viewport.
- t is the parameter, where $-\infty < t < \infty$.

We vary t to obtain various points on the straight line.

Also note that:

- If $t < 0$, then P is behind the camera.
- If $0 \leq t \leq 1$, then P lies between the camera and the viewport.
- If $t > 1$, then P is in front of the viewport.

VII. SPHERE

The sphere is the simplest geometric object and will be used as an object in the scene.

We represent a sphere as:

$$\langle P - C, P - C \rangle = r^2$$

Here:

- P is a point on the surface of the sphere.
- C is the center of the sphere.
- r is the radius of the sphere.

VIII. INTERSECTION OF RAY AND SPHERE

Let P be the point of intersection; it is a common point to both the ray and the sphere.

The equation of the ray is:

$$P = O + t\vec{D}$$

where $\vec{D} = V - O$ is the direction of the ray.

The equation of the sphere is:

$$\langle P - C, P - C \rangle = r^2$$

Substituting the value of P into the equation of the sphere gives:

$$\langle O + t\vec{D} - C, O + t\vec{D} - C \rangle = r^2$$

$$\langle t\vec{D} + \vec{CO}, t\vec{D} + \vec{CO} \rangle = r^2$$

$$\langle t\vec{D} + \vec{CO}, t\vec{D} \rangle + \langle t\vec{D} + \vec{CO}, \vec{CO} \rangle = r^2$$

$$\langle t\vec{D}, t\vec{D} \rangle + \langle t\vec{D}, \vec{CO} \rangle + \langle \vec{CO}, t\vec{D} \rangle + \langle \vec{CO}, \vec{CO} \rangle = r^2$$

Finally, we arrive at the quadratic equation:

$$at^2 + bt + c = 0$$

where:

- $a = \|\vec{D}\|^2$
- $b = 2\langle \vec{D}, \vec{CO} \rangle$
- $c = \|\vec{CO}\|^2 - r^2$

This makes sense because a ray can intersect a sphere at either 0, 1, or 2 points.

We can now find the values of t using the quadratic formula:

$$t_{1,2} = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$$

After finding the values of t , we can substitute them back into the equation of the ray to find the point P . We should be careful to take the closest point P (the one which is visible to the camera) for which $t > 1$ (indicating that the objects are in front of the viewport).

IX. LIGHTING

We use white light and assume that light has the same intensity irrespective of the distance from the source.

A. Point Lights

Point lights can be completely described by:

- The coordinate Q of their location.
- The intensity of light.

Any vector from an arbitrary point P to the light source Q will be denoted by \vec{L} .

B. Directional Lights

Directional lights are used to represent light sources that are situated very far away compared to the distances between objects in the scene. Thus, rays of light coming from these sources are essentially parallel to each other. Therefore, \vec{L} is the same for every point P in the scene.

To model these lights, we need:

- A direction vector \vec{L} .
- The intensity of light.

C. Ambient Light

In real life, an object that does not receive direct light is not completely dark because reflected light illuminates it. To avoid treating every object as a light source, we will use ambient light, which essentially illuminates every object in the scene.

To model ambient light, we need:

- The intensity of light.

D. Lighting a Point

We define a type for light as follows:

```

1 (* type for light *)
2 (* kind of light is represented by k field -
   *)
3 (* p - for point source *)
4 (* d - for directional light *)
5 (* a - for ambient source *)
6 type light =
7 {
8   k : char;
9   i : float;
10  v : point option
11 }
```

The field v stores:

- The location of the point, in the case of a point source.
- The direction vector, in the case of a directional light.
- 'None', in the case of an ambient source.

The sum of all intensities of light should equal 1. If it does not, some points in the scene will be overexposed.

X. PHONG SHADING MODEL

A. Diffuse Reflections

Now, we will calculate the illumination of a point by both a point source and a directional source through diffuse reflection.

Consider a point P in the scene, with \vec{N} representing the normal vector at P and \vec{L} as the directional vector of the light source. For directional light, \vec{L} is provided. However, for a point light source, we need to compute \vec{L} based on the position of the light source and the coordinates of point P . Let l denote the width of the light (which is proportional to its intensity) and x denote the distance over which the light spreads. The line RS is perpendicular to \vec{L} , and Q is the point of intersection.

The ratio $\frac{l}{x}$ represents the drop in intensity after the light reflects off the surface. For instance, if the light is perpendicular to point P on the surface, then $\frac{l}{x} = 1$.

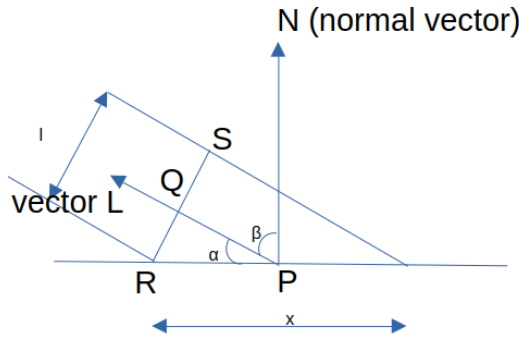


Fig. 6. Intensity after diffuse reflection

Consider the right triangle $\triangle PQR$. It is evident that

$$\angle PRQ = \beta$$

Using the definition of cosine in a right triangle, we have:

$$\cos \beta = \frac{l/2}{x/2} = \frac{l}{x}$$

We also know that the dot product of the vectors can be expressed as:

$$\langle \vec{L}, \vec{N} \rangle = \|\vec{L}\| \|\vec{N}\| \cos \beta.$$

Thus, we can rewrite $\cos \beta$ as:

$$\cos \beta = \frac{\langle \vec{L}, \vec{N} \rangle}{\|\vec{L}\| \|\vec{N}\|}.$$

If $\beta > 90^\circ$, the point is illuminated from the backside of the surface, and in this case, we treat the intensity as zero.

To obtain the intensity after reflection, we multiply the intensity of the light by the factor $\cos \beta$.

In a scene, a single point can be illuminated by multiple light sources, which may vary in type. We simply sum the intensities of all the light sources that reach the point and then

multiply the total intensity by the color of the point, which effectively adjusts the brightness.

Mathematically, this can be expressed as:

$$I_{\text{diffused}} = \sum I_i \frac{\langle \vec{L}_i, \vec{N} \rangle}{\|\vec{L}_i\| \|\vec{N}\|} \quad (1)$$

where

- I_{diffused} is the intensity due to diffuse reflection,
- I_i is the intensity of the i^{th} light source, which can be either a point source or a directional light,
- \vec{L}_i is the directional vector of the i^{th} light source.

B. Specular reflection

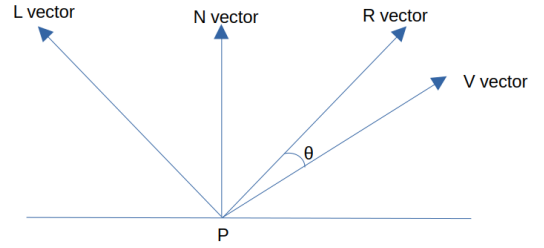


Fig. 7. Intensity after specular reflection

To model it, we use the simple fact that the angle made by the incident ray with the normal is equal to the angle made by the reflected ray with the normal.

Here

- \vec{L} is the directional vector of light pointing from point P on the surface to the point light source.
- \vec{R} is the reflected light, which makes the same angle with \vec{N} as \vec{L} .
- \vec{V} is the view vector from point P on the surface to the camera.
- \vec{N} is the normal vector.

First, we express \vec{R} in terms of \vec{L} and \vec{N} . We know that \vec{L} can be split into two components: one parallel to \vec{N} and another perpendicular to it.

$$L_{\parallel} \vec{N} = \langle \vec{L}, \hat{N} \rangle \hat{N}$$

Since

$$L_{\parallel} \vec{N} + L_{\perp} = \vec{L}$$

we have

$$\Rightarrow L_{\perp} \vec{N} = \vec{L} - \langle \vec{L}, \hat{N} \rangle \hat{N}$$

Thus, we can express \vec{R} as:

$$\vec{R} = L_{\parallel} \vec{N} - L_{\perp} \vec{N}$$

$$\Rightarrow \vec{R} = 2\langle \vec{L}, \hat{N} \rangle \hat{N} - \vec{L}$$

Now we determine the intensity due to specular reflection. For $\theta = 0$, the reflected vector and the view vector align,

so brightness is at its maximum. If $\theta = 90^\circ$, no component of the reflected ray is in the direction of the view vector, so brightness is zero in this case.

For $0 \leq \theta \leq 90^\circ$, we use $\cos \theta$ to model brightness. We define s to vary the shininess of different objects.

Then, the final expression for brightness due to specular reflection is:

$$I_{\text{specular}} = \sum I_i (\cos \theta)^s = \sum I_i \left(\frac{\langle \vec{R}_i, \vec{V}_i \rangle}{\|\vec{R}_i\| \|\vec{V}_i\|} \right)^s \quad (2)$$

Note that if $\theta > 90^\circ$, we might get a negative total intensity, which doesn't make sense. In those cases, we set the term to zero as before.

Combining both diffuse and specular brightness, along with ambient light, we get:

$$I_{\text{total}} = I_a + \sum I_i \left[\frac{\langle \vec{L}_i, \vec{N} \rangle}{\|\vec{L}_i\| \|\vec{N}\|} + \left(\frac{\langle \vec{R}_i, \vec{V} \rangle}{\|\vec{R}_i\| \|\vec{V}\|} \right)^s \right] \quad (3)$$

Here

- I_{total} is the total intensity at the point.
- I_a is the intensity of ambient light.
- I_i is the intensity of the i -th light, which could be a point source or directional.
- \vec{L}_i is the incident vector of the i -th light.
- \vec{R}_i is the reflected vector of the i -th light.
- \vec{N} is the normal vector at the point.
- \vec{V} is the view vector at the point.

C. Shadows

A shadow is the obstruction of light by one or more objects at a point.

What this means for us is that we need to check if a particular light source, such as a point source or directional light, intersects any other object in the scene before reaching the point for which we are calculating illumination due to that light. If that intersection occurs, then we do not include that light's intensity in the calculation of the total intensity; otherwise, we do.

In order to do that, we shoot a ray from the point of interest (for which we want to know whether it is shadowed) in the direction of the light source.

That ray is represented as

$$\vec{R} = P + t\vec{L}$$

Here

- P is the point of interest.
- \vec{L} is the ray of light from P toward the light source.
- t is the parameter. For a point source, $t \in [0^+, 1]$ (at $t = 1$, we reach the light source itself), and for directional light, $t \in [0^+, \infty)$. We know that a point on a sphere will definitely intersect itself, so practically we take 0.0001 instead of 0^+ . (We tried using ϵ_{float} and the results were not visually appealing, so to speak.)

Next, we find the closest intersection of the sphere with the ray in the suitable range. If a sphere exists, then only ambient light illuminates that point; otherwise, both specular and diffuse reflections also occur.

D. Mirror Reflection

After a ray from a camera intersects at a point on an object, we find the reflected ray at that point (including its color) and then shoot a reflected ray to determine if it intersects another object and calculate the resulting color. We can do this as many times as we like (but usually, after three reflections, we observe diminishing returns per computation). Eventually, we take a weighted sum of all the colors, utilizing the 'rfl' property of spheres.

XI. TOTAL INTENSITY OF LIGHT

Here, we combine all the effects, such as different kinds of reflection, into a single function, except for mirror reflection, which is handled through the ray tracing code itself.

```

1  let rec til_inner normal p o s light_l
    sphere_l intensity =
2  match light_l with
3  {k;i;v}::t ->
4    let c_intensity = ref 0. in
5    let tmax = ref 0. in
6    if k = 'a' then c_intensity := i +. !
      c_intensity
7    else
8      begin
9        let l = ref {x=0.;y=0.;z=0.} in
10       (if k = 'p' then
11         (l := (sub3 (bare v) p);
12          tmax := 1.)
13       else
14         (l := bare v;
15          tmax := infinity));
16       let shadow_s,_ = closest_sphere p !l
17         0.0001 !tmax sphere_l in
18       match shadow_s with
19       Some ss -> ()
20       |_ ->
21         let unitnormal = scale (1. /. (norm
22           normal)) normal in
23         c_intensity := specular_i o p
24           unitnormal !l i s !c_intensity;
25         c_intensity := diffuse_i normal !l i
26           !c_intensity
27       end;
28     til_inner normal p o s t sphere_l (
29       intensity +. !c_intensity)
30   | _ -> intensity;;
31
32 let til normal p o s light_l sphere_l =
33   til_inner normal p o s light_l sphere_l
34   0.;;
```

XII. RAY TRACING CODE

```

1  let rec rtx_inner o d tmin tmax sl ll limit =
2  let objintensity = ref 0. in
3  let otherintensity = ref 0. in
4  let s, t = closest_sphere o d tmin tmax sl
    in
```

```

5 begin
6   match s, t with
7     Some sphere, Some t_parameter ->
8       let p = add3 o (scale t_parameter d) in
9       let normal = sub3 p sphere.c in
10      objintensity := til normal p o sphere.s
11      ll sl;
12      if sphere.rfl > 0. && limit > 0 then
13        begin
14          let reflected = reflected_ray (scale
15            (1. /. norm normal) normal) (
16            scale (- 1.) d) in
17          otherintensity := extract (rtx_inner
18            p reflected 0.001 infinity sl
19            ll (limit-1));
20          objintensity := !objintensity *. (1.
21            -. sphere.rfl) +. sphere.rfl *.
22            !otherintensity;
23        end
24      | _, _ -> ()
25    end;
26  s, !objintensity;;
27
28 let rtx o d tmin tmax sl ll = sphere_color (
29   rtx_inner o d tmin tmax sl ll 3)

```

Now there two ways to start the ray tracing:

In functional style:

```

1 let create_tuple min_x max_x min_y max_y =
2   let rec range start stop a =
3     if start > stop then a
4     else range (start + 1) stop (start::a)
5   in
6   let list_x = range min_x max_x [] in
7   let list_y = range min_y max_y [] in
8   let a l x = List.map (fun y -> (x, y)) l
9   in
10  List.flatten (
11    List.map (a list_y) list_x
12  )
13
14 let shader o tmin tmax ls ll (x, y) =
15   let v = g_to_viewport x y in
16   let d = sub3 v o in
17   let color = rtx o d tmin tmax ls
18   ll in
19   plotc x y color
20
21 let coordinates = create_tuple (-gw/2) (gw/2)
22   (-gh/2) (gh/2) in
23 List.iter (shader o 1. infinity ls ll )
24   coordinates

```

And in imperative style:

```

1 for y = - gh/2 to gh/2 do
2   for x = - gw/2 to gw/2 do
3     let v = g_to_viewport x y in
4     let d = sub3 v o in
5     let color = rtx o d 1. infinity ls ll in
6     plotc x y color
7   done;
8 done

```

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