Fluid Thesis

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Abstract Just so I don't forget that there is an abstract environment...

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0.1 Introduction

0.2 Derivation of the Squire-Long equation

Squire-long / Bragg-Hawthorne equation for the stream function of axisymmetric inviscid fluid, using cylindrical coordinates

$$\frac{\partial^2 \Psi}{\partial z^2} + \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} = r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi}$$

radial component u, azimuthal (swirl) is v, axial component w stream function satisfies

 $\nabla \cdot u = 0 \longrightarrow \text{streamfunction exists}$

Remember for cylindrical coordinates:

$$u = \frac{1}{r} \frac{\partial \Psi}{\partial z}, \quad w = -\frac{1}{r} \frac{\partial \Psi}{\partial r}$$

 Ψ is the stream function

r is the radius

$$C = rr$$

$$H = \frac{p}{a} + \frac{1}{2}(u^2 + v^2 + w^2)$$

 $H = \frac{p}{\rho} + \frac{1}{2}(u^2 + v^2 + w^2)$ H is conserved on stream surfaces

C is conserved on stream surfaces

vorticity

$$w = w_r e_r + w_\theta e_\theta + w_z e_z$$

where w_r, w_θ, w_z can be written in terms of the velocity

Considering cylindrical coordinates (z,r,θ) with corresponding velocity (u,v,w), vorticity components (w_z, w_r, w_θ) . Axisymmetric flow as:

$$\omega_z = \frac{1}{r} \frac{\partial rv}{\partial r}, \quad \omega_r = -\frac{\partial rv}{\partial z}, \quad \omega_\theta = \frac{\partial u}{\partial z} - \frac{\partial w}{\partial r}$$

The continuity equation (conservation of mass) is satisfied by setting

$$w = \frac{1}{r} \frac{\partial \Psi}{\partial r}, \quad u = -\frac{1}{r} \frac{\partial \Psi}{\partial z}$$

Where Ψ is the stream function This gives the azimuthal component for w_{θ} :

$$\omega_{\theta} = \frac{\partial u}{\partial z} - \frac{\partial w}{\partial r}$$

$$= -\frac{1}{r} \frac{\partial^{2} \Psi}{\partial z^{2}} - \frac{1}{r} \frac{\partial^{2} \Psi}{\partial r^{2}} + \frac{1}{r^{2}} \frac{\partial \Psi}{\partial r}$$

$$= -\frac{1}{r} \left(\frac{\partial^{2} \Psi}{\partial z^{2}} + \frac{\partial^{2} \Psi}{\partial r^{2}} - \frac{1}{r} \frac{\partial \Psi}{\partial r} \right)$$

Use the vorticity equation

$$w \times v - \frac{\partial w}{\partial t} = \nabla H$$

Where

$$H = \frac{1}{2}(w^2 + u^2 + v^2) + \frac{p}{\rho}$$

This gives:

$$u\omega_{\theta} - v\omega_{r} - \frac{\partial w}{\partial t} = \frac{\partial H}{\partial x}$$
$$v\omega_{z} - w\omega_{\theta} - \frac{\partial u}{\partial t} = \frac{\partial H}{\partial r}$$
$$w\omega_{r} - u\omega_{z} - \frac{\partial v}{\partial t} = 0$$

The last one is equivalent to the material derivative of rw set to 0:

$$\frac{D(rv)}{Dt} = 0$$

From the Bernoulli equation:

$$rv = C(\Psi)$$
$$\frac{\partial \Psi}{\partial t} + \frac{1}{2}|\mathbf{w}|^2 + \frac{p}{\rho} = H(\Psi)$$

Where $H(\Psi)$ and $C(\Psi)$ are arbitrary functions.

Rewriting ω :

$$\omega_z = w \frac{dC}{d\Psi}, \quad \omega_r = u \frac{dC}{d\Psi}$$

Giving

$$\frac{\omega_{\theta}}{r} = \frac{v\omega_{r}}{ru} + \frac{1}{ru}\frac{dH}{d\Psi}\frac{\partial\Psi}{\partial z} = \frac{C}{r^{2}}\frac{dC}{d\Psi} - \frac{dH}{d\Psi}$$

Which is the form taken by the second of the dynamic equations. Now, combining this last statement with the equation for ω_{θ} :

$$\frac{\partial^2 \Psi}{\partial z^2} + \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} = r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi}$$

Taken from Batchelor's An Introduction to Fluid Dynamics

Considering the flow far upstream where there is constant uniform axial velocity and rotates with angular velocity Ω

$$\Psi_{\text{upstream}} = \frac{1}{2}Wr^2$$
$$v = \Omega r, w = W$$

And

$$C = rv = \frac{v^2}{\Omega} = \Omega r^2 = 2\Omega \Psi / W$$
$$\frac{dC}{d\Psi} = 2\Omega / W$$

Since the flow is steady, the radial equation of motion yields:

$$\frac{1}{\rho}\frac{dp}{dr} = \frac{w^2}{r} = \frac{C^2}{r^3}$$

$$\begin{split} H &= \frac{1}{2}(u^2 + v^2 + w^2) + \frac{p}{\rho} \\ &= \frac{1}{2}(\Omega^2 r^2 + W^2) + \frac{p}{\rho} \\ &= \frac{\Omega^2 \Psi}{W} + \frac{1}{2}W^2 + \frac{p}{\rho} \\ &= \frac{\Omega^2 \Psi}{W} + \frac{1}{2}W^2 + \int \frac{1}{\rho} \frac{dp}{dr} dr \\ &= \frac{\Omega^2 \Psi}{W} + \frac{1}{2}W^2 + \int \frac{C^2}{r^3} dr \\ &= \frac{\Omega^2 \Psi}{W} + \frac{1}{2}W^2 + \int \frac{\Omega^2 r^4}{r^3} dr \\ &= \frac{\Omega^2 \Psi}{W} + \frac{1}{2}W^2 + \int \Omega^2 r dr \\ &= \frac{\Omega^2 \Psi}{W} + \frac{1}{2}W^2 + \frac{1}{2}\Omega^2 r^2 \\ &= \frac{2\Omega^2 \Psi}{W} + \frac{1}{2}W^2 \end{split}$$

$$\frac{dH}{d\Psi} = \frac{\partial \frac{2\Omega^2 \Psi}{W}}{\partial \Psi}$$
$$= \frac{2\Omega^2}{W}$$

$$\begin{split} \frac{\partial^2 \Psi}{\partial z^2} + \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi} \\ \frac{\partial^2 \Psi}{\partial z^2} + \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= \frac{2r^2 \Omega^2}{W} - \frac{4\Omega^2}{W^2} \Psi \end{split}$$

Or in a more 'standard' form

$$\frac{\partial^2 \Psi}{\partial z^2} + \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} + \frac{4\Omega^2}{W^2} \Psi = \frac{2r^2\Omega^2}{W}$$

0.2.1 Homogeneous ODE

Considering the case where Ψ is just a function of the radius, r. So Ψ does not depend on z, and $\frac{\partial^2 \Psi}{\partial z^2} = 0$

To simplify it into a homogeneous ODE, a change of variables is used:

$$\Psi = \frac{1}{2}Wr^2 + \psi = \frac{1}{2}Wr^2 + rF$$

$$\frac{\partial \Psi}{\partial r} = Wr + F + r \frac{\partial F}{\partial r}$$
$$\frac{\partial^2 \Psi}{\partial r^2} = W + 2 \frac{\partial F}{\partial r} + r \frac{\partial^2 F}{\partial r^2}$$

$$\frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} + \Psi (\frac{4\Omega^2}{W^2} - \frac{1}{r^2}) = 0$$

$$r^{2}\frac{d^{2}F}{dr^{2}} - r\frac{dF}{dr} + F(r^{2}k^{2} - 1) = 0$$

Letting $k = \frac{2\Omega}{W}$ If we take x = kr, $\frac{dF}{dr} = \frac{dF}{dx}\frac{dx}{dr} = k$ and $\frac{d^2F}{dr^2} = k^2\frac{d^2F}{dx^2}$

$$\frac{x^2}{k^2}k^2\frac{d^2F}{dx^2} - \frac{x}{k}k\frac{dF}{dx} + F(\frac{x^2}{k^2}k^2 - 1) = 0$$
$$x^2\frac{d^2F}{dx^2} - x\frac{dF}{dx} + F(x^2 - 1) = 0$$

Which is the form of a bessel differential equation of order $\nu = 1$, giving solutions

$$F = AJ_1(kr) + BY_1(kr)$$

Returning to the streamfunction:

$$\Psi = \frac{1}{2}Wr^{2} + r(AJ_{1}(kr) + BY_{1}(kr))$$

And hence

$$w = \frac{1}{r} \frac{\partial \Psi}{\partial r} = W + AkJ_0(kr) + BkY_0(kr)$$

A, and B rely on boundary conditions. In this case, it is necessary forthe streamlines to be the same as at the inlet along the boundary. Also introduce a vortex breakdown condition in the core of the stream, i.e. a region $0 < r < r_*$ where the streamfunction becomes zero:

$$\Psi(R) = \frac{1}{2}WR^2$$

$$\Psi(r_*) = 0$$

Consider it as a matrix system

$$\begin{pmatrix} r_*J_1(kr_*) & r_*Y_1(kr_*) \\ RJ_1(kR) & RY_1(kR) \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix} = \begin{pmatrix} -\frac{1}{2}Wr_*^2 \\ 0 \end{pmatrix}$$

Giving

$$\begin{pmatrix} A \\ B \end{pmatrix} = \frac{1}{r_* R \left(J_1(kr_*) Y_1(kR) - Y_1(kr_*) J_1(kR) \right)} \begin{pmatrix} R Y_1(kR) & -r_* Y_1(kr_*) \\ -R J_1(kR) & r_* J_1(kr_*) \end{pmatrix} \begin{pmatrix} -\frac{1}{2} W r_*^2 \\ 0 \end{pmatrix}$$

$$A = \frac{-\frac{1}{2}RWr_*^2Y_1(kR)}{r_*R\left(J_1(kr_*)Y_1(kR) - Y_1(kr_*)J_1(kR)\right)}$$
$$B = \frac{\frac{1}{2}RWr_*^2J_1(kR)}{r_*R\left(J_1(kr_*)Y_1(kR) - Y_1(kr_*)J_1(kR)\right)}$$

And hence

$$A = \frac{-\frac{1}{2}Wr_*Y_1(kR)}{(J_1(kr_*)Y_1(kR) - Y_1(kr_*)J_1(kR))}$$
$$B = \frac{\frac{1}{2}Wr_*J_1(kR)}{(J_1(kr_*)Y_1(kR) - Y_1(kr_*)J_1(kR))}$$

With the requirement that $r_* \neq R$ so as to not divide by zero.

Using

$$w = \frac{1}{r} \frac{\partial \Psi}{\partial r}$$

Gives

$$w = W + k(AJ_0(kr) + BY_0(kr))$$

Solving this for a given k (or alternatively a desired r_*) is done numerically using MATLAB. The set of valid solutions to this problem are those which satisfy the constraint

$$w(r_*) = W + k(AJ_0(kr_*) + BY_0(kr_*)) = 0$$

The plot figure 0.2.1 shows the k, r_* combinations which satisfy the constraint.

Clearly this can only occur for values of kR > 3.8.

The first branch of this (extending from $kR \approx 3.8$) corresponds to natural solutions, whereas further branches give unwanted behaviour, which introduce reversed flow.

Code - homogeneousODE.m

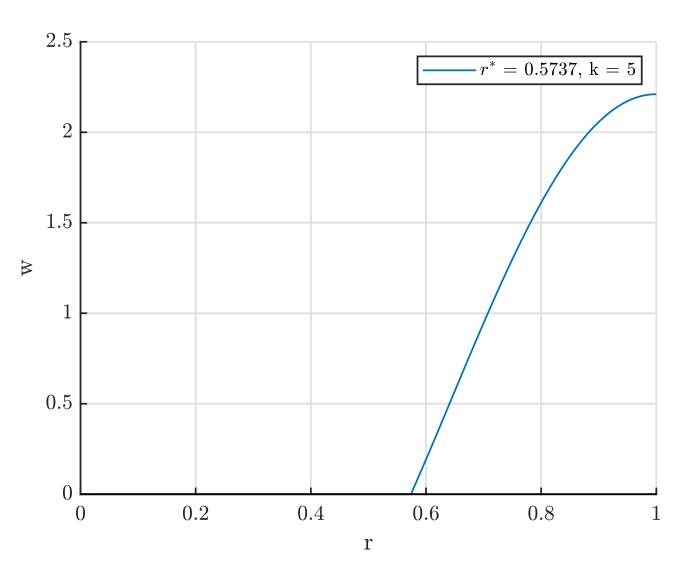


Figure 1: An example solution plot

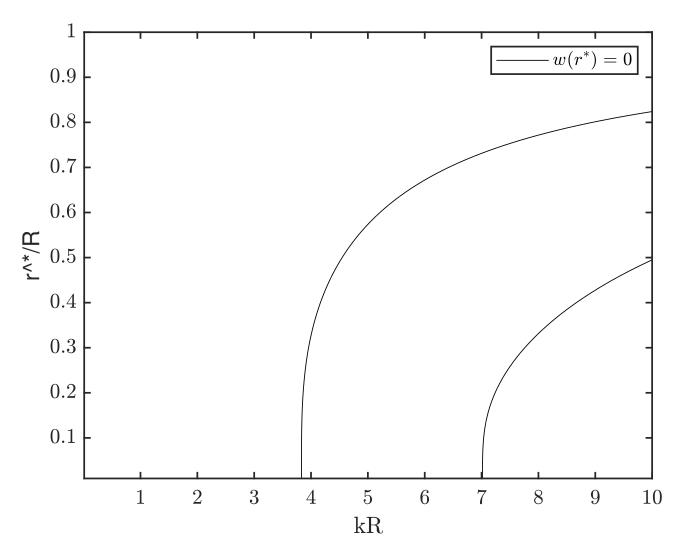


Figure 2: Solution set for the simplified problem

have to assume things for outside of the region for Ψ . I.e. if we go above the maximum input value then some assumption, and if we go below the minimum then it is a stagnation point

see if we can do it for the wall stagnation zones (i.e. psi goes to 0 near R) so when $\Psi > \frac{1}{2}WR^2$ Plug it into H and C

$$H = (\Omega R)^2 + \frac{1}{2}W^2$$

$$\frac{\partial H}{\partial \psi} = 0$$

$$C = \Omega R^2$$

$$\frac{\partial C}{\partial \Psi} = 0$$

Which then yields the separable first order ODE

$$\frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} = 0$$

And hence

$$\frac{\partial \Psi}{\partial r} = Ar$$

$$\Psi = \frac{1}{2}Ar^2 + B$$

our left hand side could be written as

$$r\frac{\partial}{\partial r} \left(\frac{1}{r} \frac{\partial \Psi}{\partial r} \right)$$

using staggered grid

$$\frac{\partial^2 \Psi}{\partial r^2} = \frac{1}{r} \frac{\partial \Psi}{\partial r}$$

at the boundary r=0

0.2.2 Numerics

Solving the ODE numerically:

$$\frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} = r^2 \frac{\partial H}{\partial \Psi} + C \frac{\partial C}{\partial \Psi}$$

finite difference - divide r as a grid of N intervals. So our grid spaces over R,

$$r_i = \Delta r_i, \quad \Delta = \frac{R}{N}$$

So (check this)

$$\frac{\partial^2 \Psi}{\partial r^2} = \frac{\Psi_{i+1} - 2\Psi_i + \Psi_{i-1}}{\Delta^2}$$
$$\frac{\partial \Psi}{\partial r} = \frac{\Psi_{i+1} - \Psi_{i-1}}{2\Delta}$$
$$\Psi_0 = 0, \quad \Psi_N = \frac{1}{2}WR^2$$

Which should work for the index i until we reach the bifurcations/stagnations Should end up with a matrix equation

$$\begin{pmatrix} 1 & 0 & \dots & 0 & 0 \\ & \mathbf{A} & & \\ 0 & 0 & \dots & 0 & 1 \end{pmatrix} \begin{pmatrix} \Psi_0 \\ \mathbf{\Psi} \\ \Psi_N \end{pmatrix} = \begin{pmatrix} 0 \\ \mathbf{f} \\ \frac{1}{2}WR^2 \end{pmatrix}$$

A should be the finite difference version of

$$\frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} = 0$$

I.e. for the i^{th} row of **A**

$$A(i) = \frac{A(i+1) - 2 * A(i) + A(i-1)}{\Delta^2} - \frac{A(i+1) - A(i-1)}{2r(i)\Delta}$$

$$A_{ij} = \begin{cases} 1 & j = i = 1\\ 1/\Delta^2 + 1/(2r_i\Delta) & j = i - 1\\ 2/\Delta^2 & j = i\\ 1/\Delta^2 - 1/(2r_i\Delta) & j = i + 1\\ 1 & j = i = N\\ 0 & otherwise \end{cases}$$

For the full equation

$$\begin{split} \frac{\partial^2 \Psi}{\partial z^2} + \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} + \Psi (\frac{4\Omega^2}{W^2} - \frac{1}{r^2}) &= 0 \\ \Psi = \frac{1}{2} W r^2 + r F \\ F = \frac{\Psi}{r} - \frac{1}{2} W r \end{split}$$

Boundary conditions for F relate to those for Ψ

$$\Psi(R) = \frac{1}{2}WR^2 \implies F(R) = 0$$

$$\Psi(r_*) = 0 \implies F(r_*) = \frac{1}{2}Wr_*^2$$

when we look at the vortex breakdown problem, introduce a coordinate transformation

$$\begin{split} \eta &= \frac{r-r_*}{R-r_*} \\ \eta &= 0, r = r_*, \eta = 1 r = R \\ \frac{\partial \Psi}{\partial r} &= \frac{\partial \Psi}{\partial \eta} \frac{\partial \eta}{\partial r} = \frac{1}{R-r_*} \frac{\partial \Psi}{\partial \eta} \\ \frac{\partial^2 \Psi}{\partial r^2} &= \frac{1}{(R-r_*)^2} \frac{\partial^2 \Psi}{\partial \eta^2} \end{split}$$

use the same conditions we have used anyway where $\Psi(r_*) = w(r_*) = 0$ Rankine body problem: At some point on the radius r_0 , we get $v = K/r_0$ for some constant K find $K = \Omega r_0^2$?

0.2.3 Rankine Body

w = W,

$$v = \begin{cases} \frac{\Gamma}{2\pi r}, & r > r_0\\ \Omega r, & r \le r_0 \end{cases}$$

Where the second condition was the previous solution. Since the velocity profile is now piecewise defined, the streamfunction must also be, i.e. it is necessary to split the streamfunction into 2 regions to solve this problem. The upstream regions:

$$\begin{cases} \Psi_{inner}, & 0 \le r \le r_0 \\ \Psi_{outer}, & r_0 \le r \le R \end{cases}$$

Note that r_0 is defined upstream, so the position of the region may have moved downstream to a new radius, \hat{r} , and hence, downstream, these regions will become around \hat{r} instead of r_0 . We enforce some similar conditions as to the normal problem:

$$\Psi(r_*) = 0,$$

$$\Psi(R) = \frac{1}{2}WR^2,$$

$$w(r_*) = 0$$

With the added condition that Ψ must remain continuous around \hat{r} I.e.

$$\lim_{r^- \to \hat{r}} \Psi(r^-) = \lim_{r^+ \to \hat{r}} \Psi(r^+)$$

And

$$\lim_{r^- \to \hat{r}} v(r^-) = \lim_{r^+ \to \hat{r}} v(r^+)$$

Where $\Psi(r^-)$ is Ψ defined for $r \leq \hat{r}$ and $\Psi(r^+)$ is defined in the region $r \geq \hat{r}$. The region for $\Psi(r)$ with $r \in [0, r_0]$ will be the same as before, i.e.

$$\Psi(r) = \frac{1}{2}Wr^2 + r(AJ_1(kr) + BY_1(kr))$$

For the region $r_0 < r < R$ the problem must be resolved from the SL equation

$$\begin{split} \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi} \\ C &= rv = \frac{\Gamma}{2\pi} \\ \frac{dC}{d\Psi} &= 0 \end{split}$$

$$\begin{split} H &= \frac{1}{2}(u^2 + v^2 + w^2) + \frac{p}{\rho} \\ &= \frac{1}{2}(0 + \frac{\Gamma^2}{4\pi^2r^2} + W^2) + \int \frac{C^2}{r^3} dr \\ &= \frac{1}{2}(\frac{\Gamma^2}{4\pi^2r^2} + W^2) + \int \frac{\Gamma^2}{4\pi^2r^3} dr \\ &= \frac{1}{2}(\frac{\Gamma^2}{4\pi^2r^2} + W^2) - \frac{\Gamma^2}{8\pi^2r^2} \\ &= \frac{W^2}{2} \\ \frac{dH}{d\Psi} &= 0 \end{split}$$

And hence the SL equation gives

$$\begin{split} \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi} \\ \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi} \\ \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= 0 \end{split}$$

Which results in:

$$\Psi = Cr^2 + D, \quad r \ge \hat{r}$$
$$w = \frac{1}{r} \frac{\partial \Psi}{\partial r} = 2C$$

With the requirement that there is no discontinuity at \hat{r} , i.e.

$$\Psi = \frac{1}{2}W\hat{r}^2 + \hat{r}(AJ_1(k\hat{r}) + BY_1(k\hat{r})) = C\hat{r}^2 + D$$

And using the same for w

$$w(\hat{r}) = W + k(AJ_0(k\hat{r}) + BY_0(k\hat{r})) = 2C$$

And lastly the wall condition

$$\Psi(R) = \frac{1}{2}WR^2 = C\hat{r}^2 + D$$

With

$$w(r_*) = 0$$

$$\frac{\Gamma}{2\pi r_0} = \Omega r_0 \implies \Omega = \frac{\Gamma}{2\pi r_0^2}$$

$$k_{outer} = \frac{2\Gamma}{2\pi W r_0^2} = \frac{\Gamma}{\pi W r_0^2}$$

Noting that the values for A and B are obtained from the r_* condition.

The coefficients for Ψ have to be resolved, since the condition $\Psi_{inner}(R) = \frac{1}{2}WR^2$ cannot be imposed.

Parameters

$$r_0, \hat{r}, r_*, R, k, \Gamma, W, A, B, C, D$$

We can fix r_0 , R, k, W and Γ . This is 11 parameters, where 5 are fixed. Require 6 conditions. Impose:

- 1). $w(r_*) = 0$ (as before)
- 2). $\Psi_{inner}(r_*) = 0$ (as before)
- 3). Since at the wall Ψ must remain the same, this applies to where v is changed, i.e. $\Psi_{inner}(\hat{r}) = \frac{1}{2}Wr_0^2$
- 4). For continuity, $\Psi_{outer}(\hat{r}) = \frac{1}{2}Wr_0^2$

5).
$$w_{outer}(\hat{r}) = w_{inner}(\hat{r})$$

6).
$$\Psi_{outer}(R) = \frac{1}{2}WR^2$$

Redo the problem instead getting A, B from 2) and 3)

$$\Psi_{inner}(r_*) = 0$$

$$\Psi_{inner}(\hat{r}) = \frac{1}{2}Wr_0^2$$

Use this for A, B

$$\Psi_{inner}(r_*) = \frac{1}{2}Wr_*^2 + r_*(AJ_1(kr_*) + BY_1(kr_*)) = 0$$

$$= r_*(AJ_1(kr_*) + BY_1(kr_*)) = -\frac{1}{2}Wr_*^2$$

$$\Psi_{inner}(\hat{r}) = \frac{1}{2}W\hat{r}^2 + \hat{r}(AJ_1(k\hat{r}) + BY_1(k\hat{r})) = \frac{1}{2}Wr_0^2$$

$$= \hat{r}(AJ_1(k\hat{r}) + BY_1(k\hat{r})) = \frac{1}{2}W(r_0^2 - \hat{r}^2)$$

This gives the matrix system for A, B below. Note that the system relies on the unknowns r_* and \hat{r} .

$$\begin{pmatrix} r_* J_1(kr_*) & r_* Y_1(kr_*) \\ \hat{r} J_1(k\hat{r}) & \hat{r} Y_1(k\hat{r}) \end{pmatrix} \begin{pmatrix} A \\ B \end{pmatrix} = \begin{pmatrix} -\frac{1}{2}Wr_*^2 \\ \frac{1}{2}W(r_0^2 - \hat{r}^2) \end{pmatrix}$$
$$\begin{pmatrix} A \\ B \end{pmatrix} = \frac{1}{\det} \begin{pmatrix} \hat{r} Y_1(k\hat{r}) & -r_* Y_1(kr_*) \\ -\hat{r} J_1(k\hat{r}) & r_* J_1(kr_*) \end{pmatrix} \begin{pmatrix} -\frac{1}{2}Wr_*^2 \\ \frac{1}{2}W(r_0^2 - \hat{r}^2) \end{pmatrix}$$

$$A = \frac{1}{\det} \left(\hat{r} Y_1(k\hat{r}) \left(-\frac{1}{2} W r_*^2 \right) - r_* Y_1(kr_*) \left(\frac{1}{2} W (r_0^2 - \hat{r}^2) \right) \right)$$

$$B = \frac{1}{\det} \left(-\hat{r} J_1 \left(-\frac{1}{2} W r_*^2 \right) + r_* J_1(kr_*) \left(\frac{1}{2} W (r_0^2 - \hat{r}^2) \right) \right)$$

Where

$$\det = \hat{r}r_*Y_1(k\hat{r})J_1(kr_*) - \hat{r}r_*J_1(k\hat{r})Y_1(kr_*)$$
$$= \hat{r}r_*(Y_1(k\hat{r})J_1(kr_*) - J_1(k\hat{r})Y_1(kr_*))$$

This for r_*

$$w_{inner}(r_*) = W + k(AJ_0(kr_*) + BY_0(kr_*)) = 0$$

Get C from:

$$w_{outer}(\hat{r}) = w_{inner}(\hat{r})$$
$$2C = W + k(AJ_0(k\hat{r}) + BY_0(k\hat{r}))$$
$$C = \frac{1}{2}(W + k(AJ_0(k\hat{r}) + BY_0(k\hat{r})))$$

Get D here:

$$\Psi_{outer}(R) = CR^2 + D = \frac{1}{2}WR^2$$
$$D = \frac{1}{2}WR^2 - C$$

Hence get \hat{r} from

$$\Psi_{outer}(\hat{r}) = C\hat{r}^2 + D = \frac{1}{2}Wr_0^2$$

$$C\hat{r}^2 + \frac{1}{2}WR^2 - C = \frac{1}{2}Wr_0^2$$

$$\left(\frac{1}{2}(W + k(AJ_0(k\hat{r}) + BY_0(k\hat{r})))\right)(\hat{r}^2 - 1) = \frac{1}{2}W(r_0^2 - R^2)$$

$$(AJ_0(k\hat{r}) + BY_0(k\hat{r}))(\hat{r}^2 - 1) = \frac{1}{k}W(r_0^2 - R^2 - 1)$$

For physically valid solutions, we must impose the condition of no net change on the momentum from upstream to downstream on the momentum (Escudier, Keller). The momentum is defined as

$$s = 2\pi \int_0^{r_t} \left(\rho w^2 + p\right) r dr$$

Which comes to:

$$\Delta s = \frac{\pi}{4} \rho U^2 k^2 r_c^2 \left[-r_b^2 + \frac{1}{4} \left(\frac{r_b^4 - r_a^4}{r_c^2} \right) + \frac{3}{4} r_c^2 + \frac{1}{2} r_c^2 \log \left(\frac{r_b^2}{r_c^2} \right) \right] = 0$$

Figure 4 shows the solution set for the problem. It displays the same results as those found in (Escudier, Keller), with the same asymptote $kr_0 \to \sqrt{2}$ as $r_0 \to 0$. The results in the two figures come from rhatrstarmomentum.m and numericalSolutionSetRankine.m respectively

0.2.4 Lamb-Oseen Vortex

Q-vortex without a Jet. Start with

$$w = W$$

$$v = \frac{\Gamma}{2\pi r} \left(1 - e^{-r^2/\delta^2} \right)$$

$$\frac{d^2 \Psi}{dr^2} - \frac{1}{r} \frac{d\Psi}{dr} = r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi}$$

Solve from r_* to R numerically.

Generate grid from r_* to R.

Boundary conditions as normal

$$\Psi(R) = \frac{1}{2}WR^2$$

$$\Psi(r_*) = 0$$

$$w(r_*) = 0$$

And the standard upstream flow

$$\Psi(r) = \frac{1}{2}Wr^2$$

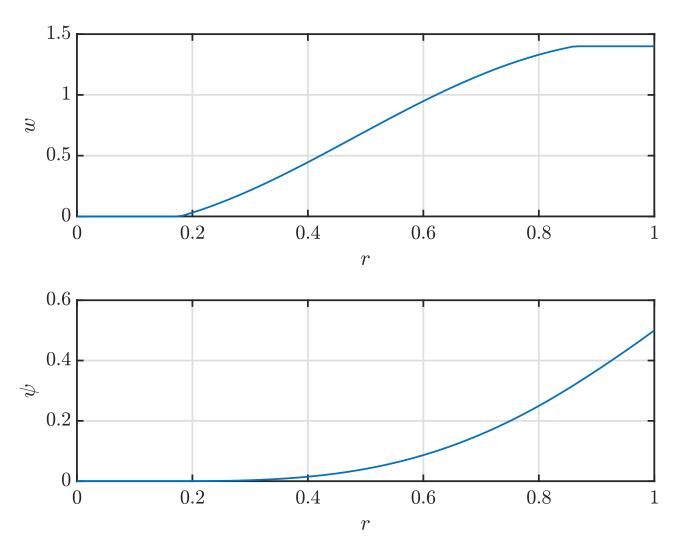


Figure 3: A solution of w and Ψ for the Rankine problem with 0 net momentum, $k = 3.8961, r_0 = 0.8619, r_* = 0.1784$. Obtained using

Non-dimensional parameter may be something like $\frac{\Gamma}{WR}$ (we can probably relate this to kr_0 for the rankine problem)

Eventually do the same thing as before with s and Δs .

Use

$$s = \int_0^R (\rho w^2 + p) r dr = \int_0^{r_*} p(r_*) r dr + \int_{r_*}^R (\rho w^2 + p) r dr$$
$$\frac{1}{\rho} \frac{\partial p}{\partial r} = \frac{v^2}{r} = \frac{\Gamma^2}{4\pi^2 r^3 +} \left(1 - e^{-r^2/\delta^2}\right)^2$$
$$\Psi = \frac{1}{2} W r^2 \implies r = \sqrt{\frac{2\Psi}{W}}$$

Rankine body solution space

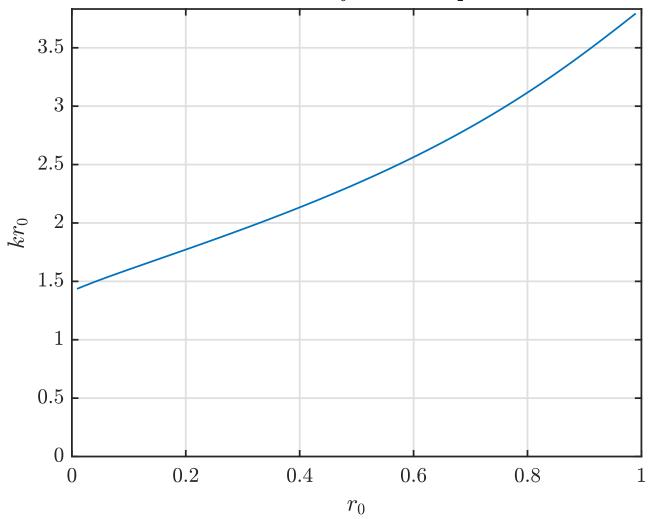


Figure 4: Solution space for the Rankine body problem

$$C = rv = \frac{\Gamma}{2\pi} \left(1 - e^{-r^2/\delta^2} \right)$$

$$\frac{\partial C}{\partial \Psi} = \frac{\Gamma}{2\pi} \frac{\partial}{\partial \Psi} \left(1 - e^{-r^2/\delta^2} \right)$$

$$= \frac{-\Gamma}{2\pi} \frac{\partial}{\partial \Psi} \left(e^{-r^2/\delta^2} \right)$$

$$= \frac{-\Gamma}{2\pi} \frac{\partial}{\partial \Psi} \left(e^{-2\Psi/W\delta^2} \right)$$

$$= \frac{\Gamma}{W\delta^2\pi} \left(e^{-2\Psi/W\delta^2} \right)$$

$$= \frac{\Gamma}{W\pi\delta^2} e^{-r^2/\delta^2}$$

$$\begin{split} &\frac{dH}{d\Psi} = \frac{dH}{dr} \frac{dr}{d\Psi} \\ &= \frac{dr}{d\Psi} \frac{d}{dr} \left(\frac{1}{2} \left(u^2 + v^2 + w^2 \right) + \frac{p}{\rho} \right) \\ &= \frac{1}{\sqrt{2W\psi}} \frac{d}{dr} \left(\frac{1}{2} v^2 + \int \frac{C^2}{r^3} dr \right) \\ &= \frac{1}{Wr} \left(\frac{1}{2} \frac{dv^2}{dr} + \frac{v^2}{r} \right) \\ &= \frac{1}{Wr} \left(\frac{\Gamma^2}{4r\pi^2} \left(\frac{-1}{r^2} + 2e^{-r^2/\delta^2} \left(\frac{1}{r^2} + \frac{1}{\delta^2} \right) - e^{-2r^2/\delta^2} \left(\frac{1}{r^2} + \frac{2}{\delta^2} \right) \right) + \frac{\Gamma^2}{4r\pi^2} \left(\frac{1 - 2e^{-r^2/\delta^2} + e^{-2r^2/\delta^2}}{r^2} \right) \right) \\ &= \frac{\Gamma^2}{4Wr^2\pi^2} \left(2e^{-r^2/\delta^2} \left(\frac{1}{r^2} + \frac{1}{\delta^2} \right) - e^{-2r^2/\delta^2} \left(\frac{1}{r^2} + \frac{2}{\delta^2} \right) + \frac{-2e^{-r^2/\delta^2} + e^{-2r^2/\delta^2}}{r^2} \right) \\ &= \frac{\Gamma^2}{2Wr^2\delta^2\pi^2} \left(e^{-r^2/\delta^2} - e^{-2r^2/\delta^2} \right) \\ &= \frac{\Gamma^2}{4\Psi\delta^2\pi^2} \left(e^{-2\Psi/W\delta^2} - e^{-4\Psi/W\delta^2} \right) \end{split}$$

How I got the middle term:

$$\begin{split} \frac{dv^2}{dr} &= \frac{d}{dr} \left(\frac{\Gamma}{2\pi r} \left(1 - e^{-r^2/\delta^2} \right) \right)^2 \\ &= \frac{\Gamma^2}{4\pi^2} \frac{d}{dr} \left(\frac{1 - 2e^{-r^2/\delta^2} + e^{-2r^2/\delta^2}}{r^2} \right) \\ &= \frac{\Gamma^2}{4\pi^2} \left(\frac{-2}{r^3} - 2 \left(-\frac{2e^{-r^2/\delta^2}}{r^3} - \frac{2e^{-r^2/\delta^2}}{r\delta^2} \right) + \left(-\frac{2e^{-2r^2/\delta^2}}{r^3} - \frac{4e^{-2r^2/\delta^2}}{r\delta^2} \right) \right) \\ &= \frac{\Gamma^2}{2r\pi^2} \left(\frac{-1}{r^2} + 2e^{-r^2/\delta^2} \left(\frac{1}{r^2} + \frac{1}{\delta^2} \right) - e^{-2r^2/\delta^2} \left(\frac{1}{r^2} + \frac{2}{\delta^2} \right) \right) \end{split}$$

$$\begin{split} \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= r^2 \frac{dH}{d\Psi} - C \frac{dC}{d\Psi} \\ &= \frac{r^2 \Gamma^2}{4\Psi \delta^2 \pi^2} \left(e^{-2\Psi/W \delta^2} - e^{-4\Psi/W \delta^2} \right) - \left(\frac{\Gamma}{2\pi} \left(1 - e^{-2\Psi/W \delta^2} \right) \right) \left(\frac{\Gamma}{W \delta^2 \pi} \left(e^{-2\Psi/W \delta^2} \right) \right) \\ &= \frac{r^2 \Gamma^2}{4\Psi \delta^2 \pi^2} \left(e^{-2\Psi/W \delta^2} - e^{-4\Psi/W \delta^2} \right) - \frac{\Gamma^2}{2W \delta^2 \pi^2} \left(1 - e^{-2\Psi/W \delta^2} \right) \left(e^{-2\Psi/W \delta^2} \right) \\ &= \frac{\Gamma^2}{2W \delta^2 \pi^2} \left(\frac{r^2 W}{2\Psi} \left(e^{-2\Psi/W \delta^2} - e^{-4\Psi/W \delta^2} \right) - \left(e^{-2\Psi/W \delta^2} - e^{-4\Psi/W \delta^2} \right) \right) \\ \frac{\partial^2 \Psi}{\partial r^2} - \frac{1}{r} \frac{\partial \Psi}{\partial r} &= \frac{\Gamma^2}{2W \delta^2 \pi^2} \left(\left(\frac{r^2 W}{2\Psi} - 1 \right) \left(e^{-2\Psi/W \delta^2} - e^{-4\Psi/W \delta^2} \right) \right) \end{split}$$

Giving the system

$$\begin{split} \Psi_1' &= \Psi_2 \\ \Psi_2' &= \frac{1}{r} \Psi_2 + \frac{\Gamma^2}{2W\delta^2 \pi^2} \left(\left(\frac{r^2 W}{2\Psi_1} - 1 \right) \left(e^{-2\Psi_1/W\delta^2} - e^{-4\Psi_1/W\delta^2} \right) \right) \end{split}$$

Take limit as $\Psi \to 0$ in matlab and use that in some suff area.

Could use byp5c and break it into a system of first order odes.

So if we change the boundaries - using a 'Linear Lagrange interpolating polynomial' So that the end points are fixed.

$$\eta = \frac{r - r_*}{R - r_*}$$

$$r - r_* = \eta(R - r_*)$$
$$r = \eta(R - r_*) + r_*$$

Such that $\eta \in [0,1]$. In the function we can get r from η

For finite differences we can just use a newton iteration, or use something like fsolve.

I SHOULD PLOT DOWNSTREAM v for all our equations also

To use the r_* and w parts, we can use $w = \frac{1}{r} \frac{\partial \Psi}{\partial r}$

Try putting in the homogeneous solution to the solver to see if it works (i.e. the one with $v = \Omega r$ and w = W).

We should expect that the Lamb-Oseen vortex should be a more smooth version of the Rankine problem - so we should be able to compare the two.

May want to find $v_{max} = r_0$ to compare to the previous problem. Should expect the circulation goes to $\Gamma/2\pi$ as $r \to \infty$.

Of course we have to have

$$\lim_{r\to 0} \frac{1}{r} \frac{\partial \Psi}{\partial r} < \infty$$

So use l'hopital's rule

$$\lim_{r\to 0} \frac{1}{r} \frac{\partial \Psi}{\partial r} = \frac{\partial^2 \Psi}{\partial r^2}$$

But we know $\psi = 0$ at r = 0 (so we can ignore this)

So we will just use

$$\frac{\partial^2 \Psi}{\partial r^2}|_i - \frac{1}{r} \frac{\partial \Psi}{\partial r}|_i = f(r_i, \Psi_i)$$

With

$$\Psi_1 = 0$$

And

$$\Psi_n = \frac{1}{2}WR^2$$

We will have 1 based indexing Alternatively get the derivatives at the mid points

$$\frac{\partial \Psi}{\partial r}\Big|_{i+1/2} = \frac{\Psi_{i+1} - \Psi_i}{h_r}$$

Where h_r is the step.

Alternatively

Can rewrite the left hand side as

$$r\left(\frac{\partial}{\partial r}\left(\frac{1}{r}\frac{\partial\Psi}{\partial r}\right)\right)$$

And plug in the last thing

When we do finite differences, grab a computational variable (call it η for now)

$$\eta = \frac{r - r_*}{R - r_*}$$

So now we are computing in $0 < \eta < 1$. So try plugging it into the ODE:

$$d\eta = \frac{dr}{R - r_*}$$

$$\frac{\partial \Psi}{\partial r} = \frac{\partial \Psi}{\partial \eta} \frac{\partial \eta}{\partial r} = \frac{1}{R - r_*} \frac{\partial \Psi}{\partial \eta}$$

If we don't know r_* the η vector becomes

$$\begin{pmatrix} \eta_0 \\ \vdots \\ \eta_{N-1} \\ r_* \end{pmatrix}$$

This will be a non-linear problem so we will need to solve using fsolve.

To get guesses could use rankine vortex stuff -

For a linear flow we could use We know that $\psi(r_*) = 0$ and $\psi(R) = \frac{1}{2}WR^2$. We could guess that ψ is constant, $\psi = Ar^2 + B$.

Try plotting $w(r_*)$ for various r_* to help find guesses (require it to be 0)

If using η , the ODE becomes $\psi(\eta=0)=0$ and $\psi(\eta=1)=\frac{1}{2}WR^2$

$$\begin{split} \frac{\partial \Psi}{\partial r} &= \frac{1}{R - r_*} \frac{\partial \Psi}{\partial \eta} \\ \frac{\partial^2 \Psi}{\partial r^2} &= \frac{1}{(R - r_*)^2} \frac{\partial^2 \Psi}{\partial \eta^2} \\ r &= \eta (R - r_*)) + r_* \end{split}$$

DE becomes

$$\begin{split} \frac{1}{(R-r_*)^2} \frac{\partial^2 \Psi}{\partial \eta^2} - \frac{1}{\eta(R-r_*)) + r_*} \frac{1}{R-r_*} \frac{\partial \Psi}{\partial \eta} &= 0 \\ \frac{1}{R-r_*} \frac{\partial^2 \Psi}{\partial \eta^2} - \frac{1}{\eta(R-r_*)) + r_*} \frac{\partial \Psi}{\partial \eta} &= 0 \end{split}$$

0.2.5 Outer vortex breakdown

Considering the initial problem for vortex breakdown, except perhaps the breakdown is a pocket expanding from R rather than 0. I.e. the breakdown occurs about the wall rather than the center. So assuming r^{\dagger} is our outer vortex breakdown radius

This simply means obtaining a new A, B and k.

$$\Psi(r) = \frac{1}{2}Wr^2 + r(AJ_1(kr) + BY_1(kr))$$
$$w(r) = W + k(AJ_0(kr) + BY_0(kr))$$

Such that

$$w(r^{\dagger}) = 0, \qquad \Psi(0) = 0, \quad and \quad \Psi(r^{\dagger}) = 0$$

To enforce $\Psi(0)=0$ note that $\lim_{r\to 0}\frac{Y_1(kr)}{r}=-\infty$. Hence it is necessary to set B=0.

$$\Psi(r) = \frac{1}{2}Wr^2 + rAJ_1(kr), \quad w(r) = W + kAJ_0(kr)$$

And to enforce $\Psi(r^{\dagger}) = 0$

$$\implies Ar^{\dagger}J_1(kr^{\dagger}) = -\frac{1}{2}Wr^{\dagger 2}$$
$$A = \frac{-Wr^{\dagger}}{2J_1(kr^{\dagger})}$$

And obtain k using

$$w(r^{\dagger}) = 0$$
$$kAJ_0(kr) = -W$$

$$\Psi(r^{\dagger}) = \Psi(R) = \frac{1}{2}WR^2$$

0.3 Appendix

0.3.1 Supplementary Materials

This is where all the basic fluid mechanics knowledge should be (definitions, etc.)

0.3.2 Resources

Books: An Introduction to Fluid Dynamics Batchelor

Swirling flow states in finite-length diverging or contracting circular pipes Zvi Rusak Wall-separation and vortex-breakdown zones in a solid-body rotation flow in a rotating finite-length straight circular pipe Zvi Rusak, and Shixiao Wang