Golden Gate Invitational Astronomy C Key

	Team Name/Number: <u>Astronomy</u>	
	Raw Score: 89.5/89.5	
	Grading Instructions:	
	 (specified below) For questions with more than one ar 12a, 13, 14a, 14b, 19, 20) Each part of the answer is 0. For example, question 4 has (and as a whole the question The total point values for the questions Questions with longer answers included the main parts of an answer of the answer are noted by breather than 15 miles and 15 miles are 15 miles are 15 miles and 15 miles are 15 miles are	the answer B, E, so B and E are each worth 0.5 pts is thus worth 1 point) se questions are noted in parentheses next to the de 28, 29a needed as well as the associated points for that part
	Section A	a. <u>Ultraviolet or UV</u>
1.	1, 5, 6, 8, 16, 21, 23, 25, 27 (4.5)	b. <u>Precession</u>
2.	<u>6, 16, 23</u> (1.5)	9. Hen 2-428
3.	9, 10, 13, 14 (2)	a. The Chandrasekhar limit
4.	B, E (1)	b. Type Ia SN
5.	18, 7, 14, 3, 23 (2.5)	c. AM CVn
6.	HR Diagram or color-magnitude diagram	10. Post-AGB
	a. <u>K</u>	a. Optical/Visible
	b. <u>H</u>	
	c. <u>D</u>	· · · · · · · · · · · · · · · · · · ·
	d. <u>E</u>	1122, differentiated by characteristically long
	e. <u>L</u>	period(1
7.	Asymptotic Giant branch or AGB	a. <u>Left, Stefan-Boltzmann Law</u> (1
	aFirst dredge-up_	b. X-ray
	b. William Herschel	122

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a. Red Dwarf, White Dwarf (1) b.	_ <u>Isochrones</u>
b. <u>GPE / Gravitational Potential Energy</u>	<u>y</u> 21. <u>Mess</u>	ier 15 or M15
13. <u>26, Sirius (or Sirius A and B)</u> (1	a.	PN (Planetary Nebula)
a. <u>1</u>	b.	_They may require more massive
b. <u>Only Balmer lines (no He I or</u>		progenitors (or systems) to form, which
metals)	_	is rarer OR planetary nebulae are
14 <u>AM CVn</u>	<u> </u>	dispersed quickly/have short lifetimes
a. Less, more	(1)	
b. More, more	(1)	
15. <u>16</u>		
a. <u>X-ray</u>	<u> </u>	
b. <u>Gravitational waves</u>	_	
16. <u>13</u>	<u> </u>	
a. <u>Nickel or Ni-56</u>		
17. <u>SNR 0509-67.5</u>	<u> </u>	
a. <u>Fe or Iron</u>	_	
b Double-degenerate or two white dw	<u>arf</u>	
binary system		
18 Tycho's SNR		
a Silicon or Si		
b. <u>Shell-type</u>	<u> </u>	
c Cosmic rays		
19. 20, Sagittarius (1)	
a <u>Delayed detonation (a slow wavefro</u>	<u>ont</u>	
followed by a much faster one)		
b. <u>Electrons</u>		
20. <u>12, NGC 1846</u> ((1)	
a Multiple stellar populations		

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Section B

22. 0.112 kpc

23. 0.886 lv

_		,		
24	4 RR Lyrae stars			
	a.	1.714 kpc		
251	1,900	0 Lsun		
26. <u>-90.1 km/s</u>				
	a.	364 km/s		
	b.	_375 km/s		
	c.	79.1 AU		
27. <u>87.71 yr</u>				
	a.	0.976"		
	b.	0.191 solar masses		
	c.	2.27 * 10 ⁻⁹ rad/s		
	d.	_1.45 * 10 ⁻⁷ rad/s		
28. Occurred ~23.2 million years ago [1 pt]. We				
would be unlikely to see a trace (the SN would				
be dispersed). [1 pt](2				
29. <u>3.31</u> * 10 ⁴⁷ ergs				
	a.	Type Ia SN in general should explode		
		with consistent amounts of energy based		
		on the Chandrasekhar limit [1 pt]. Since		
		the kinetic energy was lower, the ejecta		
		may have been relatively heavy [1 pt].		
		This makes sense with a double		
		degenerate progenitor (a Type Ia SN		

from two white dwarfs) because the

companion of a single white dwarf

progenitor system would be less

evolved/would only fuse up to lighter

elements by the time the SN occurred

(OR that a system of white dwarfs

would be more evolved/made of heavier

elements) [2 pt]. (4)

30. 9700 K

- a. _ The top, green curve____
- b. The bottom, blue curve
- c. False
- d. False
- e. <u>True</u>

Tiebreaker list (in order):

6, 30, 3, 29, 4, 1, 28, 14, 26, 2, 10, 17, 21, 24, 8, 11, 22, 20, 5, 13, 15, 23, 25, 7, 27, 18, 16, 19, 9, 12