Table of Contents

- 1. StartUp
- 2. Getting Help
- 3. Vectors
- 3.1. Some Basics
- 3.2. Construction of STL Vectors
- 3.2.1. Referencing, memory allocation, indexing, slicing
- 3.2.2. Vector arithmetic
- 4. File I/O
- 4.1. Opening and Closing a CR Data File
- 4.2. Setting and retrieving metadata (parameters)
- 4.3. Reading in Data
- 5. Fourier Transforms (FFT)
- 6. Basic Plotting
- 6.1. Mathgl
- 6.2. Matplotlib
- 7. Coordinates
- 8. Appendix: Listing of all Functions:

pycrtools TUTORIAL

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1. StartUp

First one needs to load the library. This can be done by "from hftools import *" which makes all the c++ functions and their python wrappers available.

In addition there are some helpful definitions in python as well, which are stored in "pycrtools.py". This file actually imports hittools for you. Hence, all you actually need to do is (make sure the file is in your search path)

>>> from pycrtools import *

The most convenient way is probably to define an alias in your .bashr or .profile files, like

alias pycr='/sw/bin/python2.6 -i LOFARSOFT/src/CR-Tools/implement/Pypeline/pycrinit.py'

then you can simply start the pyrtools with 'pycr'.

The file pycrinit.py also allows you to read in a second file to be executed, like this one here. So, you can run the tutorial with

pycr -i tutorial.py

from the UNIX prompt.

2. Getting Help

==========

The pycrtools have a built-in help system, which can be accessed with:

>>> help()

To get information on a specific function or method type help(func). This will essentially print the docstring of the python object and list all its methods. Hence, help will work on any decently implemented python object, including the standard ones.

For example,

>>> help(IntVec)

will give documentation on the integer vector, while

>>> help(IntVec.sin)

will give the documentation on the "sin" method associated with it.

For a listing of all available functions in the pycrtools type

>>> help(all).

3. Vectors

========

3.1. Some Basics

The fundamental data structure we use is a standard c++ vector defined in the c++ standard template library (STL). This is wrapped and exposed to python using the Boost Python system.

(NB: Unfortunately different systems provide different python data structures. Hence a function exposed to python with SWIG or SIP is not directly able to accept a BOOST PYTHON wrapped vector as input or vice versa. If you want to do this you have to provide extra conversion routines.)

In line with the basic python philosophy, vectors are passed as references. Since we are working with large sets of data processing time is as important as convenience. Hence the basic principle is that we try to avoid copying large chunks of data as much as possible.

The basic functions operating on the data in c++ either take STL iterators as inputs (i.e. pointers to the begin and end of the memory where the data in the STL vector is stored) or casa::Vectors, which are created with shared memory (i.e. their physical memory is the same as that of the STL vector).

For that reason MEMORY ALLOCATION is done almost exculsively in the Python layer. The fast majority of c++ functions is not even able to allocate or free any memory. This allows for very efficient memory management and processing, but also means that the user is responsible for providing properly sized vectors as inputs AND output vectors. I.e.: you need to know beforehand what sized vector you expect in return!

That may be annoying, but forces you to think carefully about how to use memory. For example, if there is a processing done multiple times using a scratch vector of fixed size, you reuse the same vector over and over again, thus avoiding a lot of unnecessary allocation and deallocation of memory and creation of vectors.

Also, we are currently not supporting multidimensional data structures. Mulit-dimensionally data is simply written sequentially into the memory - you need to know how your data is organized. On the other hand, in many case you can then easily process an entire matrix at once.

3.2. Construction of STL Vectors

A number of vector types are provided: bool, int, float, complex, and str.

To creat a vector most efficently, use the original vector constructors:

```
- BoolVec()
- IntVec()
- FloatVec()
- ComplexVec()
- StringVec()

e.g.
>>> v=FloatVec()
>>> v
Vec(0)=[]
```

will create a floating point vector of size 0.

The vector can be filled with a python list or tuple, by using the extend attrib ute:

```
>>> v.extend([1,2,3,4])
>>> v
Vec(4)=[1.0,2.0,3.0,4.0]
```

Note, that python has automatically converted the integers into floats, since the STL vector does not allow any automatic typing.

The stl vector can be converted back to a python list by using the python list creator:

```
>>> list(v)
[1.0, 2.0, 3.0, 4.0]
```

However, the basic Boost Python STL vector constructor takes no arguments and is a bit cumbersome to use in the long run. Here we provide a wrapper function that is useful for interactive use.

Usage:

Vector(Type) - will create an empty vector of type "Type", where Type is

a basic Python type, i.e. bool, int, float, complex, str.

Vector(Type, size) - will create an vector of type "Type", with length "size".

Vector(Type,size,fill) - will create an vector of type "Type", with length
"size" and initialized with the value "fill"

Vector([1,2,3,...]) or Vector((1,2,3,...)) - if a list or a tuple is provided as first argument then a vector is created of the type of the first element in the list or tuple (here an integer) and filled with the contents of the list or tuple.

So, what we will now use is:

```
>>>v=Vector([1.,2,3,4])
>>> v
Vec(4)=[1.0,2.0,3.0,4.0]
```

Note, that size and fill take precedence over the list and tuple input. Hence if you create a vector with Vector([1,2,3],size=2) it will contain only [1,2]. Vector([1,2,3],size=2,fill=4) will give [4,4].

As the latest addition some simple support for multiple dimensions has been implemented, using the methods:

```
vector.setDim([n1,n2,..])
vector.getDim()
vector.elem(n)
```

This will be described later

3.2.1. Referencing, memory allocation, indexing, slicing

Following basic python rules, the STL vector will persist in memory as long as there is a python reference to it. If you destroy v also the c++ memory will disappear. Hence, if you keep a pointer to the vector in c++ and try to work on it after the python object was destroyed, your programme may crash. That's why by default memory management is done ONLY on one side, namely the python side!

To illustrate how Python deals with references, consider the following example:

```
>>> x=v
>>> x[0]=3
>>> v
Vec(4)=[3.0,2.0,3.0,4.0]
```

Hence, the new python object x is actually a reference to the same c++ vector that was created in v. Modifying elements in x modifies elements in v. If you destroy v or x, the vector will not be destroyed, since there is still a reference to it left. Only if you destroy x and v the memory will be freed.

As note above, this vector is subscriptable and sliceable, using the

```
standard python syntax.
>>> v[1:3]
Vec(2) = [2.0, 3.0]
We can also resize vectors and change their memory allocation:
>>> v1=Vector([0.0,1,2,3,4,5])
>>> v2=Vector(float,len(v1),2.0)
>>>v1
Vec(6) = [0.0, 1.0, 2.0, 3.0, 4.0, 5.0]
>>>v2
Vec(6) = [2.0, 2.0, 2.0, 2.0, 2.0, 2.0]
With the resize attribute you allocate new memory while keeping the
data. It is not guaranteed that the new memory actually occupies the
same physical space.
>>> v2.resize(8)
>>> v2
Vec(8) = [2.0, 2.0, 2.0, 2.0, 2.0, 2.0, 0.0, 0.0]
Resize a vector and fill with non-zero values:
>>> v2.resize(10,-1)
>>> v2
Resize a vector to same size as another vector:
>>> v2.resize(v1)
>>> v2
Vec(6) = [2.0, 2.0, 2.0, 2.0, 2.0, 2.0]
Make a new vector of same size and type as the original one:
>>> v3=v2.new()
>>> v3
Vec(6) = [0.0, 0.0, 0.0, 0.0, 0.0, 0.0]
Fill a vector with values
>>> v3.fill(-2)
Vec(6) = [-2.0, -2.0, -2.0, -2.0, -2.0, -2.0]
3.2.2. Vector arithmetic
The vectors have a number of mathematical functions attached to
them. A full list can be seen by typing
>>> dir(Vector(float))
['__add__', '__class__', '__contains__', '__delattr__', '__delitem__', '__dict__
', '__div__', '__doc__', '__format__', '__getattribute__', '__getitem__', '__has
h__', '__iadd__', '__idiv__', '__imul__', '__init__', '__instance_size__', '__is
ub__', '__iter__', '__len__', '__module__', '__mul__', '__new__', '__reduce__',
'__reduce_ex__', '__repr__', '__setattr__', '__setitem__', '__sizeof__', '__str_
_', '__sub__', '__subclasshook__', '__weakref__', 'abs', 'acos', 'add', 'addadd'
```

, 'addaddconv', 'append', 'asin', 'atan', 'ceil', 'convert', 'copy', 'cos', 'cos

h', 'div', 'divadd', 'divaddconv', 'downsample', 'exp', 'extend', 'extendflat', 'fft', 'fill', 'findgreaterequal', 'findgreaterequalabs', 'findgreaterthan', 'fi ndgreaterthanabs', 'findlessequal', 'findlessequalabs', 'findlessthan', 'findles sthanabs', 'findlowerbound', 'floor', 'iadd', 'idiv', 'imul', 'isub', 'log', 'lo g10', 'mean', 'median', 'mul', 'muladd', 'muladdconv', 'new', 'norm', 'normalize ', 'resize', 'sin', 'sinh', 'sort', 'sortmedian', 'sqrt', 'square', 'stddev', 's ub', 'subadd', 'subaddconv', 'sum', 'tan', 'tanh'] Some of the basic arithmetic is available in an intuitve way. You can add a scalar to a vector by >>> v1+3 Vec(6) = [3.0, 4.0, 5.0, 6.0, 7.0, 8.0]This will actually create a new vector (and destroy it right away, since no reference was kept). The original vector is unchanged. You can also add two vectors: >>>v1+v2 Vec(6) = [2.0, 3.0, 4.0, 5.0, 6.0, 7.0]

In order to change the vector, you can use the "in place" operators +=, -=, /=, *=:

Adding a vector in place: >>>v1+=v2; >>>v1 Vec(6)=[2.0,3.0,4.0,5.0,6.0,7.0]

now v1 was actually modified such that v2 was added to the content of v1 and the results is stored in v1.

Similarly you can do

- v1-=v2
- -v1*=v2
- -v1/=v2

Here is a list of functions you can use

Mean: v1.mean() = 4.5 Median: v1.median() = 5.0 Summing: v1.sum() = 27.0 Standard Deviation: v1.stddev() = 1.87082869339

- 4. File I/O
- 4.1. Opening and Closing a CR Data File

Let's see how we can open a file. First define a filename, e.g.:

- >>> LOFARSOFT=os.environ["LOFARSOFT"]
- >>> filename=LOFARSOFT+"/data/lopes/example.event"
- >>> #filename=LOFARSOFT+"/data/lofar/rw_20080701_162002_0109.h5"
- > '/Users/falcke/LOFAR/usg/data/lopes/sample.event'

We can create a new file object, using the "crfile" class, which is an

interface to the LOFAR CRTOOLS datareader class and was defined in pycrtools.py.

```
The following will open a data file:
>>> file=crfile(filename)
>>> file.set("Blocksize",1024)
[hftools.tmp.cc,3758]: Opening LOPES File=/Users/falcke/LOFAR/usg/data/lopes/exa
mple.event
[DataReader::summary]
 -- blocksize ..... = 65536
-- FFT length ..... = 32769
-- file streams connected? 0
 -- shape(adc2voltage) ... = [65536, 8]
 -- shape(fft2calfft) .... = [32769, 8]
 -- nof. antennas ..... = 8
 -- nof. selected antennas = 8
 -- nof. selected channels = 32769
 -- shape(fx) ..... = [65536, 8]
 -- shape(voltage) ..... = [65536, 8]
 -- shape(fft) ..... = [32769, 8]
 -- shape(calfft) ..... = [32769, 8]
The crfile class stores a pointer to the data reader object, which can
be retrieved with
>>> file.iptr
> 17209856
and similarly the filename with
>>> file.filename
> '/Users/falcke/LOFAR/usg/data/lopes/example.event'
The file will be automatically closed (and the DataReader object be
destroyed), whenever the crfile object is deleted, e.g. with
"file=0". One can also explicitly close the file with "file.close()".
4.2. Setting and retrieving metadata (parameters)
Now we need to access the meta data in the file. This is done using a
single method "get". This method actually calls the function
"hFileGetParameter" defined in the c++ code.
Which observatory did we actually use?
>>> obsname=file.get("Observatory")
> 'LOPES'
There are more keywords, of course. A list of implemented parameters
we can access is obtained by
>>> keywords=file.get("help")
> hFileGetParameter - available keywords: nofAntennas, nofSelectedChannels, nofS
electedAntennas, nofBaselines, block, blocksize, stride, fftLength, nyquistZone,
 sampleInterval, referenceTime, sampleFrequency, antennas, selectedAntennas, sel
ectedChannels, positions, increment, frequencyValues, frequencyRange, Date, Obse
rvatory, Filesize, dDate, presync, TL, LTL, EventClass, SampleFreq, StartSample,
AntennaIDs, help
```

Note, that the results are returned as PythonObjects. Hence, this

```
makes use of the power of python with automatic typing. For, example
>>> file.get("frequencyRange")
> Vec(2) = [40000000.0, 80000000.0]
actually returns a vector.
Here no difference is made where the data comes frome. The keyword
Obervatory accesses the header record in the data file while the
frequencRange accesses a method of the DataReader.
Now we will define a number of useful variables that contain essential
parameters that we later will use.
>>> obsdate =file.get("Date"); p_("obsdate")
>>> filesize =file.get("Filesize"); p_("filesize")
>>> blocksize =file.get("blocksize"); p_("blocksize")
>>> nAntennas =file.get("nofAntennas"); p_("nAntennas")
>>> antennas =file.get("antennas"); p_("antennas")
>>> antennaIDs=file.get("AntennaIDs"); p_("antennaIDs")
>>> selectedAntennas=file.get("selectedAntennas"); p_("selectedAntennas")
>>> nofSelectedAntennas=file.get("nofSelectedAntennas"); p_("nofSelectedAntennas
")
>>> fftlength =file.get("fftLength"); p_("fftlength")
>>> sampleFrequency =file.get("sampleFrequency"); p_("sampleFrequency")
>>> nBlocks=filesize/blocksize; p_("nBlocks")
> obsdate = 1067339149
> filesize = 65536
> blocksize = 1024
> nAntennas = 8
> antennas = Vec(8)=[0,1,2,3,4,5,6,7]
> antennaIDs = Vec(8)=[10101,10102,10201,10202,20101,20102,20201,20202]
> selectedAntennas = Vec(8) = [0,1,2,3,4,5,6,7]
> nofSelectedAntennas = 8
> fftlength = 513
> sampleFrequency = 80000000.0
> blocksize = 1024
We can also change parameters in a very similar fashion, using the
"set" method, which is an implementation of the "hFileSetParameter"
function. E.g. changing the blocksize we already did before. This is
simply
>>> blocksize=1024
>>> file.set("Blocksize",blocksize)
again the list of implemented keywords is visible with using
>>> file.set("help",0)
> hFileSetParameter - available keywords: Blocksize, StartBlock,
  Block, Stride, SampleOffset, NyquistZone, ReferenceTime,
  SampleFrequency, Shift, SelectedAntennas, help
The set method actually returns the crfile object. Hence you can
append multiple set commands after each other.
>>> file.set("Block",2).set("SelectedAntennas",[0,2])
> crfile</Users/falcke/LOFAR/usg/data/lopes/example.event>
Note, that we have now reduced the number of antennas to two: namely
antenna 0 and 2 and the number of selected antennas is
>>>file.get("nofSelectedAntennas")
 2
```

```
However, now we want to work on all antennas again:
>>> file.set("SelectedAntennas",range(nAntennas))
4.3. Reading in Data
```

The next step is to actually read in data. This is done with the read method (accessing "hFileRead"). The data is read flatly into a 1D vector. This is also true if mutliple antennas are read out at once. Here simply the data from the antennas follow each other.

Also, by default memory allocation of the vectors has to be done in python before calling any of the functions. This improves speed and efficiency, but requires one to programme carefully and to understand the data structrue.

First we create a FloatVec, which is BoostPython wrapped standard

```
(STL) vector of doubles.
>>> fxdata=Vector()
and resize it to the size we need
>>> fxdata.setDim([nofSelectedAntennas,blocksize])
This is now a large vector filled with zeros.
>>> fxdata
Now we can read in the raw time series data, using "file.read" and a
keyword. Currently implemented are the data fields "Fx", "Voltage",
"FFT", "CalFFT", "Time", "Frequency".
So, let us read in the raw time series data, i.e. the electric field
in the antenna as digitized by the ADC. This is provided by the
keyword "Fx" (means f(x) ).
>>> file.read("Fx",fxdata)
>>>fxdata
 Vec(2048) = [212.0, -278.0, 196.0, 236.0, 4.0, -94.0, -192.0, 208.0, -66.0, -62.0, ...]
and voila the vector is filled with time series data from the data file..
Access the various antennas through slicing
>>> ant1data=fxdata[0:blocksize]
>>> ant2data=fxdata.elem(1)
>>> ant2data
Vec(1024) = [-94.0, -165.0, -6.0, 35.0, -310.0, -23.0, -128.0, 97.0, 239.0, 289.0, ...]
which makes a copy of the data vector if used in this way, while
>>> ant1data[0:3]=[0,1,2];
>>> antldata;
Vec(2046) = [0.0, 1.0, 2.0, -94.0, -192.0, 208.0, -66.0, -62.0, -157.0, -120.0, ...]
actually modifies the original data vector.
To get the x -Axis we create a second vector
>>> timedata=Vector(float,blocksize)
>>> file.read("Time",timedata)
> Vec(1024)=[-0.000384,-0.0003839875,-0.000383975,-0.0003839625,-0.00038395,-0.0
003839375,-0.000383925,-0.0003839125,-0.0003839,-0.0003838875,...]
```

This is the time relative to the trigger in seconds. So, let's have

```
that in microseconds, by multiplying with one million.
>>> timedata *= 10**6
>>> timedata
Vec(1024)=[-384.0,-383.9875,-383.975,-383.9625,-383.95,-383.9375,-383.925,-383.9
125, -383.9, -383.8875, ...]
We do the same now for the frequency axis, which we convert to MHz. As
length we have to take the length of the Fourier transformed time
block (which is blocksize/2+1).
>>> freqdata=Vector(float,fftlength)
>>> file.read("Frequency", freqdata)
>>> freqdata/=10**6
>>> freqdata
Vec(513)=[40.0,40.078125,40.15625,40.234375,40.3125,40.390625,40.46875,40.546875
,40.625,40.703125,...]
```

5. Fourier Transforms (FFT)

We can make a FFT of a float vector. This function will only return the non-redundant part of the FFT (i.e., just one half). Again we need to provide a properly sized output vector (input length/2+1). We also have to specify as a second parameter in which NyquistZone the data was take.

Nyquist sampling means that one needs, for example, 200 MHz sampling rate to digitize a bandwidth of 100 MHz. The first Nyquist zone is

```
then 0-100 MHz, and the second is 100-200 MHz.
So, let's do the transform:
>>> fftdata=Vector(complex,fftlength)
>>> fxdata[0:blocksize].fft(fftdata,1)
>>> fftdata
Vec(513)=[(6078+0j),(99.3936739874-28.663986893j),(-93.6321366929-4.95059820124j
),(82.9590664565+28.8729314743j),(-83.6744655239+4.46573054789j),(169.1861864-61
.2949652607j),(-118.623662378+53.2694320202j),(75.764787806-74.6606191354j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.2694320202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),(-118.623662378+53.269420202)j),j),(-118.623662378+53.269420202)j),(-118.623662378+53620202)j),
5.629434646+29.4373842905j),(98.0844400537-16.0574421952j),...]
Here we have used the fft method of the float vector, which is just a
call to the stand-alone function hFFT defined in hftools.cc.
to get the power, we have to square the complex data and convert it to
floats. This can be done using the complex function "norm"
>>> spectrum=Vector(float,fftlength)
>>> fftdata.norm(spectrum)
We can now try to calulcate the average spectum of the data set for
one antenna, by looping over all blocks.
>>> avspectrum=Vector(float)
>>> avspectrum.setDim([nofSelectedAntennas,fftlength])
>>> fftall=Vector(complex)
>>> fftall.setDim([nofSelectedAntennas,fftlength])
>>> for block in range(nBlocks):
>>>
                 print block, ", ",
                  file.set("Block", block).read("FFT", fftall)
>>>
                  fftall.spectralpower(avspectrum)
>>>
(NOTE: This doesn't work yet due to a bug in file.read("FFT",...) !!!!)
```

6. Basic Plotting

In order to plot the data, we can use external packages. Two packages are being provided here: matplotlib and mathgl. The former is specifically designed for python and thus slightly easier to use interactively. Since version 0.99 it is supposed to bea capable of 3D plots (at the time of writing we use 0.98). Mathgl is perhaps a little faster and convenient for real time applications and hence is used in our GUI programming.

6.1. Mathgl

Here is a simple example on how to use mathgl code (here without a widget)

```
widget)
from mathgl import *
width=800
height=600
size=1024
gr=mglGraph(mglGraphPS, width, height)
y=mglData(size)
y.Modify("cos(2*pi*x)")
x=mglData(size)
x.Modify("x*1024");
ymax=y.Maximal()
ymin=y.Minimal()
gr.Clf()
gr.SetRanges(0,0.5,ymin,ymax)
gr.Axis("xy")
gr.Title("Test Plot x")
gr.Label("x","x-Axis",1)
gr.Label("y","y-Axis",1)
gr.Plot(x,y);
gr.WriteEPS("test-y.eps","Test Plot")
6.2. Matplotlib
Now we import matplotlib
>>> import matplotlib.pyplot as plt
>>> print "\n!! A plot window should pop up somehwere (in the background?) !!"
>>> plt.show()
>>> plt.title("Average Spectrum for Two Antennas")
>>> plt.plot(freqdata,avspectrum.elem(0))
>>> plt.plot(freqdata,avspectrum.elem(1))
>>> plt.ylabel("Power of Electric Field [ADC counts]$^2$")
>>> plt.xlabel("Frequency [MHz]")
and a window should pop up. (NB: At least on a Mac the window likes
to stubbornly hide behind other windows, so search your screen
carefully if no window pops up.)
to plot time series data, use:
plt.clf()
plt.plot(timedata,fxdata[0:blocksize])
plt.ylabel("Electric Field [ADC counts]")
plt.xlabel("Time [$\mu$s]")
```

7. Coordinates

We also have access to a few functions dealing with astronomical coordinates. Assume, we have a source at an Azmiut/Elevation position of (178 Degree, 28 Degree) and we want to convert that into Cartesian coordinates (which, for example, is required by our beamformer).

We first turn this into std vector and create a vector that is supposed to hold the Cartesian coordinates. Note that the AZEL vector is actually AzElRadius, where we need to set the radius to unity.

```
>>> azel=FloatVec()
>>> azel.extend((178,28,1))
>>> cartesian=azel.new()
>>> azel
Vec(3) = [178.0, 28.0, 1.0]
>>> cartesian
Vec(3) = [0.0, 0.0, 0.0]
```

We then do the conversion, using

>>> hCoordinateConvert(azel,CoordinateTypes.AzElRadius,cartesian,CoordinateTypes .Cartesian, True)

yielding the following output vector:

```
>>> cartesian
```

Vec(3) = [0.0308144266055, -0.882409725042, 0.469471562786]

>>> >>>

8. Appendix: Listing of all Functions:

SECTION: Administrative Vector Function

```
______
```

hFill(vec, fill_value) - Fills a vector with a constant value. hNew(vec) - Make and return a new vector of the same s

ize and type as the input vector.

hConvert(vec1, vec2) - Copies and converts a vector to a vector o

f another type.

- Copies a vector to another one. hCopy(vec, outvec)

SECTION: Math Function

- Returns the squared value of the parameter square(val)

hPhase(frequency, time) - Returns the interferometer phase in radian s for a given frequency and time.

funcGaussian(x, sigma, mu) - Implementation of the Gauss function.

hExp(vec) - Take the exp of all the elements in the ve

hExp(vec, vecout)

- Take the exp of all the elements in the ve

ctor and return results in a second vector.

hLog(vec) - Take the log of all the elements in the ve

ctor.

ctor.

hLog(vec, vecout) - Take the log of all the elements in the ve ctor and return results in a second vector.

hLog10(vec) - Take the log10 of all the elements in the

vector.

```
hLog10(vec, vecout)
                                    - Take the log10 of all the elements in the
vector and return results in a second vector.
                                     - Take the sin of all the elements in the ve
hSin(vec)
ctor.
hSin(vec, vecout)
                                     - Take the sin of all the elements in the ve
ctor and return results in a second vector.
                                     - Take the sinh of all the elements in the v
hSinh(vec)
ector.
hSinh(vec, vecout)
                                     - Take the sinh of all the elements in the v
ector and return results in a second vector.
                                     - Take the sqrt of all the elements in the \boldsymbol{v}
hSqrt(vec)
ector.
                                     - Take the sqrt of all the elements in the v
hSqrt(vec, vecout)
ector and return results in a second vector.
                                    - Take the square of all the elements in the
hSquare(vec)
vector.
hSquare(vec, vecout)
                                     - Take the square of all the elements in the
 vector and return results in a second vector.
                                     - Take the tan of all the elements in the ve
hTan(vec)
ctor.
                                     - Take the tan of all the elements in the ve
hTan(vec, vecout)
ctor and return results in a second vector.
                                     - Take the tanh of all the elements in the v
hTanh(vec)
ector.
                                     - Take the tanh of all the elements in the \boldsymbol{v}
hTanh(vec, vecout)
ector and return results in a second vector.
hAbs(vec)
                                     - Take the abs of all the elements in the ve
ctor.
hAbs(vec, vecout)
                                     - Take the abs of all the elements in the ve
ctor and return results in a second vector.
                                    - Take the cos of all the elements in the ve
hCos(vec)
ctor.
                                     - Take the cos of all the elements in the ve
hCos(vec, vecout)
ctor and return results in a second vector.
                                     - Take the cosh of all the elements in the v
hCosh(vec)
ector.
                                     - Take the cosh of all the elements in the \boldsymbol{v}
hCosh(vec, vecout)
ector and return results in a second vector.
                                     - Take the ceil of all the elements in the v
hCeil(vec)
ector.
                                     - Take the ceil of all the elements in the v
hCeil(vec, vecout)
ector and return results in a second vector.
                                     - Take the floor of all the elements in the
hFloor(vec)
vector.
                                     - Take the floor of all the elements in the
hFloor(vec, vecout)
vector and return results in a second vector.
                                     - Take the acos of all the elements in the v
hAcos(vec)
ector.
hAcos(vec, vecout)
                                     - Take the acos of all the elements in the v
ector and return results in a second vector.
hAsin(vec)
                                     - Take the asin of all the elements in the v
ector.
                                     - Take the asin of all the elements in the v
hAsin(vec, vecout)
ector and return results in a second vector.
hAtan(vec)
                                     - Take the atan of all the elements in the v
ector.
hAtan(vec, vecout)
                                    - Take the atan of all the elements in the v
ector and return results in a second vector.
hiSub(vec1, vec2)
                                     - Performs a Sub between the two vectors, wh
```

ich is returned in the first vector. If the second vector is shorter it will be

```
applied multiple times.
hiSub(vec1, val)
                                    - Performs a Sub between the vector and a sc
alar (applied to each element), which is returned in the first vector.
hSub(vec1, vec2, vec3)
                                    - Performs a Sub between the two vectors, wh
ich is returned in the third vector.
hSubAdd(vec1, vec2, vec3)
                                    - Performs a Sub between the two vectors, an
d adds the result to the output (third) vector.
hSubAddConv(vec1, vec2, vec3)
                                   - Performs a Sub between the two vectors, an
d adds the result to the output (third) vector - automatic casting is done.
hSub(vec1, val, vec2)
                                    - Performs a Sub between the vector and a sc
alar, where the result is returned in the second vector.
hiMul(vec1, vec2)
                                    - Performs a Mul between the two vectors, wh
ich is returned in the first vector. If the second vector is shorter it will be
applied multiple times.
hiMul(vec1, val)
                                    - Performs a Mul between the vector and a sc
alar (applied to each element), which is returned in the first vector.
hMul(vec1, vec2, vec3)
                                   - Performs a Mul between the two vectors, wh
ich is returned in the third vector.
hMulAdd(vec1, vec2, vec3)
                                    - Performs a Mul between the two vectors, an
d adds the result to the output (third) vector.
hMulAddConv(vec1, vec2, vec3)
                                    - Performs a Mul between the two vectors, an
d adds the result to the output (third) vector - automatic casting is done.
hMul(vec1, val, vec2)
                                    - Performs a Mul between the vector and a sc
alar, where the result is returned in the second vector.
hiAdd(vec1, vec2)
                                    - Performs a Add between the two vectors, wh
ich is returned in the first vector. If the second vector is shorter it will be
applied multiple times.
                                    - Performs a Add between the vector and a sc
hiAdd(vec1, val)
alar (applied to each element), which is returned in the first vector.
hAdd(vec1, vec2, vec3)
                                    - Performs a Add between the two vectors, wh
ich is returned in the third vector.
hAddAdd(vec1, vec2, vec3)
                                    - Performs a Add between the two vectors, an
d adds the result to the output (third) vector.
hAddAddConv(vec1, vec2, vec3)
                                  - Performs a Add between the two vectors, an
d adds the result to the output (third) vector - automatic casting is done.
                                    - Performs a Add between the vector and a sc
hAdd(vec1, val, vec2)
alar, where the result is returned in the second vector.
hiDiv(vec1, vec2)
                                    - Performs a Div between the two vectors, wh
ich is returned in the first vector. If the second vector is shorter it will be
applied multiple times.
                                    - Performs a Div between the vector and a sc
hiDiv(vec1, val)
alar (applied to each element), which is returned in the first vector.
hDiv(vec1, vec2, vec3)
                                    - Performs a Div between the two vectors, wh
ich is returned in the third vector.
hDivAdd(vec1, vec2, vec3)
                                    - Performs a Div between the two vectors, an
d adds the result to the output (third) vector.
hDivAddConv(vec1, vec2, vec3)
                                    - Performs a Div between the two vectors, an
d adds the result to the output (third) vector - automatic casting is done.
hDiv(vec1, val, vec2)
                                    - Performs a Div between the vector and a sc
alar, where the result is returned in the second vector.
hConj(vec)
                                    - Calculate the complex conjugate of all ele
ments in the complex vector.
hCrossCorrelateComplex(vec1, vec2) - Multiplies the elements of the first vecto
r with the complex conjugate of the elements in the second and returns the resul
ts in the first.
hReal(vec, vecout)
                                    - Take the real of all the elements in the c
omplex vector and return results in a float vector.
                                    - Take the arg of all the elements in the co
mplex vector and return results in a float vector.
hImag(vec, vecout)
                                    - Take the imag of all the elements in the c
```

```
omplex vector and return results in a float vector.
                                    - Take the norm of all the elements in the c
hNorm(vec, vecout)
omplex vector and return results in a float vector.
hNegate(vec)
                                    - Multiplies each element in the vector with
 -1 in place, i.e. the input vector is also the output vector.
                                    - Performs a sum over the values in a vector
hSum(vec)
 and returns the value.
hNorm(vec)
                                    - Returns the lengths or norm of a vector (i
.e. Sart(Sum i(xi*+2))).
hNormalize(vec)
                                    - Normalizes a vector to length unity.
hMean(vec)
                                    - Returns the mean value of all elements in
a vector.
                                    - Sorts a vector in place.
hSort(vec)
hSortMedian(vec)
                                    - Sorts a vector in place and returns the me
dian value of the elements.
hMedian(vec)
                                    - Returns the median value of the elements.
hStdDev(vec, mean)
                                    - Calculates the standard deviation around a
 mean value.
                                    - Calculates the standard deviation of a vec
hStdDev(vec)
tor of values.
hFindLessEqual(vec, threshold, vecout) - Find the samples that are LessEqual a c
ertain threshold value and returns the number of samples found and the positions
 of the samples in a second vector.
hFindLessEqualAbs(vec, threshold, vecout) - Find the samples whose absolute valu
es are LessEqual a certain threshold value and returns the number of samples fou
nd and the positions of the samples in a second vector.
hFindGreaterThan(vec, threshold, vecout) - Find the samples that are GreaterThan
 a certain threshold value and returns the number of samples found and the posit
ions of the samples in a second vector.
hFindGreaterThanAbs(vec, threshold, vecout) - Find the samples whose absolute va
lues are GreaterThan a certain threshold value and returns the number of samples
 found and the positions of the samples in a second vector.
hFindGreaterEqual(vec, threshold, vecout) - Find the samples that are GreaterEqu
al a certain threshold value and returns the number of samples found and the pos
itions of the samples in a second vector.
hFindGreaterEqualAbs(vec, threshold, vecout) - Find the samples whose absolute v
alues are GreaterEqual a certain threshold value and returns the number of sampl
es found and the positions of the samples in a second vector.
```

hFindLessThan(vec, threshold, vecout) - Find the samples that are LessThan a cer tain threshold value and returns the number of samples found and the positions o f the samples in a second vector.

hFindLessThanAbs(vec, threshold, vecout) - Find the samples whose absolute value s are LessThan a certain threshold value and returns the number of samples found and the positions of the samples in a second vector.

hDownsample(vec1, vec2) - Downsample the input vector to a smaller o utput vector.

 $\label{lownsample} \mbox{hDownsample(vec, downsample_factor) - Downsample the input vector by a cetain factor and return a new vector.} \\$

hFindLowerBound(vec, value) - Finds the location (i.e., returns integer) in a monotonically increasing vector, where the input search value is just above or equal to the value in the vector.

hFlatWeights(wlen) - Returns vector of weights of length len wi th constant weights normalized to give a sum of unity. Can be used by hRunningAv erageT.

hLinearWeights(wlen) - Returns vector of weights of length wlen w ith linearly rising and decreasing weights centered at len/2.

hGaussianWeights(wlen) - Returns vector of weights of length wlen w ith Gaussian distribution centered at len/2 and sigma=len/4 (i.e. the Gaussian e xtends over 2 sigma in both directions).

hWeights(wlen, wtype) - Create a normalized weight vector.

hRunningAverage(idata, odata, weights) - Calculate the running average of an input vector using a weight vector.

hRunningAverage(idata, odata, wlen, wtype) - Overloaded function to automatically calculate weights.

SECTION: RF (Radio Frequency) Function

hGeometricDelayFarField(antPosition, skyDirection, length) - Calculates the time delay in seconds for a signal received at an antenna position relative to a phase center from a source located in a certain direction in farfield (based on L. Bahren).

hGeometricDelayNearField(antPosition, skyPosition, distance) - Calculates the ti me delay in seconds for a signal received at an antenna position relative to a p hase center from a source located at a certain 3D space coordinate in nearfield (based on L. Bahren).

hGeometricDelays(antPositions, skyPositions, delays, farfield) - Calculates the time delay in seconds for signals received at various antenna positions relative to a phase center from sources located at certain 3D space coordinates in near or far field.

hGeometricPhases(frequencies, antPositions, skyPositions, phases, farfield) - Ca lculates the phase gradients for signals received at various antenna positions r elative to a phase center from sources located at certain 3D space coordinates i n near or far field and for different frequencies.

hGeometricWeights(frequencies, antPositions, skyPositions, weights, farfield) - Calculates the phase gradients as complex weights for signals received at variou s antenna positions relative to a phase center from sources located at certain 3 D space coordinates in near or far field and for different frequencies.

hSpectralPower(vec, outvec) - Calculates the power of a complex spectrum and add it to an output vector.

SECTION: I/O Function (DataReader)

hFileClose(iptr) - Function to close a file with a datareader object providing the pointer to the object as an integer.

hFileOpen(Filename) - Function to open a file based on a filenam e and returning a pointer to a datareader object as an integer.

hFileGetParameter(iptr, keyword) - Return information from a data file as a P ython object.

hFileSetParameter(iptr, keyword, pyob) - Set parameters in a data file with a Py thon object as input.

hFileRead(iptr, Datatype, vec) - Read data from a Datareader object (pointe r in iptr) into a vector, where the size should be pre-allocated.

hCalTable(filename, keyword, date, pyob) - Return a list of antenna positions fr om the CalTables - this is a test.

SECTION: Coordinate Conversion (VectorConversion.cc)

hCoordinateConvert(source, sourceCoordinate, target, targetCoordinate, anglesInD egrees) - Converts a 3D spatial vector into a different Coordinate type (e.g. Sp herical to Cartesian).

hReadFileOld(vec, iptr, Datatype, Antenna, Blocksize, Block, Stride, Shift) - Re ad data from a Datareader object (pointer in iptr) into a vector.