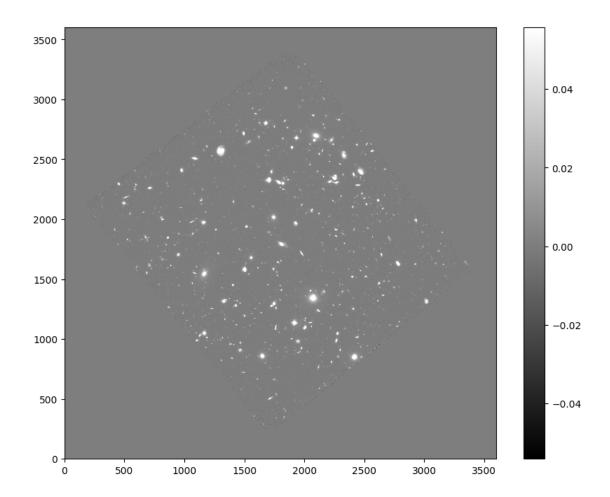
Final-Project

March 18, 2024

```
[5]: import numpy as np
     from astropy.io import fits
     import sep
     import matplotlib.pyplot as plt
     from matplotlib import rcParams
     %matplotlib inline
     rcParams['figure.figsize'] = [10., 8.]
[6]: # read image into standard 2-d numpy array
     image = fits.open("f105.fits")
     image.info()
     data = image[0].data
     print(type(data))
    Filename: f105.fits
           Name
                     Ver
    No.
                            Туре
                                      Cards
                                               Dimensions
                                                            Format
                                               (3600, 3600)
      O PRIMARY
                       1 PrimaryHDU
                                        359
                                                              float32
    <class 'numpy.ndarray'>
[7]: # show the image
     m, s = np.mean(data), np.std(data)
     plt.imshow(data, interpolation='nearest', cmap='gray', vmin=m-s, vmax=m+s,__
      ⇔origin='lower')
     plt.colorbar();
     plt.savefig('image.png',bbox_inches='tight',dpi=600)
```



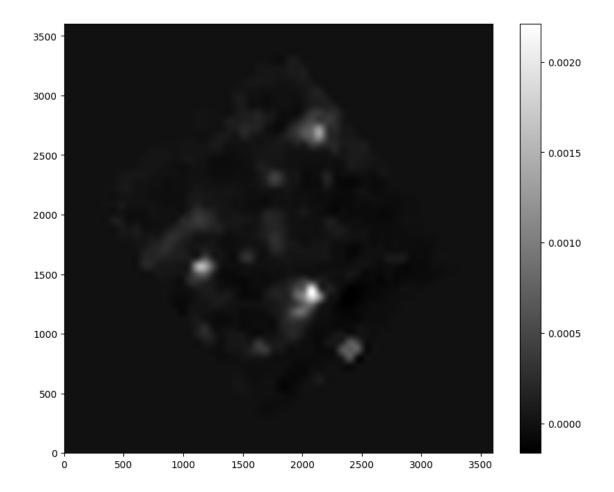
```
[8]: # measure a spatially varying background on the image
    data = data.byteswap().newbyteorder()
    bkg = sep.Background(data)

[9]: # get a "global" mean and noise of the image background:
    print(bkg.globalback)
    print(bkg.globalrms)

0.0
    0.0005398219218477607

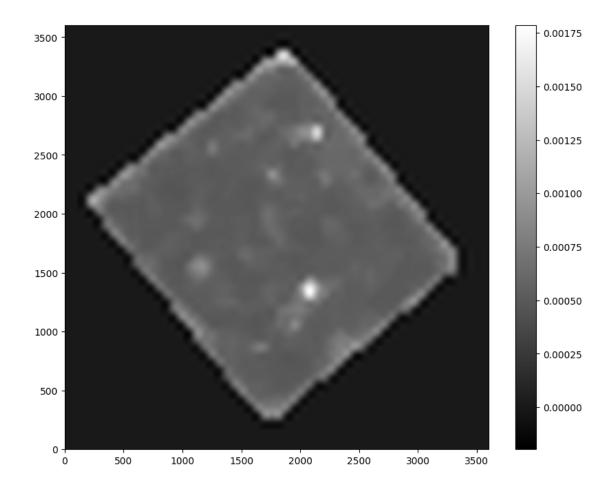
[10]: # evaluate background as 2-d array, same size as original image
    bkg_image = bkg.back()
    # bkg_image = np.array(bkg) # equivalent to above

[11]: # show the background
    plt.imshow(bkg_image, interpolation='nearest', cmap='gray', origin='lower')
    plt.colorbar();
    plt.savefig('background.png',bbox_inches='tight',dpi=600)
```

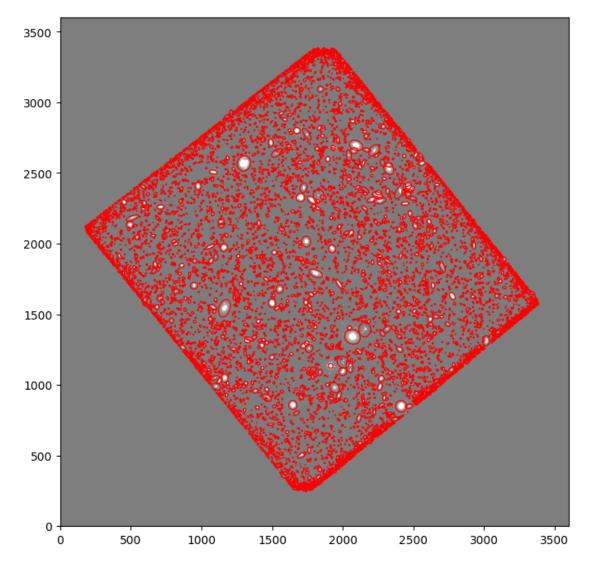


```
[12]: # evaluate the background noise as 2-d array, same size as original image
    bkg_rms = bkg.rms()

[13]: # show the background noise
    plt.imshow(bkg_rms, interpolation='nearest', cmap='gray', origin='lower')
    plt.colorbar();
    plt.savefig('backgroundNoise.png',bbox_inches='tight',dpi=600)
```



[14]: # subtract the background

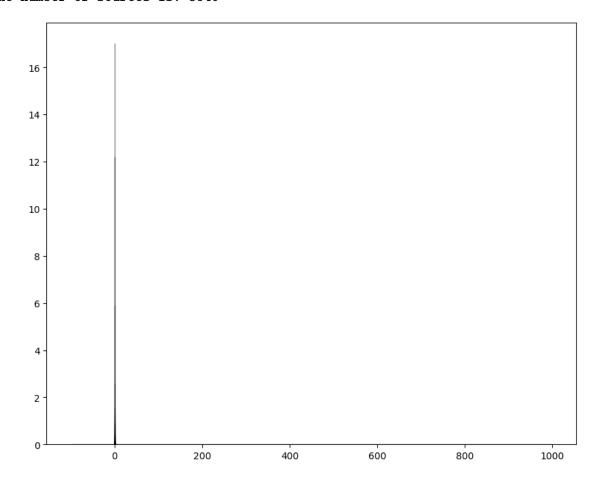


```
[18]: # available fields
objects.dtype.names
```

```
[18]: ('thresh',
       'npix',
       'tnpix',
       'xmin',
       'xmax',
       'ymin',
       'ymax',
       'x',
       'y',
       'x2',
       'y2',
       'xy',
       'errx2',
       'erry2',
       'errxy',
       'a',
       'b',
       'theta',
       'cxx',
       'cyy',
       'cxy',
       'cflux',
       'flux',
       'cpeak',
       'peak',
       'xcpeak',
       'ycpeak',
       'xpeak',
       'ypeak',
       'flag')
[19]: flux, fluxerr, flag = sep.sum_circle(data_sub, objects['x'], objects['y'],
                                            3.0, err=bkg.globalrms, gain=1.0)
[21]: # show the first 10 objects results:
      for i in range(10):
          print("object {:d}: flux = {:f} +/- {:f}".format(i, flux[i], fluxerr[i]))
     object 0: flux = 0.031282 +/- 0.176890
     object 1: flux = 0.031018 +/- 0.176142
     object 2: flux = -0.024388 + /- 0.002883
     object 3: flux = 0.001947 + - 0.044219
     object 4: flux = 0.012457 + - 0.111649
     object 5: flux = -0.011228 +/- 0.002875
     object 6: flux = 0.029368 + /- 0.171394
     object 7: flux = -0.009126 +/- 0.002875
     object 8: flux = 0.048023 +/- 0.219161
     object 9: flux = 0.027840 + - 0.166877
```

```
[50]: # number of sources:
    print("The number of sources is:",flux.size)
    # histogram of sources
    width = 0.02
    my_bins = np.arange(-100,1000,width)
    plt.hist(flux,bins=my_bins,alpha=0.5,edgecolor='black',density='True')
    plt.show()
```

The number of sources is: 8640



```
[57]: # mean, median, and std deviation
print("The mean of the fluxes is",np.mean(flux))
print("The median of the fluxes is",np.median(flux))
print("The standard deviation of the fluxes is",np.std(flux))
# outlier
outlier = 0
find = 0
for i in range(flux.size):
    if outlier < flux[i]:</pre>
```

```
outlier = flux[i]
        find = i
print("The largest outlier is:",outlier)
# show where the largest outlier is
fig, ax = plt.subplots()
m, s = np.mean(data_sub), np.std(data_sub)
im = ax.imshow(data_sub, interpolation='nearest', cmap='gray',
               vmin=m-s, vmax=m+s, origin='lower')
e = Ellipse(xy=(objects['x'][find], objects['y'][find]),
                width=6*objects['a'][find],
                height=6*objects['b'][find],
                angle=objects['theta'][find] * 180. / np.pi)
e.set_facecolor('none')
e.set_edgecolor('red')
ax.add_artist(e)
# sigma of outlier
sigma = (outlier-np.mean(flux)/np.std(flux))
print("The largest outlier is", sigma, "standard deviations away from the mean")
```

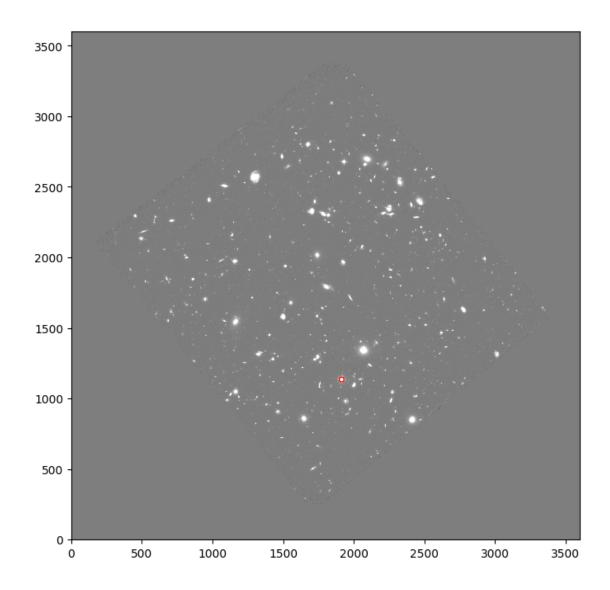
The mean of the fluxes is 0.36185728037707154

The median of the fluxes is 0.030960064365426664

The standard deviation of the fluxes is 9.243528029706706

The largest outlier is: 807.2972835731507

The largest outlier is 807.2581364770188 standard deviations away from the mean



```
[63]: # three color false image
import matplotlib.colors as colors
image2 = fits.open("f125.fits")
image.info()
data2 = image[0].data

image3 = fits.open("f160.fits")
image.info()
data3 = image[0].data

h = data3
s = data2
v = data
hsv_image = np.zeros((3600,3600,3))
```

```
hsv_image[:,:,0] = h
hsv_image[:,:,1] = s
hsv_image[:,:,2] = v
rgb_convert = colors.hsv_to_rgb(hsv_image)
f = plt.figure(figsize=(7,7))
plt.imshow(rgb_convert,origin='lower')
```

Filename: f105.fits

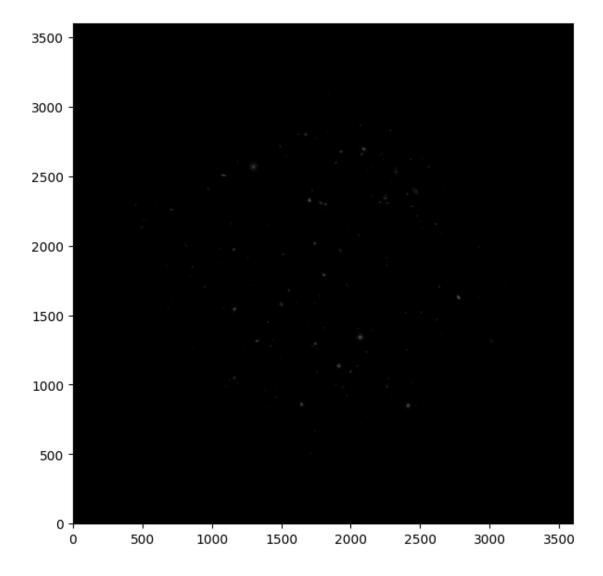
No. Name Ver Type Cards Dimensions Format 0 PRIMARY 1 PrimaryHDU 359 (3600, 3600) float32

Filename: f105.fits

No. Name Ver Type Cards Dimensions Format
O PRIMARY 1 PrimaryHDU 359 (3600, 3600) float32

Clipping input data to the valid range for imshow with RGB data ([0..1] for floats or [0..255] for integers).

[63]: <matplotlib.image.AxesImage at 0x1f66102c590>



[]:[