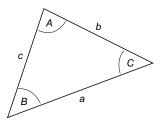
Trigonometry:

For a triangle as shown,

$$\frac{a}{\sin A} = \frac{b}{\sin B} = \frac{c}{\sin C};$$
$$a^2 = b^2 + c^2 - 2bc \cos A.$$



For any two angles A and B,

$$\sin A + \sin B = 2 \sin \left(\frac{A+B}{2}\right) \cos \left(\frac{A-B}{2}\right);$$

$$\sin A - \sin B = 2 \cos \left(\frac{A+B}{2}\right) \sin \left(\frac{A-B}{2}\right);$$

$$\cos A + \cos B = 2 \cos \left(\frac{A+B}{2}\right) \cos \left(\frac{A-B}{2}\right);$$

$$\cos A - \cos B = -2 \sin \left(\frac{A+B}{2}\right) \sin \left(\frac{A-B}{2}\right).$$

$$2 \sin A \cos B = \sin (A + B) + \sin (A - B);$$

$$2 \cos A \cos B = \cos (A - B) + \cos (A + B);$$

$$2 \sin A \sin B = \cos (A - B) - \cos (A + B).$$

$$\sin^{2} A + \cos^{2} A = 1;$$

$$1 + \cot^{2} A = \csc^{2} A;$$

$$1 + \tan^{2} A = \sec^{2} A.$$

$$\cos 2A = \cos^{2} A - \sin^{2} A;$$

$$\sin 2A = 2 \sin A \cos A;$$

$$\sin^{2} A = \frac{1 - \cos 2A}{2};$$

$$\cos^{2} A = \frac{1 + \cos 2A}{2}.$$

 $\sin (A \pm B) = \sin A \cos B \pm \cos A \sin B;$ $\cos (A \pm B) = \cos A \cos B \mp \sin A \sin B.$

Maclaurin series:

$$f(x) = f(0) + \frac{x}{1!}f'(0) + \frac{x^2}{2!}f''(0) + \dots + \frac{x^r}{r!}f^{(r)}(0) + \dots$$

Taylor series (one variable):

$$f(x) = f(a) + \frac{(x-a)}{1!}f'(a) + \frac{(x-a)^2}{2!}f''(a) + \dots + \frac{(x-a)^r}{r!}f^{(r)}(a) + \dots$$

The binomial theorem:

If n is a positive integer,

$$(1+x)^n = 1 + \frac{nx}{1!} + \frac{n(n-1)x^2}{2!} + \dots + x^n.$$

When n is negative or fractional the series is infinite, and converges when -1 < x < 1.

Some power series expansions:

$$e^x = 1 + \frac{x}{1!} + \frac{x^2}{2!} + \frac{x^3}{3!} + \dots$$
 for all x ;

$$\sin x = x - \frac{x^3}{3!} + \frac{x^5}{5!} - \frac{x^7}{7!} + \dots$$
 for all x ;

$$\cos x = 1 - \frac{x^2}{2!} + \frac{x^4}{4!} - \frac{x^6}{6!} + \dots$$
 for all x ;

$$\ln(1+x) = x - \frac{x^2}{2} + \frac{x^3}{3} - \frac{x^4}{4} + \dots \qquad -1 < x \le 1.$$

Green's theorem in the plane:

$$\oint\limits_C \left(P\,dx + Q\,dy\right) = \iint\limits_R \left(\frac{\partial Q}{\partial x} - \frac{\partial P}{\partial y}\right) dx\,dy.$$

Vector calculus:

If f(x, y, z) is a scalar field and

 $\mathbf{A}(x, y, z) = A_x \hat{\mathbf{e}}_x + A_y \hat{\mathbf{e}}_y + A_z \hat{\mathbf{e}}_z \text{ is a vector field,}$

$$\nabla^2 f = \frac{\partial^2 f}{\partial x^2} + \frac{\partial^2 f}{\partial y^2} + \frac{\partial^2 f}{\partial z^2};$$

$$\nabla^2 \mathbf{A} = \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}\right) \left(A_x \hat{\mathbf{e}}_x + A_y \hat{\mathbf{e}}_y + A_z \hat{\mathbf{e}}_z\right).$$

If f and g are scalar fields and \mathbf{A} and \mathbf{B} are vector fields.

$$\operatorname{grad}(fg) = f \operatorname{grad} g + g \operatorname{grad} f;$$

$$\operatorname{div}(f\mathbf{A}) = f \operatorname{div} \mathbf{A} + \mathbf{A} \cdot \operatorname{grad} f;$$

$$\operatorname{curl}(f\mathbf{A}) = f \operatorname{curl} \mathbf{A} + (\operatorname{grad} f) \times \mathbf{A};$$

$$\operatorname{curl} \operatorname{grad} f = 0;$$

$$\operatorname{div} \operatorname{curl} \mathbf{A} = 0;$$

$$\operatorname{curl} \operatorname{curl} \mathbf{A} = \operatorname{grad} \operatorname{div} \mathbf{A} - \nabla^2 \mathbf{A};$$

$$\operatorname{grad}(\mathbf{A} \cdot \mathbf{B}) = (\mathbf{B} \cdot \operatorname{grad}) \mathbf{A} + (\mathbf{A} \cdot \operatorname{grad}) \mathbf{B}$$

$$+ \mathbf{B} \times (\operatorname{curl} \mathbf{A}) + \mathbf{A} \times (\operatorname{curl} \mathbf{B});$$

$$\operatorname{div}(\mathbf{A} \times \mathbf{B}) = \mathbf{B} \cdot (\operatorname{curl} \mathbf{A}) - \mathbf{A} \cdot (\operatorname{curl} \mathbf{B});$$

$$\operatorname{curl}(\mathbf{A} \times \mathbf{B}) = (\mathbf{B} \cdot \operatorname{grad}) \mathbf{A} - (\mathbf{A} \cdot \operatorname{grad}) \mathbf{B}$$

$$+ \mathbf{A} \operatorname{div} \mathbf{B} - \mathbf{B} \operatorname{div} \mathbf{A}.$$

Stokes' theorem:

$$\oint_C \mathbf{A} \cdot d\mathbf{l} = \iint_S (\operatorname{curl} \mathbf{A}) \cdot d\mathbf{S}.$$

The divergence theorem:

$$\iint\limits_{S} \mathbf{A} \cdot d\mathbf{S} = \iiint\limits_{V} (\operatorname{div} \mathbf{A}) \, dV.$$

Spherical polar co-ordinates:

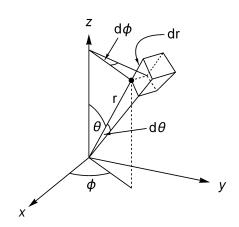


Diagram shows spherical polar co-ordinates (r, θ, ϕ) . Volume element $dV = r^2 \sin \theta \, dr \, d\theta \, d\phi$.

If f is a scalar field and

 $\mathbf{A} = A_r \hat{\mathbf{e}}_r + A_\theta \hat{\mathbf{e}}_\theta + A_\phi \hat{\mathbf{e}}_\phi$ is a vector field,

$$\operatorname{grad} f = \frac{\partial f}{\partial r} \hat{\mathbf{e}}_r + \frac{1}{r} \frac{\partial f}{\partial \theta} \hat{\mathbf{e}}_{\theta} + \frac{1}{r \sin \theta} \frac{\partial f}{\partial \phi} \hat{\mathbf{e}}_{\phi};$$

$$\operatorname{div} \mathbf{A} = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 A_r \right) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} \left(A_{\theta} \sin \theta \right) + \frac{1}{r \sin \theta} \frac{\partial A_{\phi}}{\partial \phi};$$

$$\operatorname{curl} \mathbf{A} = \frac{1}{r^2 \sin \theta} \begin{vmatrix} \hat{\mathbf{e}}_r & r \hat{\mathbf{e}}_{\theta} & r \sin \theta \hat{\mathbf{e}}_{\phi} \\ \frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & \frac{\partial}{\partial \phi} \\ A_r & r A_{\theta} & r \sin \theta A_{\phi} \end{vmatrix};$$

$$\nabla^2 f = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial f}{\partial r} \right) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial f}{\partial \theta} \right)$$

Cylindrical polar co-ordinates:

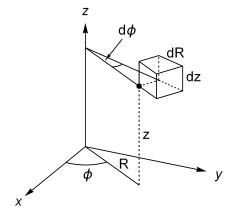


Diagram shows cylindrical polar co-ordinates (R, ϕ, z) . Volume element $dV = R dR d\phi dz$.

$$\left. \begin{array}{ll} x & = & R\cos\phi; \\ y & = & R\sin\phi; \\ z & = & z. \end{array} \right\} \qquad \left. \begin{array}{ll} R\geqslant 0; \\ 0\leqslant \phi < 2\pi; \\ -\infty < z < \infty. \end{array}$$

If f is a scalar field and

 $\mathbf{A} = A_R \hat{\mathbf{e}}_R + A_\phi \hat{\mathbf{e}}_\phi + A_z \hat{\mathbf{e}}_z$ is a vector field,

$$\operatorname{grad} f = \frac{\partial f}{\partial R} \hat{\mathbf{e}}_{R} + \frac{1}{R} \frac{\partial f}{\partial \phi} \hat{\mathbf{e}}_{\phi} + \frac{\partial f}{\partial z} \hat{\mathbf{e}}_{z};$$
$$\operatorname{div} \mathbf{A} = \frac{1}{R} \frac{\partial}{\partial R} (R A_{R}) + \frac{1}{R} \frac{\partial A_{\phi}}{\partial \phi} + \frac{\partial A_{z}}{\partial z};$$

$$\operatorname{curl} \mathbf{A} = \frac{1}{R} \begin{vmatrix} \hat{\mathbf{e}}_{R} & R \hat{\mathbf{e}}_{\phi} & \hat{\mathbf{e}}_{z} \\ \frac{\partial}{\partial R} & \frac{\partial}{\partial \phi} & \frac{\partial}{\partial z} \\ A_{R} & R A_{\phi} & A_{z} \end{vmatrix};$$

$$\nabla^{2} f = \frac{1}{R} \frac{\partial}{\partial R} \left(R \frac{\partial f}{\partial R} \right) + \frac{1}{R^{2}} \frac{\partial^{2} f}{\partial \phi^{2}} + \frac{\partial^{2} f}{\partial z^{2}}.$$

Note. The following alternative notations are sometimes used for grad, div, and curl:

$$\begin{aligned} & \operatorname{grad} f & \to & \nabla f; \\ & \operatorname{div} \mathbf{A} & \to & \nabla \cdot \mathbf{A}; \\ & \operatorname{curl} \mathbf{A} & \to & \nabla \times \mathbf{A}. \end{aligned}$$

Revised summer 2019 to match the conventions used in the School of Mathematics and Statistics rather than MathCentre.ac.uk. This version formatted in LATEX by CAH; last updated 6.ix.19.

PHYSICS & ASTRONOMY DATA CARD

Physical Constants

•		
Gravitational constant	G	$6.673 \times 10^{-11} \ \mathrm{N \ m^2 \ kg^{-2}}$
Standard acceleration of gravity	g	$9.807~{ m m}~{ m s}^{-2}$
Speed of light in vacuum	c	$2.998 \times 10^{8} \ \mathrm{m \ s^{-1}}$
Permeability of vacuum	μ_0	$4\pi imes 10^{-7}~{ m H~m^{-1}}$
Permittivity of vacuum	$\varepsilon_0 = 1/\mu_0 c^2$	$8.854 \times 10^{-12} \mathrm{F m^{-1}}$
	$(4\pi\varepsilon_0)^{-1}$	$8.988 \times 10^9 \ \mathrm{N \ m^2 \ C^{-2}}$
Elementary charge (electron charge = $-e$)	e	$1.602 \times 10^{-19} \mathrm{C}$
Electron mass	$m_{ m e}$	$9.109 \times 10^{-31} \text{ kg}$
Electron charge-to-mass ratio	$e/m_{ m e}$	$1.759 imes 10^{11}~{ m C~kg^{-1}}$
Proton mass	$m_{ m p}$	$1.673 \times 10^{-27} \text{ kg}$
Neutron mass	$m_{ m n}$	$1.675 \times 10^{-27} \text{ kg}$
Planck's constant	h	$6.626 \times 10^{-34} \mathrm{J s}$
	$\hbar = h/2\pi$	$1.055 \times 10^{-34} \mathrm{J s}$
Avogadro's constant	$N_{ m A}$	$6.022 \times 10^{23} \ \mathrm{mol^{-1}}$
Boltzmann's constant	$k_{ m B}$	$1.381 \times 10^{-23} \ \mathrm{J \ K^{-1}}$
Molar gas constant	$R = N_{\rm A} k_{\rm B}$	$8.314~{ m J}~{ m K}^{-1}~{ m mol}^{-1}$
Rydberg constant	R_{∞}	$1.097 \times 10^7 \; \mathrm{m}^{-1}$
Bohr radius	a_0	$5.292 \times 10^{-11} \text{ m}$
Bohr magneton	$\mu_{ m B} = e\hbar/2m_{ m e}$	$9.274 \times 10^{-24} \mathrm{J}\mathrm{T}^{-1}$
Nuclear magneton	$\mu_{ m N} = e\hbar/2m_{ m p}$	$5.051 \times 10^{-27} \mathrm{J}\mathrm{T}^{-1}$
Stefan-Boltzmann constant	σ	$5.670 \times 10^{-8} \ \mathrm{W \ m^{-2} \ K^{-4}}$
Wien's constant	$b = \lambda_{\max} T$	$2.898 imes 10^{-3} \ \mathrm{m \ K}$
Fine structure constant	$\alpha = e^2/4\pi\varepsilon_0\hbar c$	7.297×10^{-3}
	α^{-1}	137.04

Astronomical constants

Solar mass	M_{\odot}	$1.989 imes 10^{30}~\mathrm{kg}$
Solar radius	R_{\odot}	$6.960 imes 10^5~\mathrm{km}$
Solar luminosity	L_{\odot}	$3.90 \times 10^{26} \text{ W}$
Astronomical Unit	AU	$1.496 imes 10^{11} \mathrm{\ m}$
Light year	LY	$9.460 imes 10^{15} \mathrm{\ m}$
Parsec	pc	$206265 \text{ AU} = 3.262 \text{ LY} = 3.086 \times 10^{16} \text{ m}$

Conversion factors and definitions

$k_{\rm B}T$ at room temperature	$0.0258~\mathrm{eV}$
Electron volt	$1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$
Rydberg energy	$1 \mathrm{\ Ry} = 13.6 \mathrm{\ eV}$
Angstrom	$1 \text{ Å} = 10^{-10} \text{ m}$
Micron	$1~\mu\mathrm{m}=10^{-6}~\mathrm{m}$
Atomic mass unit, $u = m(^{12}C)/12$	$1 u = 931.5 \mathrm{MeV/c^2}$
Standard atmosphere	$1 \text{ atm} = 1.013 \times 10^5 \text{ Pa}$
Bar	$1 \text{ bar} = 10^5 \text{ Pa}$

Prefixes

Factor	10^{12}	10^{9}	10^{6}	10^{3}	10^{-3}	10^{-6}	10^{-9}	10^{-12}	10^{-15}	10^{-18}
Prefix	tera	giga	mega	kilo	milli	micro	nano	pico	femto	atto
Symbol	T	G	M	k	m	μ	n	p	f	a