## Weekly work summary

Guangzhi Ren

December 29, 2019

#### content

- comparision between different algorithm
- revisiting to former simulation result
- modification of energy transfer equation
- ZF/RHF/GAM

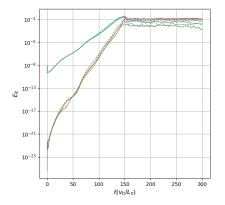


Figure: time evolution of energy in LW/RK/AB

3/25

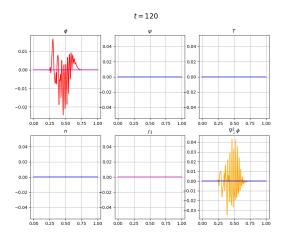


Figure: radial mode structure with AB-256

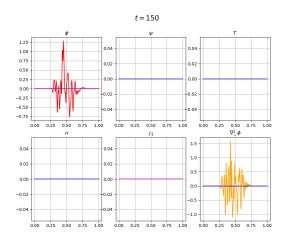


Figure: radial mode structure with LW-256

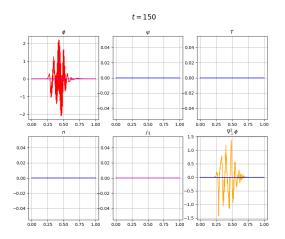


Figure: radial mode structure with AB-256

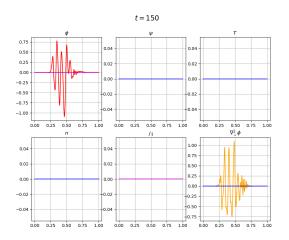


Figure: radial mode structure with AB-512

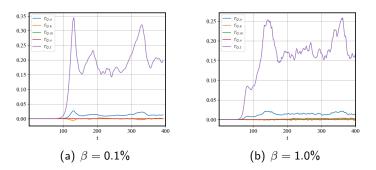


Figure: Temporal evolution of radial tansport flux

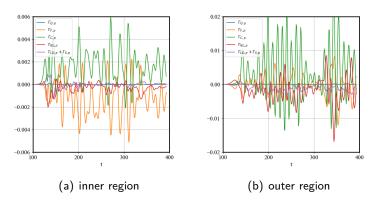


Figure: Temporal evolution of  $p_{1,0}$  energy drives with  $\beta=0.1\%$ 

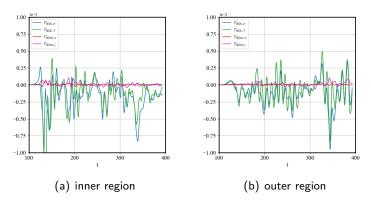


Figure: Temporal evolution of the component of  $\Gamma_{NL,p}$ 

10 / 25

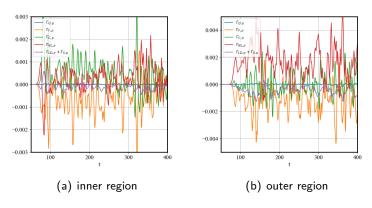


Figure: Temporal evolution of  $p_{1,0}$  energy drives with  $\beta = 0.1\%$ 

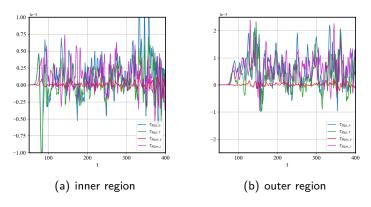


Figure: Temporal evolution of the component of  $\Gamma_{NL,p}$ 

12 / 25

modification of energy transfer equation

$$< f> = \frac{\int f dV}{\int dV} = \frac{\int f r dr d\theta d\zeta}{\int r dr d\theta d\zeta}$$
 (1)

for zonal flow energy  $E_k = \frac{1}{2} < v_E^2 >$ :

$$\frac{\partial}{\partial t}E_k = \Gamma_{NLk,k} + \Gamma_{NLm,k} + \Gamma_{C,k} \tag{2}$$

for pressure  $E_p = \frac{1}{2} < p^2 >$ :

$$p = p_i + p_e = n_0 T + (1 + \tau) T_0 n \tag{3}$$

$$\frac{\partial}{\partial t}E_{p} = \Gamma_{Q,p} + \Gamma_{F,p} + \Gamma_{C,p} + \Gamma_{LD,p} + \Gamma_{NL,p} + \Gamma_{S,p} \tag{4}$$

Guangzhi Ren

## modification of energy transfer equation

flux surface average

$$\langle f \rangle = \frac{\int f dS}{\int dS} = \frac{\int f d\theta d\zeta}{\int d\theta d\zeta}$$
 (5)

for zonal flow

$$\frac{\partial}{\partial t} < v_E > = -\frac{1}{r^2} \frac{\partial}{\partial r} r^2 < v_\theta v_r > + \frac{\beta}{n_{eq}} \frac{1}{r^2} \frac{\partial}{\partial r} r^2 < B_\theta B_r > -\frac{2\epsilon}{n_{eq}}$$
(6)

and in the code, we use the alternative expression,

$$\frac{\partial}{\partial t} < v_E > = \frac{1}{r} < \phi \frac{\partial \Omega}{\partial \theta} > + \frac{\beta}{n_{eq}} \frac{1}{r} < A \frac{\partial A}{\partial \theta} > - \frac{2\epsilon}{n_{eq}}$$
 (7)

with the Fourier expansion,

$$\frac{\partial}{\partial t} < v_E > = \sum_{m1+m2=0} \phi_{m1} \Omega_{\theta,m2} + \sum_{m1+m2=0} A_{m1} j_{\theta,m2} - \frac{i\epsilon}{n_{eq}} (p_{1,0} - p_{-1,0})$$

(8)

# modification of energy transfer equation

as for equation for thr evolution of

$$\frac{\partial}{\partial t} = < (R_Q + R_F + R_C + R_{NL} + R_{LD} + R_S) \sin \theta > \qquad (9)$$

$$< R_{Q} \sin \theta > = a[(1+\tau)T_{eq}\frac{\partial n_{eq}}{\partial r} + n_{eq}\frac{\partial T_{eq}}{\partial r}] < \phi \cos \theta >$$

$$+ (1+\tau)\beta T_{eq}\frac{\partial j_{0}}{\partial r} < A\cos \theta >$$

$$< R_{F} \sin \theta > = (\Gamma + \tau)n_{eq}T_{eq}\frac{\epsilon}{q} < v\cos \theta > -(1+\tau)T_{eq}\frac{\epsilon}{q} < j\cos \theta >$$

$$< R_{C} \sin \theta > = \epsilon(\Gamma + \tau)p_{eq} < v_{\theta} >$$

$$+ \epsilon(2\Gamma - 1)p_{eq} < \frac{\partial T}{\partial r} > + \epsilon((\Gamma - 1) - \tau(\tau + 1))T_{eq}^{2} < \frac{\partial n}{\partial r} >$$

(10)

Guangzhi Ren Wee

 $< R_{LD}\sin heta> = -(\Gamma-1)\sqrt{rac{8T_{eq}}{\pi}n_{eq}rac{\epsilon}{\sigma}} < T\sin heta>$ 

- gyrofluid simulation [Beer,Waltz]: a total collisionless decay of poloidal rotation
- gyrokinetic analysis by [RH 1998]: linear collisionless kinetic mechanisms do not damp the zonal flows completely
- verification by various gyrokinetic codes
- considering shaping effect [Xiao 2006]
- modification of zonal flow closures in gyrofluid model[Mandell,Sugama]

Considering the polarization drift in plasma,

$$v_{pj} = \frac{m_j}{e_i B^2} \frac{d\mathbf{E}}{dt} \tag{11}$$

this gives rise to the polarization current,

$$j^{cl} = \sum_{j} e_{j} n_{j} \mathbf{v}_{pj} = \epsilon_{0} \epsilon^{cl} \frac{d\mathbf{E}}{dt}$$
 (12)

where  $\epsilon^{cl} = m_i n_i / \epsilon_0 B^2 = (\omega_{pi}/\Omega_i)^2 = (k_{Di}\rho_i)^2 >> 1$ 

And in tokamak, considering the toroidal effect, we should include the traped and passing particals. Some fraction of charged particals( $f_t \sim \sqrt{\epsilon}, \epsilon = r/R$ ) are trapped by the magnetic mirror and have a radial excursion by

$$\Delta_t = \sqrt{\epsilon} \rho_{pi} = \frac{q\rho_i}{\sqrt{\epsilon}} \tag{13}$$

and as for passing partical the excursion is

$$\Delta_p = q\rho_i \tag{14}$$

So during this trapped partial orbit motion, we alse have the similar polarization effect if we have radial electric field  $E_r$ ,

$$j^{nc} = \epsilon_0 \epsilon^{nc} \frac{dE_r}{dt} \tag{15}$$

$$\epsilon^{nc} = \sqrt{\epsilon} k_{Di}^2 \Delta_t^2 = \frac{q^2}{\sqrt{\epsilon}} \epsilon^{cl}$$
 (16)

and Hinton-Rosenbluth give a expression for the polarization as,

$$\epsilon = \epsilon^{cl} + \epsilon^{nc} = \frac{\omega_{pi}^2}{\Omega_i^2} (1 + \frac{1.6q^2}{\sqrt{\epsilon}})$$
 (17)

The factor 1.6 comes from detialed kinetic calculation including passing partical contribution.

Then we consider the continuity equation of the polarization current,

$$\frac{\partial \rho_{p}}{\partial t} + \nabla \cdot \boldsymbol{j}_{p} = S \tag{18}$$

S is the external source density. Taking the flux surface average and Fourier expansion in space,

$$\frac{\partial}{\partial t} < \rho_{p}(\mathbf{k}) > + \langle i\mathbf{k}_{\perp} \cdot \mathbf{j}_{p}(\mathbf{k}) \rangle = \langle S(\mathbf{k}) \rangle$$
 (19)

$$\epsilon_0 \epsilon_p < k_\perp^2 > \Phi_k = <\rho_p > -\int < S(\mathbf{k}) > dt$$
 (20)

Consider an initial source perturbation  $< S(\mathbf{k}) >= \delta_{\mathbf{k}}(0)\delta(t)$ . when the time scale is in a few gyro motion and much shorter than the bounce time of trapped partical.we have,

$$\epsilon_0 \epsilon^{cl} \Phi_k(t=+0) = -e_i \delta n_k(0) \tag{21}$$

when the time scale is longer than the bounce time of trapped particals, the electrostatic potential is further shielded by the addition of the neoclassical polarization, then we have,

$$\epsilon_0(\epsilon^{cl} + \epsilon^{nc})\Phi_k(t = +\infty) = -e_i\delta n_k(0)$$
 (22)

Therefore, the ratio of the long term zonal flow potential to the initial zonal flow potential is given by,

$$\frac{\Phi_k(t=\infty)}{\Phi_k(t=0)} = \frac{\epsilon^{cl}}{\epsilon^{cl} + \epsilon^{nc}}$$
 (23)

Using the Hinton formula, we have,

$$\frac{\Phi_k(t=\infty)}{\Phi_k(t=0)} = \frac{1}{1 + 1.6q^2/\sqrt{\epsilon}} \tag{24}$$

- first prediction by [Winsor 1968]
- Fully forgotten between 1968 and 1996
- dispersion relation, frequency, radial structure and propagation
- close relation to alfven eigen mode
- experimental observations of GAM, H1-heliac,[Shats PRL 2002];
   DIII- D, [McKee PoP 2003]

The starting equations are the follows:

$$\rho_{0} \frac{\partial \mathbf{v}}{\partial t} = \frac{1}{c} \mathbf{J} \times \mathbf{B} - \nabla p$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho_{0} \mathbf{v} = 0$$

$$\nabla \phi = \frac{1}{c} \mathbf{b} \times \mathbf{B}$$

$$\nabla \cdot \mathbf{J} = 0$$

$$\rho_{0}^{-\gamma} - \gamma p_{0} \rho_{0}^{-\gamma - 1} \frac{\partial \rho}{\partial t} + \mathbf{v} \cdot \nabla (p_{0} \rho_{0}^{-\gamma}) = 0$$
(25)

finally, we can deduce the dispersion relation:

$$\omega^{2} \int |\rho|^{2} \mathcal{J} dS = \frac{\gamma p_{0}}{\rho_{0}} \left( |\int \rho_{0} \frac{\mathbf{B} \times \nabla \psi \cdot \nabla B^{2}}{B^{4}} \mathcal{J} dS |^{2} / \int \frac{|\nabla \psi|^{2}}{B^{2}} \mathcal{J} dS + \int \frac{|\mathbf{B} \cdot \nabla \rho|^{2}}{B^{2}} \mathcal{J} dS \right)$$
(26)

The first term is due to motion in the  $\mathbf{B} \times \nabla \psi$  direction.It is associated with geodesic curvature, i.e., the surface component of the magnetic field line curvature. And the second ordinary sound propagation propagating along the field lines.

In the limit of circular cross section with large aspect ratio  $(r \ll R)$ , the dispersion relation becomes,

$$\omega^2 = \frac{2\gamma \rho_0}{\rho_0 R^2} [1 + 1/(2q^2)] = 2\frac{C_s^2}{R^2} (1 + 1/(2q^2))$$
 (27)

with the defination of sound wave velocity in neutral gas  $C_s = (\gamma p_0/\rho_0)^{1/2}$  .

question: relation between RHF and GAM