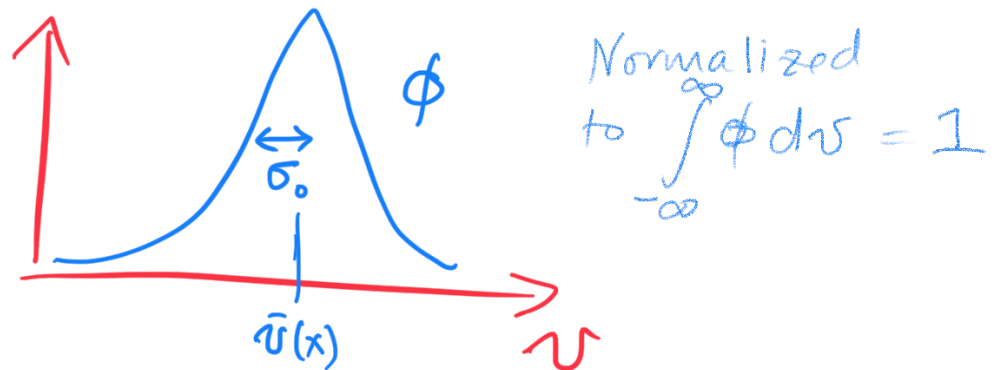


Effects of seeing on structure function

Simple case: assume uniform brightness:
 $I(x, v) = I_0 \phi[v; \bar{v}(x), \sigma_0]$ (and width)



That is, the only quantity that varies with position x is the centroid velocity $\bar{v}(x)$. In general, x is a 2d position on plane of sky.

The seeing acts on each v -slice of the velocity cube: Even though $I_0(x)$ is constant, $I(x, v)$ is not constant with x if fixed v because $\bar{v}(x)$ is varying.

$$\tilde{I}(x, v) = I(x, v) \otimes K(x, s_0)$$

where \otimes is convolution and $K(x, s_0)$ is the seeing profile with width s_0 .

For example Gaussian: $K(v, s_0) = \frac{1}{\sigma_0 \sqrt{2\pi}} \exp\left(-\frac{v^2}{2s_0^2}\right)$

for example, Gaussian: $K(x, s_0) = \frac{1}{\sqrt{2\pi}s_0} e^{-\frac{x^2}{2s_0^2}}$
 with $\text{FWHM} \approx 2.3 s_0$.

Two spatial points:

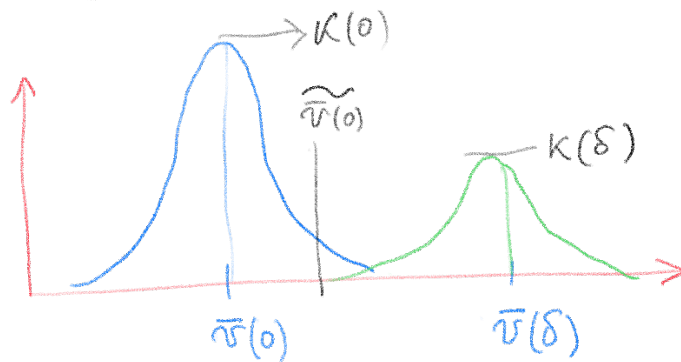
If only two points exist: $x = \{0, \delta\}$

then the convolution is just a sum:

$$\tilde{I}(0, v) = I(0, v) K(0, s_0) + I(\delta, v) K(\delta, s_0)$$

$$\tilde{I}(\delta, v) = I(\delta, v) K(0, s_0) + I(0, v) K(\delta, s_0)$$

We don't need to worry about the normalization of the convolution much, but perhaps we should also divide by $K(0, s_0) + K(\delta, s_0)$



So, the mean velocity in the blurred map

$$\tilde{v}(0) = \frac{K(0) \bar{v}(0) + K(\delta) \bar{v}(\delta)}{K(0) + K(\delta)}$$

$$\tilde{v}(\delta) = \frac{K(0) \bar{v}(\delta) + K(\delta) \bar{v}(0)}{K(0) + K(\delta)}$$

The difference in velocity is

$$\Delta v = \bar{v}(0) - \bar{v}(\delta) \quad \text{in the original map}$$

$$\tilde{\Delta v} = \tilde{\bar{v}}(0) - \tilde{\bar{v}}(\delta) \quad \text{in the blurred map}$$

$$\tilde{\Delta v} = \frac{[K(0) - K(\delta)] \bar{v}(0) - [K(0) - K(\delta)] \bar{v}(\delta)}{K(0) + K(\delta)}$$

$$\Rightarrow \tilde{\Delta v} = \left[\frac{K(0) - K(\delta)}{K(0) + K(\delta)} \right] \Delta v$$

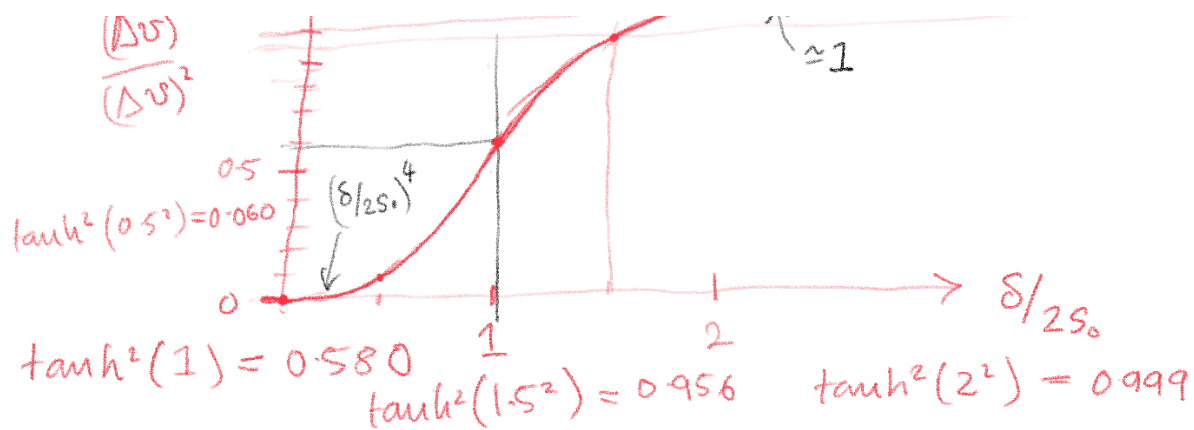
$$(\tilde{\Delta v})^2 = \left[\frac{1 - e^{-\delta^2/2s_0^2}}{1 + e^{-\delta^2/2s_0^2}} \right]^2 (\Delta v)^2$$

$$\text{or } (\tilde{\Delta v})^2 = \tanh^2 \left[\left(\frac{\delta}{2s_0} \right)^2 \right] (\Delta v)^2$$

Limit of small separations: $\delta/s_0 \rightarrow 0$

$$\Rightarrow e^{-\delta^2/2s_0^2} \simeq 1 - \frac{\delta^2}{2s_0^2}$$

$$\Rightarrow [\dots]^2 \simeq \left[\frac{1}{2} \frac{\delta^2}{2s_0^2} \right]^2 \simeq \left(\frac{\delta}{2s_0} \right)^4$$



Application to structure function.

In reality, there are more than two points

But given that $\Delta v(\Delta x)$ increases with Δx , we can assume that blurring of scales smaller than Δx has a negligible effect on $\Delta v \rightarrow \tilde{\Delta v}$

[I do not know how to prove this rigorously]

The assumption of constant I_0 can be justified as follows:

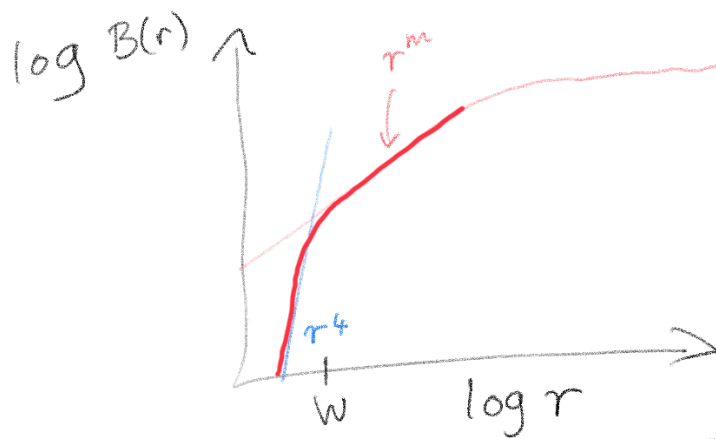
The spatial variations in I have largest amplitude for scales $\gtrsim l_0$. So long as $s_0 \ll l_0$ then the variations in I on scales affected by seeing should be small.

So, if we have an ideal structure function $B(r)$ then the effects of seeing will give a corrected

$$\tilde{B}(r) = \tanh^2\left[\left(\frac{r}{2s_0}\right)^2\right] B(r)$$

Note that s_0 is RMS seeing width, so FWHM is

$$W = 2(2\ln 2)^{1/2} s_0 \approx 2.355 s_0, \text{ so } r \approx 0.85 \frac{r}{s_0}$$



The following bit doesn't really go anywhere ...

Some of our sources may not satisfy $s_0 \ll t_0$, so it is worth thinking what the effect of I variations would be: regions where I is large would have \bar{v} less affected by seeing than in the constant I case, but low- I regions would have \bar{v} more affected. With luck, the net change would be small.

For instance, in the 2-point calculation, assume that

$$I_0(\delta) = A I_0(0)$$

$$\Rightarrow \tilde{\bar{v}}(0) = \frac{K(0) \bar{v}(0) + A K(\delta) \bar{v}(\delta)}{K(0) + A K(\delta)}$$

$$\tilde{\bar{v}}(\delta) = \frac{A K(0) \bar{v}(\delta) + K(\delta) \bar{v}(0)}{A K(0) + K(\delta)}$$

$$= \frac{K(0) \bar{v}(\delta) + A^{-1} K(\delta) \bar{v}(0)}{K(0) + A^{-1} K(\delta)}$$

$$\text{so } \Delta \tilde{\bar{v}} = \tilde{\bar{v}}(0) - \tilde{\bar{v}}(\delta)$$

$$\text{Define } R(\delta) \equiv \frac{K(0) + A K(\delta)}{K(0) + A^{-1} K(\delta)}$$

$$\tilde{\bar{v}}(\delta) = \frac{R(\delta) K(0) \bar{v}(\delta) + A^{-1} R(\delta) K(\delta) \bar{v}(0)}{K(0) + A K(\delta)}$$

... we have the same denominator, so we can subtract:

Now

$$\begin{aligned}\Delta \tilde{v} &= \frac{K(0) \tilde{v}(0) + A K(\delta) \tilde{v}(\delta) - R(\delta) K(0) \tilde{v}(\delta) - A^{-1} R(\delta) K(\delta) \tilde{v}(0)}{K(0) + A K(\delta)} \\ &= \frac{[K(0) - A^{-1} R(\delta) K(\delta)] \tilde{v}(0) - [R(\delta) K(0) - A K(\delta)] \tilde{v}(\delta)}{K(0) + A K(\delta)}\end{aligned}$$

Unfortunately, the two terms [...] are not the same, so we cannot put $\Delta \tilde{v} \propto \Delta v$

What happens when $A \rightarrow \infty$?

$$\text{In that case: } \left. \begin{array}{l} \tilde{v}(0) \rightarrow \tilde{v}(\delta) \\ \tilde{v}(\delta) \rightarrow \tilde{v}(\delta) \end{array} \right\} \Delta \tilde{v} \rightarrow 0$$

This doesn't help much...