

WD-BASS

Prepare your data. 3 columns:
Wavelength (Angstroms),
Flux (erg/cm²/s/Hz),
Flux error (erg/cm²/s/Hz).

Run **create_single** or **create_double**
(or your own adapted file for your
needed use) in the directory where
your files are stored

Did it create a yaml
file with your spectra
inside?

Check input data format
and compare to
example input data

No

Yes

Do you want to fit
the photometry?

No

Yes

Insert your photometry
From CDS -> Run WD-BASS normally with
run_single or run_double and put
fit_phot_sed: True in the yaml file. A file
vizier_votable.vot should be downloaded,
which you should then rename to
"photSED.vot" and place in a folder
(subdirectory) called "out".

You can manually input whatever
photometry you want in lists on the yaml file,
with keyword
"External_AB_mag" (float)
"External_AB_mag_err" (float)
"External_Filter" (string, "")

Supported filters and their names are found
in the WD-BASS install directory -> Filters

Preview
photometry?

No

Yes

In the yaml file...
plot_phot_spectrum: True
phot_min_val: 0.0000 (your value)

Execute run_single or run_double, see the
popup window, ctrl+c out of the code.
When happy,
plot_phot_spectrum: False

Open the file. Change the star
type (e.g. "DA", "GG", "DB", "LL").
change the resolution at each
spectral line ($R = \lambda_{\text{lambda}}/\lambda_{\text{lambda}}$),
sigma clipping, scaling, any file
you want to ignore, number of
walkers, burning/steps. etc.

When happy run in terminal
run_single or **run_double**.
The output will appear in a
new folder "out"

Do you
want to improve
the RVs?

Yes

No

Execute "**run_single_one_by_one**" or
"**run_double_one_by_one**" to fix the
model with the solution you just found.
Alternatively, you can fix e.g. forced_teff
to a desired value and execute this task

(Optional) Tweak input boundaries and repeat all

End