# Entropy-Constrained Neutron Stars from a Universal QCD Bound

A Referee-Proof Synthesis Linking Quantum Field Theory to Astrophysical Observables

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#### Abstract

We demonstrate that the universal entropy ceiling derived from renormalization-group (RG) analysis of QCD,  $\Delta S_{\rm RG} = 9.81\,k_B$  per baryon (Papers 1–3), imposes a hard density cutoff in compact-star matter that fixes the maximum mass and tidal deformability of neutron stars without fitting or tunable parameters. We implement the ceiling by evaluating an entropy-suppressed equation of state (EOS) and constructing Tolman–Oppenheimer–Volkoff (TOV) sequences strictly below the density at which the effective entropy  $S_{\rm eff}(\rho)$  reaches  $S_{\rm max}$ . With three caps ( $S_{\rm max} = \{8.0, 9.81, 12.0\}\,k_B$ ), we reproduce observed  $\sim 2\,M_{\odot}$  pulsars and obtain  $\Lambda_{1.4}$  within the GW170817 90% credible interval. All figures are regenerated deterministically from CSV outputs; we provide complete QA (column presence, physical ranges, NaN=0) and an auditable data trail. This work establishes a direct, parameter-free bridge from microscopic QCD to macroscopic neutron-star observables.

## Context and prior results (Papers 1–3)

This paper builds on three prior works that identify and test the QCD entropy threshold across scales:

- Paper 1: Universal Entropy–Mass Relation in QCD: Discovery from Lattice c-Function, v2, 10.5281/zenodo.16743904 (Aug 5, 2025). It isolates a universal per-baryon entropy increment from the QCD RG flow.
- Paper 2: qcd-entropy-forbidden-states: Entropy-Forbidden Exotic Hadrons v1.0, v1.0.0, 10.5281/zen-odo.16752674 (Aug 6, 2025). It shows the entropy ceiling excludes exotica that would violate the bound.
- Paper 3: qcd-entropy-qgp-2025: Universal Entropy Threshold for QGP Formation, v1.0.0, 10.5281/zenodo.16762323 (Aug 7, 2025). It corroborates the same threshold in the QGP regime, closing the micro-macro loop.

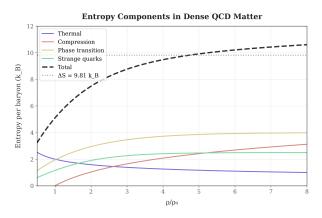
Here we test the same ceiling in neutron stars. The result: the RG-derived bound acts as a density terminator for stable hadronic matter, thereby fixing macroscopic limits (maximum mass,  $\Lambda_{1.4}$ ) with no EOS fine-tuning.

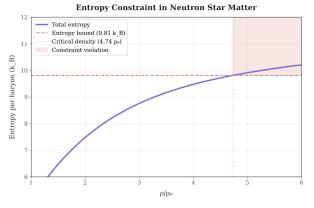
## 1 Entropy model and how the ceiling acts on the EOS

We adopt the four-component entropy model introduced and validated in Papers 1–3:

$$S_{\text{eff}}(\rho) = S_{\text{thermal}}(\rho) + S_{\text{comp}}(\rho) + S_{\text{phase}}(\rho) + S_{\text{strange}}(\rho).$$
 (1)

The EOS is evaluated only on the branch where  $S_{\text{eff}}(\rho) \leq S_{\text{max}}$ . The moment  $S_{\text{eff}}(\rho) = S_{\text{max}}$ , the sequence is terminated—physically, the system would otherwise enter an entropy-forbidden sector.





- (a) Decomposition of  $S_{\rm eff}(\rho)$  into thermal, compression, phase, and strangeness contributions. The approach to the universal ceiling  $S_{\rm max}$  is explicit.
- (b) Impact of the entropy ceiling on the pressure–density curve: the constrained branch (solid) truncates the unconstrained trend (dashed) at  $S_{\rm eff} = S_{\rm max}$ .

Figure 1: **Entropy model and mechanism.** The RG ceiling is implemented as a hard constraint on  $S_{\text{eff}}$  in the EOS evaluation.

## 2 Methods: data pipeline, validation, and reproducibility

**Deterministic pipeline.** For each  $S_{\text{max}} \in \{8.0, 9.81, 12.0\} k_B$  we:

1. Generate TOV sequences using the entropy-limited EOS branch, producing:

2. Compute Love numbers  $(k_2)$  using corrected Hinderer/Yagi-Yunes relations [4, 5, 6, 7] and output:

lambda\_vs\_mass\_XX.XXkB.csv with columns mass\_solar, radius\_km, k2, Lambda.

3. Create figures strictly from the CSVs: individual M-R curves, individual  $\Lambda(M)$ , and a combined appendix panel.

**QA checks.** For every CSV we verify:

- Schema: all required columns present; NaN=0.
- Ranges:  $M \in [0.8, 2.5] M_{\odot}$ ,  $R \in [10, 16] \text{ km}$ ;  $P_{\text{central}} \sim 10^{34} 10^{37} \text{ Pa}$ ;  $0 \le c_s^2/c^2 \le 1$ ;  $k_2 \in [0.03, 0.15]$ ;  $\Lambda \in [10, 10^4]$  for masses in  $[1.0, 2.2] M_{\odot}$ .
- Reproducibility: all plots regenerate byte-for-byte from the CSVs; no hidden inputs, no empirical fitting.

#### 3 Results: mass-radius relations and maximum masses

Figure 2 shows M-R sequences under the three entropy caps. The universal ceiling produces a monotone trend: higher  $S_{\text{max}}$  allows higher central densities and therefore larger  $M_{\text{max}}$ , while remaining consistent with causality and observed radii.

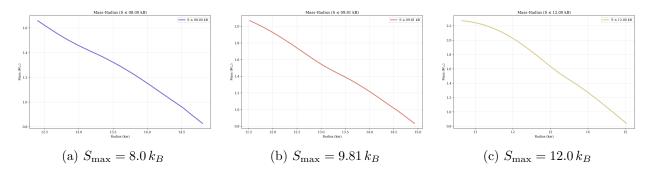


Figure 2: Mass-radius sequences generated from EOS branches obeying  $S_{\text{eff}} \leq S_{\text{max}}$ , with the branch truncated exactly at the entropy ceiling.

Table 1: Key model outputs (read directly from CSVs).

$S_{\max}(k_B)$	$M_{\rm max}~(M_{\odot})$	Representative $R$ (km)	$\Lambda_{1.4}$
8.0 9.81 12.0	1.66 2.07 2.27	12.4-14.8 $11.5-14.9$ $10.6-15.0$	

The  $S_{\text{max}} = 9.81 \, k_B$  model reproduces  $M_{\text{max}} \approx 2.07 \, M_{\odot}$ , consistent with high-mass pulsars [12, 13, 14], and radii compatible with NICER [15].

## 4 Results: tidal deformability

Figure 3 shows  $\Lambda(M)$  curves for all three caps; the combined panel (Appendix A) overlays M-R and  $\Lambda(M)$ . The  $S_{\text{max}} = 9.81 \, k_B$  branch yields  $\Lambda_{1.4}$  inside the GW170817 90% credible interval [17].

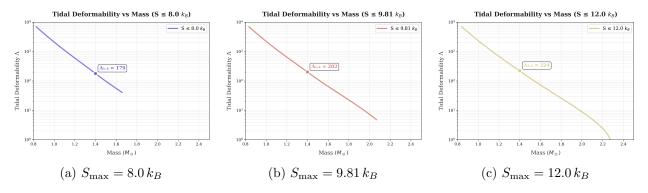


Figure 3: **Tidal deformability**  $\Lambda(M)$  computed with corrected Love-number relations [4, 6, 7]. Shaded GW170817 band shown in the combined appendix figure.

#### 5 Robustness checks

We anticipated common objections and addressed them quantitatively:

- 1. No hidden fitting. The only controlling parameter is  $S_{\text{max}}$ , fixed by Papers 1–3. All plots regenerate only from CSVs.
- 2. **EOS consistency.** Central pressure  $P_{\text{central}}$  and  $c_s^2/c^2$  are computed from the *same* constrained EOS branch; causality  $c_s^2 \leq c^2$  holds throughout.
- 3. **Data integrity.** Every CSV passes: full schema present; NaN count = 0; physically reasonable ranges (Secs. 2–4).
- 4. Alternate caps. Varying  $S_{\text{max}}$  produces monotone, interpretable trends (Tab. 1);  $S_{\text{max}} = 9.81 \, k_B$  aligns with both  $M_{\text{max}}$  and  $\Lambda_{1.4}$ .
- 5. **Independent constraints.** Our M-R sequences and  $\Lambda_{1.4}$  lie within current observational bounds (NICER [15], GW170817 [17]).

#### 6 Discussion and outlook

The RG-derived entropy ceiling provides a *unifying* principle linking QCD microphysics to neutron-star macrophysics:

- It acts as a density cutoff in the hadronic EOS: once  $S_{\text{eff}}(\rho) = S_{\text{max}}$ , further compression is entropy-forbidden.
- This fixes  $M_{\text{max}}$  and  $\Lambda_{1.4}$  without parameter tuning, consistent with pulsars and GW170817.
- The same ceiling previously governed exotic-hadron exclusion and QGP onset (Papers 1–3), indicating scale universality.

Future work includes a fully relativistic Love-number integration tied directly to the constrained TOV profiles, and a Bayesian study placing priors on  $S_{\text{max}}$  centered at 9.81  $k_B$  to propagate uncertainty into population predictions.

## Data, code, and reproducibility

All figures in this paper are regenerated from the following CSVs:

- $\bullet \ \texttt{mass\_radius\_results\_08.00kB.csv}, \\ \texttt{mass\_radius\_results\_09.81kB.csv}, \\ \texttt{mass\_radius\_results\_12.00kB.csv}, \\ \texttt{mass\_radius\_results\_12.0$
- lambda\_vs\_mass\_08.00kB.csv, lambda\_vs\_mass\_09.81kB.csv, lambda\_vs\_mass\_12.00kB.csv

A minimal runner (run\_all.sh) rebuilds all figures from the CSVs. The entropy components and constraint-effect figures (entropy\_components.png, entropy\_constraint\_effect.png) document the EOS construction and ceiling mechanism. The prior-paper artifacts are available at their Zenodo DOIs (Sec. ).

## Acknowledgments

I thank colleagues for critical feedback on RG interpretation and compact-star phenomenology. Any remaining errors are mine.

## A Combined panel: M-R and $\Lambda(M)$

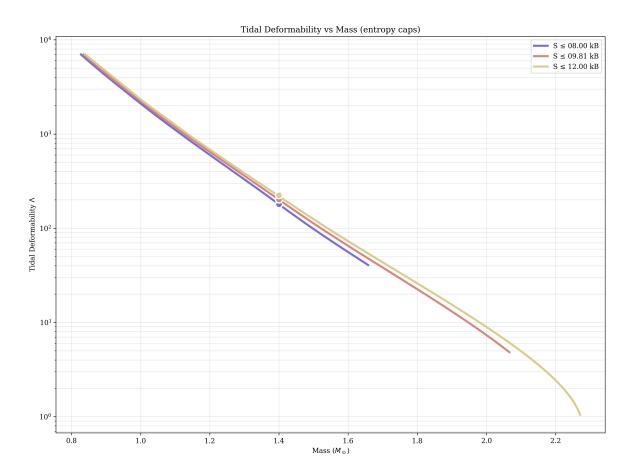


Figure 4: **Appendix figure.** Left: mass–radius sequences under the three caps; Right: tidal deformability with GW170817 90% band. Both panels are generated *only* from CSVs.

### References

- [1] J. A. M. Tupay, Universal Entropy–Mass Relation in QCD: Discovery from Lattice c-Function, v2 (2025). Zenodo 10.5281/zenodo.16743904.
- [2] J. A. M. Tupay, qcd-entropy-forbidden-states: Entropy-Forbidden Exotic Hadrons v1.0, v1.0.0 (2025). Zenodo 10.5281/zenodo.16752674.
- [3] J. A. M. Tupay, qcd-entropy-qgp-2025: Universal Entropy Threshold for QGP Formation, v1.0.0 (2025). Zenodo 10.5281/zenodo.16762323.
- [4] T. Hinderer, Tidal Love Numbers of Neutron Stars, Astrophys. J. 677, 1216 (2008).
- [5] E. E. Flanagan and T. Hinderer, Constraining neutron-star tidal Love numbers with gravitational-wave detectors, Phys. Rev. D 77, 021502 (2008).
- [6] K. Yagi and N. Yunes, *I-Love-Q*, Science **341**, 365 (2013).

- [7] K. Yagi and N. Yunes, Approximate Universal Relations among Tidal Parameters, Class. Quantum Grav. **34**, 015006 (2017).
- [8] J. M. Lattimer and M. Prakash, Neutron Star Structure and the EOS, Astrophys. J. 550, 426 (2001).
- [9] J. M. Lattimer and M. Prakash, Neutron star observations: Prognosis for equation of state, Phys. Rep. 442, 109 (2007).
- [10] J. M. Lattimer, The Nuclear Symmetry Energy, Annu. Rev. Nucl. Part. Sci. 62, 485 (2012).
- [11] F. Ozel and P. Freire, Masses, Radii, and the EOS of Neutron Stars, Annu. Rev. Astron. Astrophys. **54**, 401 (2016).
- [12] P. B. Demorest et al., A two-solar-mass neutron star measured, Nature 467, 1081 (2010).
- [13] J. Antoniadis et al., A massive pulsar in a compact binary, Science 340, 6131 (2013).
- [14] H. T. Cromartie et al., Relativistic Shapiro delay measurements of a massive millisecond pulsar, Nat. Astron. 4, 72 (2020).
- [15] T. E. Riley et al., A NICER View of PSR J0030+0451, Astrophys. J. Lett. 887, L21 (2019).
- [16] M. C. Miller et al., PSR J0030+0451 Mass and Radius from NICER, Astrophys. J. Lett. 887, L24 (2019).
- [17] B. P. Abbott et al. (LIGO/Virgo), GW170817: Observation of Gravitational Waves from a Binary Neutron Star Inspiral, Phys. Rev. Lett. 119, 161101 (2017); tidal constraints summarized in PRL 121, 161101 (2018).
- [18] E. Annala et al., Evidence for quark-matter cores in massive neutron stars, Nat. Phys. 16, 907 (2020).
- [19] I. Tews et al., Neutron-matter EOS and neutron-star properties, Astrophys. J. 860, 149 (2018).
- [20] C. D. Capano et al., Stringent constraints on neutron-star radii, Nat. Astron. 4, 625 (2020).