

# TOI-396 Planetary characterisation

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Master Thesis at The Space Research Institute in Graz (Austria)

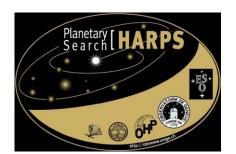
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#### The scientific case

Fundamental planetary parameters

Radial velocity method

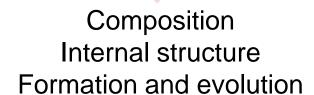


Mass

**Radius** 

Transit method

Density







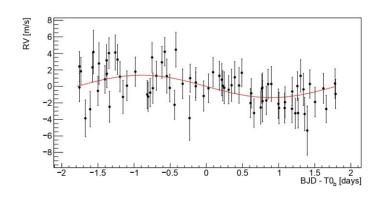
# Radial Velocity Method

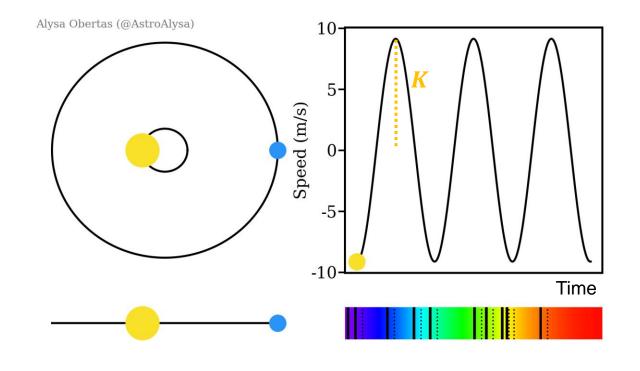


78 HARPS spectra



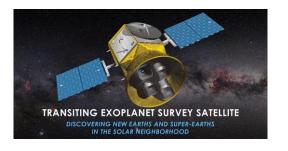
Planetary mass determination







#### **Transit Method**

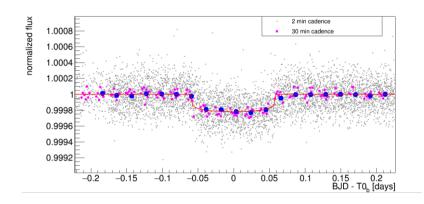


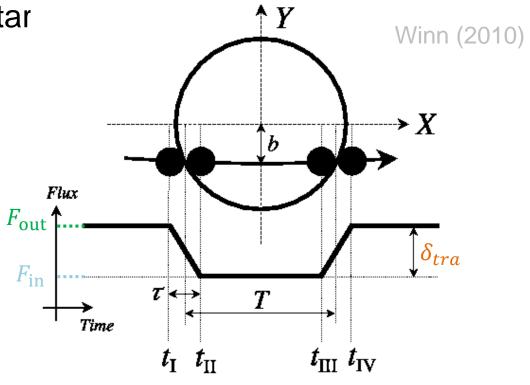
Periodic dimming of the brightness of the star as the orbiting planet transits in front of it

**TESS** transit observations



#### Planetary radius determination







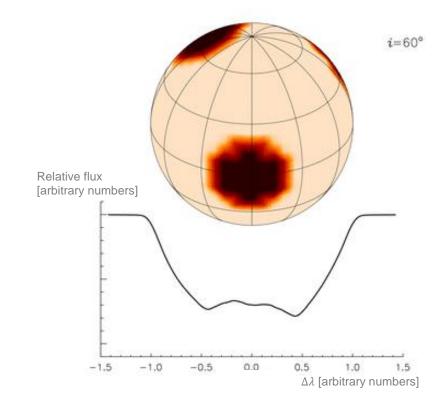
## Stellar Activity

Star spots, flares, faculae

Correlation with stellar rotational period

Potentially masking a planetary signal

RV detrending using break-point method Stellar activity is not stationary



Axel Hahlin et al. (2018)

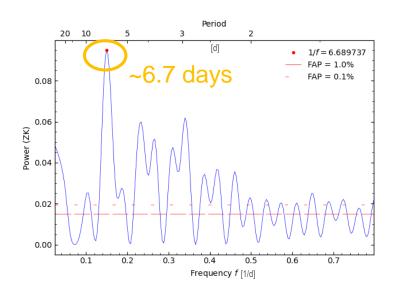


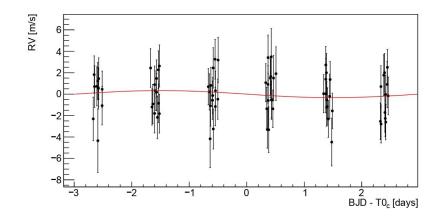
### Stellar Activity

TOI-396 c RV signal is not statistically detected

Hypothesis:  $P_{rot} \approx P_c$ 

Periodograms of activity indicators





Stellar activity covered the RV signal generated by TOI-396 c

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#### Results

• Discovery paper: Vanderburg (2019)

Radii determination using transit method

 My work: Joint analysis of RV time series and transits using a Markov Chain Monte Carlo code (Bonfanti & Gillon, 2020)

Masses and radii determination

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TESS Spots a Compact System of Super-Earths around the Naked-eye Star HR 858

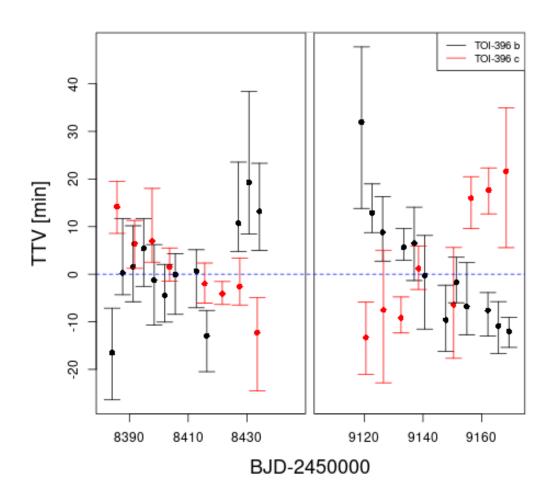
Andrew Vanderburg<sup>1,27</sup>, Chelsea X. Huang<sup>2,28</sup>, Joseph E. Rodriguez<sup>3,29</sup>, Juliette C. Becker<sup>4,30,31</sup>, George R. Ricker<sup>2</sup>, Roland K. Vanderspek<sup>2</sup>, David W. Latham<sup>3</sup>, Sara Seager<sup>2,5</sup>, Joshua N. Winn<sup>6</sup>, Jon M. Jenkins<sup>7</sup>, Brett Addison<sup>8</sup>,

TOI-396 d  $R_d \simeq 2.1 R_{\oplus}$ TOI-396 b  $\rho_d \simeq 3.8 \text{ g cm}^{-3}$  $R_b \simeq 2.1 R_{\oplus}$  $M_h \simeq 3.3 M_{\oplus}$  $\rho_h \simeq 2.0 \text{ g cm}$ TOI-396 c  $R_c \simeq 2.1 R_{\oplus}$  $M_c < 3.8 M_{\oplus}$  $\rho_{c} < 2.5 \text{ g cm}^{-3}$ 

Irene Amateis



## Transit timing variations (TTV)



Anti-correlation pattern in TTV of TOI-396 b and c



Mean motion resonance

$$\frac{P_c}{P_b} \cong 1,666213 \approx 5:3$$



Proof that TOI-396 c orbits around TOI-396



### Un unusual system

Atmospheric characterization needed

A paper on this thesis is the process of being submitted

