

An spacetime extension for the SPDE approach

Elias Teixeira Krainski

King Abdullah University of Science and Technology (KAUST)

May 2024

Outline

Ideas and motivation

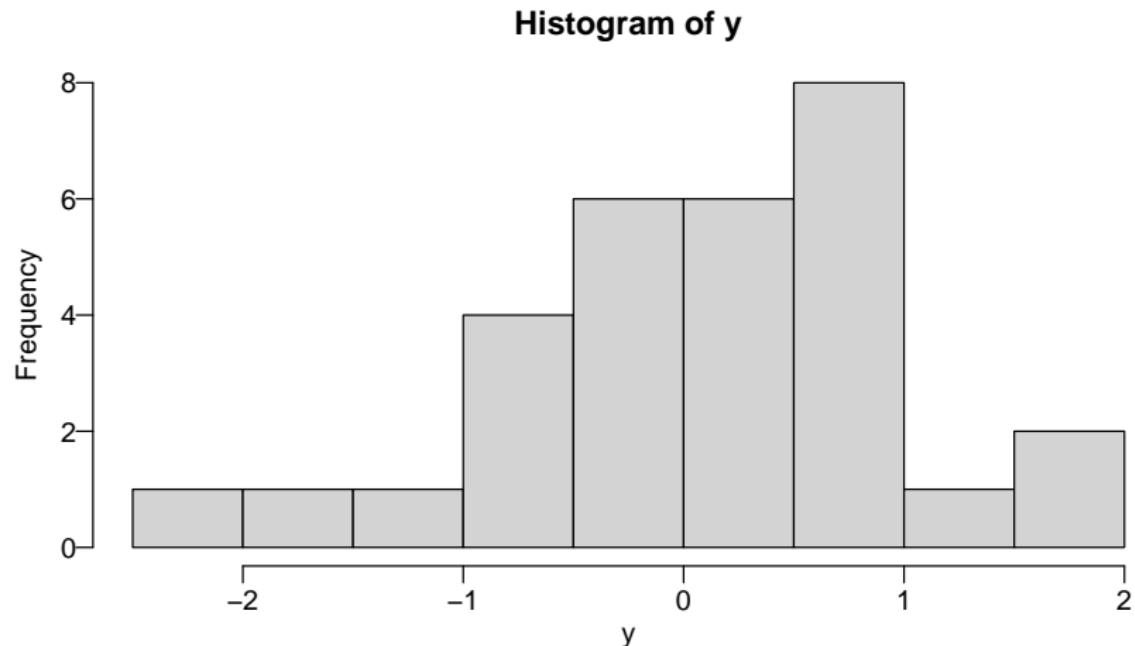
An spacetime SPDE model extension, Lindgren et al. (2024)

Implementation

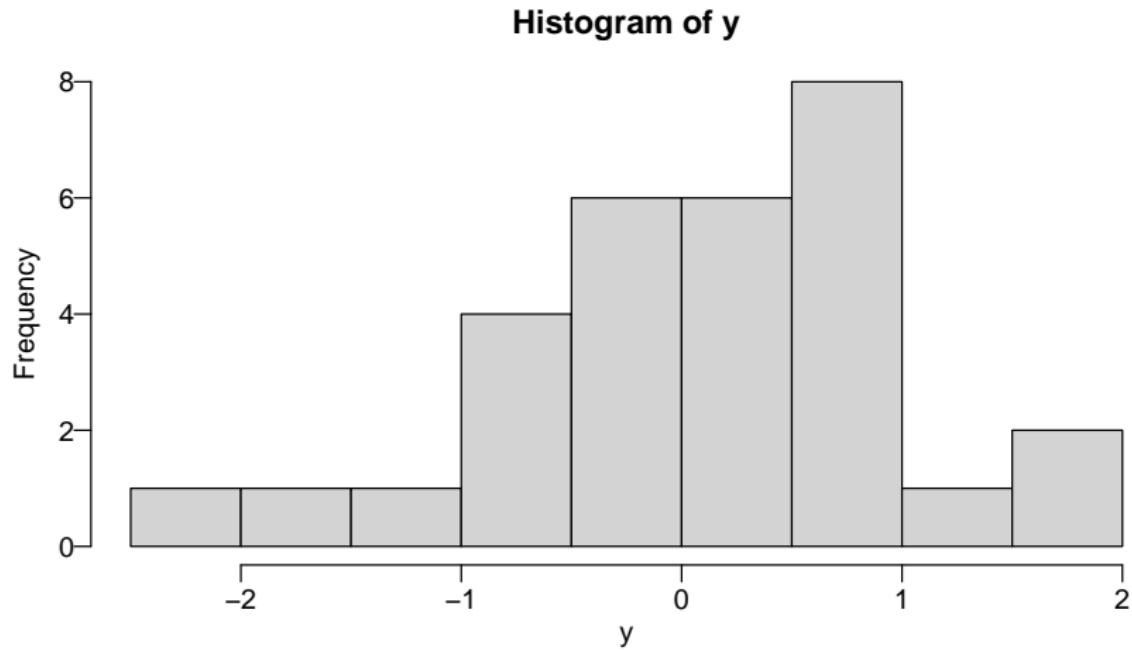
Examples

Ideas and motivation

Symmetric or assymetric?

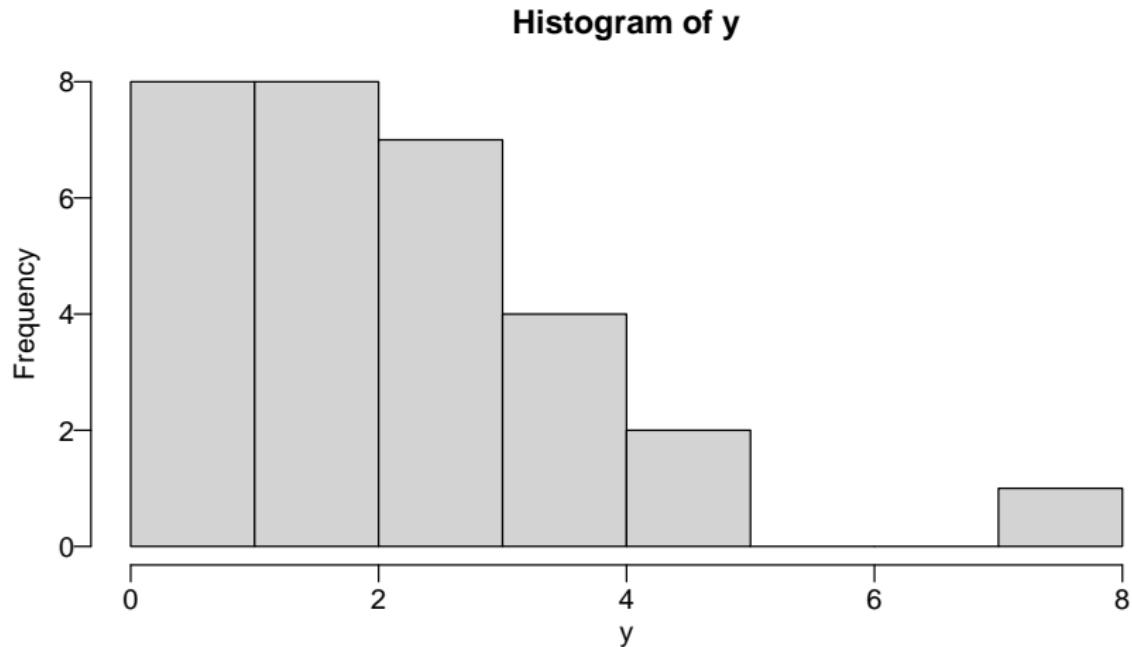


Symmetric or assymetric?

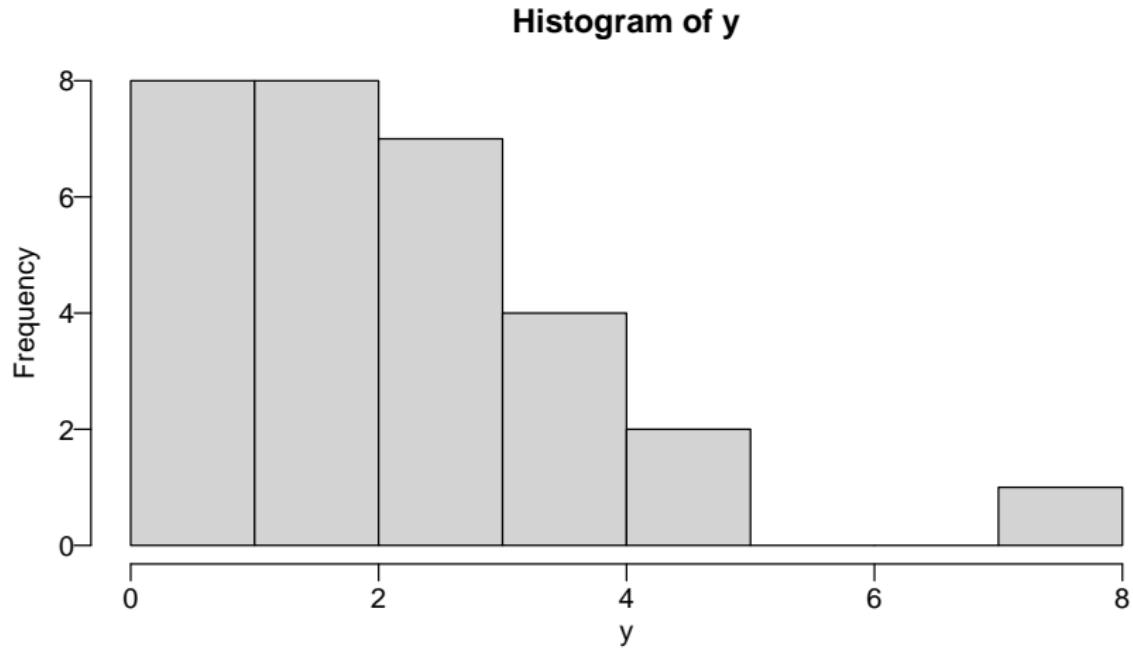


```
set.seed(1); y <- rnorm(30); hist(y)
```

What causes this to be assymetric?



What causes this to be assymetric?



```
set.seed(1); x = rexp(30); y = 1 + x + rnorm(30); hist(y)
```

Stationary and non-stationary

Stationary

- ▶ Covariance is the same over the domain
 - ▶ E.g. 1d: does not matter where, but only the lag

$$\text{Cov}[u(t_1), u(t_2)] = \text{cov}(h), \quad h = |t_1 - t_2|$$

Stationary and non-stationary

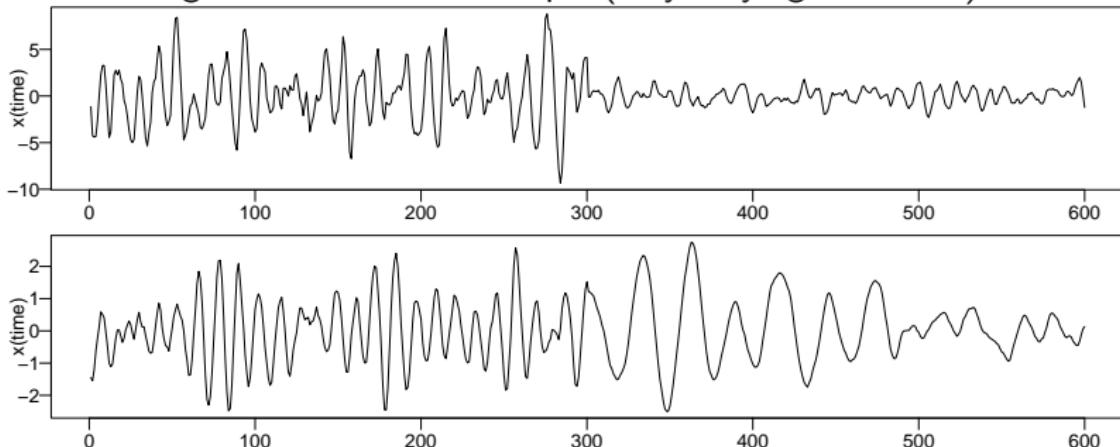
Stationary

- ▶ Covariance is the same over the domain
 - ▶ E.g. 1d: does not matter where, but only the lag

$$\text{Cov}[u(t_1), u(t_2)] = \text{cov}(h), \quad h = |t_1 - t_2|$$

Non-stationary

- ▶ Any model where precision/covariance depends on the location
 - ▶ e.g.: 1 dimensional example (only varying over time)



Stationary or non-stationary?



Spatial Statistics

Volume 14, Part C, November 2015, Pages 505-531



Does non-stationary spatial data
always require non-stationary random
fields?

Geir-Arne Fuglstad ^a  , Daniel Simpson ^a, Finn Lindgren ^b, Håvard Rue ^a

Stationary or non-stationary?



Spatial Statistics

Volume 14, Part C, November 2015, Pages 505-531



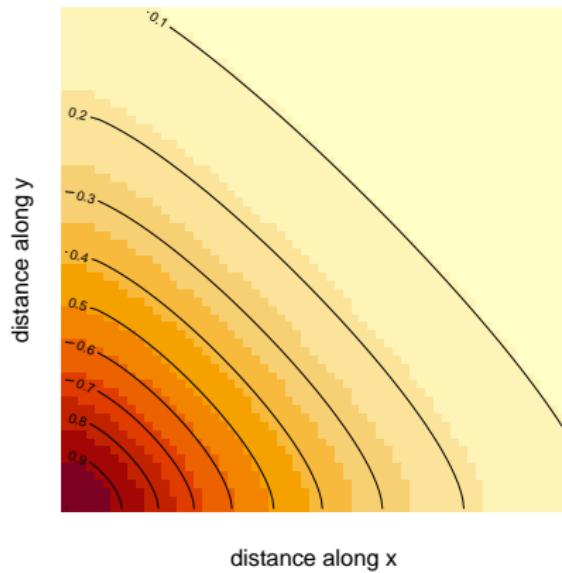
Does non-stationary spatial data
always require non-stationary random
fields?

Geir-Arne Fuglstad ^a  , Daniel Simpson ^a, Finn Lindgren ^b, Håvard Rue ^a

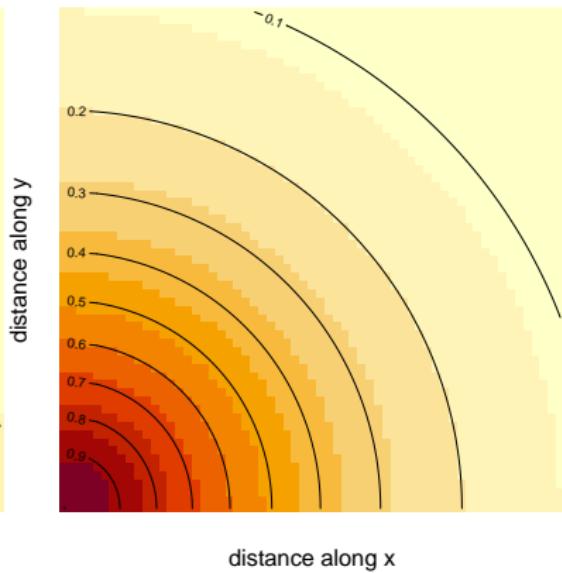
It is hard to beat simple suitable models, Fuglstad et al. (2015).

Separable or non-separable?

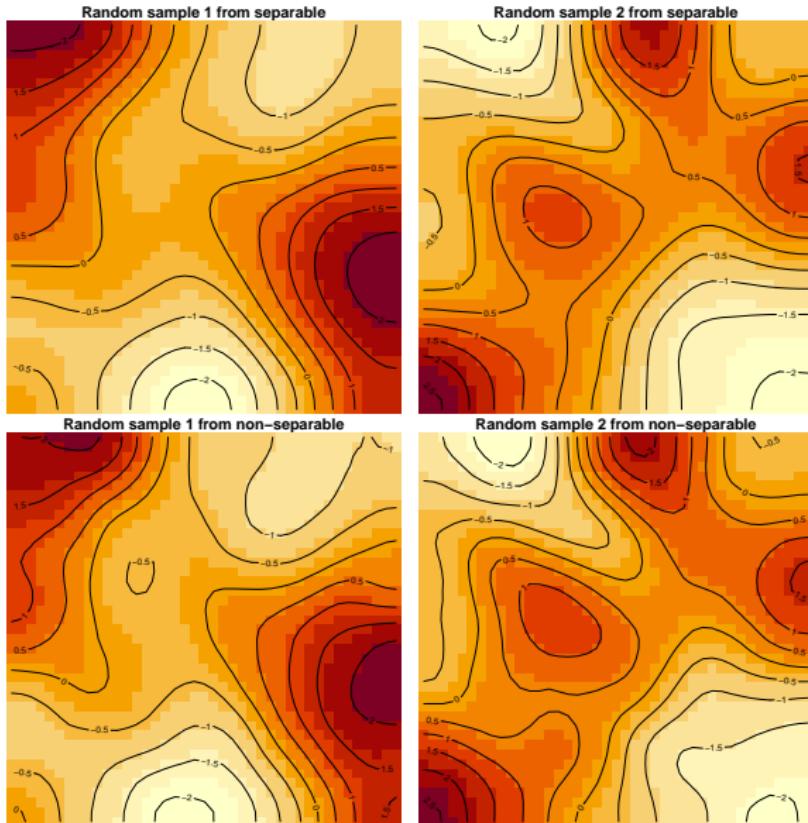
Separable correlation



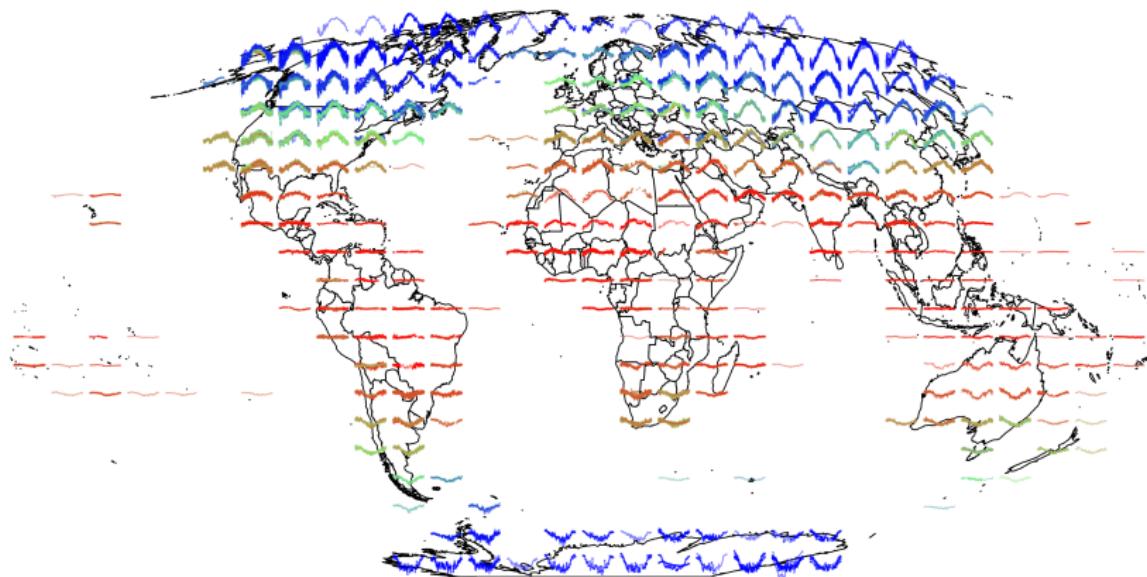
Non-separable correlation



Separable or non-separable?



Global temperature, see Lindgren et al. (2024)



Symmetric? Stationary? Separable?

**An spacetime SPDE model extension,
Lindgren et al. (2024)**

The space-time models

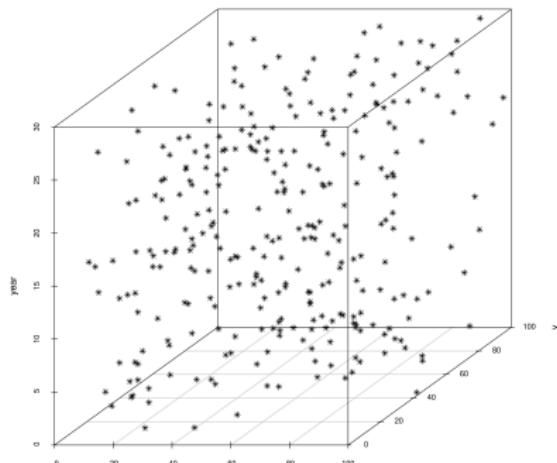
Models for processes that on the space-time domain

- ▶ Discrete space-time domain
 - ▶ $\mathbf{u} = \{u_{11}, \dots, u_{n1}, u_{12}, \dots, u_{n2}, \dots, u_{1T}, \dots, u_{nT}\}$
- ▶ Continuous space-time domain
 - ▶ $\mathbf{u}(\mathbf{s}, t), \mathbf{s} \in \mathbb{R}^d, t \in \mathbb{R}$

The space-time models

Models for processes that on the space-time domain

- ▶ Discrete space-time domain
 - ▶ $\mathbf{u} = \{u_{11}, \dots, u_{n1}, u_{12}, \dots, u_{n2}, \dots, u_{1T}, \dots, u_{nT}\}$
- ▶ Continuous space-time domain
 - ▶ $\mathbf{u}(\mathbf{s}, t), \mathbf{s} \in \mathbb{R}^d, t \in \mathbb{R}$
- ▶ Real world (locations)



Separable space-time models

- ▶ When covariance/precision can be written as Kronecker product of a purely temporal and purely spatial ones
- ▶ E.g.: Discrete space-time domain

$$\pi(\mathbf{u}) \propto (|\mathbf{Q1} \otimes \mathbf{Q2}|^{-1})^{1/2} \exp\left(-\frac{1}{2}\mathbf{u}^T \{\mathbf{Q1} \otimes \mathbf{Q2}\} \mathbf{u}\right)$$

- ▶ $\mathbf{Q1}$ has dimension equals T
- ▶ $\mathbf{Q2}$ has dimension equals n

Desired properties of a space-time process

- ▶ Three desired properties, Stein (2013)
 - ▶ **1st:** “any degree of smoothness in space and any (possibly different) degree of smoothness in time”
 - ▶ **2nd:** “smoother away from the origin than it is at the origin”
 - ▶ **3th:** covariance “computed accurately and efficiently”

Desired properties of a space-time process

- ▶ Three desired properties, Stein (2013)
 - ▶ **1st:** “any degree of smoothness in space and any (possibly different) degree of smoothness in time”
 - ▶ **2nd:** “smoother away from the origin than it is at the origin”
 - ▶ **3th:** covariance “computed accurately and efficiently”
- ▶ Separable models fails to fulfill **2nd**
- ▶ **3th** difficulty with “good” models

Desired properties of a space-time process

- ▶ Three desired properties, Stein (2013)
 - ▶ **1st:** “any degree of smoothness in space and any (possibly different) degree of smoothness in time”
 - ▶ **2nd:** “smoother away from the origin than it is at the origin”
 - ▶ **3th:** covariance “computed accurately and efficiently”
- ▶ Separable models fails to fulfill **2nd**
- ▶ **3th** difficulty with “good” models

Why SPDE?

- ▶ A solution to a linear non-fractional SPDE is a Markov process, Rozanov (1977)
 - ▶ Reciprocal of its spectral density is a polynomial
 - ▶ 1st and 2nd desired properties follows
- ▶ Allow efficient computations
 - ▶ Replace Stein's third desired property by
 - ▶ fitting process computationally efficient
- ▶ More in Vergara, Allard, and Desassis (2022)

Space-time SPDE models

- ▶ SDE in time **then** SPDE in space

$$\begin{aligned} \left(\gamma_t \frac{\partial}{\partial t} + a^2 \right) u(\mathbf{s}, \delta t) &= \xi(\mathbf{s}, \delta t) \\ (\kappa^2 - \Delta)^{\alpha_\epsilon/2} \xi(\mathbf{s}, \delta t) &= \mathcal{W}(\mathbf{s}, \delta t). \end{aligned}$$

Space-time SPDE models

- ▶ SDE in time **then** SPDE in space

$$\begin{aligned} \left(\gamma_t \frac{\partial}{\partial t} + a^2 \right) u(\mathbf{s}, \delta t) &= \xi(\mathbf{s}, \delta t) \\ (\kappa^2 - \Delta)^{\alpha_e/2} \xi(\mathbf{s}, \delta t) &= \mathcal{W}(\mathbf{s}, \delta t). \end{aligned}$$

- ▶ SPDE in time **and** in space

$$\begin{aligned} \left(\gamma_t \frac{\partial}{\partial t} + (\gamma_s^2 - \Delta)^{\alpha_s/2} \right) u(\mathbf{s}, \delta t) &= \xi(\mathbf{s}, \delta t) \\ (\gamma_s^2 - \Delta)^{\alpha_e/2} \xi(\mathbf{s}, \delta t) &= \mathcal{W}(\mathbf{s}, \delta t). \end{aligned}$$

A class of SPDE models, Lindgren et al. (2024)

- ▶ Spatial operator $L_s = \gamma_s^2 - \Delta$

$$\left(\gamma_t \frac{\partial}{\partial t} + L_s^{\alpha_s/2} \right)^{\alpha_t} u(\mathbf{s}, \delta t) = \xi(\mathbf{s}, \delta t)$$
$$L_s^{\alpha_e/2} \xi(\mathbf{s}, \delta t) = \mathcal{W}(\mathbf{s}, \delta t).$$

A class of SPDE models, Lindgren et al. (2024)

- ▶ Spatial operator $L_s = \gamma_s^2 - \Delta$

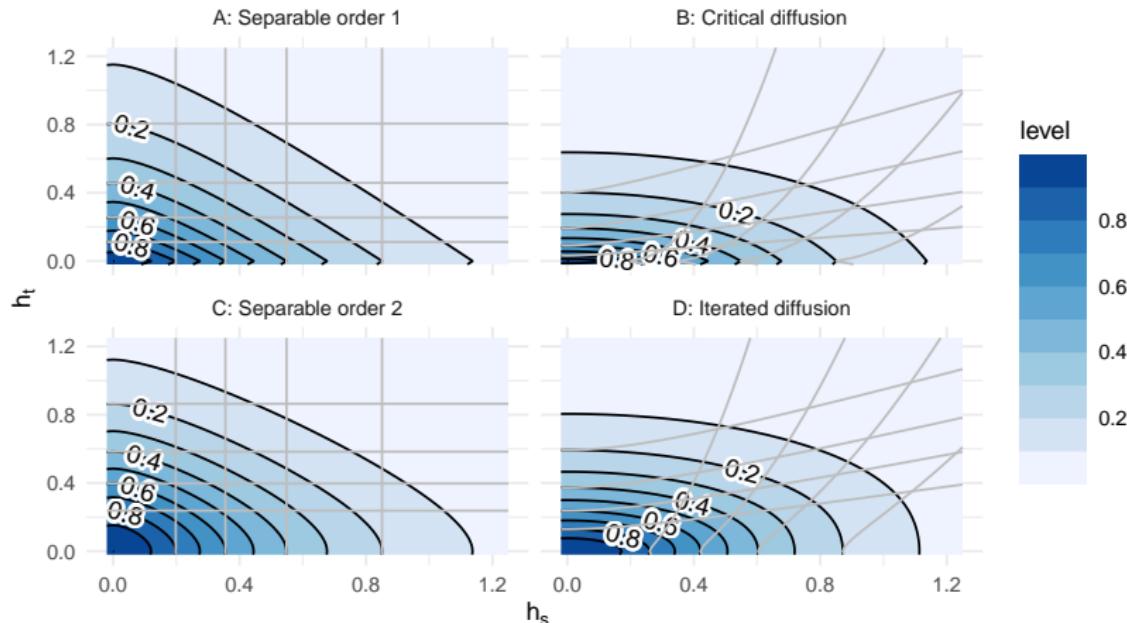
$$\left(\gamma_t \frac{\partial}{\partial t} + L_s^{\alpha_s/2} \right)^{\alpha_t} u(\mathbf{s}, \delta t) = \xi(\mathbf{s}, \delta t)$$
$$L_s^{\alpha_e/2} \xi(\mathbf{s}, \delta t) = \mathcal{W}(\mathbf{s}, \delta t).$$

Table 1: Four specific models on 2-dimensional spatial domains.

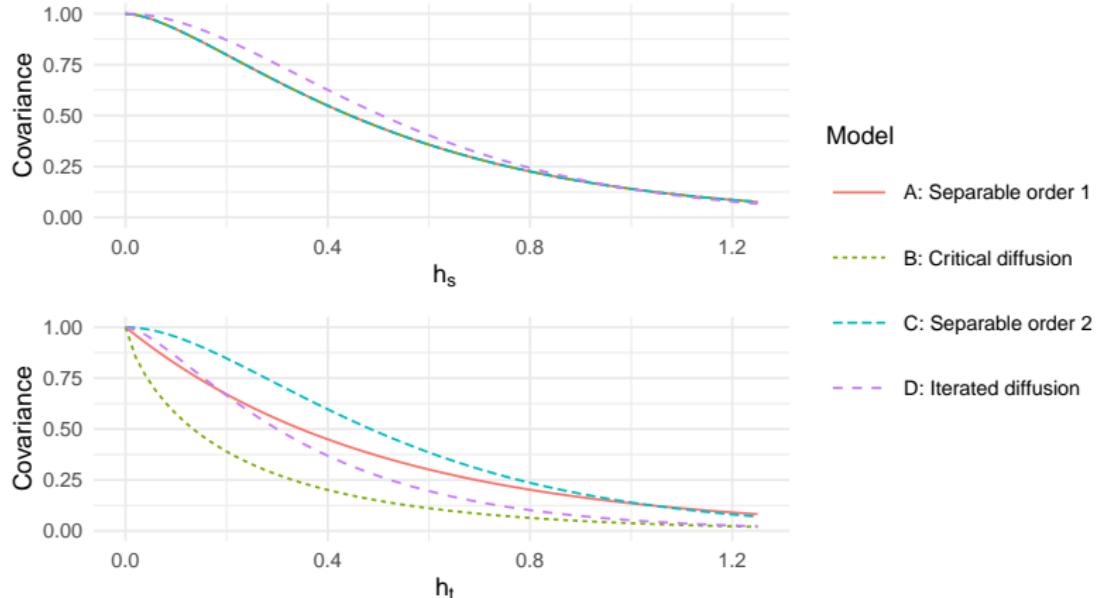
Model	α_t	α_s	α_e	Type	ν_t	ν_s	nonsep
A:(1,0,2)	1	0	2	Separable order 1	1/2	1	0.0
B:(1,2,1)	1	2	1	Critical diffusion	1/2	1	0.5
C:(2,0,2)	2	0	2	Separable order 2	3/2	1	0.0
D:(2,2,0)	2	2	0	Iterated diffusion	1	2	1.0

Separability is a function of α_t , α_s , α_e and d .

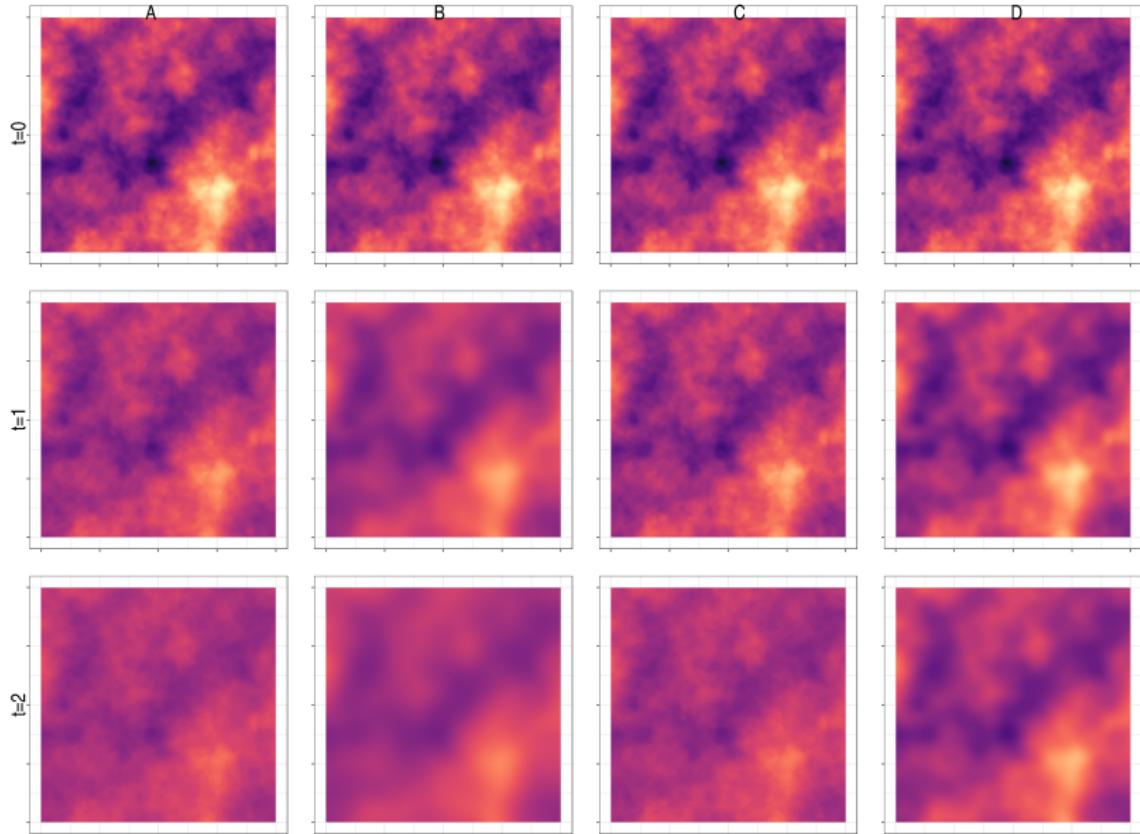
Space-time correlations



Marginal correlations



Example on predictions



Precision matrix

- ▶ A sum of Kronecker products

$$\mathbf{Q}_u(\gamma_s, \gamma_t, \gamma_s) = \gamma_e^2 \sum_{k=0}^{2\alpha_t} \gamma_t^k \mathbf{J}_{\alpha_t, k/2} \otimes \mathbf{K}_{\alpha_s}(\gamma_s)$$

where the $\mathbf{J}_{\alpha_t, k/2}$ matrices are from the temporal discretization.

Interpretable parametrization

- ▶ The parameters in the SPDE are local
- ▶ It is easier to consider the marginal parameters

Interpretable parametrization

- ▶ The parameters in the SPDE are local
- ▶ It is easier to consider the marginal parameters
- ▶ $\sigma = \frac{C(\alpha_t, \alpha_s, \alpha_e, d)}{\gamma_e^2 \gamma_t \gamma_s^{2\alpha-d}}$
- ▶ $r_s = \sqrt{8\nu} / \gamma_s^2$
- ▶ $r_t = \frac{\gamma_t \sqrt{8(\alpha_t - 0.5)}}{\gamma_s^{\alpha_s}}$

Implementation

Implementation

- ▶ Integrated Nested Laplace Approximations - INLA, Rue, Martino, and Chopin (2009)
 - ▶ **INLA** package

Implementation

- ▶ Integrated Nested Laplace Approximations - INLA, Rue, Martino, and Chopin (2009)
 - ▶ **INLA** package
- ▶ **INLA** package uses efficient direct solver algorithms for general sparse precision matrices
 - ▶ recommended: PARDISO library (or another parallel solver)

Implementation

- ▶ Integrated Nested Laplace Approximations - INLA, Rue, Martino, and Chopin (2009)
 - ▶ **INLA** package
- ▶ **INLA** package uses efficient direct solver algorithms for general sparse precision matrices
 - ▶ recommended: PARDISO library (or another parallel solver)
- ▶ New avenue: LA + VB, see van Niekerk et al. (2023)

Implementation

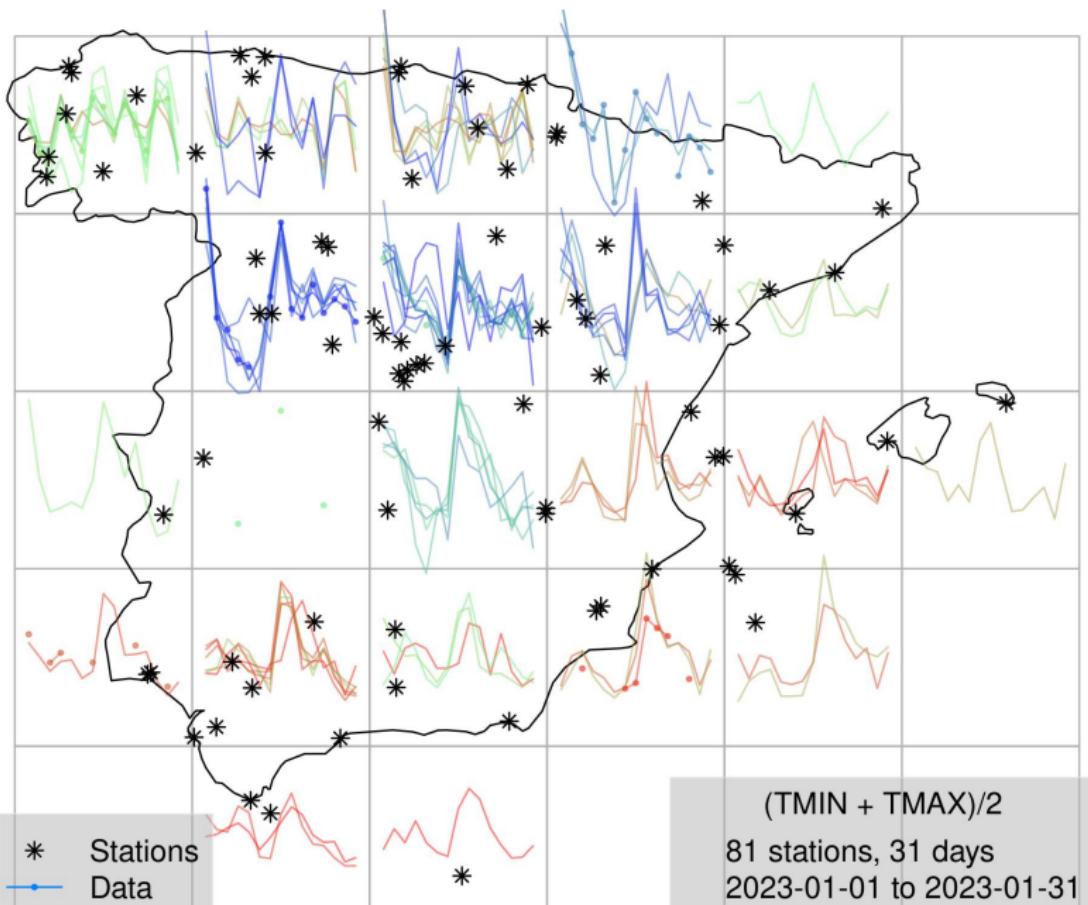
- ▶ Integrated Nested Laplace Approximations - INLA, Rue, Martino, and Chopin (2009)
 - ▶ **INLA** package
- ▶ **INLA** package uses efficient direct solver algorithms for general sparse precision matrices
 - ▶ recommended: PARDISO library (or another parallel solver)
- ▶ New avenue: LA + VB, see van Niekerk et al. (2023)
- ▶ cgeneric interface in **INLA**
 - ▶ to implement new random effect models
 - ▶ define functions to compute **Q**
 - ▶ cgeneric takes full advantage of parallel computations

Implementation

- ▶ Integrated Nested Laplace Approximations - INLA, Rue, Martino, and Chopin (2009)
 - ▶ **INLA** package
- ▶ **INLA** package uses efficient direct solver algorithms for general sparse precision matrices
 - ▶ recommended: PARDISO library (or another parallel solver)
- ▶ New avenue: LA + VB, see van Niekerk et al. (2023)
- ▶ cgeneric interface in **INLA**
 - ▶ to implement new random effect models
 - ▶ define functions to compute **Q**
 - ▶ cgeneric takes full advantage of parallel computations
- ▶ **INLAspacetime** package in
 - ▶ <https://github.com/eliaskrainski/INLAspacetime>

Examples

Daily temperature data in Spain, January 2023



The data in the long format

```
str(dataf)
## 'data.frame':    2511 obs. of  6 variables:
## $ iloc: int  1 2 3 4 5 6 7 8 9 10 ...
## $ xloc: num  442 168 367 578 925 ...
## $ yloc: num  4474 4311 4059 4795 4584 ...
## $ elev: num  0.667 0.185 0.007 0.251 0.004 ...
## $ time: int  1 1 1 1 1 1 1 1 1 ...
## $ resp: num  11.3 15.1 13.8 16.9 12.2 ...
```

Elevation (in km) to be used as covariate

Model definition, with INLAspacetime

Define the space-time model with INLAspacetime

```
stmodel <- stModel.define(  
    smesh = smesh, # spatial mesh  
    tmesh = tmesh, # temporal mesh  
    model = "121", # model (\alpha_t, \alpha_s, \alpha_e)  
    control.priors = list(  
        prs = c(100, 0.05), # P(spatial range < 100) = 0.05  
        prt = c(2, 0.05), # P(temporal range < 1) = 0.05  
        psigma = c(4, 0.05)), # P(sigma_u > 4) = 0.05  
    constr = TRUE)
```

Define the right-hand-side, inlabru

```
M <- ~ Intercept(1) + elev +  
    field(list(space = cbind(xloc, yloc),  
              time = time),  
          model=stmodel)
```

Model fit with inlabru

Set the PC-prior for the likelihood precision

```
lkprec <- list(  
  prec = list(prior = "pcprec",  
              param = c(4, 0.05)))
```

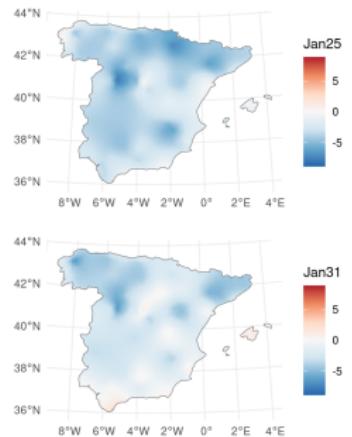
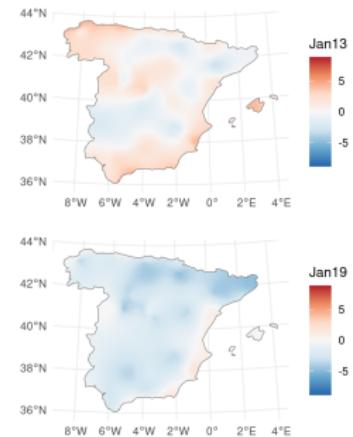
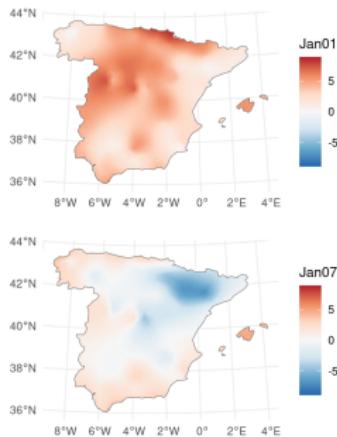
The likelihood setup in inlabru

```
lhood <- like(  
  formula = resp ~ .,  
  family = "gaussian",  
  control.family = list(  
    hyper = lkprec),  
  data = dataf)
```

Fit

```
result <- bru(M, lhood)
```

Posterior mean of $u(.)$, for some days



Jan01

Jan13

Jan25

Jan07

Jan19

Jan31

Posterior mean of $u(.)$, for some days



UK example: Daily wind speed, inla.gcv(..., m = 10)

M102, time 16

Neighbors:

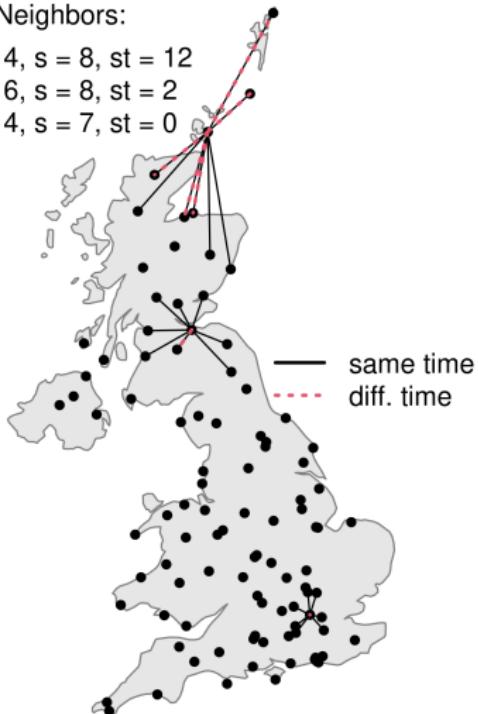
$t = 6, s = 5, st = 6$
 $t = 4, s = 5, st = 0$
 $t = 4, s = 9, st = 0$



M121, time 16

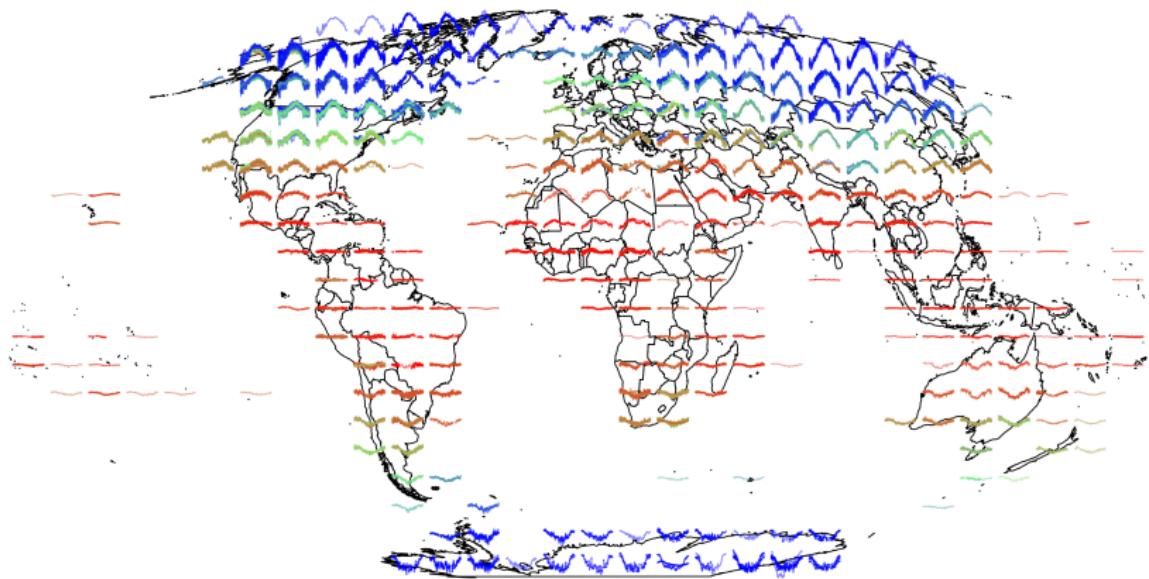
Neighbors:

$t = 4, s = 8, st = 12$
 $t = 6, s = 8, st = 2$
 $t = 4, s = 7, st = 0$

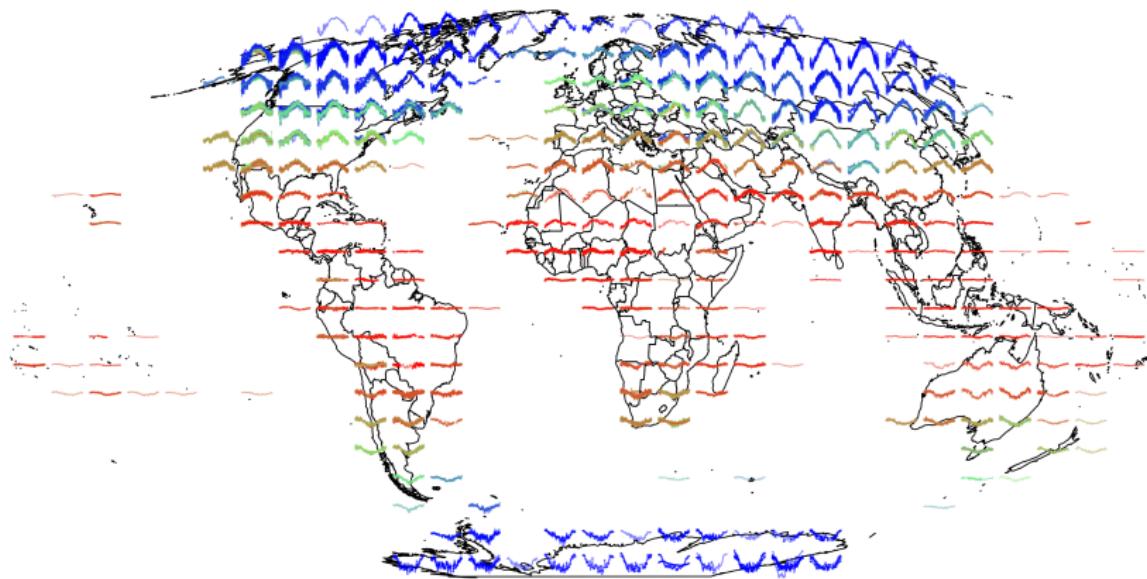


— same time
- - - diff. time

Daily temperature data, Lindgren et al. (2024)



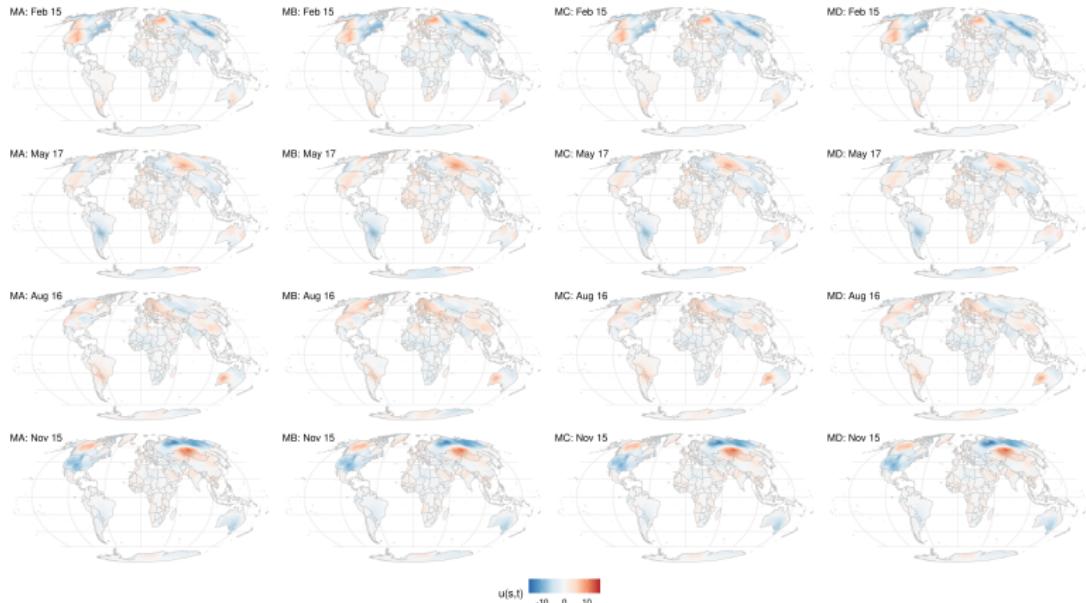
Daily temperature data, Lindgren et al. (2024)



$$\mathbf{y} = \mathbf{1}\mu + \mathbf{E}\alpha + \mathbf{B}\mathbf{b} + \mathbf{A}_v\mathbf{v} + \mathbf{A}_u\mathbf{u} + \mathbf{e}$$

elevation + seasonal,latitude + slow space-time + fast space-time

Posterior mean of $u(\cdot)$, for some days, four models



References

- Fuglstad, G-A, D. Simpson, F. Lindgren, and H. Rue. 2015. "Does Non-Stationary Spatial Data Always Require Non-Stationary Random Fields." *Spatial Statistics* 14: 505–31.
- Lindgren, F., H. Bakka, D. Bolin, E. T. Krainski, and H. Rue. 2024. "The Diffusion-Based Extension of the Matérn Field to Space-Time (with Discussion)." *SORT*. <https://raco.cat/index.php/SORT/article/view/428665>.
- Rozanov, J. A. 1977. "Markov Random Fields and Stochastic Partial Differential Equations." *Math. USSR Sbornik* 32 (4): 515–34. <http://dx.doi.org/10.1070/SM1977v032n04ABEH002404>.
- Rue, H., S. Martino, and N. Chopin. 2009. "Approximate Bayesian Inference for Latent Gaussian Models Using Integrated Nested Laplace Approximations (with Discussion)." *Journal of the Royal Statistical Society, Series B* 71 (2): 319–92.
- Stein, M. L. 2013. "On a Class of Space-Time Intrinsic Random Functions." *Bernoulli* 19 (2): 287–408.
- van Niekerk, J., E. T. Krainski, D. Rustand, and H. Rue. 2023. "A New Avenue for Bayesian Inference with INLA." <https://doi.org/10.1016/j.csda.2023.107692>.
- Vergara, Ricardo Carrizo, Denis Allard, and Nicolas Desassis. 2022. "A General Framework for SPDE-Based Stationary Random Fields." *Bernoulli* 28 (1): 1–32. <https://doi.org/10.3150/20-BEJ1317>.