Appendix

Results of mass loss

Table 2. Dust and wind properties.

model	$n_d/n_{\rm H}$	$\dot{M}$ ($M_\odot$/yr)	$f_{\rm Mg}$	a_{gr} (μm)
st28gm06n052	$3 \cdot 10^{-16}$	$5 \cdot 10^{-6}$	0.5	0.8
st28gm05n033	$3 \cdot 10^{-15}$	$5 \cdot 10^{-8}$	0.15	0.2

Notes. Listed here are the assumed seed particle abundance $n_d/n_{\rm H}$, and the resulting temporal means of the mass-loss rate $\dot{M}$, the fraction of Mg condensed into grains $f_{\rm Mg}$, and the grain radius a_{gr} at the outer boundary. When forming Mg_2SiO_4 grains in a gas of solar composition, the abundance of Mg is the limiting factor, since that element will be used up first. In the models described here, however, $f_{\rm Mg}$ is well below its maximum value of 1.

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