Photometric Stereo

Computer Vision
Fall 2019
Columbia University

Last Time: Two-View Stereo

Key Idea: use feature motion to understand shape

Today: Photometric Stereo

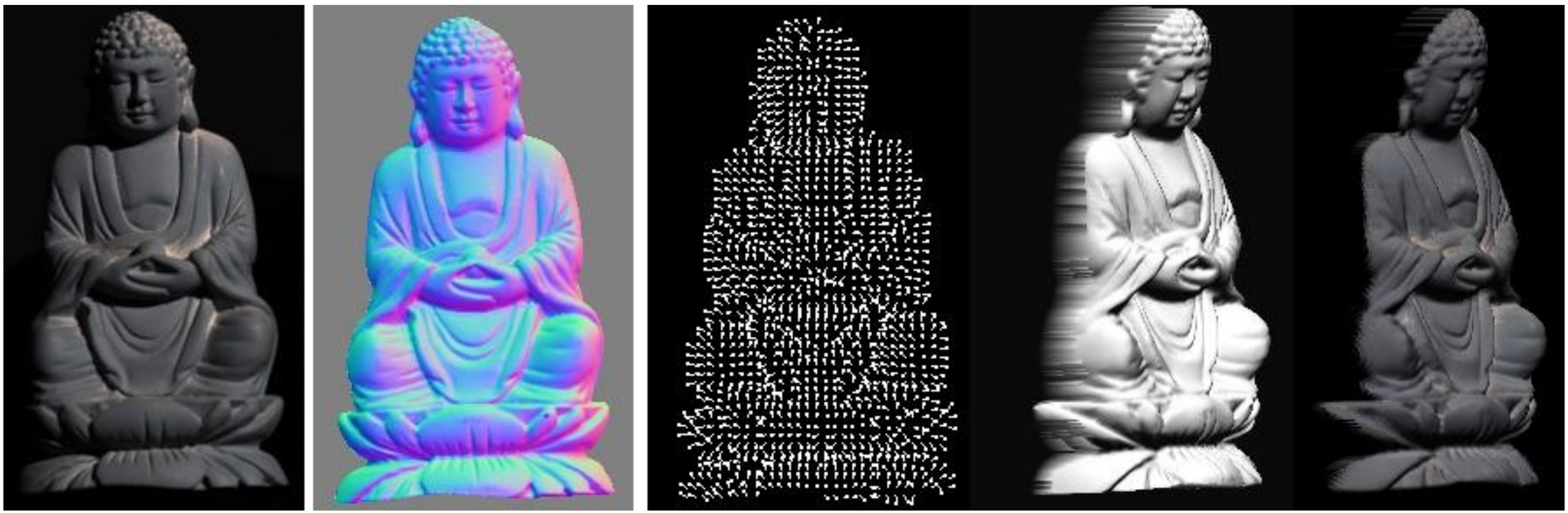
Key Idea: use pixel brightness to understand shape

Today: Photometric Stereo

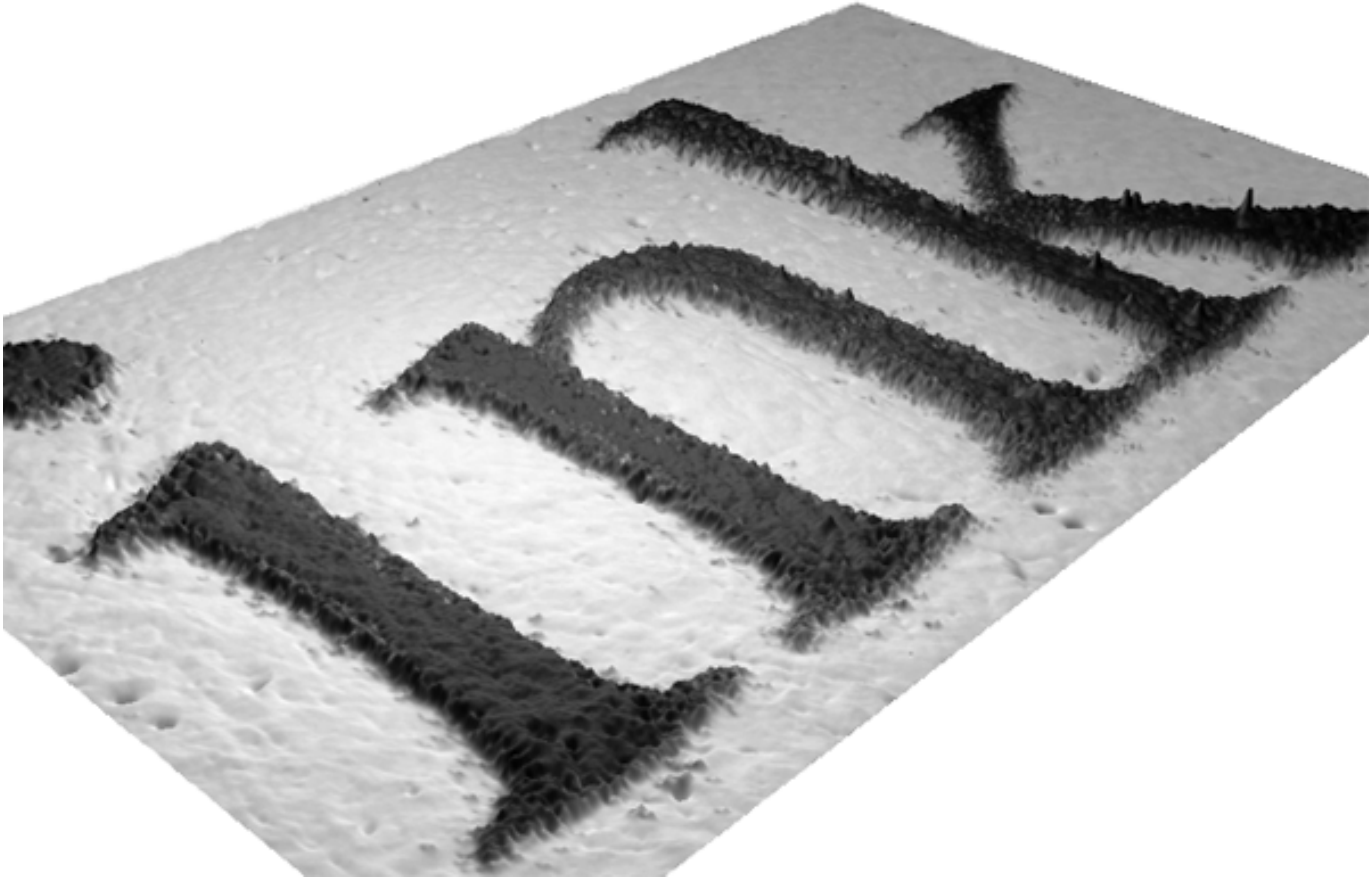
Key Idea: use pixel brightness to understand shape

Photometric Stereo

What results can you get?

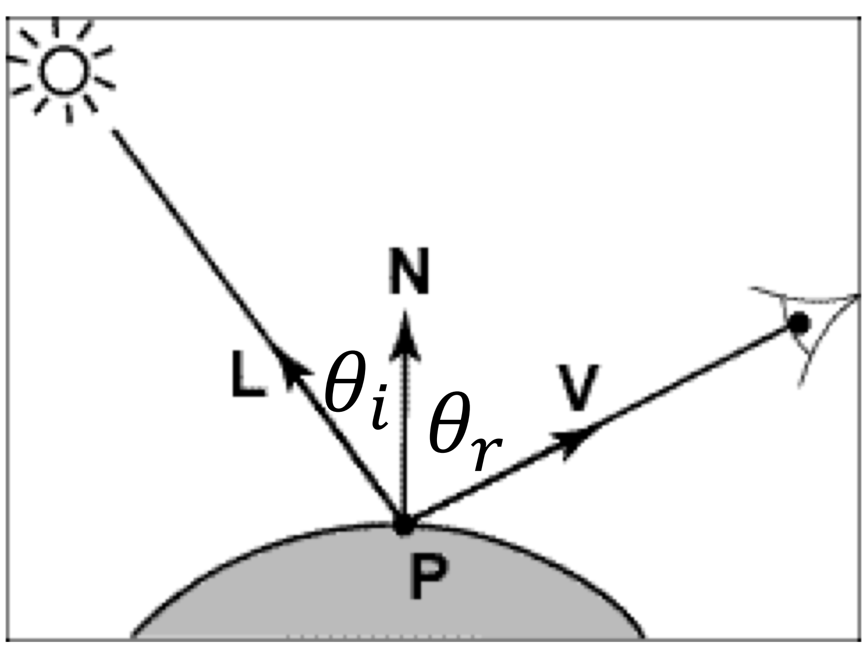


Input (1 of 12) | Normals (RGB colormap) | Normals (vectors) | Shaded 3D rendering | Textured 3D rendering


ink

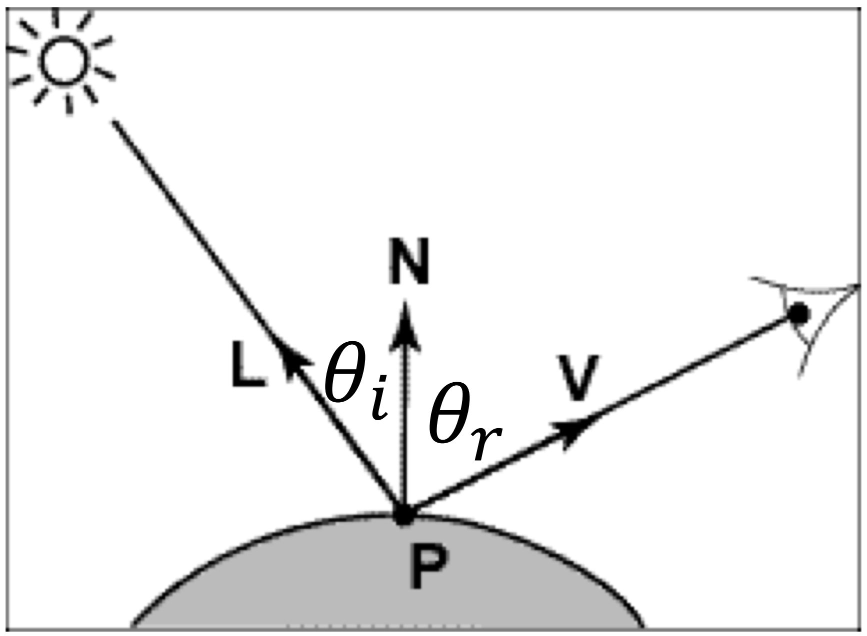
Light rays interacting with a surface

- Light of radiance L_i comes from light source at an incoming direction θ_i
- It sends out a ray of radiance L_r in the outgoing direction θ_r
- How does L_r relate to L_i ?



- **N** is surface normal
- **L** is direction of light, making θ_i with normal
- **V** is viewing direction, making θ_r with normal

Light rays interacting with a surface



- **N** is surface normal
- **L** is direction of light, making θ_i with normal
- **V** is viewing direction, making θ_r with normal

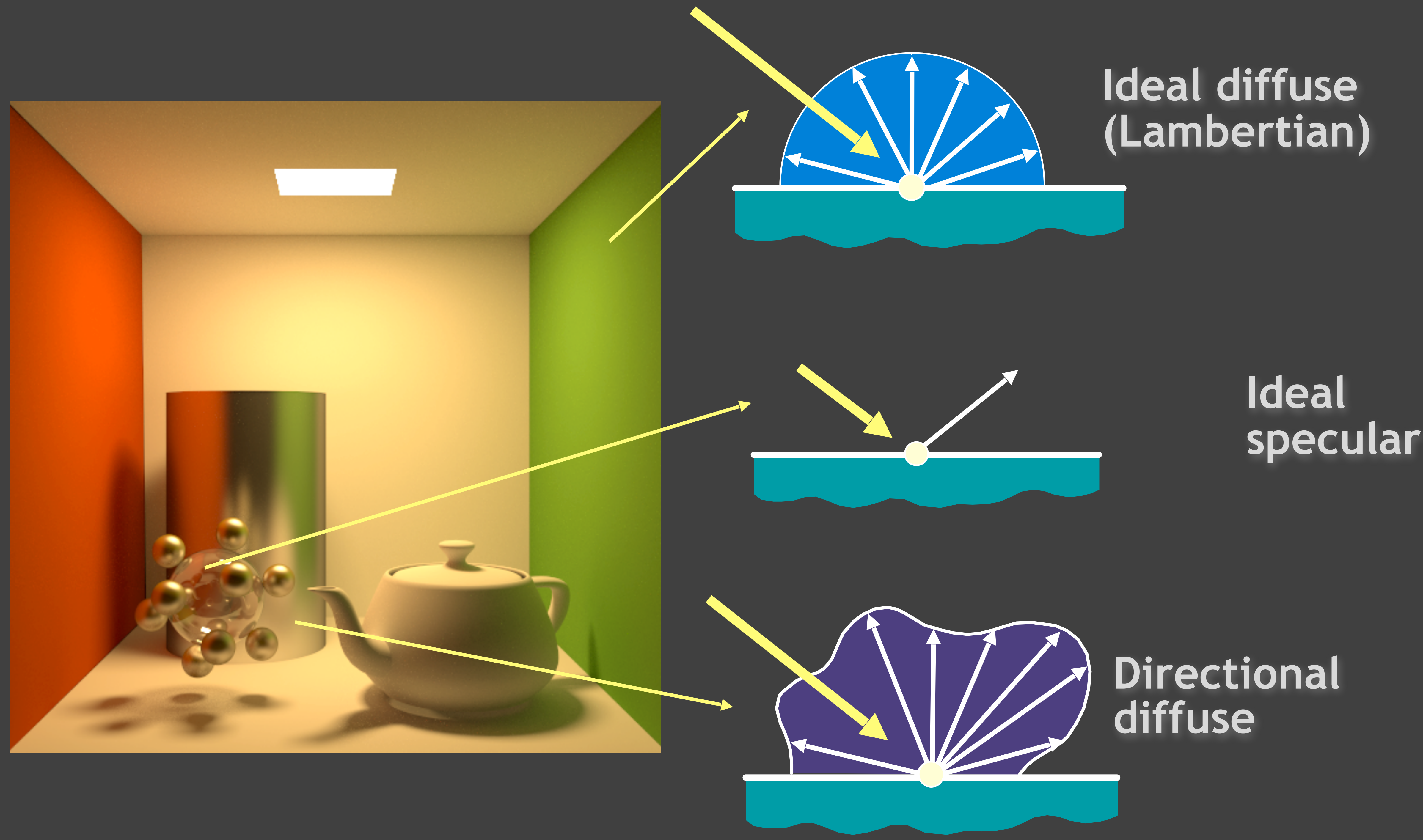
Output radiance along **V**

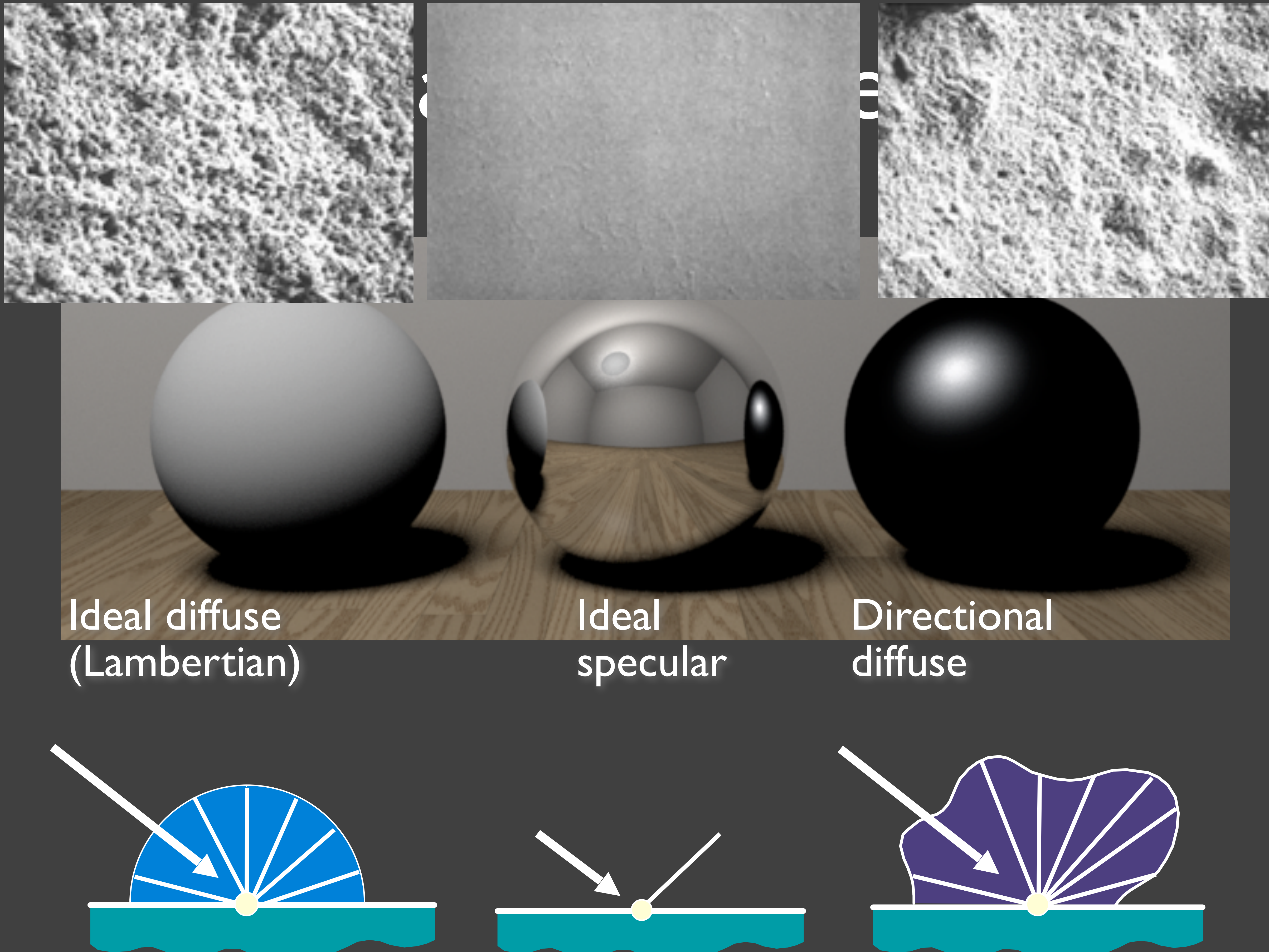
$$L_r = \rho(\theta_i, \theta_r) L_i \cos\theta_i$$

Incoming irradiance along **L**

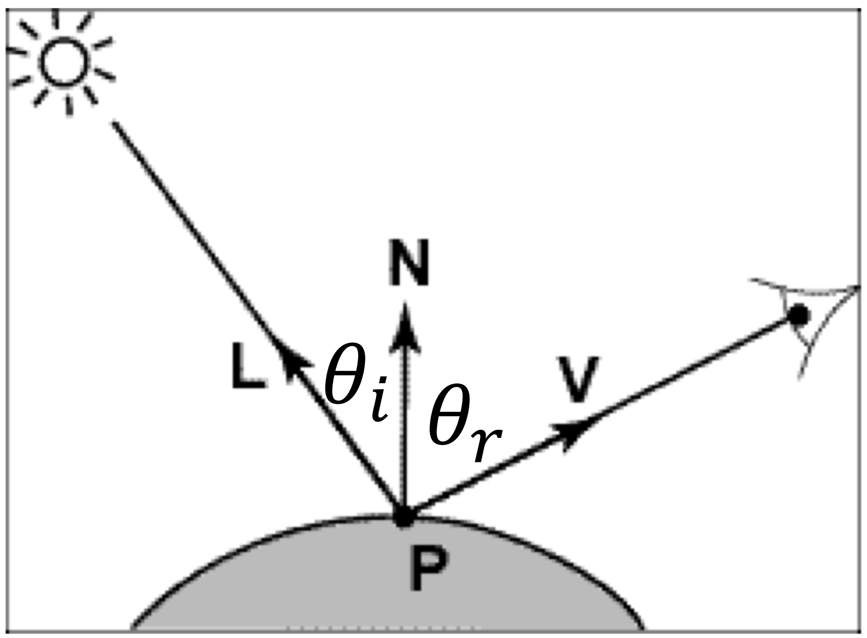
Bi-directional reflectance function (BRDF)

Materials - Three Forms




Ideal diffuse
(Lambertian)
Ideal
specular
Directional
diffuse

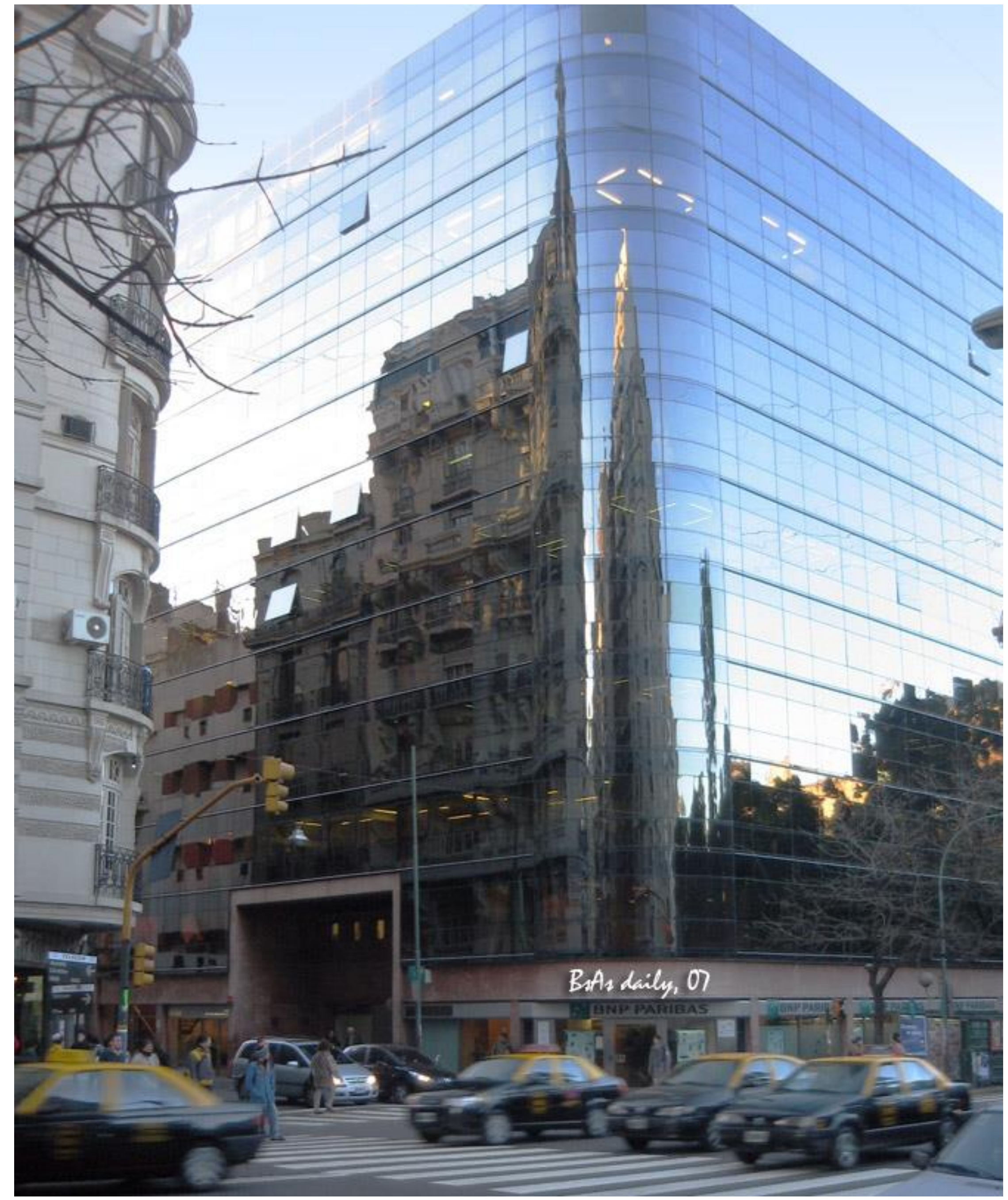
Light rays interacting with a surface



$$L_r = \rho(\theta_i, \theta_r) L_i \cos\theta_i$$

- Special case 1: Perfect mirror
  - $\rho(\theta_i, \theta_r)$ = 0 unless $\theta_i = \theta_r$
- Special case 2: Matte surface
  - $\rho(\theta_i, \theta_r)$ = ρ_0 (constant)

For now, ignore specular reflection



Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

And Refraction…

Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

And Interreflections…

Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

And Subsurface Scattering…

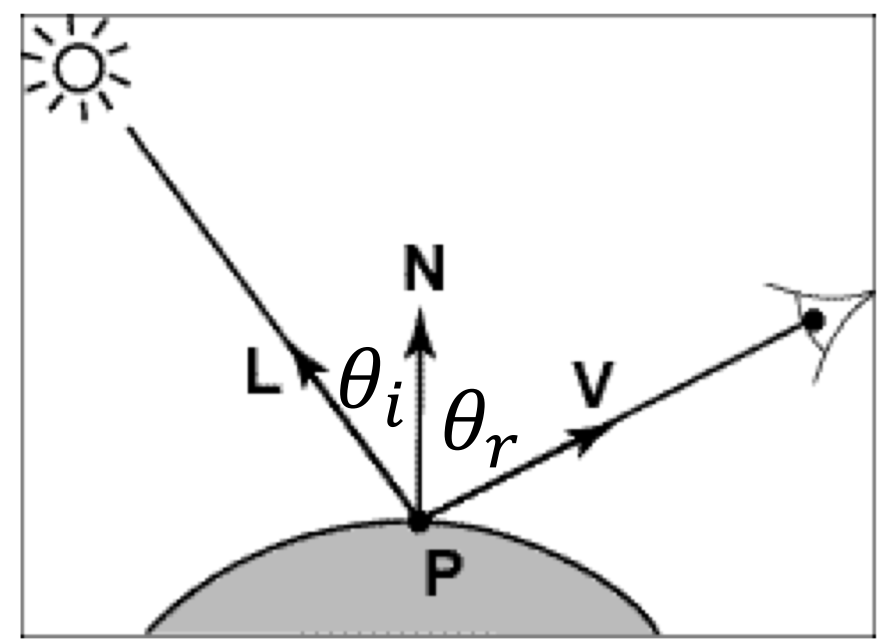
Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

Lambertian surfaces

- For a lambertian surface:

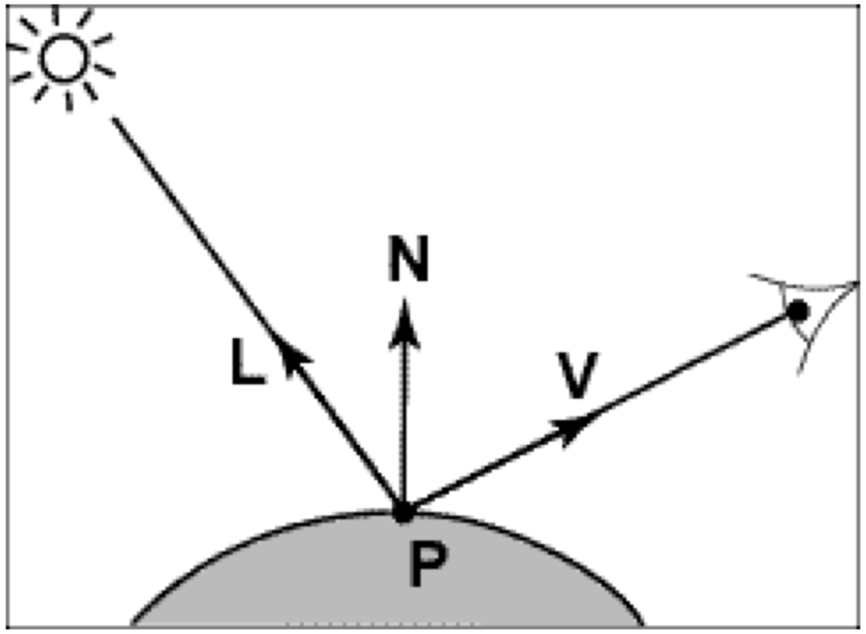
$$L_r = \rho L_i \cos \theta_i$$

$$\Rightarrow L_r = \rho L_i \mathbf{L} \cdot \mathbf{N}$$



- ρ is called *albedo*
  - Think of this as paint
  - High albedo: white colored surface
  - Low albedo: black surface
  - Varies from point to point

Lambertian surfaces



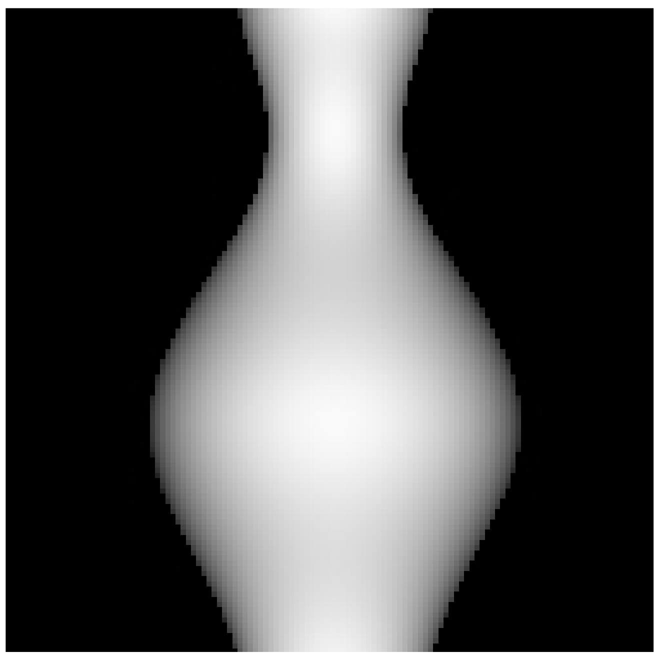
Intrinsic Image Decomposition

$$I(x,y) = \underbrace{\rho(x,y)}_{\text{Reflectance image}} \underbrace{L_i \mathbf{L} \cdot \mathbf{N}(x,y)}_{\text{Shading image}}$$

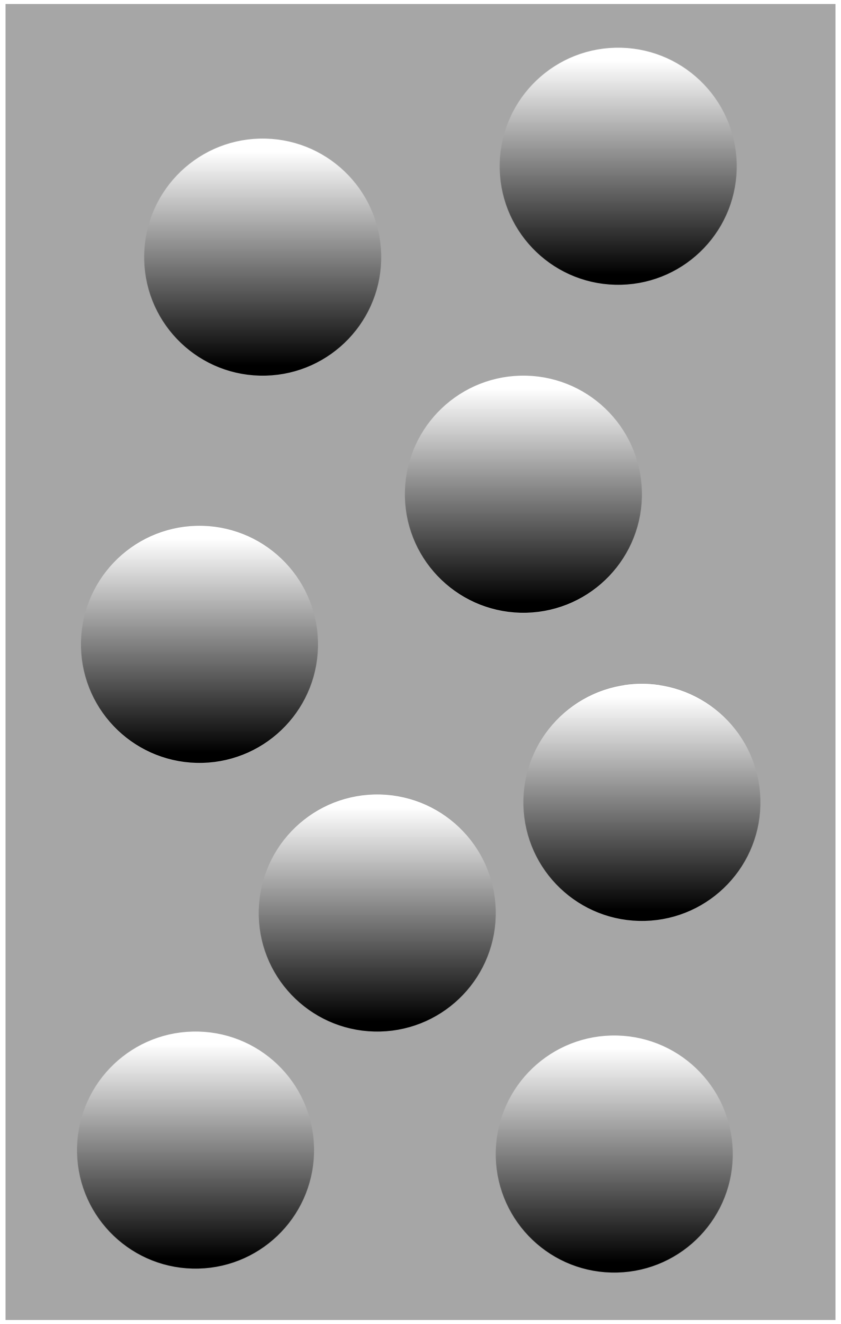
Reflectance image

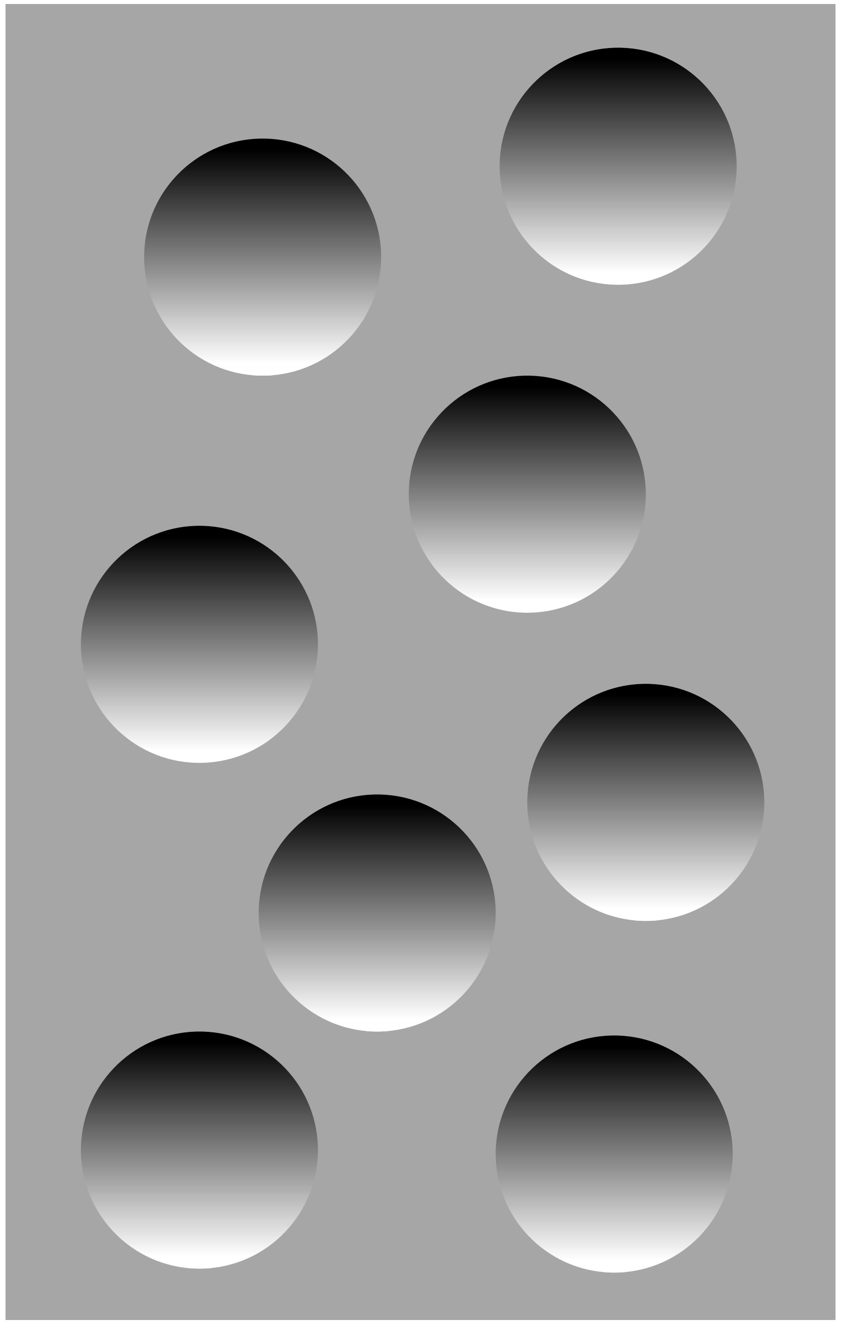
Shading image

Solution 1: Recovery from a single image

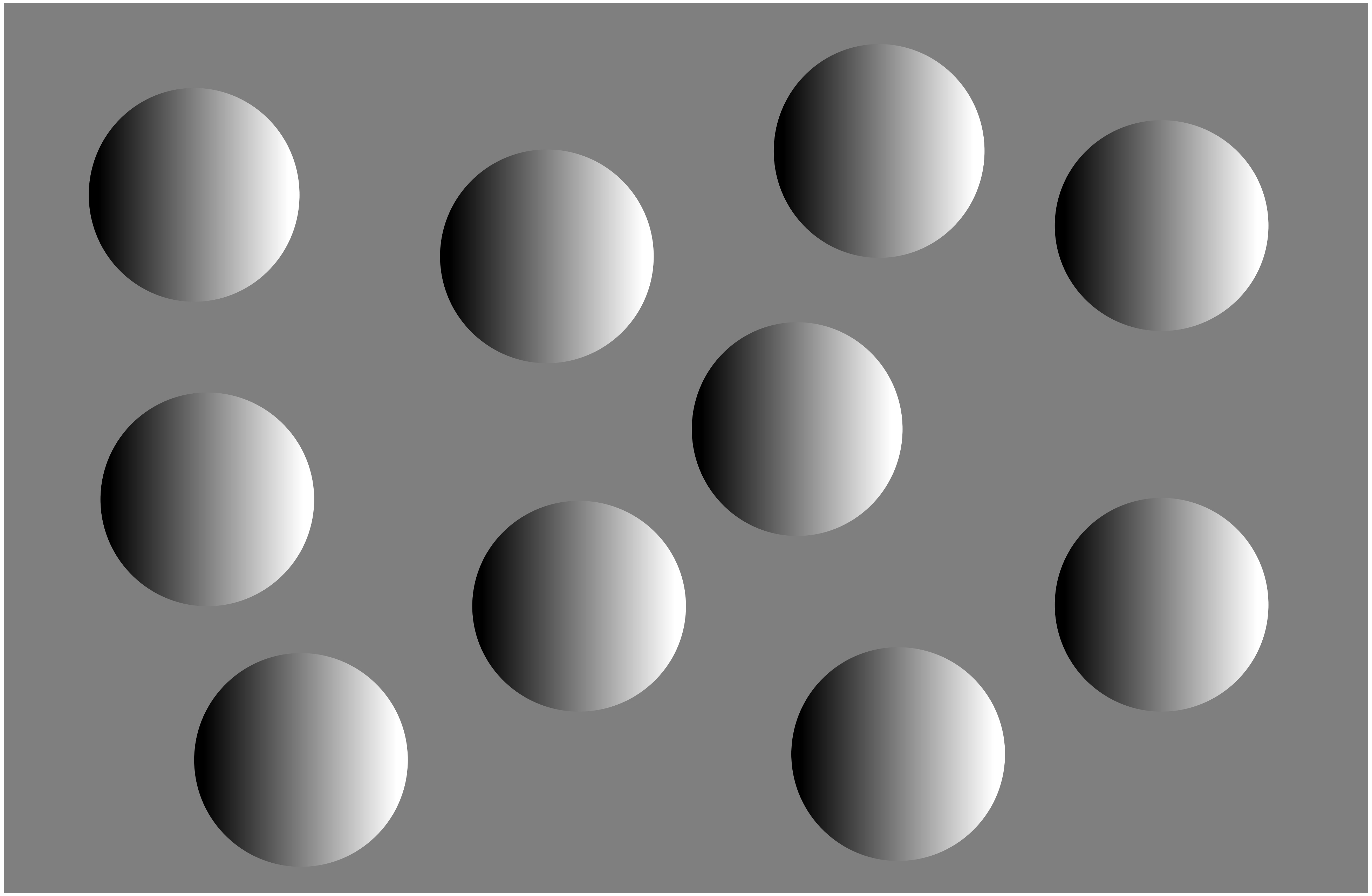


Solution 1: Recovery from a single image

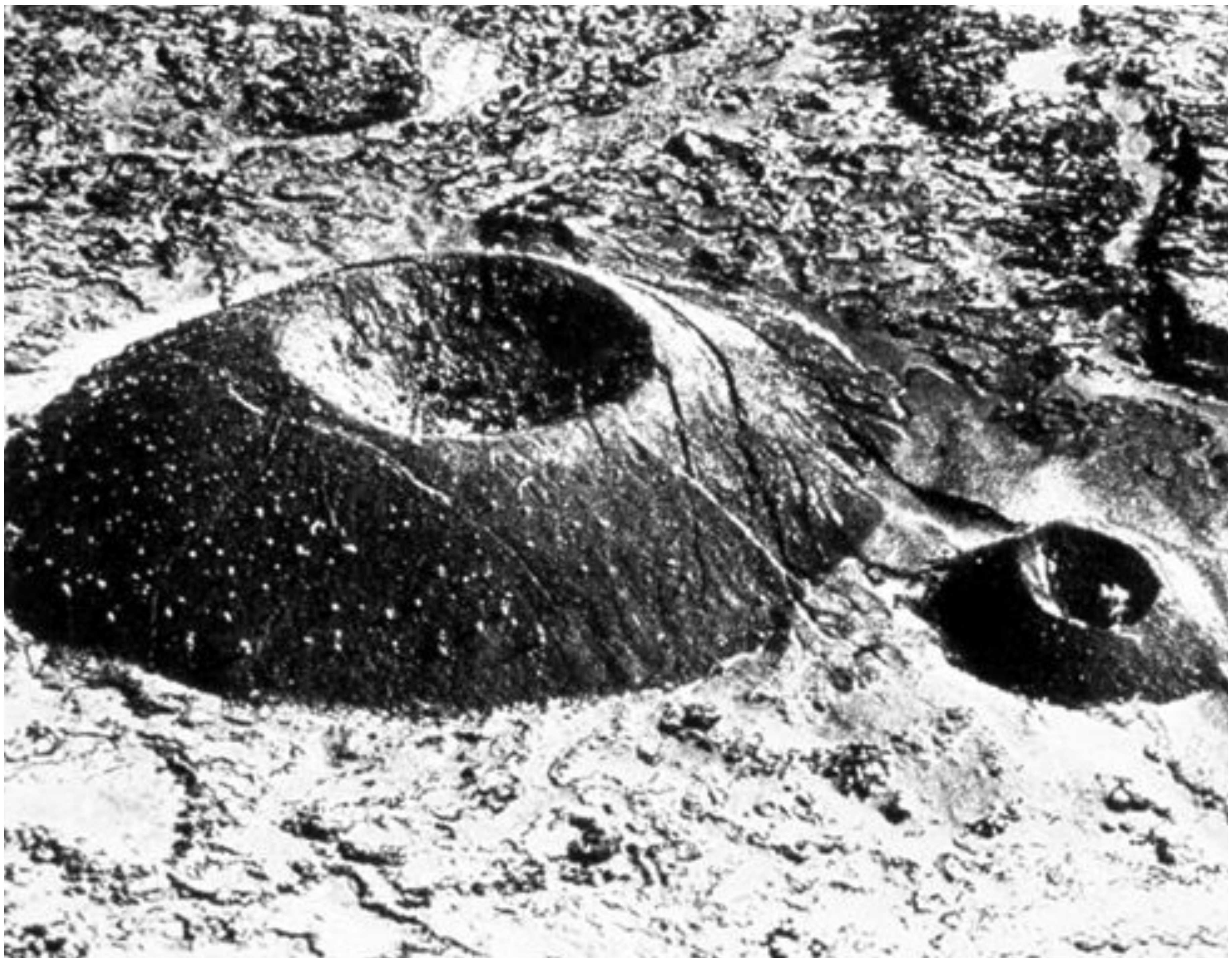




Solution 1: Recovery from a single image

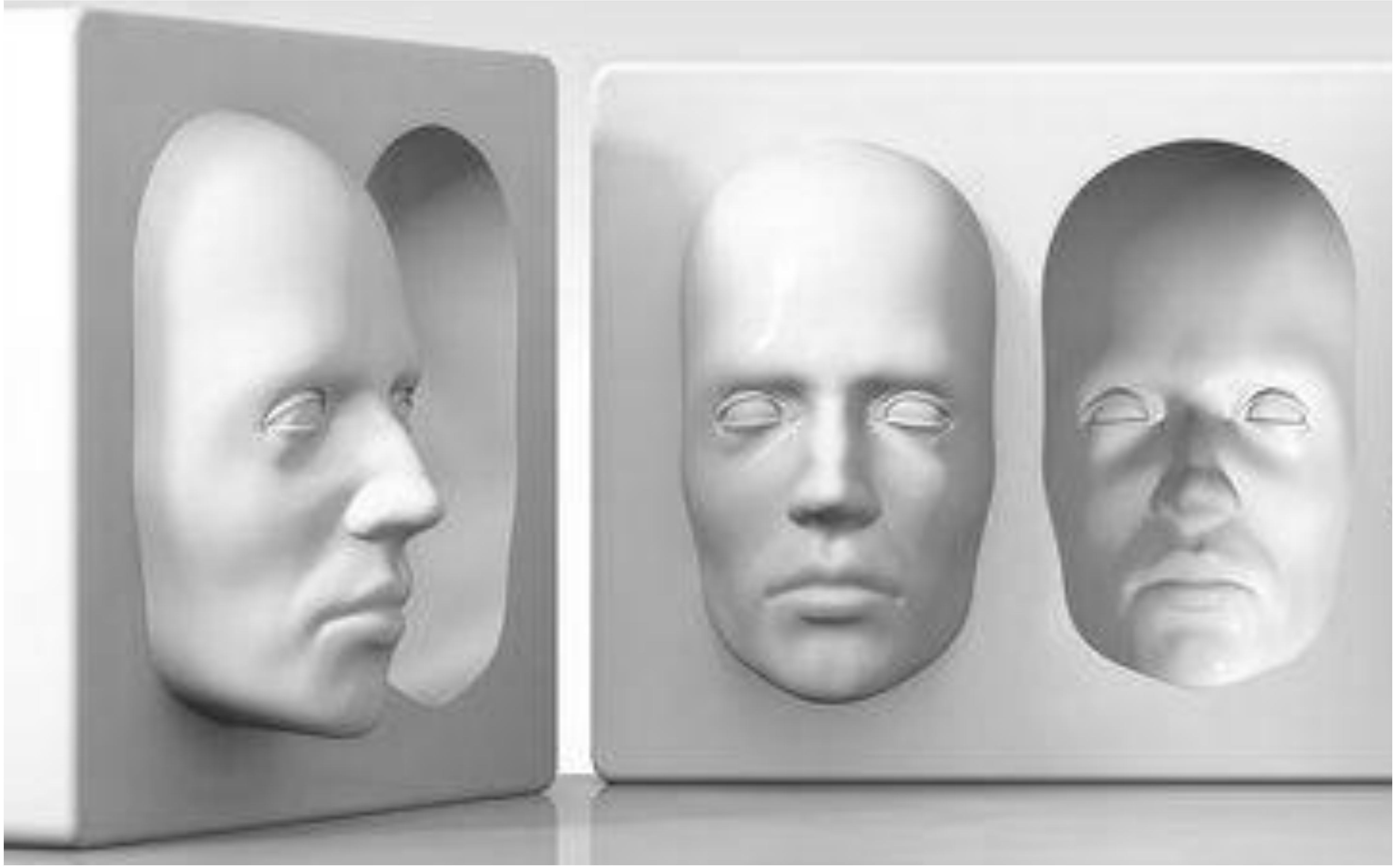


Solution 1: Recovery from a single image

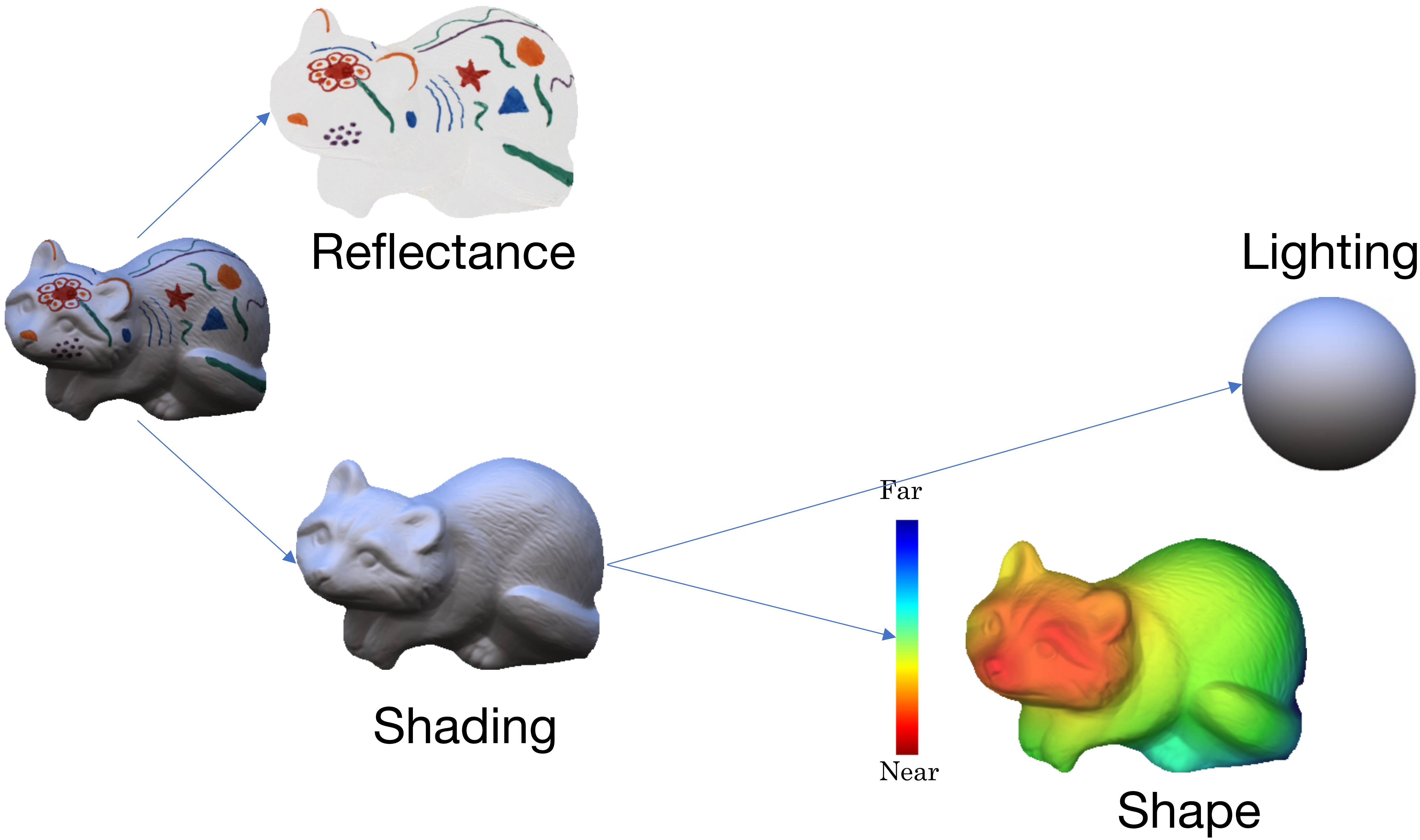


Solution 1: Recovery from a single image

Solution 1: Recovery from a single image



Solution 1: Recovery from a single image



Photometric Stereo

Key Idea: use pixel brightness to understand shape

Solution 2: Recovery from multiple images

$$I(x, y) = \rho(x, y) L_i \mathbf{L} \cdot \mathbf{N}(x, y)$$

- Represents an equation in the albedo and normals
- Multiple images give constraints on albedo and normals
- Called *Photometric Stereo*

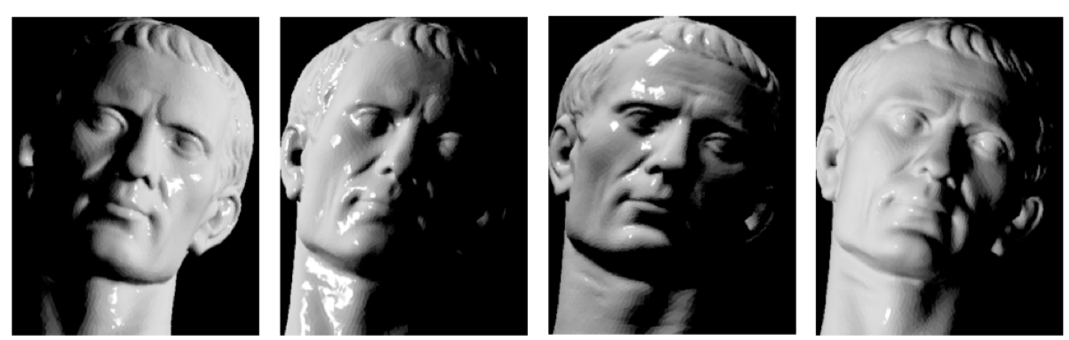


Image credit: Wikipedia

Photometric stereo - the math

$$I(x,y) = \rho(x,y) L_i \mathbf{L} \cdot \mathbf{N}(x,y)$$

- Consider single pixel
- Assume $L_i = 1$

$$I = \rho \mathbf{L} \cdot \mathbf{N}$$

$$I = \rho \mathbf{N}^T \mathbf{L}$$

- Write $\mathbf{G} = \rho \mathbf{N}$
- $\mathbf{G}$ is a 3-vector
  - Norm of **G** = ρ
  - Direction of **G** = **N**

Photometric stereo - the math

$$I = \mathbf{L}^T \mathbf{G}$$

- Multiple images with different light sources but same viewing direction?

$$I_1 = \mathbf{L}_1^T \mathbf{G}$$

$$I_2 = \mathbf{L}_2^T \mathbf{G}$$

$$\vdots$$

$$I_k = \mathbf{L}_k^T \mathbf{G}$$

Photometric stereo - the math

$$I_1 = \mathbf{L}_1^T \mathbf{G}$$

$$I_2 = \mathbf{L}_2^T \mathbf{G}$$

$$\vdots$$

$$I_k = \mathbf{L}_k^T \mathbf{G}$$

- Assume lighting directions are known
- Each is a linear equation in G
- Stack everything up into a massive linear system of equations!

Photometric stereo - the math

$$I_1 = \mathbf{L}_1^T \mathbf{G}$$

$$I_2 = \mathbf{L}_2^T \mathbf{G}$$

$$\vdots$$

$$I_k = \mathbf{L}_k^T \mathbf{G}$$

k x 3 matrix of lighting directions

$$\mathbf{I} = \mathbf{L}^T \mathbf{G}$$

k x 1 vector of intensities

3x1 vector of unknowns

Photometric stereo - the math

- How do we recover G if the problem is overconstrained?
  - More than 3 lights: more than 3 images

- Least squares

$$\min_{\mathbf{G}} \|\mathbf{I} - \mathbf{L}^T \mathbf{G}\|^2$$

Normal equations

$$\|\mathbf{I} - \mathbf{L}^T\mathbf{G}\|^2 = \mathbf{I}^T\mathbf{I} + \mathbf{G}^T\mathbf{L}\mathbf{L}^T\mathbf{G} - 2\mathbf{G}^T\mathbf{L}\mathbf{I}$$

- Take derivative with respect to **G** and set to 0

$$2\mathbf{L}\mathbf{L}^T\mathbf{G} - 2\mathbf{L}\mathbf{I} = 0$$

$$\Rightarrow \mathbf{G} = (\mathbf{L}\mathbf{L}^T)^{-1}\mathbf{L}\mathbf{I}$$

Estimating normals and albedo from $\mathbf{G}$

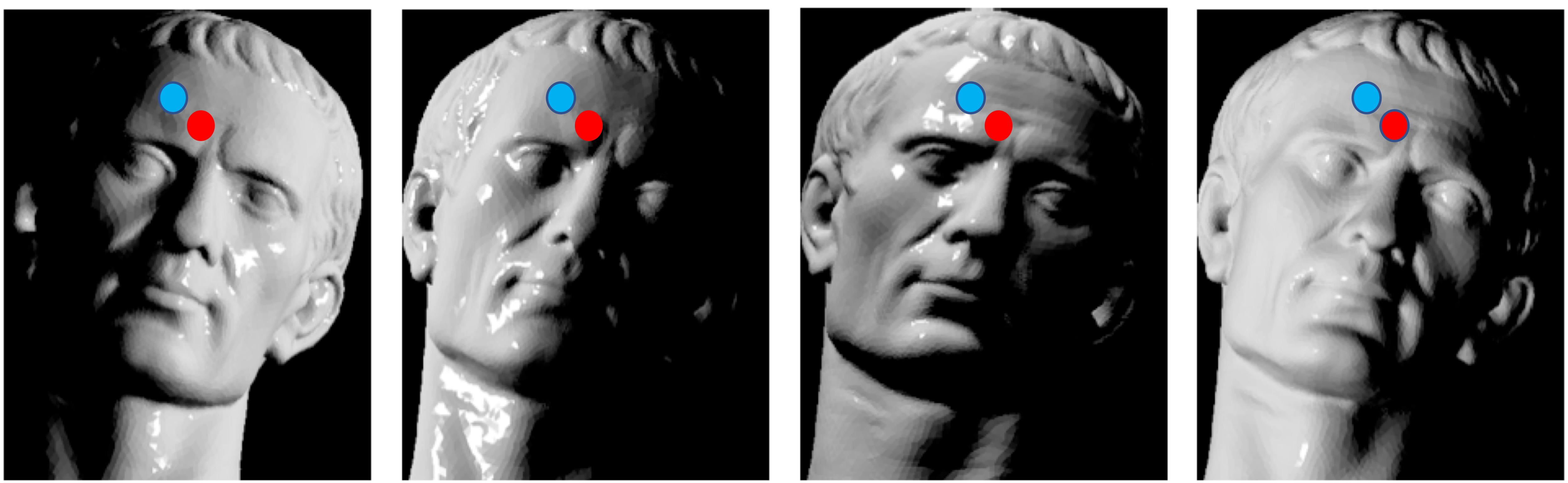
- Recall that $\mathbf{G} = \rho\mathbf{N}$

$$\|\mathbf{G}\| = \rho$$

$$\frac{\mathbf{G}}{\|\mathbf{G}\|} = \mathbf{N}$$

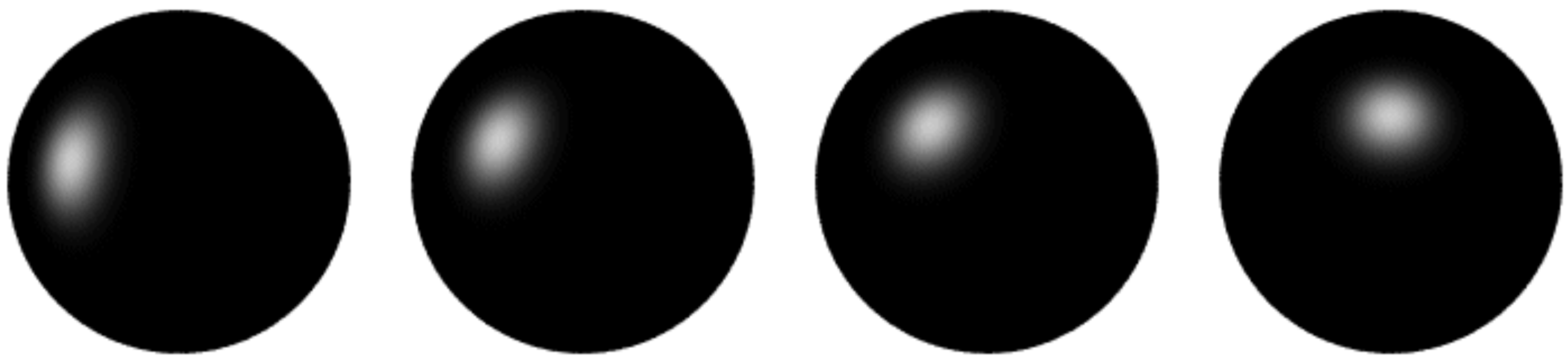
Multiple pixels

- We’ve looked at a single pixel till now
- How do we handle multiple pixels?
- Essentially independent equations!



Determining Light Directions

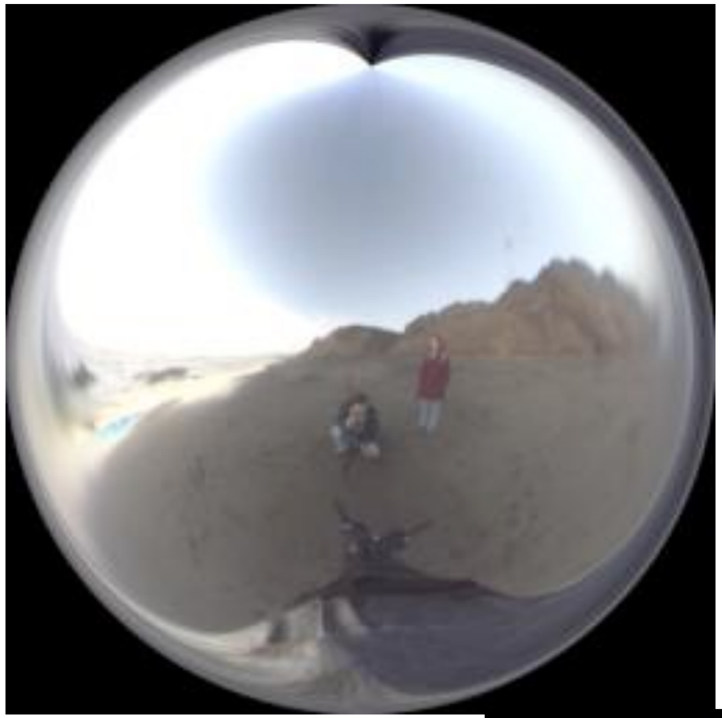
- Trick: Place a mirror ball in the scene.



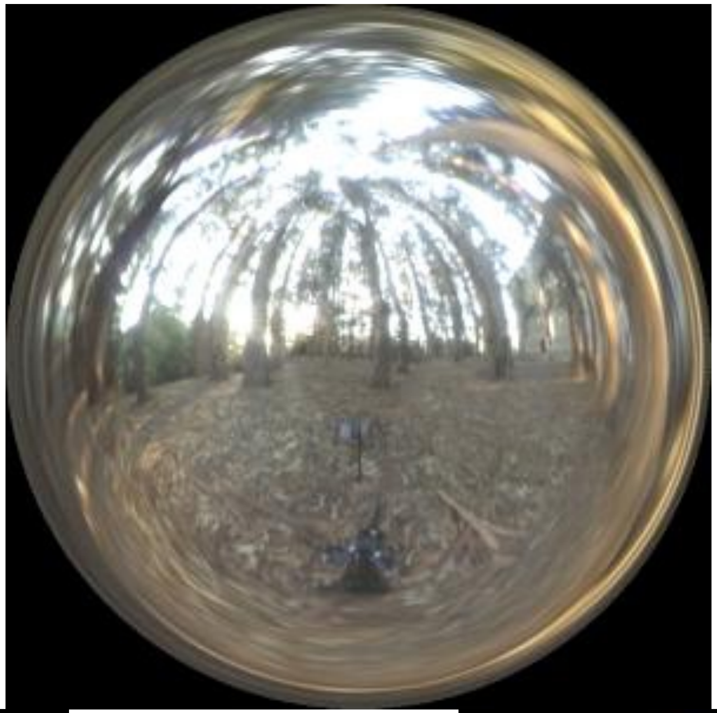
- The location of the highlight is determined by the light source direction.

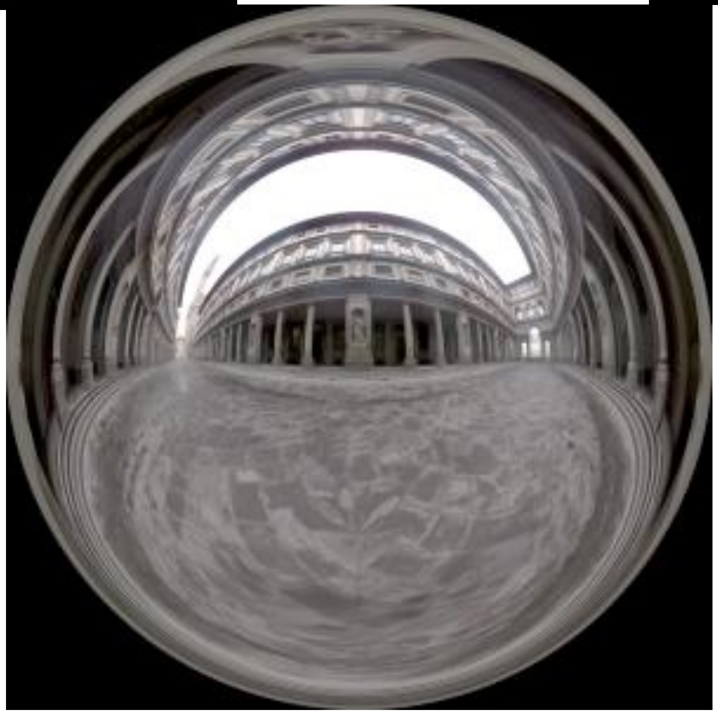
Real-World HDR Lighting Environments

Funston Beach

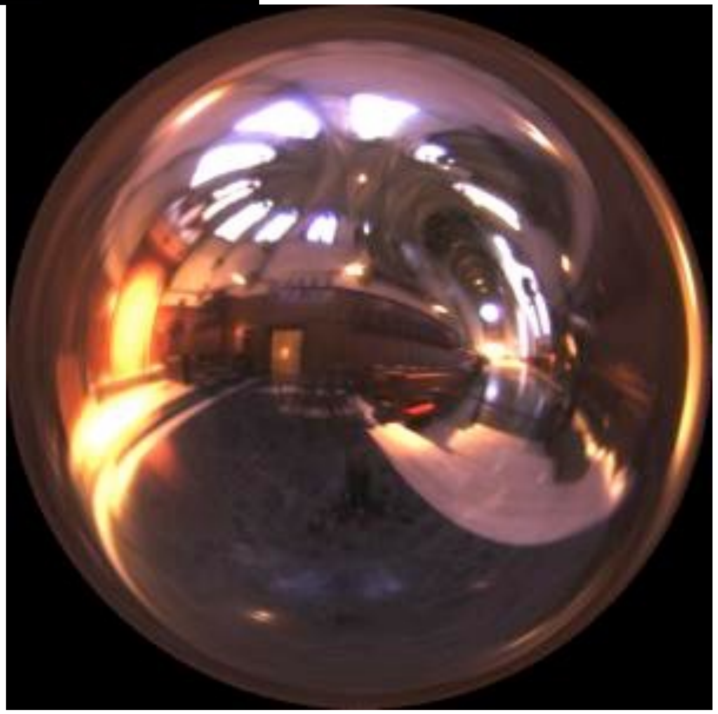


Eucalyptus Grove





Uffizi Gallery



Grace Cathedral

Lighting Environments from the Light Probe Image Gallery:
http://www.debevec.org/Probes/

Mirrored Sphere


Canon
Canon
REMOTE SWITCH

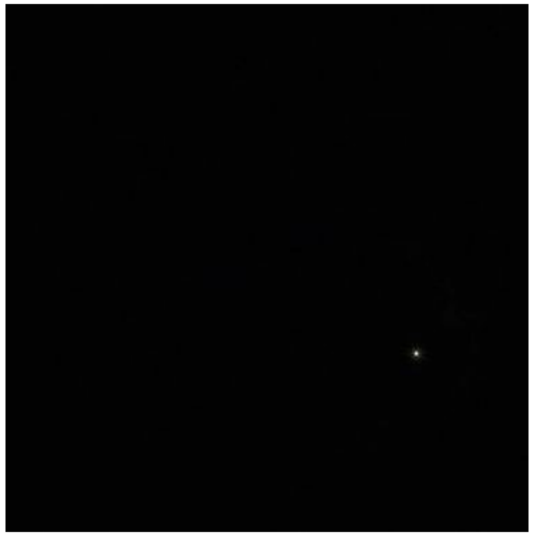
Extreme HDR Image Series

1 sec
f/4

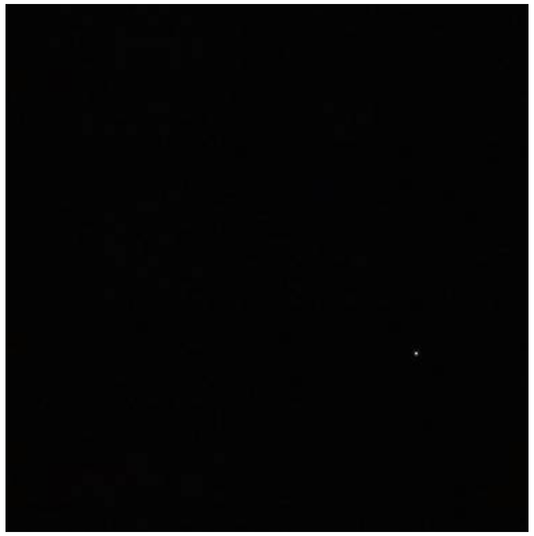
1/4 sec
f/4

1/30 sec
f/4

1/30 sec
f/16



1/250 sec
f/16



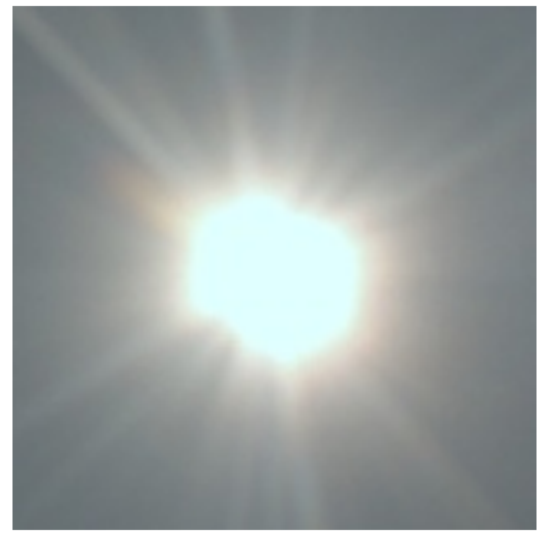
1/1000 sec
f/16

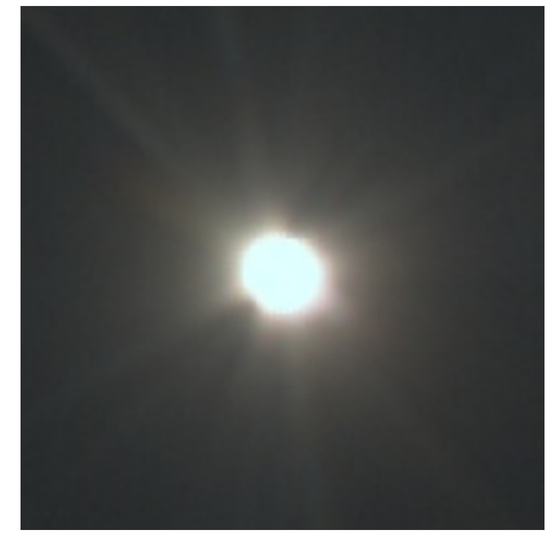
1/8000 sec f/16

Extreme HDR Image Series

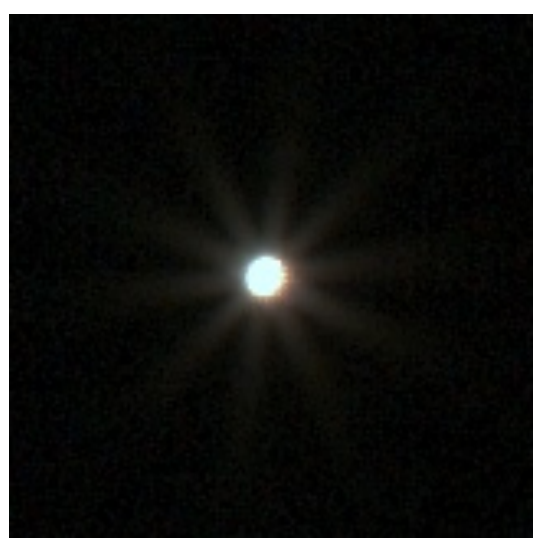
sun closeup

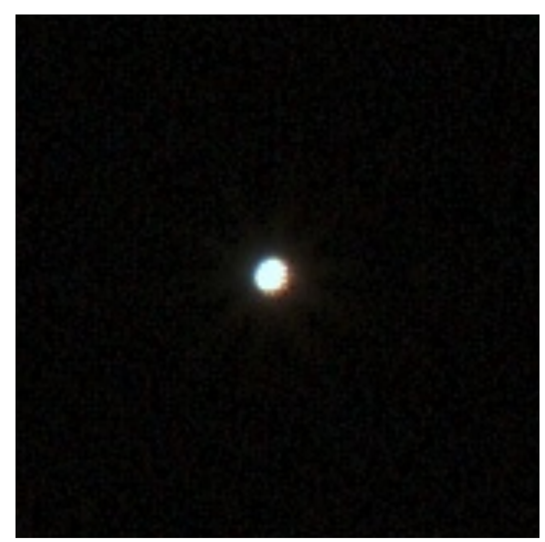
1 sec
f/4


1/4 sec
f/4


1/30 sec
f/4

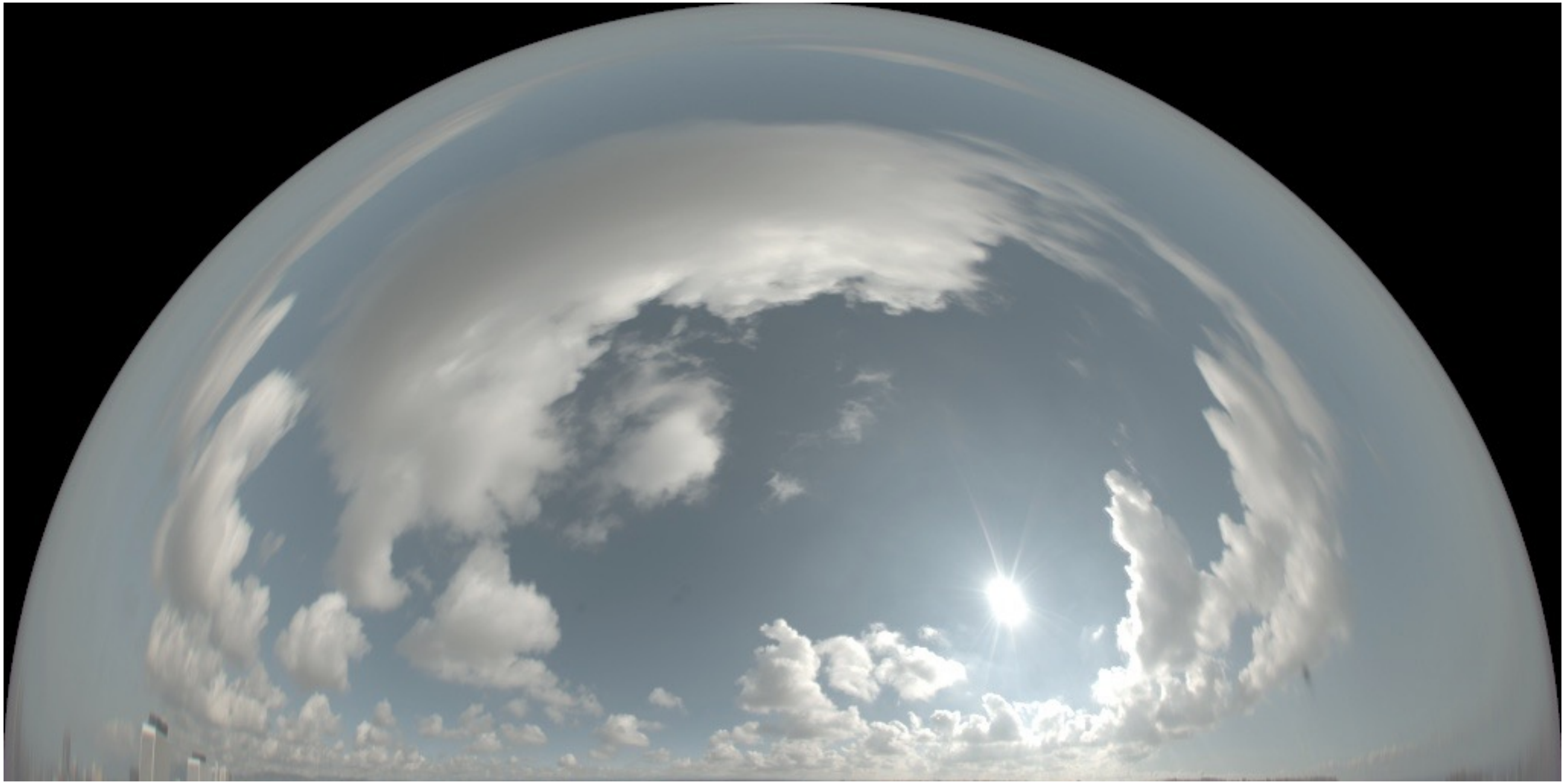

1/30 sec
f/16


1/250 sec
f/16


1/1000 sec
f/16

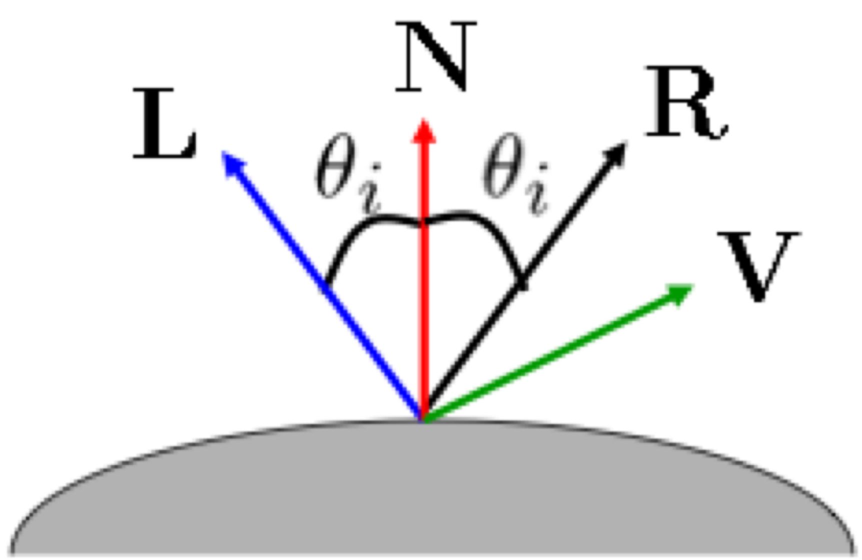

1/8000 sec f/16
only image that does not saturate!

HDRI Sky Probe



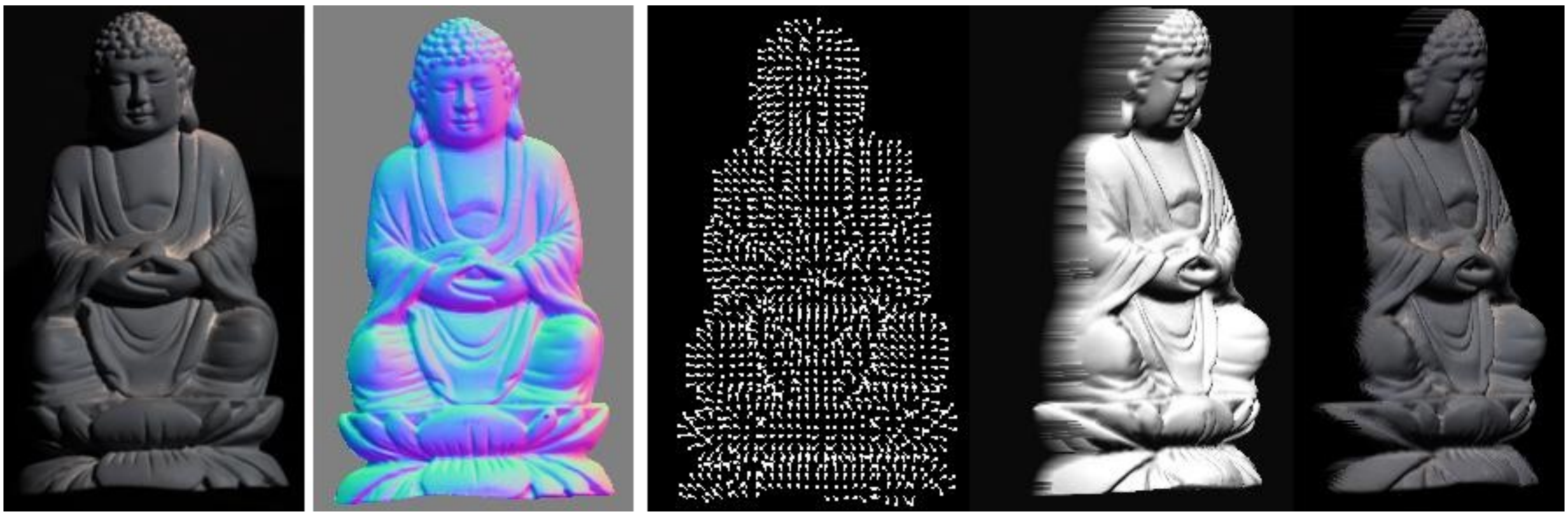
Determining Light Directions

- For a perfect mirror, the light is reflected across N:



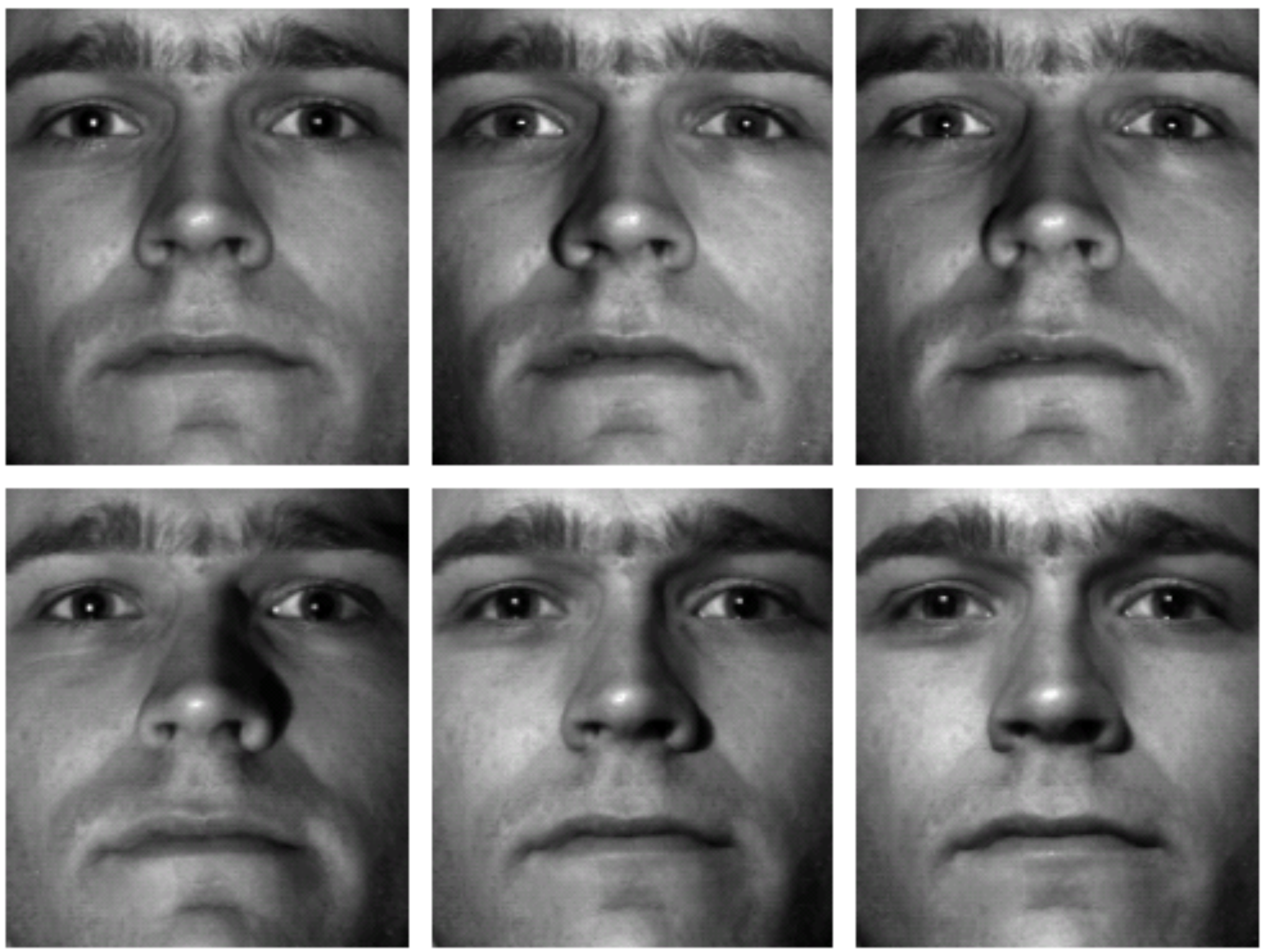
$$I_e = \begin{cases} I_i & \text{if } \mathbf{V} = \mathbf{R} \\ 0 & \text{otherwise} \end{cases}$$

Photometric Stereo



Input (1 of 12) | Normals (RGB colormap) | Normals (vectors) | Shaded 3D rendering | Textured 3D rendering

Results



from Athos Georghiades

Unknown Lighting

Surface normals Light directions

$$I = kN \cdot \ell L$$

Diffuse albedo

Light intensity

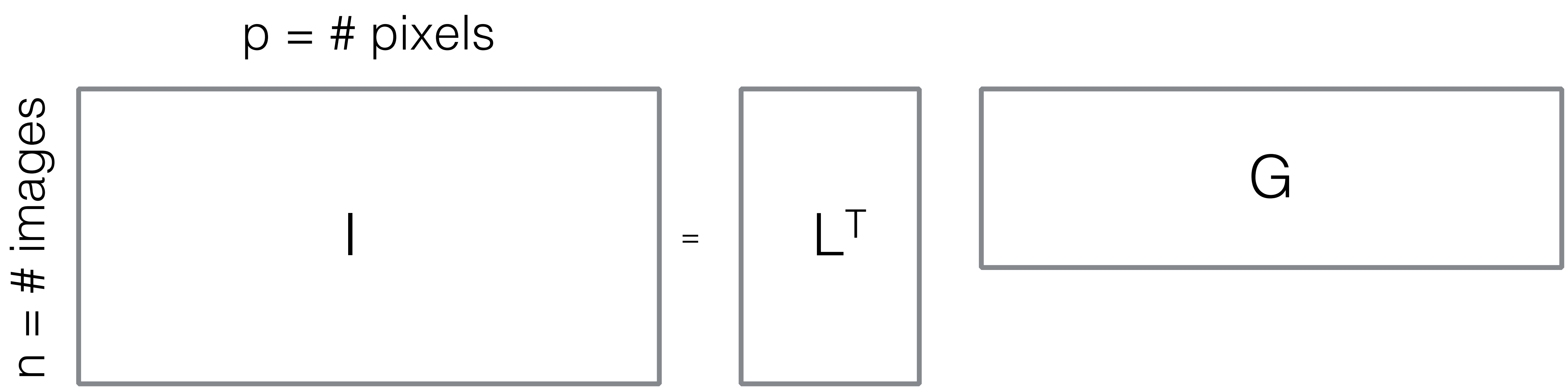
Unknown Lighting

Surface normals, scaled by albedo

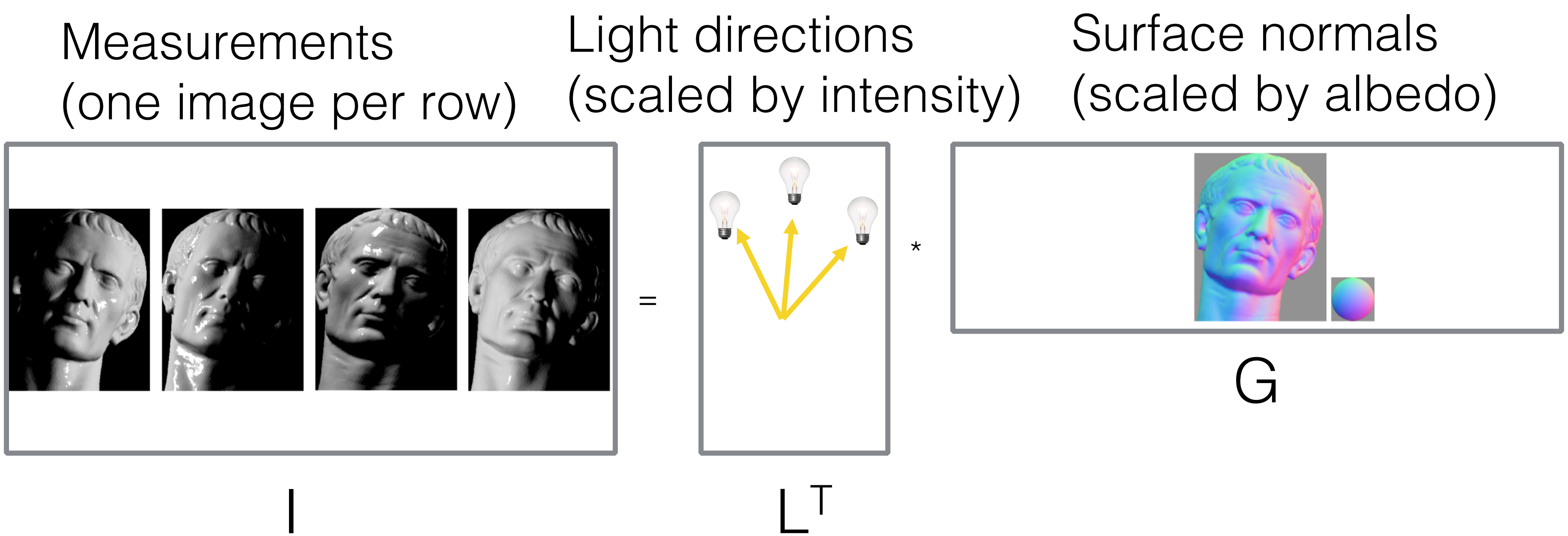
Light directions, scaled by intensity

$$I = N \cdot L$$

Unknown Lighting



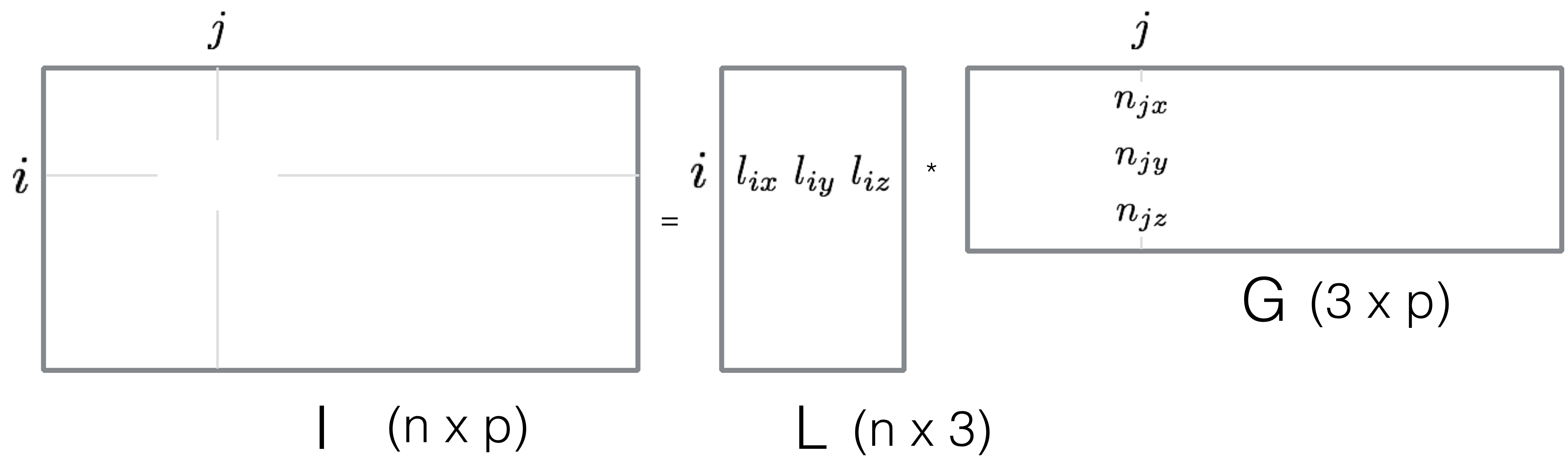
Unknown Lighting



Both L and G are now unknown!

This is a matrix factorization problem.

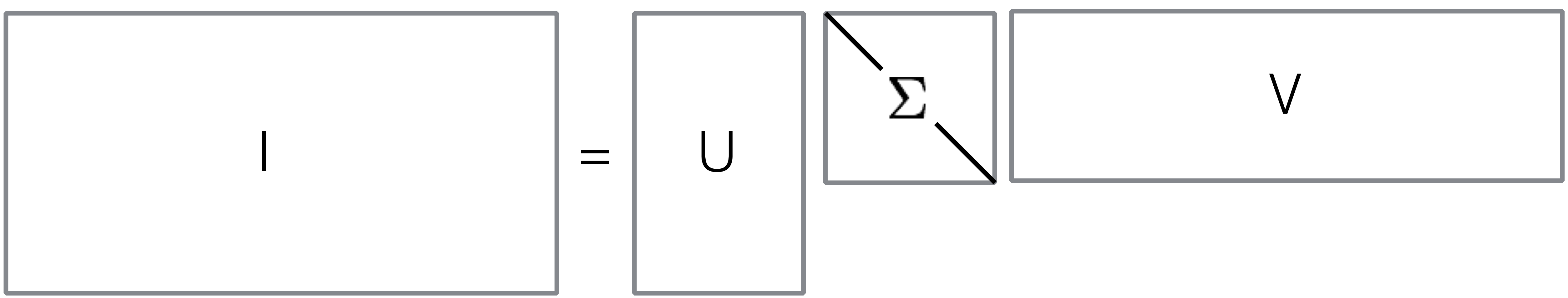
Unknown Lighting



There’s hope: We know that I is rank 3

Unknown Lighting

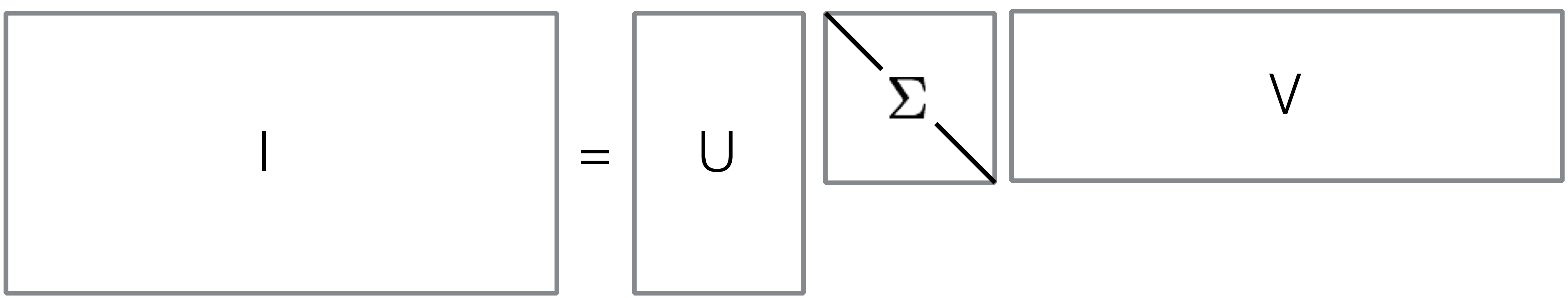
Use the SVD to decompose I:



SVD gives the best rank-3 approximation of a matrix.

Unknown Lighting

Use the SVD to decompose I:



SVD gives the best rank-3 approximation of a matrix.

What do we do with Σ?

Unknown Lighting

Use the SVD to decompose I:

$$I = U\sqrt{\Sigma} \quad \sqrt{\Sigma}V$$

Can we just do that?

Unknown Lighting

Use the SVD to decompose I:

$$I = U\sqrt{\Sigma} \; A \; A^{-1} \; \sqrt{\Sigma}V$$

Can we just do that? ...almost.

The decomposition is unique up to an invertible 3x3 A.

Unknown Lighting

Use the SVD to decompose I:

$$I = U\sqrt{\Sigma} \; A \; A^{-1} \; \sqrt{\Sigma}V$$

Can we just do that? …almost. $L = U\sqrt{\Sigma}A, G = A^{-1}\sqrt{\Sigma}V$

The decomposition is unique up to an invertible 3x3 A.

Unknown Lighting

Use the SVD to decompose I:

$$I = U\sqrt{\Sigma} \; A \; A^{-1} \; \sqrt{\Sigma}V$$

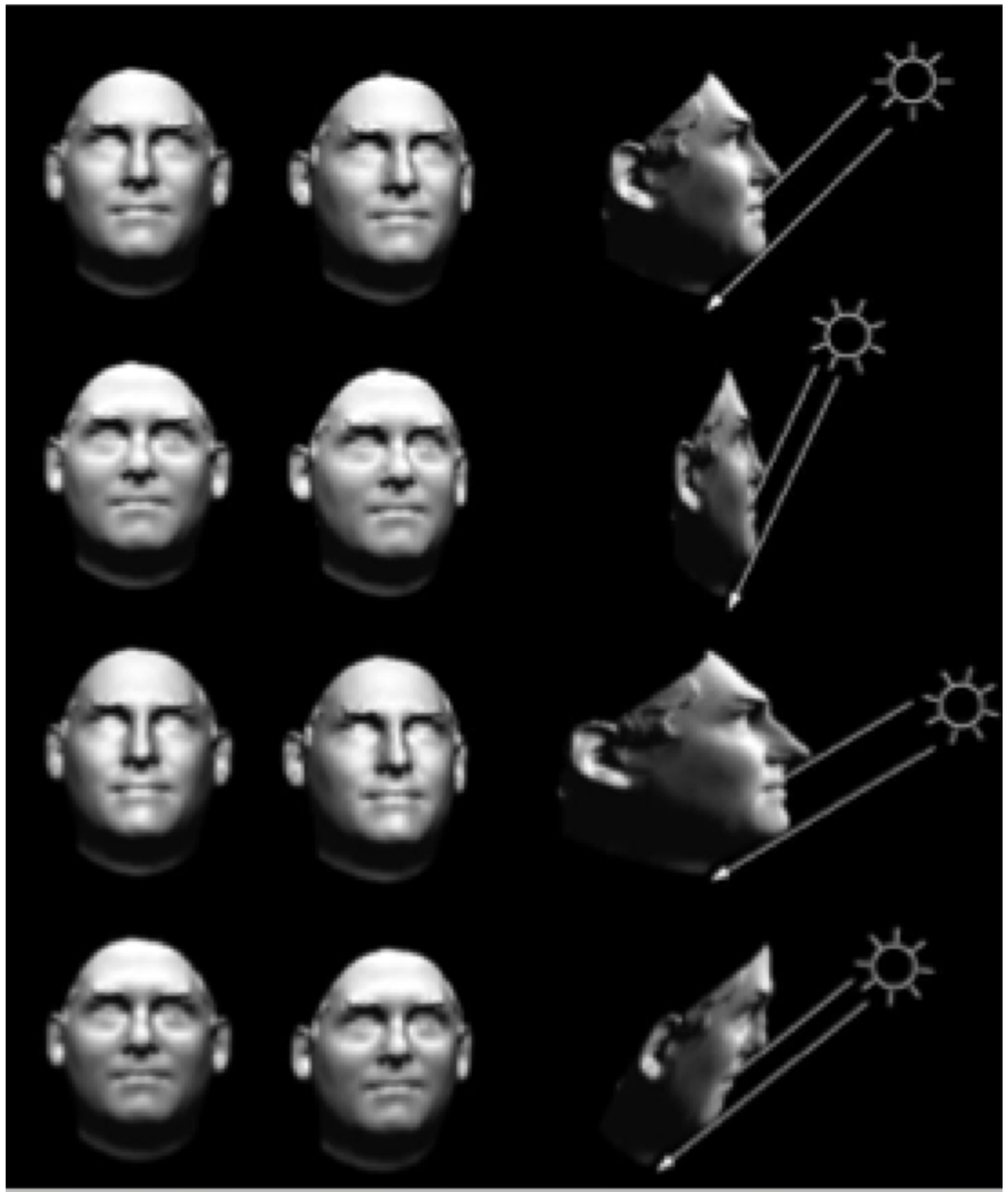
$$L = U\sqrt{\Sigma}A, G = A^{-1}\sqrt{\Sigma}V$$

You can find A if you know

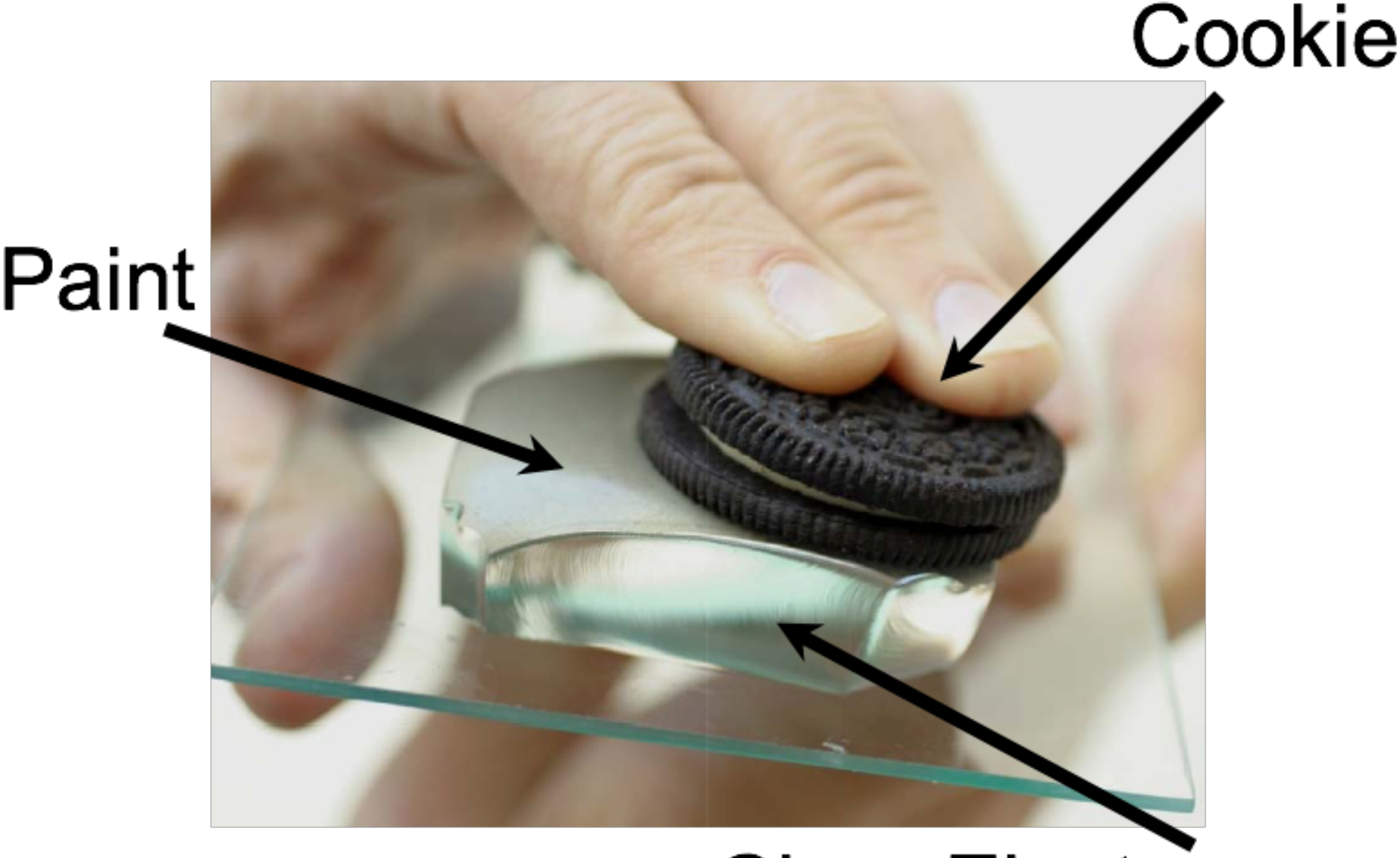
- 6 points with the same reflectance, or
- 6 lights with the same intensity.

Unknown Lighting: Ambiguities

- Multiple combinations of lighting and geometry can produce the same sets of images.
- Add assumptions or prior knowledge about geometry or lighting, etc. to limit the ambiguity.



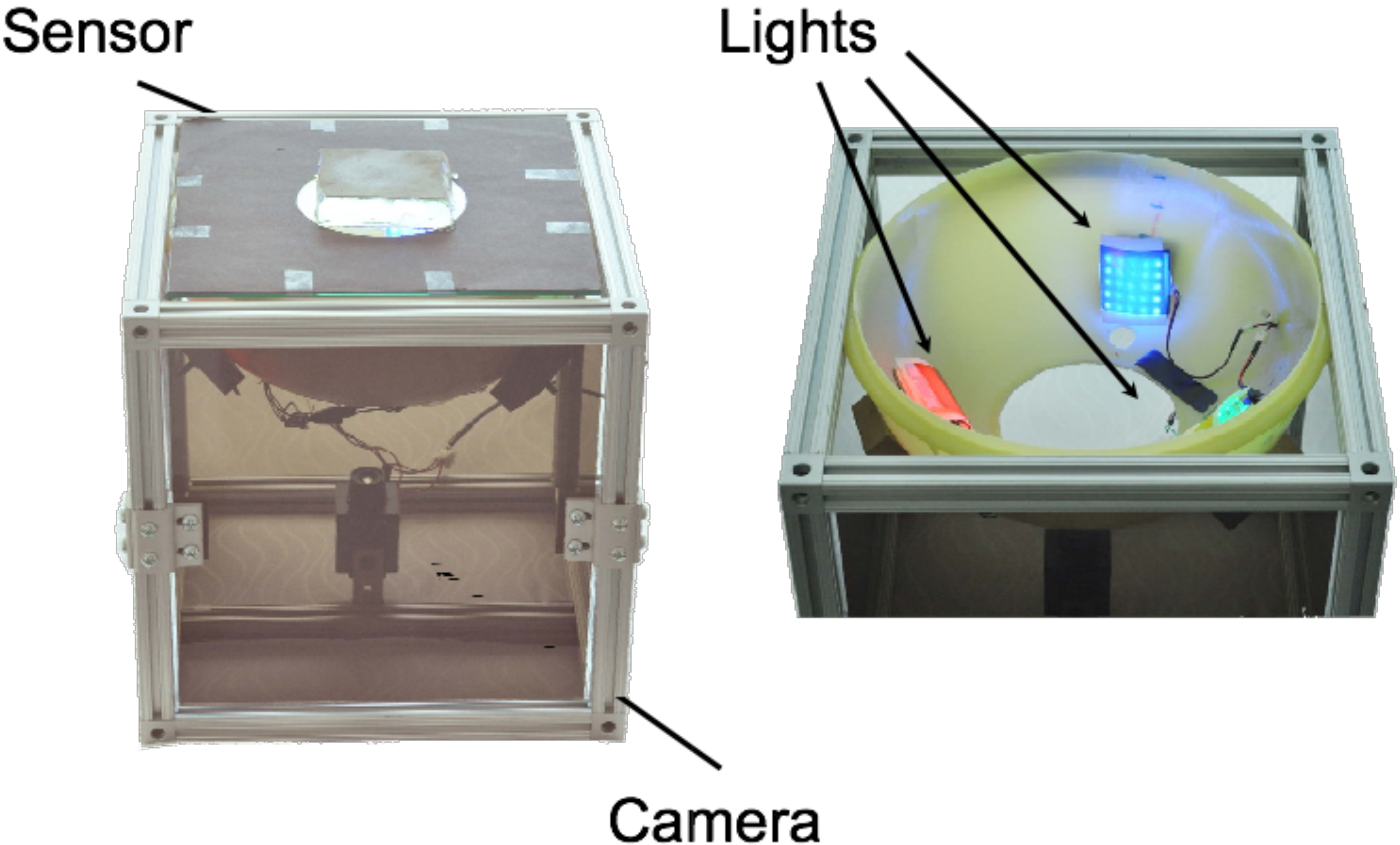
[Belhumeur et al.'97]



Johnson and Adelson, 2009


OREO

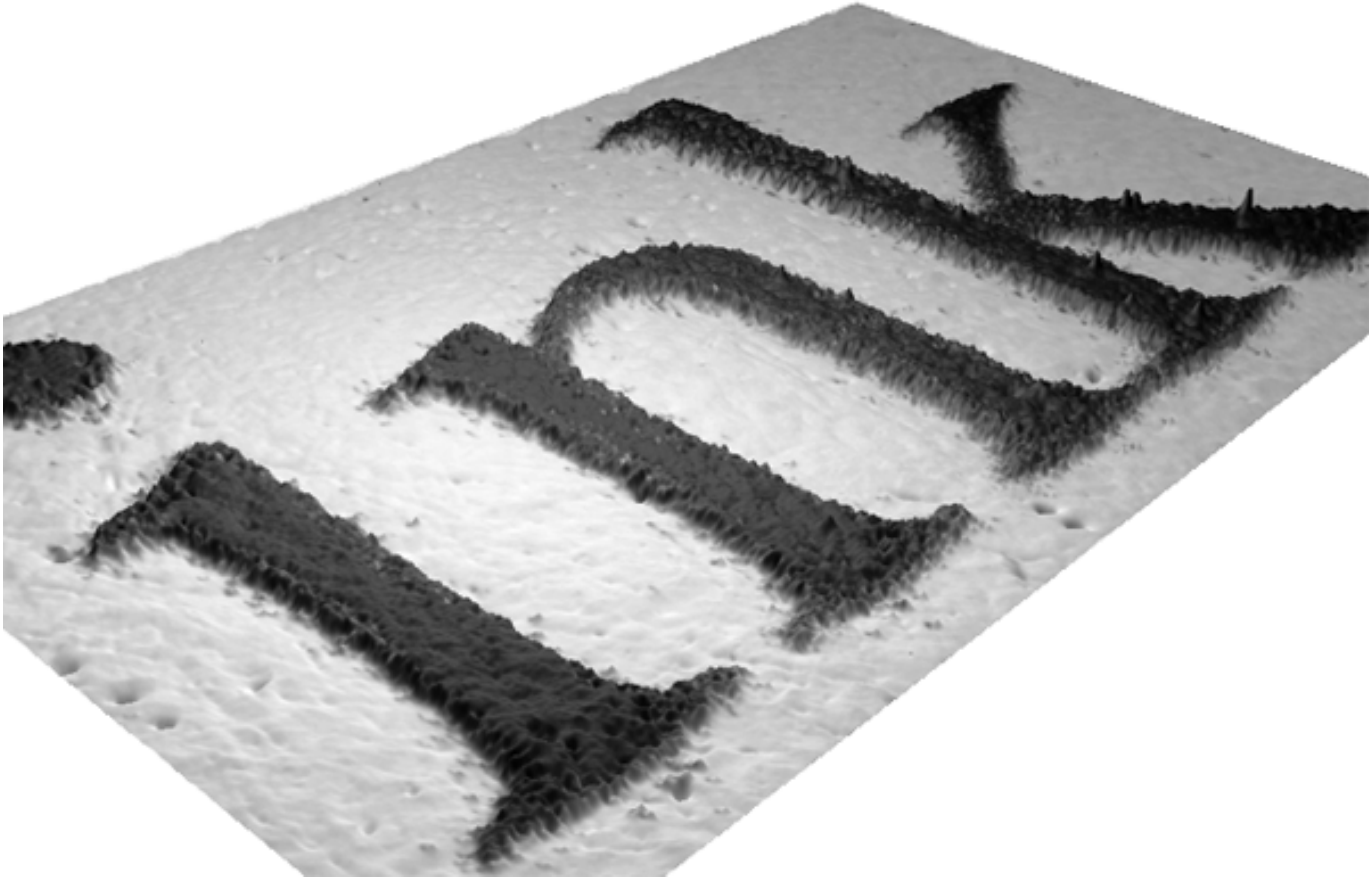
Lights, camera, action

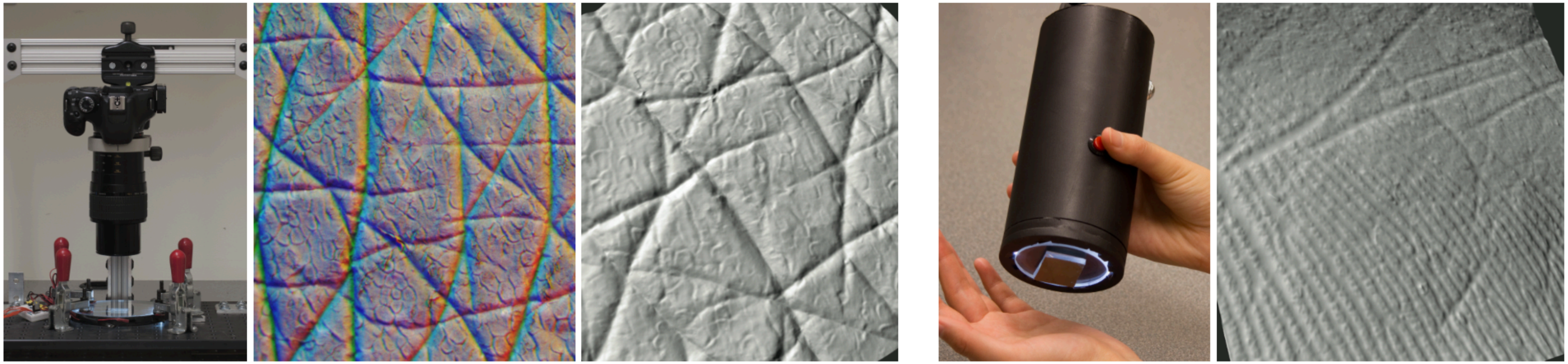


(a) (b)

Figure 2. (a) This decorative pin consists of a glass bas-relief portrait mounted in a shiny gold setting. (b) The RGB image provided by the retrographic sensor. The pin is pressed into the elastomer skin, and colored lights illuminate it from three directions.


OREO





(a) bench configuration (b) captured (c) reconstruction (d) portable configuration (e) reconstruction

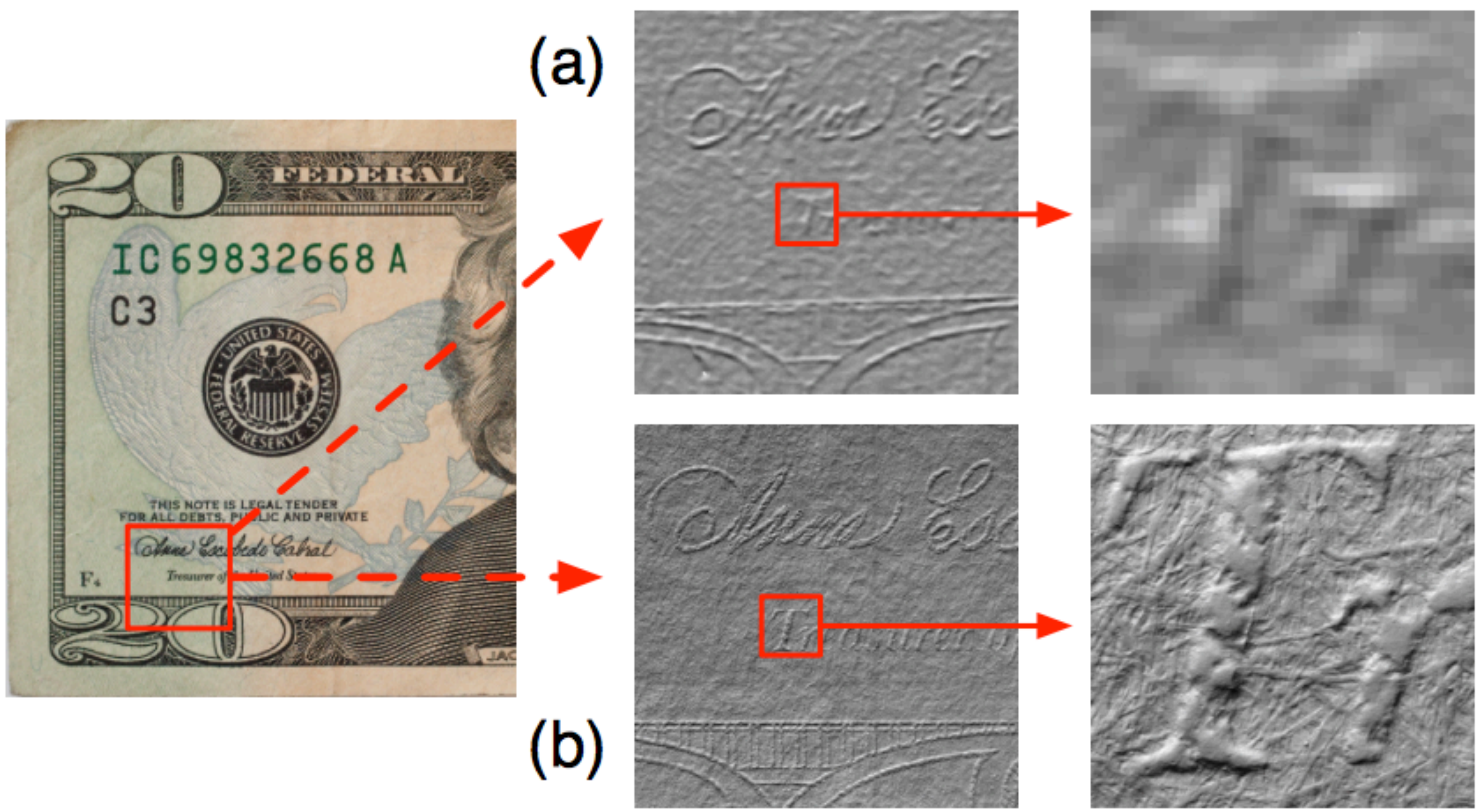


Figure 7: *Comparison with the high-resolution result from the original retrographic sensor. (a) Rendering of the high-resolution \$20 bill example from the original retrographic sensor with a close-up view. (b) Rendering of the captured geometry using our method.*

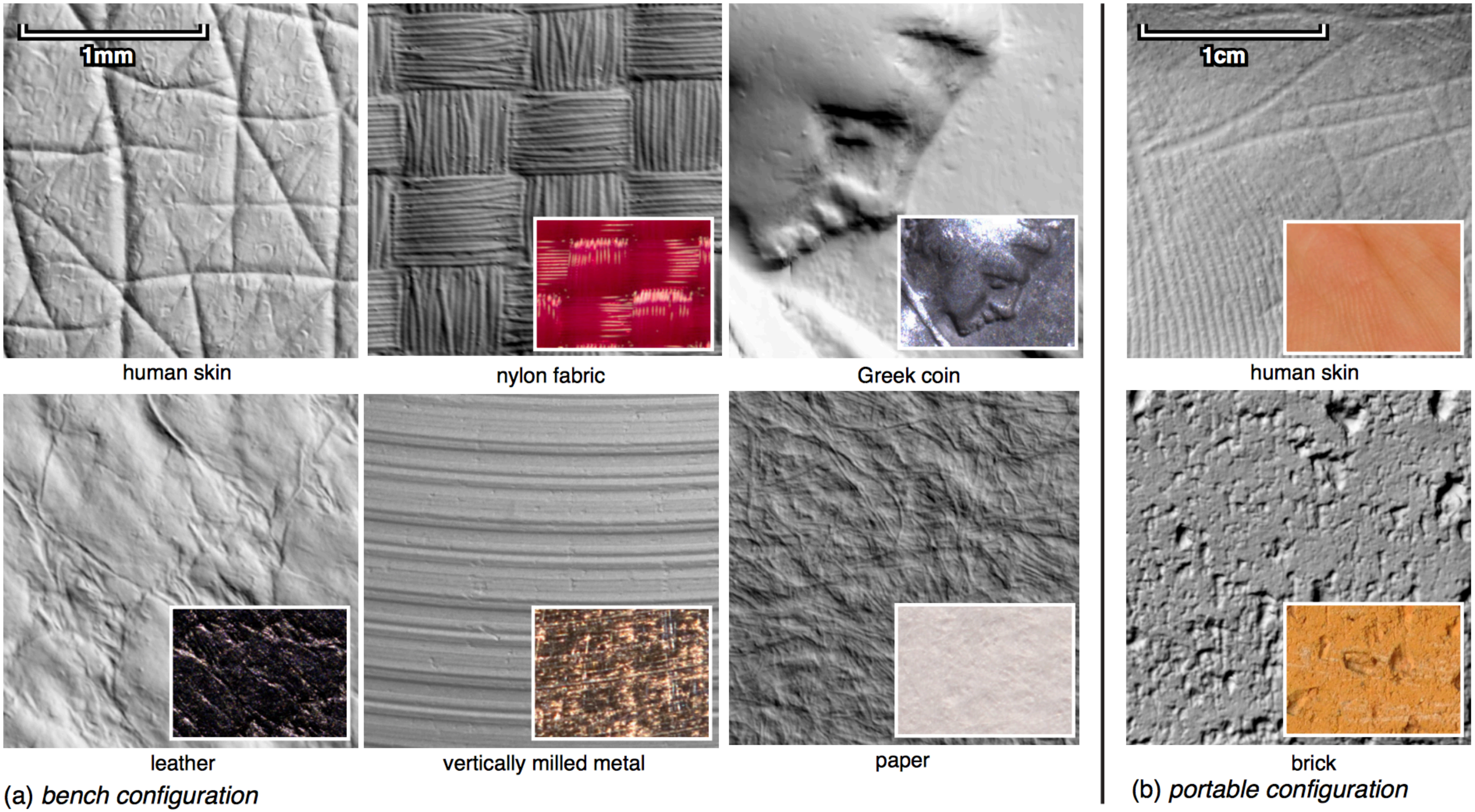


Figure 9: *Example geometry measured with the bench and portable configurations. Outer image: rendering under direct lighting. Inset: macro photograph of original sample. Scale shown in upper left. Color images are shown for context and are to similar, but not exact scale.*